

## 1 Chapter 1

- Energy  $E$  of a photon of light in eV:

$$\lambda = \frac{1.24\text{eV}}{E}$$

- Distance  $D$  between adjacent planes in cubic lattices:

$$D = \frac{a}{\sqrt{h^2 + k^2 + l^2}}$$

- Angle between 2 Miller index directions  $A$  and  $B$ :

$$\cos \theta = \frac{A \bullet B}{|A||B|}$$

## 2 Chapter 2

- Planck relationship:

$$E = hv = \left(\frac{h}{2\pi}\right)(2\pi v) = \hbar\omega$$

- Classical energy of a particle:

$$\frac{1}{2}mv^2 = \frac{1}{2} \frac{m^2 v^2}{m} = \frac{\rho^2}{2m}$$

- De Broglie:

$$\lambda = \frac{h}{\rho} = \frac{h}{mv} \Rightarrow \rho = \frac{h}{\lambda} = \frac{h}{2\pi} \frac{2\pi}{\lambda} = \hbar k$$

- Momentum or energy in terms of  $k$  can be derived by combining De Broglie with the classical energy of a particle:

$$E = \hbar\omega = \frac{\rho^2}{2m} = \frac{\hbar^2 k^2}{2m}$$

- Rydberg Constant:

$$R = 109,678\text{cm}^{-1}$$

- Lyman:

$$v = cR\left(\frac{1}{1^2} - \frac{1}{n^2}\right), n = 2, 3, 4, \dots$$

- Balmer:

$$v = cR\left(\frac{1}{2^2} - \frac{1}{n^2}\right), n = 3, 4, 5, \dots$$

- Paschen:

$$v = cR(\frac{1}{3^2} - \frac{1}{n^2}), n = 4, 5, 6, \dots$$

- Postulate for Bohr:

$$\rho_0 = n\hbar, n = 1, 2, 3, 4, \dots$$

- Finding radial forces on orbiting electron:  
(Electrical force toward nucleus) = (Equivalent force in terms of radial acceleration)

$$\begin{aligned} -\frac{q^2}{kr^2} &= -\frac{mv^2}{r} \\ \rho_0 = n\hbar = mvr &\Rightarrow mv^2 = \frac{m^2v^2}{m} = \frac{n^2\hbar^2}{mr^2} \\ \frac{q^2}{kr^2} &= \frac{1}{mr} \frac{n^2\hbar^2}{r^2} \Rightarrow r_n = \frac{kn^2\hbar^2}{mq^2} \end{aligned}$$

$|r_n$  is the radius of the nth orbit

$$-\frac{q^2}{kr^2} = -\frac{mv^2}{r} \Rightarrow \frac{n\hbar}{rm} = \frac{q^2}{kn\hbar}$$

by subbing r from above.

$\Rightarrow$  K.E. of  $e^-$  =

$$\frac{1}{2}mv^2 = \frac{mq^4}{2k^2n^2\hbar^2}$$

P.E. of  $e^-$  =

$$-\frac{q^2}{kr_n} = -\frac{mq^4}{k^2n^2\hbar^2}$$

by subbing r

Total Energy of  $e^-$  =

$$E_n = KE = PE = -\frac{mq^4}{2k^2n^2\hbar^2} = -KE$$

- Energy difference between orbits:

$$E_{n2} - E_{n1} = \frac{mq^4}{2k^2\hbar^2} \left( \frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

- Frequency of light given by a transition between orbits:

$$V_{21} = \left[ \frac{mq^4}{2k^2\hbar^2h} \right] \left( \frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

- Heisenberg uncertainty principle:

$$(\Delta x)(\Delta p_x) \geq \frac{\hbar}{2}$$

$$(\Delta E)(\Delta t) \geq \frac{\hbar}{2}$$

## 2.1 Quantum Mechanics

Classical Variable	→	Quantum Operator
$x$	→	$x$
$f(x)$	→	$f(x)$
$\rho(x)$	→	$\frac{\hbar}{j} \frac{\partial}{\partial x}$
$E$	→	$-\frac{\hbar}{j} \frac{\partial}{\partial t}$

