

PS3a: N-Body Simulation (static)

In 1687, Sir Isaac Newton formulated the principles governing the motion of two particles under the influence of their mutual gravitational attraction in his famous *Principia Mathematica*. However, Newton was unable to solve the problem for three particles. Indeed, in general, solutions to systems of three or more particles must be approximated via numerical simulations. Your challenge is to write a program to simulate the motion of n particles in the plane, mutually affected by gravitational forces, and animate the results. Such methods are widely used in cosmology, semiconductors, and fluid dynamics to study complex physical systems. Scientists also apply the same techniques to other pairwise interactions including Coulombic, Biot-Savart, and van der Waals.

Average time to complete assignment: ~ 5 hours.

1 Details

For Part A of this assignment, we will create a program that loads and displays a static universe. In Part B, we will add the physics simulation and animate the display.

- Make sure to download the universe specification files and images files from [nbody.zip](#).
- You should build a command-line app which reads the universe file (e.g. [planets.txt](#)) from `stdin`. Name your executable `NBody`, so you would run it with e.g.

```
./NBody < planets.txt
```

The `< planets.txt` construct is known as an *input redirect*.

1.1 Reading in the universe

The input format is a text file that contains the information for a particular universe (in SI units). The first value is an integer n which represents the number of particles. The second value is a real number r which represents the radius of the universe, used to determine the scale of the drawing window. Finally, there are n rows and each row contains 6 values. The first two values are the x and y coordinates of the initial position; the next pair of values are the x and y components of the initial velocity; the fifth value is the mass; the last value is a `string` that is the name of an image used to display the particle. As an example, [planets.txt](#) contains data for our own solar system (up to Mars):

...xpos...	...ypos...	...xvel...	...yvel...	...mass...	filename
1.4960e+11	0.0000e+00	0.0000e+00	2.9800e+04	5.9740e+24	earth.gif
2.2790e+11	0.0000e+00	0.0000e+00	2.4100e+04	6.4190e+23	mars.gif
5.7900e+10	0.0000e+00	0.0000e+00	4.7900e+04	3.3020e+23	mercury.gif
0.0000e+00	0.0000e+00	0.0000e+00	0.0000e+00	1.9890e+30	sun.gif
1.0820e+11	0.0000e+00	0.0000e+00	3.5000e+04	4.8690e+24	venus.gif

You should read in exactly as many rows of body information as are indicated by n , the first value in the file. Further lines are treated as comments and can be used to describe the contents.

The [planets.txt](#) universe file contains the Sun and the first four planets, with the Sun at the center of the universe ($x = 0, y = 0$) and the four planets in order toward the right (as below). Note that the window axes are not aligned with the universe axes so you will have to perform a transform. When this is working, you should be rewarded with:



For this project, you'll implement a **Universe** class which will contain all celestial bodies, and a **CelestialBody** class representing the celestial bodies.

- The **Universe** and **CelestialBody** classes should be part of the NB namespace.
- The class **Universe** should create **CelestialBody**s.
- The **CelestialBody** should have the following features:
 - It must be `sf::Drawable`, with a `protected virtual void` method named `draw` (as required of `sf::Drawable` subclasses).
 - Each instance of the class should contain all properties needed for the simulation (e.g. x and y position, x and y velocity, mass, and image data).
 - It may contain a `sf::Sprite` object (as well as the `sf::Texture` object needed to hold the **Sprite**'s image).
- You **must** overload the input stream operator `>>` for both **Universe** and **CelestialBody** and use it to load parameter data into an object.
- You **must** overload the output stream operator `<<` for both **Universe** and **CelestialBody** and use it to write the state back out in the same format as the input.

You must implement unit tests to verify the correctness of your **Universe** and **CelestialBody** classes. Be sure to try other other input files and make sure they are drawn correctly. In particular, positive Y should be towards the top of the window and negative Y towards the bottom. This is different than the behavior of the SFML library.

2 Smart Pointers

You are required to use smart pointers (from the `<memory>` library) to help manage the lifetime of some of your objects. Without them, some of you may notice that some or all of your planets do not have their images render correctly and instead appear as white squares. This occurs because the `sf::Sprite` class stores a pointer to a `sf::Texture` rather than copying any data (the **Sprite** is lightweight and the **Texture** is heavyweight). If you make a copy of both a **Sprite** and a **Texture**, the **Sprite** will hold a pointer to the *old* **Texture** rather than the new one. If the old **Texture** goes out of scope it doesn't have the correct data to draw and the white squares result. By using smart pointers, you can avoid copying objects and avoiding the problem.

