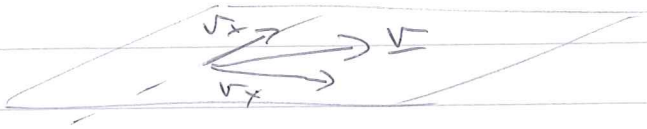


# Problem Set 3.

Note questions not in order on sheet.

1

9. The kinetic energy for a molecule confined to 2-D is



$$E = \frac{1}{2} m v_x^2 + \frac{1}{2} m v_y^2$$

In thermal equilibrium at a temperature  $T$ , the velocity will have a Boltzmann distribution

$$P(v_x, v_y) dv_x dv_y \propto e^{-\frac{1}{2} m (v_x^2 + v_y^2) / k_B T} dv_x dv_y$$

Probability velocity is  $v_x \rightarrow v_x + dv_x$   
&  $v_y \rightarrow v_y + dv_y$

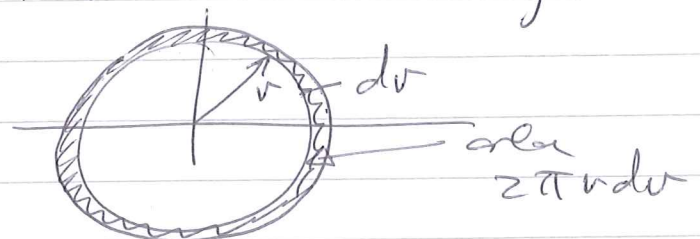
If we instead represent the velocity as a magnitude ( $v$ ) & an angle ( $\theta$ ), then integrate over all  $\theta$

$$\Rightarrow P(v) dv \propto e^{-\frac{1}{2} m v^2 / k_B T} v dv$$

Boltzmann factor

c.f.  $rd\theta$  in polar coordinates

Size of velocity space in  $v \rightarrow v+dv$  range:



2

We can normalise this as follows:

$$\int_0^{\infty} N v e^{-\frac{1}{2} m v^2 / k_B T} dv = 1$$

$$\Rightarrow N = \left( \int_0^{\infty} v e^{-\frac{1}{2} m v^2 / k_B T} dv \right)^{-1}$$
$$= \frac{2}{\sqrt{\pi}} \frac{m}{k_B T}$$

$$P(v) = \left( \frac{m}{k_B T} \right) v e^{-\frac{m v^2}{2 k_B T}}$$

The average kinetic energy is:

$$\textcircled{1} \left\langle \frac{1}{2} m v^2 \right\rangle = \int_0^{\infty} \frac{1}{2} m v^2 P(v) dv = \frac{1}{2} m \cdot \frac{m}{k_B T} \cdot \underbrace{\int_0^{\infty} v^3 e^{-\frac{m v^2}{2 k_B T}} dv}_{\frac{1}{2} \left( \frac{m}{k_B T} \right)^2 \text{ (formula book)}}$$

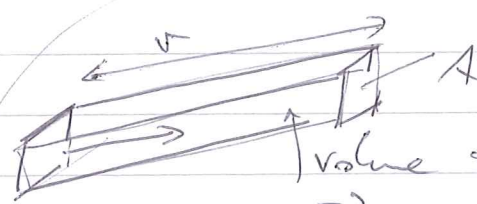
$$= \frac{1}{2} \frac{m^2}{k_B T} \cdot \frac{1}{2} \frac{4 (k_B T)^2}{m^2} = k_B T.$$

$$\textcircled{2} \left\langle \frac{1}{2} m v_x^2 \right\rangle = \int_0^{\infty} \frac{1}{2} m v_x^2 P(v_x) dv_x = \frac{1}{2} k_B T.$$

2x degrees of freedom  $\Rightarrow k_B T$  in total.

10

$$\begin{aligned}
 \int_0^N dN' = N &= \int_0^\infty \frac{1}{4} n v \cdot 4\pi \left( \frac{m}{2\pi k_B T} \right)^{\frac{3}{2}} v^2 e^{-\frac{mv^2}{2k_B T}} dv \\
 &= n\pi \left( \frac{m}{2\pi k_B T} \right)^{\frac{3}{2}} \int_0^\infty v^3 e^{-\frac{mv^2}{2k_B T}} dv \\
 &= n\pi \left( \frac{m}{2\pi k_B T} \right)^{\frac{3}{2}} \cdot \frac{1}{2} \cdot \left( \frac{2k_B T}{m} \right)^2 \\
 &= \frac{n (k_B T)^{\frac{1}{2}}}{m^{\frac{1}{2}} \pi^{\frac{1}{2}} 2^{\frac{1}{2}}} = n \sqrt{\frac{k_B T}{2\pi m}}
 \end{aligned}$$

