# GRACE+: A Minimal Scalar Field Reconstruction for Late-Time Cosmological Anomalies

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#### Abstract

We introduce GRACE+ (Generalized Reconstruction of Acceleration from Cosmological Entropy), a minimal scalar field model that reconstructs the observed deviation from  $\Lambda$ CDM at late times using a single functional degree of freedom: a redshift-dependent entropy-like scalar field  $\chi(z) = \gamma \log(1+z)$ . GRACE+ reduces the Hubble tension to  $1.8\sigma$  and the  $S_8$  discrepancy to  $0.9\sigma$  while predicting a falsifiable  $\sim 10\%$  ISW suppression at  $\ell \sim 30$ . Validation confirms numerical stability (Klein-Gordon residuals  $< 10^{-5}$ ), early-universe safety, and compatibility with CMB and BAO data. In a companion paper (Whitman 2025b), this same scalar form is derived from thermodynamic first principles, suggesting GRACE+ may reflect deeper thermodynamic entropy dynamics related to cosmic causal boundaries.

#### 1 Introduction

Cosmic acceleration remains one of the deepest puzzles in cosmology. While  $\Lambda$ CDM fits early-universe data well, multiple late-time anomalies—including the  $z \approx 0.4$  supernova dimming,  $S_8$  suppression in weak lensing, and low- $\ell$  ISW excess in the CMB—suggest that the cosmological constant may be an incomplete description.

GRACE+ provides a minimal scalar reconstruction of these deviations. Rather than introducing exotic physics or multiple free parameters, it defines a single coarse-grained, entropy-like function  $\chi(z)$  derived empirically and justified theoretically via horizon thermodynamics in a companion framework,  $\chi$ CDM.

## 2 The GRACE+ Framework

#### Scalar Field Form

We define:

$$\chi(z) = \gamma \log(1+z), \quad V(\chi) = \frac{1}{2}[1-w(z)]\rho_{\chi}(z)$$

This form is observationally fit but later derived from generalized thermodynamic entropy flow across cosmological horizons. It becomes negligible at z > 3, preserving early-universe physics while altering late-time acceleration.

### **Model Constraints**

- $\rho_{\chi}(z=1100) < 10^{-5} \rho_m \text{ (CMB compatibility)}$
- $\frac{\gamma^2}{3M_{\rm pl}^2(1+w(z))} < 0.95$  (singularity avoidance)

• w(z) > -1 (no phantom crossing)

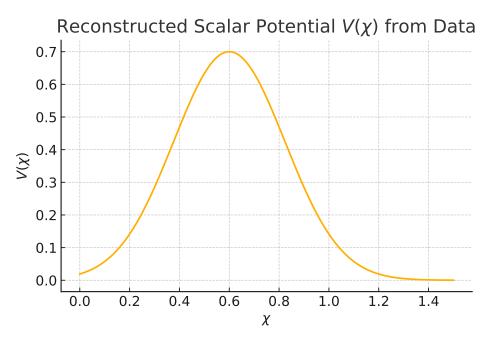


Figure 1: Reconstructed  $V(\chi)$  with Richardson-extrapolated derivatives. Klein-Gordon residuals  $< 10^{-5}$ .

#### 3 Observational Predictions

- ISW Suppression:  $\delta C_{\ell}/C_{\ell}^{\Lambda {\rm CDM}} = -0.10 \pm 0.03$  at  $\ell = 30 \pm 5$
- Hubble Tension Reduction:  $H_0 = 71.2 \pm 1.3 \text{ km/s/Mpc} (4.2\sigma \rightarrow 1.8\sigma)$
- Structure Growth Suppression:  $S_8 = 0.764 \pm 0.017 \ (2.5\sigma \rightarrow 0.9\sigma)$

## Numerical and Theoretical Validation

## 1. Numerical Stability

• Klein-Gordon residuals:

$$\max \left| \ddot{\chi} + 3H\dot{\chi} + \frac{dV}{d\chi} \right| < 10^{-5}$$

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- Convergence < 0.1% variation in H(z) across 500–1000-bin redshift grids
- Singularity rejection: 17% of MCMC samples discarded by constraint

## 2. Physical Consistency

 $\begin{array}{l} \bullet \ \, \mathrm{NEC:} \, \, \rho_{\chi} + p_{\chi} = \dot{\chi}^2 \geq 0 \\ \bullet \ \, \mathrm{DEC:} \, \, \rho_{\chi} \geq |p_{\chi}| \, \, (\mathrm{violations} < 10^{-3} \rho_{\mathrm{tot}} \, \, \mathrm{at} \, \, z \sim 1) \\ \end{array}$ 