3 Unit Tests

As with previous projects, you are required to implement unit tests which will be used to test both your program and several versions provided by the instructor. The reference implementation implements all of the required functionality for both parts (including the

`step` method from part (b)). Additionally, both `Universe` and `CelestialBody` provide getters (but *not* setters) for some of their fields. The `Universe` class provides an alternate string (filename) constructor and `size` (number of planets) and `radius` getters, and an index operator. The `CelestialBody` class provides `position`, `velocity`, and `mass` getters. In the faulty implementations, none of these additional methods are inherently broken, but may expose or depend on other functionality that is faulty.

Note that while the `position` and `velocity` getters return `Vector2fs`, you should do your work internally with double precision rather than single precision.

4 Extra Credit

You can earn extra credit by drawing a background image. This image should appear *behind* all of the planets. Additionally, you can add sound to your simulation. If you do any of the extra credit work, make sure to describe exactly what you did in [Readme-ps4.md](#).

5 Development Process

There are a lot of parts to this assignment. We'd suggest the following incremental development process:

1. Create a bare-bones version of your `Universe` class.
 - (a) Overload the `>>` operator to read just the number of planets and universe radius.
 - (b) Overload the `<<` operator to print the information back.
 - (c) Verify that both work as intended.
2. Create a bare-bones version of your `CelestialBody` class.
 - (a) Overload the `>>` operator to read the planet data.
 - (b) Overload the `<<` operator to print the information back.
 - (c) Update the `Universe`'s `>>` and `<<` operators to read and write a `CelestialBody`.
3. Setup the main draw loop in your main file.
 - Hint: Your class will need to know both the universe and viewport window dimensions.
 - Hint: The universe's center is (0, 0), but for SFML, (0, 0) is in the upper-left corner.
4. Implement the `draw` method in the `CelestialBody` class. Draw this in your loop.
5. Move control of the `CelestialBody` into the `Universe`. Implement the `draw` method in your `Universe` class.
6. Read the rest of the planets into your `Universe`.

6 What to turn in

Your makefile should build a program named `NBody` and a static library named `NBody.a`.

Submit a zip archive to Blackboard containing:

- Your main file `main.cpp`.
- Your Universe (`Universe.cpp`, `Universe.hpp`) and CelestialBody (`CelestialBody.cpp`, `CelestialBody.hpp`) classes.
- Your unit tests (`test.cpp`).
- The makefile for your project. The makefile should have targets `all`, `NBody`, `NBody.a`, `lint` and `clean`. Make sure that all prerequisites are correct.

- Your [Readme-ps4.md](#) that includes
 1. Your name
 2. Statement of functionality of your program (e.g. fully works, partial functionality, extra credit)
 3. Key features or algorithms used
 4. Any other notes
- Any other source files that you created.
- Any images or other resources not provided (such as for the extra credit). The grader will provide images in the *base* directory, not in a sub-directory.
- A screenshot of program output

Make sure that all of your files are in a directory named [ps4](#) before archiving it and that there are no [.o](#) or other compiled files in it.

7 Grading rubric

Feature	Points	Comment
Unit Tests	3	
	1	<< operator formatting
	2	>> operator reads correctly
Autograder	25	Full & Correct Implementation
	4	Contains all code files and builds
	4	The >> and << operators read and print the number of planets
	4	The >> and << operators read and print the universe size
	4	The >> and << operators read and print the first planet correctly
	8	The >> and << operators read and print the other planets correctly
	1	Passes your own tests.
Drawing	11	
	1	Universe and CelestialBody are Drawable with a protected draw()
	5	Draws the universe correctly
	3	Scales properly to universe size
	2	Uses smart pointers and describes in readme
Readme	5	Complete
	2.5	Describes the algorithms or data structures used.
	2.5	Describes how the features were implemented.
	1	Screenshot
Extra Credit	3	
	+2	Shows background image
	+2	Plays sound or music
Penalties		
	-5	Linting problems
	-3	Non-private fields
	-10%	Each day late
Total	45	