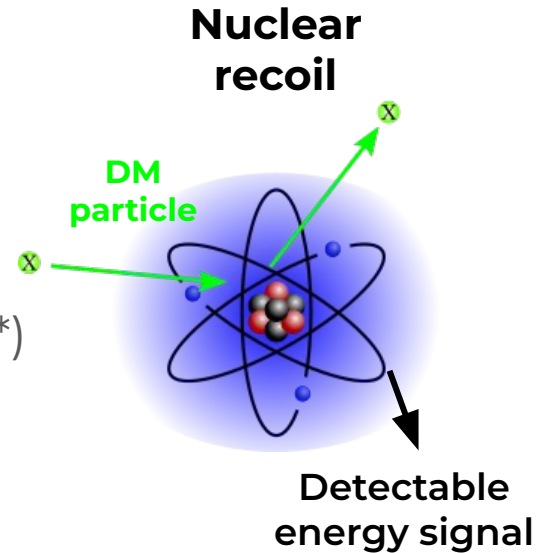


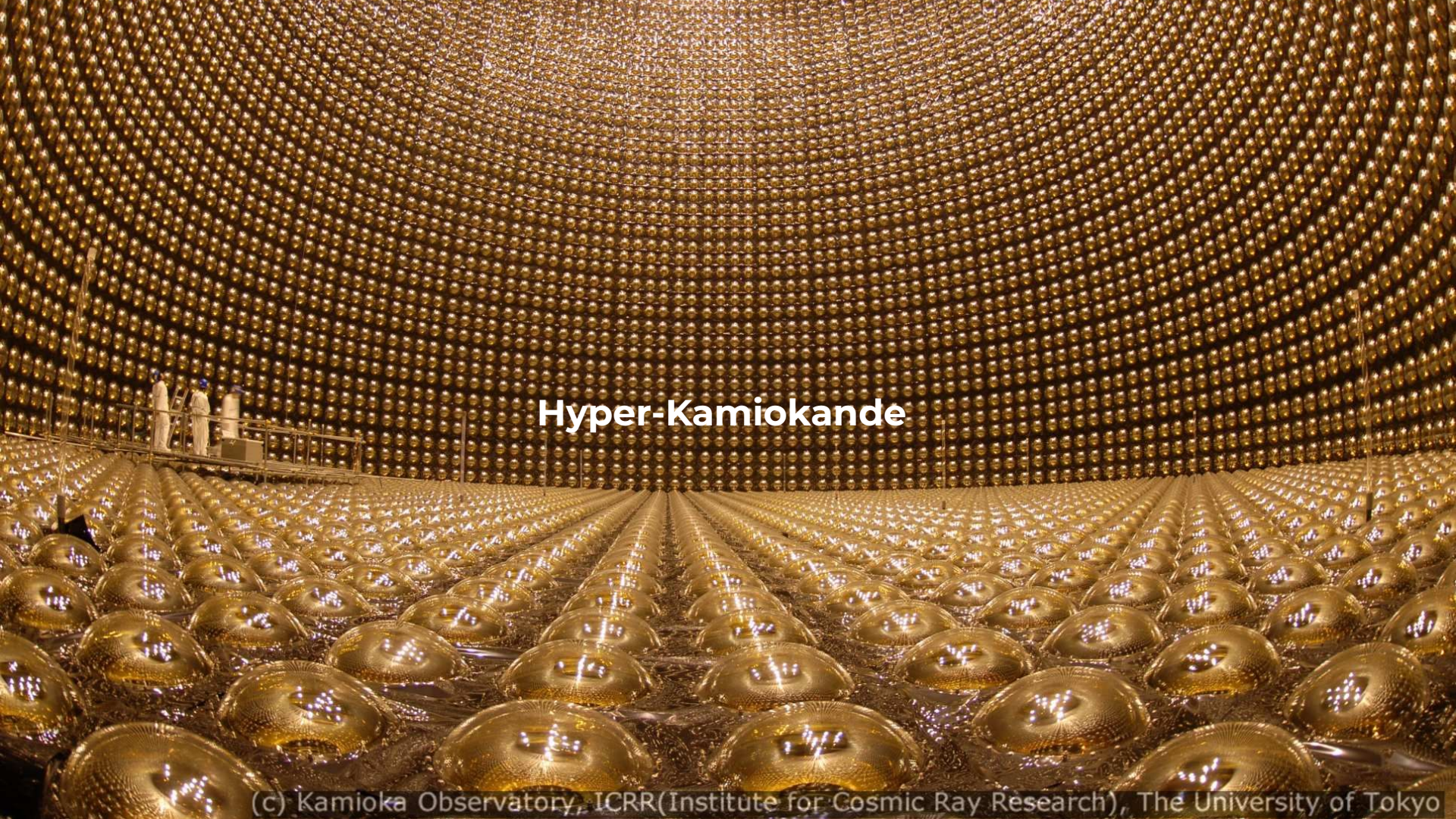
Rare events searches (RES) as counting experiments