- Normalization of the probability density (the wave function is the probability density):

$$\int_{-\infty}^{\infty} \Psi^* \Psi dx dy dz = 1 \quad (1)$$

- Time averaged expectation of the particle state

$$\langle Q \rangle = \int_{-\infty}^{\infty} \Psi^* Q_{op} \Psi dx dy dz \quad (2)$$

- Classical energy of a particle:

$$KE + PE = E \Rightarrow \frac{1}{2}mv^2 + V = E$$

$$\frac{1}{2}mv^2 = \frac{1}{2} \frac{(mv)^2}{m} = \frac{1}{2} \frac{\rho^2}{m} = \frac{\rho^2}{2m} \Rightarrow \frac{\rho^2}{2m} + V = E \quad (4)$$

$$\rho \rightarrow \frac{\hbar}{j} \frac{\partial}{\partial x}, E \rightarrow -\frac{\hbar}{j} \frac{\partial}{\partial t}$$

$$\Rightarrow \frac{-1}{2m} \hbar \frac{\partial^2}{\partial x^2} \Psi(x, t) + V(x) \Psi(x, t) = \frac{-\hbar}{j} \frac{\partial \Psi(x, t)}{\partial t}$$

where

$$\left(\frac{\partial}{\partial x}\right)^2 \rightarrow \frac{\partial^2}{\partial x^2}, j^2 = -1$$

- Wave function in 3D then:

$$\frac{-\hbar}{2m} \nabla^2 \Psi + V \Psi = \frac{-\hbar}{j} \frac{\partial \Psi}{\partial t} \Rightarrow \nabla^2 \Psi = \frac{\partial^2 \Psi}{\partial x^2} + \frac{\partial^2 \Psi}{\partial y^2} + \frac{\partial^2 \Psi}{\partial z^2} \quad (5)$$

- Separation of variables:

$$\frac{-\hbar}{2m} \frac{\partial^2 \Psi(x, t)}{\partial x^2} + V(x) \Psi(x, t) = \frac{-\hbar}{j} \frac{\partial \Psi(x, t)}{\partial t} \quad (6)$$

$\Rightarrow$

$$-\frac{\hbar}{2m} \frac{\partial^2 \psi(x)}{\partial x^2} \phi(t) + V(x) \psi(x) \phi(t) = -\frac{\hbar}{j} \psi(x) \frac{\partial \phi}{\partial t} \quad (7)$$

$\Rightarrow$

$$\frac{d\phi(t)}{dt} + \frac{j}{\hbar} E \phi(t) = 0 \quad (8)$$

(time dependent portion)

$$-\frac{\hbar}{2m} \frac{d^2 \psi(x)}{dx^2} + V(x) \psi(x) = E \psi(x) \quad (9)$$

(time independent portion)

- $E \equiv$  equivalent constant, corresponds to total energy of the particle
- Wave function as linear combination of various eigenfunctions

$$\psi(x, t) = \sum_n C_n \Psi_n e^{-j \frac{E_n}{\hbar} t} \ni E_n \equiv nth \text{ prefactor} \quad (10)$$

- Infinite potential well

$$V(x) = \begin{cases} 0 & , x \neq 0 \text{ and } x \neq L \\ \infty & , x = 0 \text{ or } x = L \end{cases}$$

$\Rightarrow$

$$\left( \frac{\hbar^2}{2m} \frac{d^2}{dx^2} + V(x) \right) \psi(x) = E \psi(x) \Rightarrow \frac{d^2 \psi(x)}{dx^2} = \frac{2m}{\hbar^2} E \psi(x) \quad (11)$$

$\Rightarrow$

$$\frac{d^2 \psi(x)}{dx^2} + \frac{2m}{\hbar^2} E \psi(x) = 0 \quad (12)$$

$\Rightarrow$

$$\psi(x) = \{ \sin(kx), \cos(kx) \} \ni k = \frac{\sqrt{2mE}}{\hbar} \quad (13)$$

### 3 Chapter 3

- Equilibrium number of EHP's in pure Si at room temp:

$$10^{10} \frac{EHP}{cm^3}$$

- Si atom density in pure Si at room temp:

$$5 * 10^{22} \frac{atoms}{cm^3}$$