Note:  $dN = \frac{1}{4} n v P(v) dv$  # per area  
 $\triangleright 4$  fraction  $v \rightarrow dv + v$   
  
 value that can scribe area  
 $\Rightarrow$  A.v. n total # particles  
 only  $\frac{1}{2}$  are to the left &  $\frac{1}{2}$  have +ve velocity

$$\begin{aligned}
 &\int_0^\infty \frac{1}{2} m v^2 \frac{1}{4} n v \cdot 4\pi \left( \frac{m}{2\pi k_B T} \right)^{\frac{3}{2}} v^2 e^{-\frac{mv^2}{2k_B T}} dv \\
 &= \frac{m n \pi}{2} \left( \frac{m}{2\pi k_B T} \right)^{\frac{3}{2}} \int_0^\infty v^5 e^{-\frac{mv^2}{2k_B T}} dv \\
 &= \frac{m n \pi}{2} \cdot \left( \frac{m}{2\pi k_B T} \right)^{\frac{3}{2}} \left( \frac{2k_B T}{m} \right)^3 = \sqrt{\frac{2}{\pi}} n \frac{(k_B T)^{\frac{3}{2}}}{m^{\frac{1}{2}}}
 \end{aligned}$$

Hence  $\left\langle \frac{1}{2} m v^2 \right\rangle = \frac{\sqrt{\frac{2}{\pi}} \cdot n \cdot \frac{(k_B T)^{3/2}}{m^{1/2}}}{n \cdot \sqrt{\frac{k_B T}{2\pi m}}}$

$= 2 k_B T.$

The fast particles are more likely to escape, & hence push up the average kinetic energy of the particles that leave.

1. The temperature of the gas follows from the average kinetic energy:

$\frac{1}{2} eV = \frac{3}{2} k_B T \Rightarrow T \approx 7730 \text{ K}.$   
or  $(k_B T = \frac{2}{3} \text{ eV})$

$P(n=3) = \frac{1}{Z} e^{-\frac{E_3}{k_B T}} \cdot (2 \cdot 3^2)$

$P(n=1) = \frac{1}{Z} e^{-\frac{E_1}{k_B T}} \cdot (2 \cdot 1^2).$

$\Rightarrow \frac{P(n=3)}{P(n=1)} = 9 \cdot e^{-\frac{(E_3 - E_1)}{k_B T}}.$

$= 9 \cdot \exp \left\{ - \frac{13.6 \text{ eV}}{\frac{2}{3} \text{ eV}} \cdot \left( -\frac{1}{3^2} + \frac{1}{1^2} \right) \right\}$

$= 9 \cdot \exp(-18.1) = 1.2 \times 10^{-7}.$

2  $\epsilon$  —————  $P(\epsilon) \propto e^{-\epsilon/k_B T}$

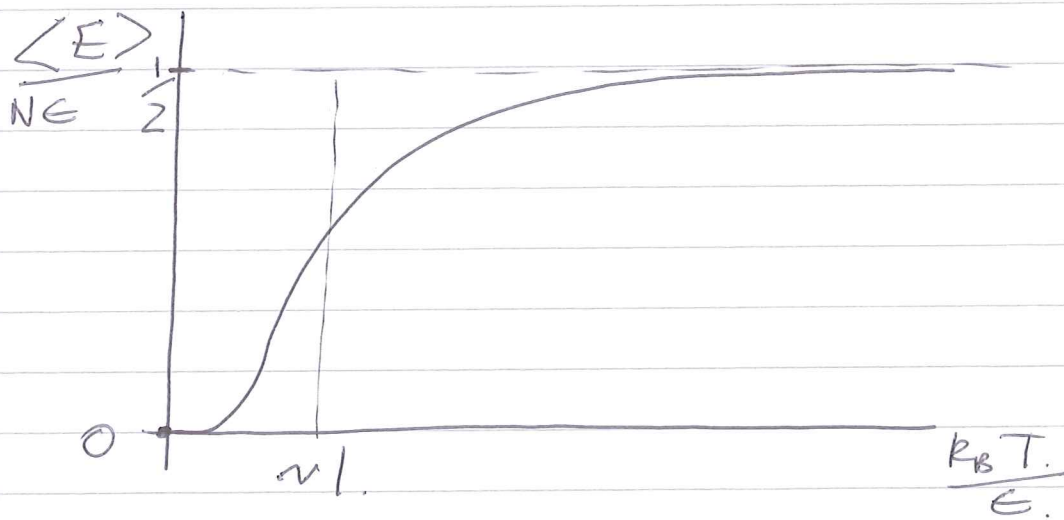
0 —————  $P(0) \propto 1$  (Boltzmann factor)