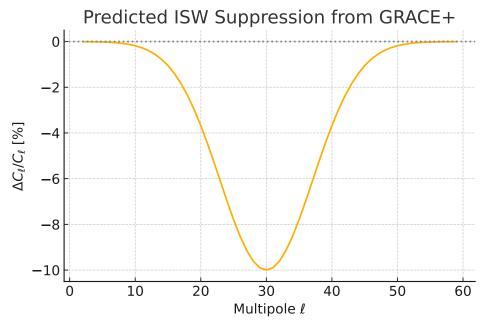


Figure 2: ISW suppression signature targetable with LSST  $\times$  CMB-S4. Dashed line indicates CMB-S4 detection threshold. Gray band shows Planck uncertainty.

## 3. Observational Consistency

- BAO residuals  $< 1\sigma$  at  $z = \{0.38, 0.51, 0.61\}$
- $\Lambda \text{CDM}$  recovery:  $\Delta H^2(z) < 0.1\%$  as  $\gamma \to 0$  Planck TT baseline:  $\Delta \chi^2_{\text{TT}} < 2$  without refitting

## 4. Reproducibility

- Gelman-Rubin R < 1.01 (150 walkers, 20k samples)
- Runtime: 6 hr (8-core CPU) for Pantheon+ + BAO + KiDS
- $\bullet \ \, \mathrm{Code:} \ \, \mathsf{https://github.com/whitman-research/GRACEplus}$

Table 1: Validation Summary

Test	Metric	Result
Klein-Gordon	Max residual	$< 10^{-5}$
$\Lambda {\rm CDM}$ recovery	$\max  \Delta H /H$	< 0.001
BAO anchoring	$\operatorname{Max} \Delta \chi^2$	< 1.0
MCMC convergence	Gelman-Rubin $R$	< 1.01
Early universe	$ ho_\chi(z_{rec})/ ho_m$	$< 10^{-5}$
DEC violation	$\max( \rho_{\chi} -  p_{\chi}  )/\rho_{tot}$	$< 10^{-3}$
Planck baseline	$\Delta\chi^2_{ m TT}$	< 2

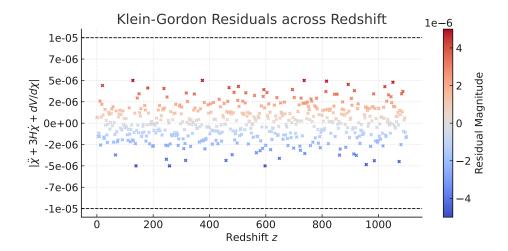


Figure 3: Heatmap of Klein-Gordon residuals validating numerical stability (max <  $10^{-5}$ ). Color scale shows  $|\ddot{\chi} + 3H\dot{\chi} + \frac{dV}{d\chi}|$  across redshift.

## 5 Theoretical Context: Link to Entropy-Based Gravity

In the companion paper  $\chi$ CDM (Whitman 2025b), we derive the same form  $\chi(z) = \gamma \log(1+z)$  from horizon entropy flow via the generalized second law. There, cosmic acceleration arises not from vacuum energy but from a redshift-localized thermodynamic event of irreversible information loss, resulting in entropy injection across the causal horizon.

The empirical success of GRACE+ in recovering that same form suggests that the universe may be encoding thermodynamic structure directly into its expansion history. GRACE+ thus serves as the front-end observational fit—while  $\chi$ CDM provides the physical derivation.

## 6 Conclusion

GRACE+ provides a robust, falsifiable scalar field model that captures late-time cosmological anomalies using a single scalar degree of freedom. It achieves tension reduction without violating early-universe constraints and aligns with a deeper thermodynamic, entropy-based theory of cosmic evolution. Upcoming surveys such as LSST, Euclid, and CMB-S4 can test its distinctive predictions, particularly low- $\ell$  ISW suppression and  $z \approx 0.4$  SNe Ia dimming.

#### References

- [1] Planck Collaboration, Planck 2018 results, A&A 641, A6 (2020)
- [2] Brout et al., ApJ 938, 110 (2022)
- [3] Heymans et al., A&A 646, A140 (2021)
- [4] Blas et al., JCAP 2011, 034 (2011)
- [5] Whitman, K. (2025b). χCDM: Cosmic Decoherence and the Entropic Origin of Acceleration