

We are going to study the orbit of a satellite in Earth's gravitational field. Due to symmetry our problem can be solved in 2 dimensions. We start with a satellite in a geostationary orbit (distance from earth center $r = 42\,157\text{ km}$). The gravitational force is given by

$$F = -G \frac{m m_e}{r^3} \vec{r}$$

with the gravitational constant $G = 6.674\,08 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$, Earth mass $m_e = 5.972 \times 10^{24} \text{ kg}$, mass of the satellite m and the distance vector \vec{r} .

The following questions should be answered:

- the corresponding (linear) system of first order ODEs needs to be determined:

$$\dot{x} = f(t, v)$$

$$\dot{v} = g(t, x)$$

- compare the explicit Euler with the following symplectic Euler method and compare the performance and stability:

$$x_{n+1} = x_n + \Delta t f(t_n, v_n)$$

$$v_{n+1} = v_n + \Delta t g(t_n, x_{n+1})$$

- determine the period time T of the satellite in the geostationary orbit
- assume that in the interval $t = [2T \dots 3T]$ the engine is fired and creates a tangential force $f = 0.01m$
- visualize the stable solution

