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In [2]: import numpy as np
import matplotlib.pyplot as plt
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In [4]: # Task 1
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Define the probability density function (PDF) for the nearest star distance.

For a uniform Poisson process in 3D, the probability that no star is found within a sphere of radius R is:

$$P(\text{no stars in sphere}) = \exp\left(-\frac{4}{3}\pi nR^3\right)$$

Therefore, the cumulative probability that the nearest star is within distance R is:

$$ext{CDF}(R) = 1 - \expigg(-rac{4}{3}\pi nR^3igg)$$

Differentiating the CDF with respect to R gives the PDF:

$$ext{PDF}(R) = rac{d}{dR}igg[1-\expigg(-rac{4}{3}\pi nR^3igg)igg] = 4\pi nR^2\expigg(-rac{4}{3}\pi nR^3igg)$$

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In [5]: n = 0.01 # You can adjust this value as needed

def nearest_star_pdf(R, n):
    return 4 * np.pi * n * R**2 * np.exp(- (4/3) * np.pi * n * R**3)

R_values = np.linspace(0, 50, 500) # Adjust the range as needed
pdf_values = nearest_star_pdf(R_values, n)
plt.figure(figsize=(8, 6))
plt.plot(R_values, pdf_values, label=r'$P(R) = 4 \pi n R^2 e^{-\frac{4}{3} \pi c} n R^3}')
plt.xlabel('Distance R')
plt.ylabel('Probability Density')
plt.title('Probability Density Function for the Nearest Star Distance')
plt.legend()
plt.grid(True)
plt.show()
```

Probability Density Function for the Nearest Star Distance $\begin{array}{c|c} 0.40 & & & & & & & & & \\ 0.35 & & & & & & & & & & \\ 0.30 & & & & & & & & & & \\ 0.30 & & & & & & & & & & \\ 0.25 & & & & & & & & & & \\ 0.10 & & & & & & & & & & \\ 0.10 & & & & & & & & & & & \\ 0.05 & & & & & & & & & & & \\ 0.00 & & & & & & & & & & & \\ \end{array}$

Task 2

Consider the driven, damped harmonic oscillator:

10

$$rac{d^2x}{dt^2} + \gammarac{dx}{dt} + \omega_0^2x = Fe^{i\omega_ft}$$

20

30

Distance R

40

50

Taking the Fourier transform yields:

0

$$(-\omega_f^2+i\gamma\omega_f+\omega_0^2) ilde x=F$$

Thus, the Fourier amplitude is:

$$ilde{x}=rac{F}{\omega_0^2-\omega_f^2+i\gamma\omega_f}$$

Defining the susceptibility as

$$\chi(\omega_f) = rac{1}{\omega_0^2 - \omega_f^2 + i \gamma \omega_f},$$

its imaginary part is:

$$ext{Im}[\chi(\omega_f)] = rac{\gamma \omega_f}{(\omega_0^2 - \omega_f^2)^2 + \gamma^2 \omega_f^2}$$

Therefore, the energy absorbed per cycle is given by:

$$E = F\pi \; ext{Im}[\chi(\omega_f)] = rac{F\pi\gamma\omega_f}{(\omega_0^2 - \omega_f^2)^2 + \gamma^2\omega_f^2}$$

In [7]: # Task 3

Question 1: Define the transition matrix P for the 3-spin system. Label the states as follows:

 $1: \mid \uparrow \uparrow \uparrow
angle,$

 $2: |\uparrow\uparrow\downarrow
angle,$

 $3: |\uparrow\downarrow\uparrow\rangle,$

 $4: \mid \uparrow \downarrow \downarrow
angle,$

 $5: |\downarrow\uparrow\uparrow