Experiments looking for neutrinos (ν) or dark matter (DM) are usually:

- 🏰 Huge (higher mass \rightarrow higher exposure)
- ♟️ Extremely sensitive (ν and DM do not interact a lot*)
- 🏰 Extremely quiet (to measure these feeble signals)

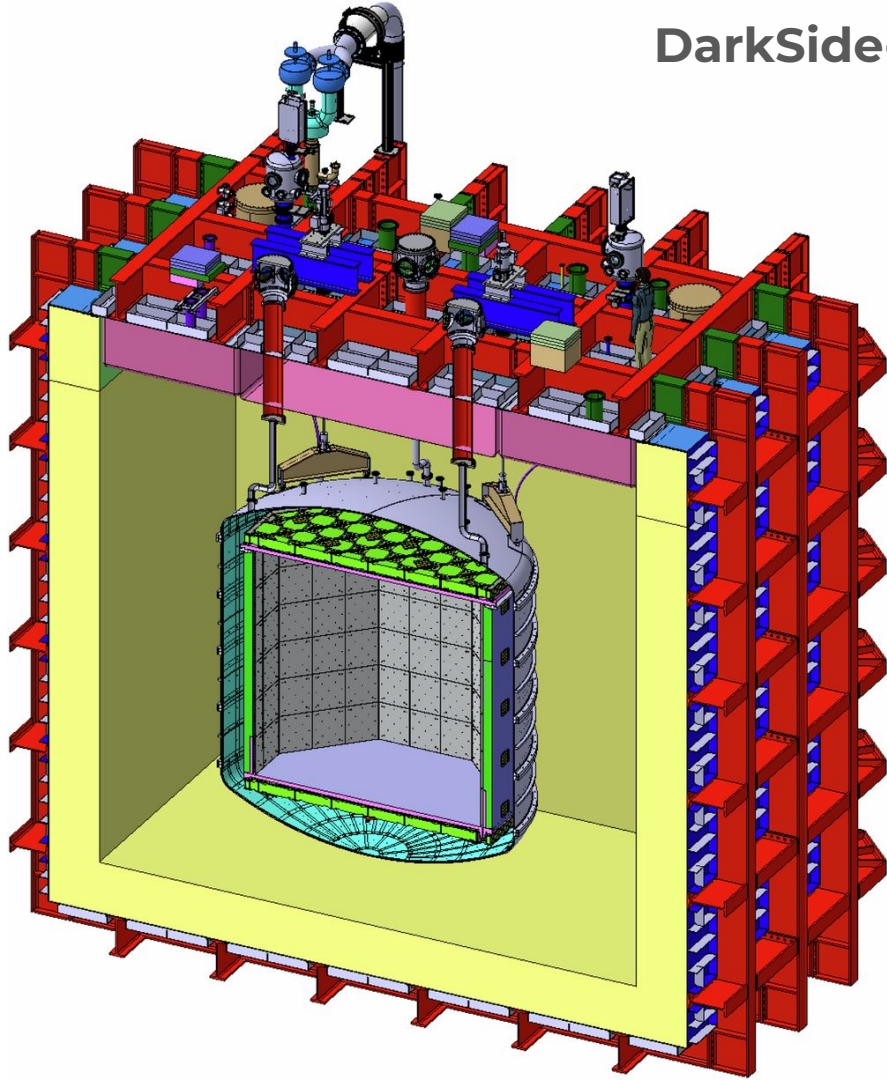


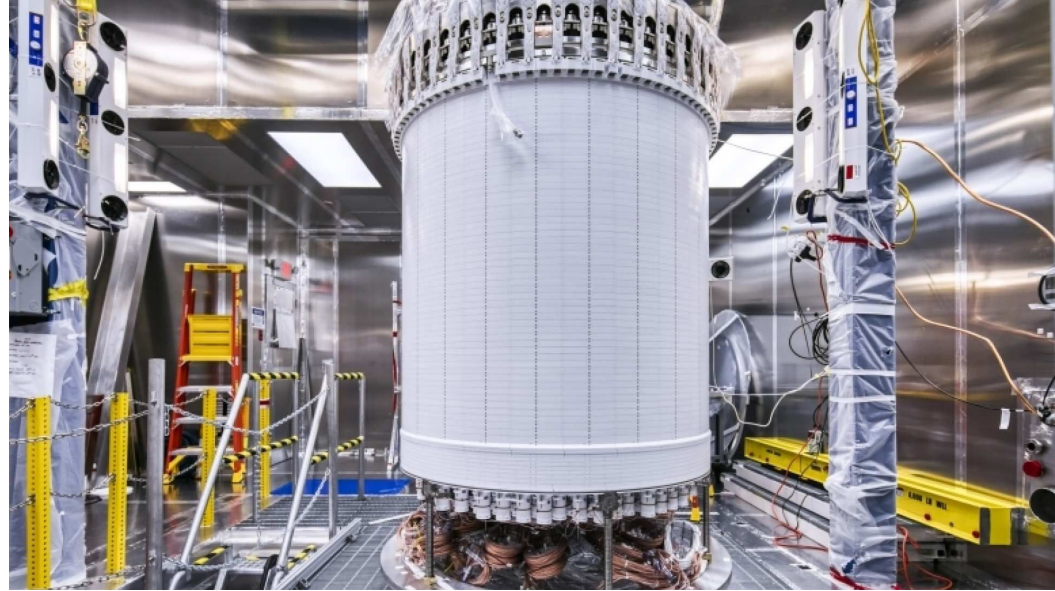
* especially DM...



Hyper-Kamiokande

DarkSide-20k

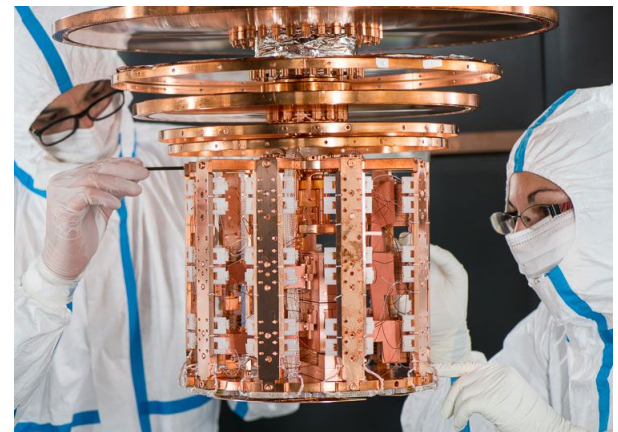




LZ

CRESST

Xenon nT



Some of them measure DM interacting with nuclei...

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Some with electrons, shell electrons, conducting electrons, valence electrons...

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They measure ionization, phonons, light, charge transportation, light propagation, they use photomultipliers, superconductors, semiconductors, non-conductors, gas, solids, fluids, superfluids, sometimes just water, sometimes just plastic, some of them will even go to space...

But they all measure one thing:

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$$P(n|\mu) = (s)^n \exp(-s) / n!$$

Background subtraction

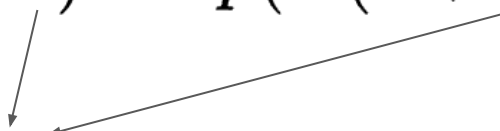
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But what if I know my background? *

Background subtraction

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But what if I know my background? *