$$\therefore Z = 1 + e^{-\epsilon/k_B T}$$

$$\& \quad P(0) = \frac{1}{1 + e^{-\epsilon/k_B T}} \quad P(\epsilon) = \frac{1}{1 + e^{-\epsilon/k_B T}} \cdot e^{-\epsilon/k_B T}$$

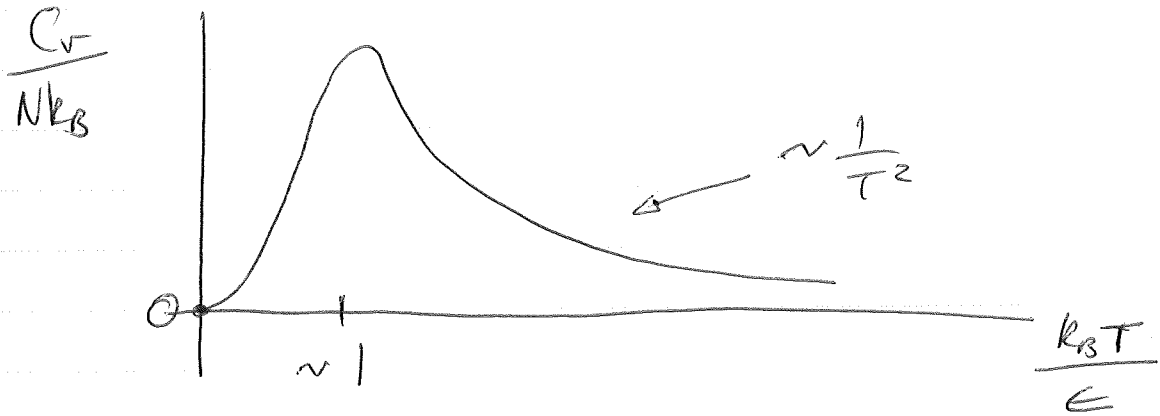
$$a) \quad \langle E \rangle = N(P(0) \cdot 0 + P(\epsilon) \cdot \epsilon)$$

$$= \frac{N \epsilon e^{-\epsilon/k_B T}}{1 + e^{-\epsilon/k_B T}} \equiv \frac{N \epsilon}{e^{\epsilon/k_B T} + 1}$$



$$b) \quad C_v = \left. \frac{\partial \langle E \rangle}{\partial T} \right|_v$$

$$= N \epsilon \cdot \frac{1}{(e^{\epsilon/k_B T} + 1)^2} \cdot \frac{\epsilon}{k_B T^2} = N k_B \left( \frac{\epsilon}{k_B T} \right)^2 \cdot \frac{1}{(e^{\epsilon/k_B T} + 1)^2}$$



Energy — for  $k_B T \ll \epsilon$  there is not enough thermal energy to excite any of the particles  $\Rightarrow \langle E \rangle \approx 0$ . When  $k_B T \sim \epsilon$  the particles have enough thermal energy for some of them to occupy the excited state.  $k_B T \gg \epsilon$  the particles are evenly distributed between 0 &  $\epsilon$  levels.

Heat capacity —

- $k_B T \ll \epsilon$  — all particles in lowest level & can't absorb heat energy
- $k_B T \sim \epsilon$  — particles excited across gap of  $\epsilon \Rightarrow$  large heat capacity (Schottky anomaly)
- $k_B T \gg \epsilon$  — particles evenly distributed & no further excitation possible  $\Rightarrow C_v \rightarrow 0$ .