angle,$

 $6: |\downarrow\uparrow\downarrow\rangle,$

 $7: |\downarrow\downarrow\uparrow\rangle,$

 $8: |\downarrow\downarrow\downarrow\rangle$

Only pairs of neighboring spins that are anti-aligned (i.e. one up and one down) can "flip" via the hopping operators. For example, if sites i and i+1 form the pair (\uparrow,\downarrow) they can flip to (\downarrow,\uparrow) and vice-versa. With periodic boundary conditions the pairs are (1,2), (2,3) and (3,1). Assign equal probability to each allowed move. Then the transition matrix is:

$$P = egin{pmatrix} 1 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \ 0 & 0 & rac{1}{2} & 0 & rac{1}{2} & 0 & 0 & 0 \ 0 & rac{1}{2} & 0 & 0 & rac{1}{2} & 0 & 0 & 0 \ 0 & 0 & 0 & 0 & rac{1}{2} & rac{1}{2} & 0 \ 0 & rac{1}{2} & rac{1}{2} & 0 & 0 & 0 & 0 \ 0 & 0 & 0 & rac{1}{2} & 0 & 0 & 0 & rac{1}{2} & 0 \ 0 & 0 & 0 & rac{1}{2} & 0 & rac{1}{2} & 0 & 0 \ 0 & 0 & 0 & rac{1}{2} & 0 & rac{1}{2} & 0 & 0 \ 0 & 0 & 0 & 0 & 0 & 0 & 0 & 1 \ \end{pmatrix}.$$

Question 2: The stationary distribution π satisfies

$$\pi P=\pi, \quad \sum_{i=1}^8 \pi_i=1.$$

Notice that states 1 and 8 have no allowed moves (they are "frozen"), while the remaining states form two connected blocks:

Block A: states 2, 3, 5

• Block B: states 4, 6, 7

Within each block the transitions are symmetric, so the stationary probabilities in each block are uniform. In other words:

$$\pi_2 = \pi_3 = \pi_5, \qquad \pi_4 = \pi_6 = \pi_7.$$

Alternatively, if energy differences are incorporated (via Boltzmann factors at temperature T), then

$$\pi_i = rac{e^{-E_i/(k_BT)}}{Z}, \qquad Z = \sum_{i=1}^8 e^{-E_i/(k_BT)}.$$

Question 3: Power iteration is given by

$$\pi^{(k+1)} = \pi^{(k)} P.$$

For the following initial guesses:

lf

$$\pi^{(0)} = (1, 0, 0, 0, 0, 0, 0, 0)$$

the chain remains in state 1.

lf

$$\pi^{(0)} = \Big(0, rac{1}{2}, 0, 0, rac{1}{2}, 0, 0, 0\Big),$$

the iteration is confined to Block A and converges to

$$\pi_2=\pi_3=\pi_5=rac{1}{3} \quad ext{(within Block A)}.$$

lf

$$\pi^{(0)} = \left(\frac{1}{8}, \frac{1}{8}, \frac{1}{8}, \frac{1}{8}, \frac{1}{8}, \frac{1}{8}, \frac{1}{8}, \frac{1}{8}, \frac{1}{8}\right)$$

the iteration converges to a mixture of the stationary distributions in each disjoint block.

Question 4: Markov Chain in Magnon Basis

For a chain of N=3 spins, the allowed magnon momenta are quantized as

$$p_k=rac{2\pi k}{3},\quad k=0,1,2,$$

with corresponding energies

$$E_k = 2J\sin^2\!\left(rac{\pi k}{3}
ight).$$

Thus,

$$E_0 = 0, \quad E_1 = 2J\sin^2\Bigl(rac{\pi}{3}\Bigr) = rac{3J}{2}, \quad E_2 = rac{3J}{2}.$$

Assuming Boltzmann-type transitions between magnon states, we take the transition probability from state i to state j as

$$P_{ij} \sim e^{-rac{E_j-E_i}{k_BT}}.$$

Normalizing each row, we define:

For state 0:

$$P_{0j} = rac{e^{-rac{E_{j}-0}{k_{B}T}}}{Z_{0}}, \quad Z_{0} = e^{0} + 2e^{-rac{3J}{2k_{B}T}} = 1 + 2e^{-rac{3J}{2k_{B}T}}.$$

For state 1 (and similarly state 2):

$$P_{1j}=rac{e^{-rac{E_{j}-E_{1}}{k_{B}T}}}{Z_{1}},\quad Z_{1}=e^{rac{3J}{2k_{B}T}}+1+1=e^{rac{3J}{2k_{B}T}}+2.$$

Thus, the transition matrix in the magnon basis is:

$$P = \left(egin{array}{cccc} rac{1}{1+2e^{-rac{3J}{2k_BT}}} & rac{e^{-rac{3J}{2k_BT}}}{1+2e^{-rac{3J}{2k_BT}}} & rac{e^{-rac{3J}{2k_BT}}}{1+2e^{-rac{3J}{2k_BT}}}
ight) \ rac{e^{rac{3J}{2k_BT}}}{1+2e^{-rac{3J}{2k_BT}}} & rac{1}{1+2e^{-rac{3J}{2k_BT}}}
ight) \ rac{e^{rac{3J}{2k_BT}}}{e^{rac{3J}{2k_BT}}+2} & rac{1}{e^{rac{3J}{2k_BT}}+2} & rac{1}{e^{rac{3J}{2k_BT}}+2}
ight) \ .$$

The key difference relative to the site-basis transition matrix is that here the probabilities are weighted by energy differences (via Boltzmann factors), reflecting the thermal population of the magnon states.

Question 5: Stationary Distribution in the Magnon Basis

The stationary distribution π satisfies

$$\pi P=\pi,\quad \sum_{k=0}^2\pi_k=1.$$

Because the system obeys detailed balance, the stationary distribution is given by the Boltzmann distribution:

$$\pi_k = rac{e^{-rac{E_k}{k_BT}}}{\sum_{k'=0}^2 e^{-rac{E_{k'}}{k_BT}}}.$$

For N=3 with $E_0=0$ and $E_1=E_2=rac{3J}{2}$, we have:

$$\pi_0 = rac{1}{1+2e^{-rac{3J}{2k_BT}}}, \quad \pi_1 = \pi_2 = rac{e^{-rac{3J}{2k_BT}}}{1+2e^{-rac{3J}{2k_BT}}}.$$

Question 6: Power Iteration in the Magnon Basis

Power iteration evolves the distribution by

$$\pi^{(k+1)}=\pi^{(k)}P.$$

Consider three initial guesses:

Initial Guess 1: (\pi^{(0)} = (0,1,0)) (i.e. starting in state ($|k=1\rangle$). Iteration will drive the distribution toward the stationary state:

$$\pi_0 = rac{1}{1+2e^{-rac{3J}{2k_BT}}}, \quad \pi_1 = \pi_2 = rac{e^{-rac{3J}{2k_BT}}}{1+2e^{-rac{3J}{2k_BT}}}.$$

Initial Guess 2:

$$\pi^{(0)} = \left(0, rac{1}{2}, rac{1}{2}
ight).$$

The iteration remains in the subspace of states (1) and (2) and converges to the same stationary distribution.

Initial Guess 3:

$$\pi^{(0)} = \left(rac{1}{3}, rac{1}{3}, rac{1}{3}
ight)$$

.

Again, repeated multiplication by (P) leads to convergence to the Boltzmann stationary distribution.

Question 7: Master Equation Evolution

The continuous-time master equation is given by:

$$rac{d\pi_i}{dt} = \sum_{j} \left(Q_{ji}\pi_j - Q_{ij}\pi_i
ight).$$

To relate the discrete-time transition matrix P to the continuous-time transition rate matrix Q, we use:

$$Ppprox e^{Q\Delta t}\quad\Longrightarrow\quad Qpproxrac{1}{\Delta t}{
m ln}(P).$$

Once Q is determined, the master equation in matrix form reads:

$$\frac{d\pi}{dt} = \pi Q,$$

with an initial condition such as $\pi(0)=(0,1,0)$ (corresponding to starting in $|k=1\rangle$). This system can be solved numerically using standard ODE integrators (for example, with Python's scipy.integrate.solve_ivp or odeint).