$$P(n|\mu) = (s + b)^n \exp(-(s + b))/n!$$


I can subtract it from my signal, and enhance my exclusion limits

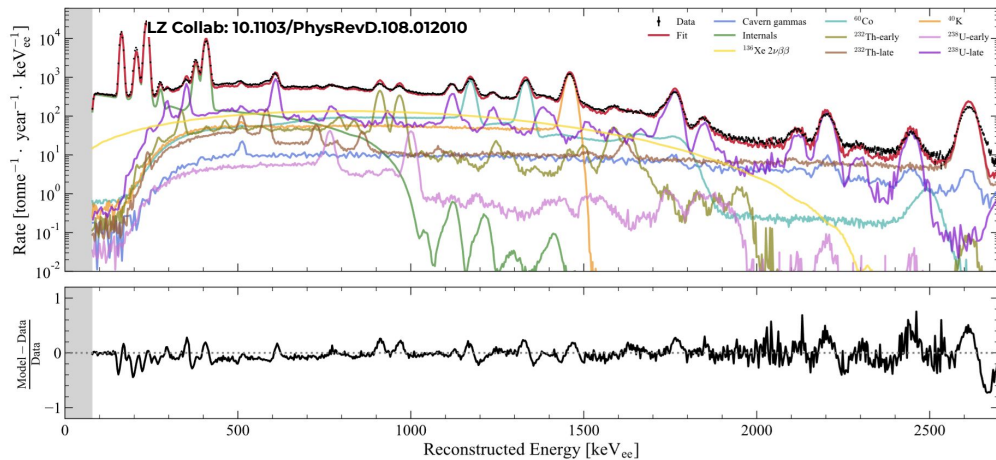
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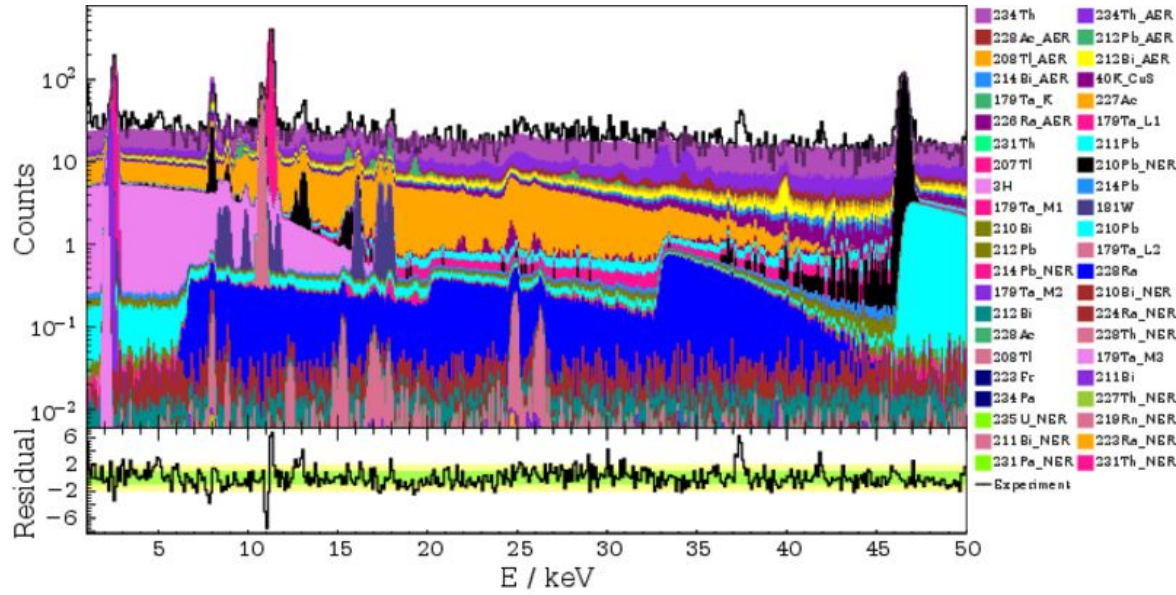
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* See 10.1103/PhysRevD.57.3873

Background subtraction

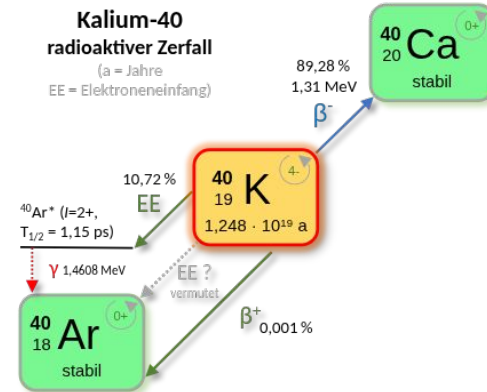
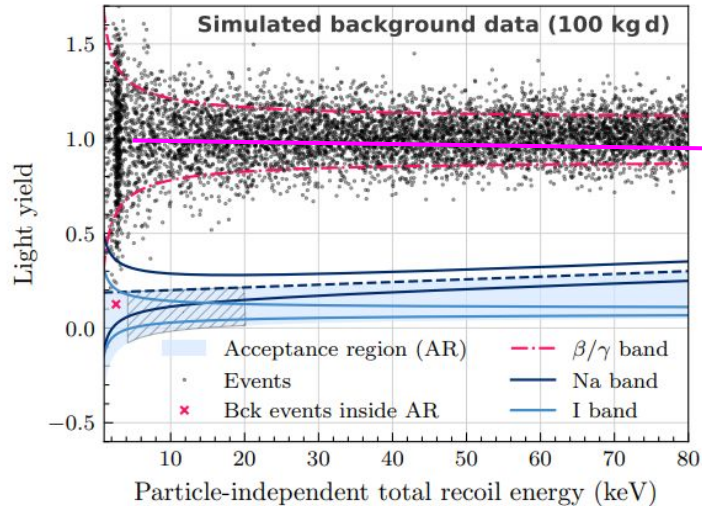
Moreover, this has already been done in our group for CRESST