4.

$$Z_q = c \int_{-\infty}^{\infty} dp \int_{-\infty}^{\infty} dx e^{-\left(\frac{p^2}{2m} + \frac{fx^2}{2}\right) \frac{1}{k_B T}}$$

$$= c \sqrt{\frac{\pi}{\left(\frac{1}{2mk_B T}\right)}} \sqrt{\frac{\pi}{\left(\frac{f}{2k_B T}\right)}}$$

$$= c \pi 2k_B T \sqrt{\frac{m}{f}} \quad \text{NB } \omega^2 = \frac{f}{m}$$

$$= c \pi 2k_B T \frac{1}{\omega}$$

The quantum version (from lectures) is

$$Z_{qm} = \frac{e^{-\frac{\hbar\omega}{2} \frac{1}{k_B T}}}{1 - e^{-\hbar\omega/k_B T}}$$

$Z_{cl}$  &  $Z_{qm}$  are quite different.  $Z_{qm}$  should be similar to  $Z_{cl}$  at high temperature

for large  $T$ :

$$Z_{qm} \approx \left(1 - \frac{\hbar\omega}{2k_B T}\right) \cdot \frac{k_B T}{\hbar\omega}$$

(expanding exponentials)

$$\approx \frac{k_B T}{\hbar\omega}$$

$$\therefore \text{If } 2\pi c = \frac{1}{\hbar} \quad \text{i.e. } c = \frac{1}{h}$$

$\Rightarrow$  the two versions are in agreement at high temperature.

6. If  $C_v = 5k_B$  then the system  
5 has

$$\frac{5k_B}{\frac{1}{2}k_B} = 10 \quad \text{quadratic degrees of freedom}$$

with energy spacing  $\epsilon \approx k_B T$ .

(It may have more than are 'quenched' because the energy levels are very widely spaced).

6. For an oxygen molecule

$$\left\langle \frac{1}{2} m v^2 \right\rangle = \frac{3}{2} k_B T \approx 40 \text{ meV} \quad (T = 300 \text{ K})$$

$$\Rightarrow \langle v^2 \rangle \sim 250 \times 10^3 \text{ m s}^{-1}$$

$$\therefore v \sim 500 \text{ m s}^{-1}$$

If  $T_0 \rightarrow 2T_0$

$$\sqrt{\langle v^2 \rangle} \rightarrow \sqrt{2} \sqrt{\langle v^2 \rangle}$$

3. From ketones

$$P(n=1) = \frac{e^{-\frac{3\hbar\omega}{2k_B T}}}{\left( \frac{e^{-\frac{\hbar\omega}{2k_B T}}}{1 - e^{-\frac{\hbar\omega}{k_B T}}} \right)} \quad \leftarrow \text{Boltzmann factor}$$

$$= e^{-\frac{\hbar\omega}{k_B T}} \cdot (1 - e^{-\frac{\hbar\omega}{k_B T}})$$



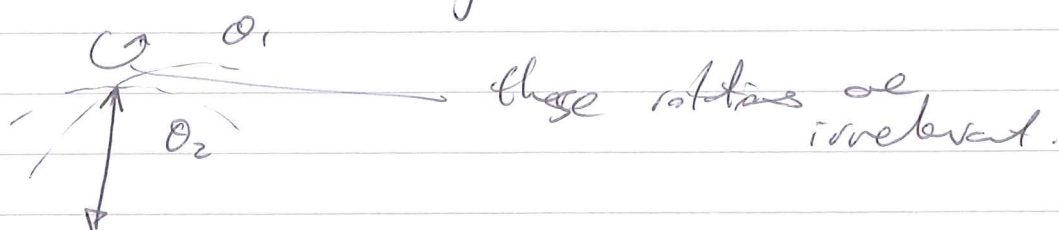
At  $T = 1\text{ K}$

$$\Rightarrow \frac{h\nu}{k_B T} = 1.014$$

$$\therefore P(n=1) \approx e^{-1} (1 - e^{-1}) = 0.231 \approx 23.1\%$$

7. If  $E$  is the energy level spacing for the rotational modes  
 $(E \approx \frac{h^2}{2I})$  in kcal/mol

Then for  $k_B T \gg E$  we can use the equipartition theorem. The diatomic molecule has 2 degrees of freedom



$\Rightarrow$  Heat capacity is  $k_B$  per molecule.

8. By the equipartition theorem

$$\begin{aligned} C_V &\approx \frac{1}{2} k_B \times \# \text{ d.o.f.} \\ &= \frac{1}{2} k_B \times (3 + 3 + 6) \\ &\quad \begin{array}{ccc} \uparrow & \uparrow & \uparrow \\ \text{translation} & \text{rotation} & \text{vibrational modes} \end{array} \\ &= 6k_B \end{aligned}$$

Typically the rotational modes are more widely spaced than translational modes. The vibrational modes are more widely spaced still.

