# **SIMPLE Documentation**

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Contents: Useful classes and functions for SIMPLE.

```
simple_lib.impact_parameter(a, e, i, w, r_star)
```

Compute the impact parameter at for a transiting planet.

Parameters a: int, float or numpy array

Semimajor axis of planet's orbit in AU

e: int, float or numpy array

Eccentricity of planet. WARNING! This function breaks down at high eccentricity (>> 0.9), so be careful!

i: int, float or numpy array

Inclination of planet in degrees. 90 degrees is edge-on.

w: int, float or numpy array

Longitude of ascending node defined with respect to sky-plane.

**r\_star**: int, float or numpy array

Radius of star in solar radii.

Returns b: float or numpy array

The impact parameter, ie transit latitude in units of stellar radius.

#### **Notes**

Using Eqn. (7), Chap. 4, Page 56 of Exoplanets, edited by S. Seager. Tucson, AZ: University of Arizona Press, 2011, 526 pp. ISBN 978-0-8165-2945-2.

#### **Examples**

```
>>> impact_parameter(1, 0, 90, 0, 1)
1.3171077641937547e-14
>>> a = np.linspace(.1, 1.5, 3)
>>> e = np.linspace(0, .9, 3)
>>> i = np.linspace(89, 91, 3)
>>> w = np.linspace(0, 360, 3)
>>> r_star = np.linspace(0.1, 10, 3)
>>> impact_parameter(a, e, i, w, r_star)
array([ 3.75401300e+00, 1.66398961e-15, 1.06989371e-01])
```

simple\_lib.inclination(fund\_plane, mutual\_inc, node)

Compute the inclination of a planet.

Uses the law a spherical cosines to compute the sky plane of a orbit given a reference plane inclination, angle from reference plane (ie mutual inclination) and a nodal angle.

#### Parameters fund\_plane: int, float or numpy array :

Inclination of of the fundamental plane of the system in degrees with respect to the sky plane 90 degrees is edge-on.

mutual\_inc : int, float or numpy array

Angle in degrees of the orbital plane of the planet with respect to the fundamental plane of the system.

```
node: int, float or numpy array
```

Rotation in degrees of the planet's orbit about the perpendicular of the reference plane. I.e. the longitude of the node with respect to the reference plane.

#### Returns i: float or numpy array

The inclination of the planet's orbit with respect to the sky plane.

#### **Notes**

See eqn. () in

#### **Examples**

### $\verb|simple_lib.semimajor_axis| (period, mass)$

Compute the semimajor axis of an object.

This is a simple implementation of the general form Kepler's Third law.

**Parameters period**: int, float or numpy array

The orbital period of the orbiting body in units of days.

mass: int, float or array-like

The mass of the central body (or mass sum) in units of solar mass.

Returns a: float or numpy array

The semimajor axis in AU.

#### **Examples**

Parameters p: int, float or numpy array

Period of planet orbit in days

a: int, float or numpy array

Semimajor axis of planet's orbit in AU

e: int, float or numpy array

Eccentricity of planet. WARNING! This function breaks down at high eccentricity (>> 0.9), so be careful!

i: int, float or numpy array

Inclination of planet in degrees. 90 degrees is edge-on.

w: int, float or numpy array

Longitude of ascending node defined with respect to sky-plane.

**b**: int, float or numpy array

Impact parameter of planet.

r\_star: int, float or numpy array

Radius of star in solar radii.

r\_planet : int, float or numpy array

Radius of planet in Earth radii

**Returns** T : float or numpy array

The Q1-Q4 (full) transit duration of the planet in hours.

#### **Notes**

Using Eqns. (15) and (16), Chap. 4, Page 58 of Exoplanets, edited by S. Seager. Tucson, AZ: University of Arizona Press, 2011, 526 pp. ISBN 978-0-8165-2945-2.

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