Background subtraction

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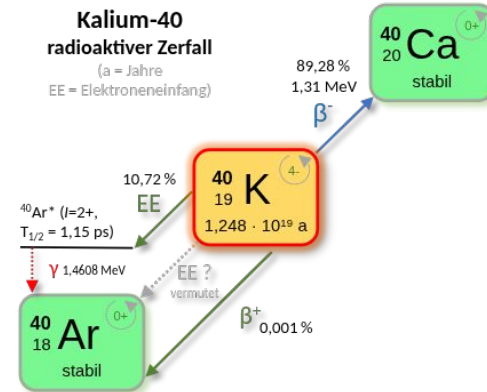
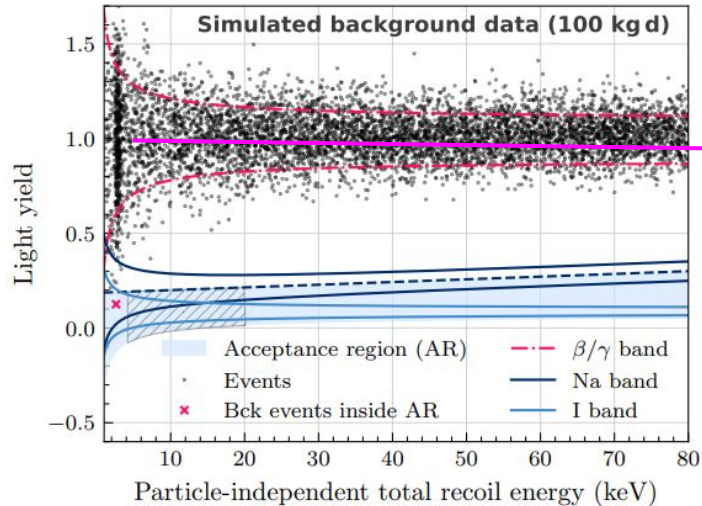
However, COSINUS is special:



Background subtraction

Moreover, this has already been done in our group for CRESST

However, COSINUS is special:



If we can fit our background and know the K-40 activity, we can greatly suppress the effect of the 3 keV line

Going blind

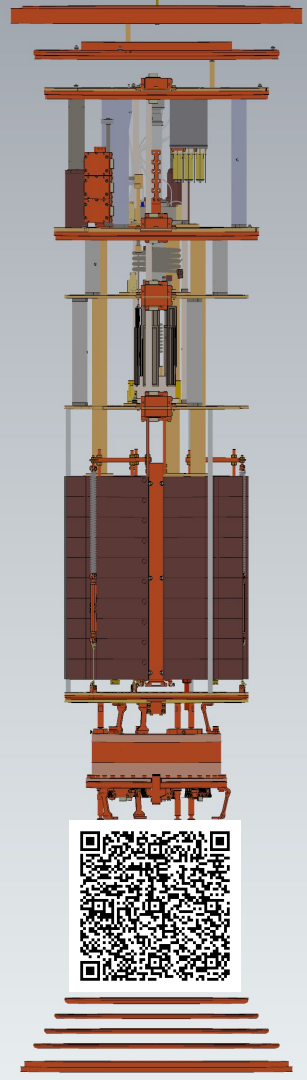
Most DM experiments go blind after some time: they take data for years, without looking at it.

Monte Carlo simulations are then crucial for understanding the backgrounds during this period as they save information of **each event, track by track**

☞ Are K-40 decays by-products expected in our detectors?
Can we use this as a tagging method?

☞ Is surface roughness something we should be aware of?

☞ Can we give analysts information on the expected background?



“A world-leading experiment requires world-leading understanding of the background via high performance simulations” (Us, 16.12.2024)

⇒ ImpCRESST (Geant4-based in-house made soft) allows us to detailedly characterize possible backgrounds that could be highly detrimental to COSINUS goals and *new particles discovery potential*

⇒ Radiogenic backgrounds (nuclear decays), cosmogenic backgrounds (muons and cosmic rays), among others

⇒ Simulations are instrumental to the development of new **cutting-edge** cryogenic-based experiments in **Austria** and **Worldwide**

Projects availability:

👑 Background model (see Fig)

🏠 Non-proportionality in NaI

🔬 Phonon propagation

🌌 Supernovae neutrino directionality capabilities

mailto: mariano.cababie@tuwien.ac.at

