## 1 Chapter 1

• Energy E of a photon of light in eV:

$$\lambda = \frac{1.24eV}{E}$$

• Distance D between adjacent planes in cubic lattices:

$$D = \frac{a}{\sqrt{h^2 + k^2 + l^2}}$$

• Angle between 2 Miller index directions A and B:

$$\cos\theta = \frac{A \bullet B}{|A||B|}$$

## 2 Chapter 2

• Planck relationship:

$$E = hv = (\frac{h}{2\pi})(2\pi v) = \hbar\omega$$

• Classical energy of a particle:

$$\frac{1}{2}mv^2 = \frac{1}{2}\frac{m^2v^2}{m} = \frac{\rho^2}{2m}$$

• De Broglie:

$$\lambda = \frac{h}{\rho} = \frac{h}{mv} \Rightarrow \rho = \frac{h}{\lambda} = \frac{h}{2\pi} \frac{2\pi}{\lambda} = \hbar k$$

• Momentum or energy in terms of k can be derived by combining De Broglie with the classical energy of a particle:

$$E = \hbar\omega = \frac{\rho^2}{2m} = \frac{\hbar^2 k^2}{2m}$$

• Rydberg Constant:

$$R = 109,678cm^{-1}$$

• Lyman:

$$v = cR(\frac{1}{1^2} - \frac{1}{n^2}), n = 2, 3, 4, \dots$$

• Balmer:

$$v = cR(\frac{1}{2^2} - \frac{1}{n^2}), n = 3, 4, 5...$$

• Paschen:

$$v = cR(\frac{1}{3^2} - \frac{1}{n^2}), n = 4, 5, 6, \dots$$

• Postulate for Bohr:

$$\rho_0 = n\hbar, n = 1, 2, 3, 4, \dots$$

• Finding radial forces on orbiting electron: (Electrical force toward nucleus) = (Equivalent force in terms of radial acceleration)

$$-\frac{q^2}{kr^2} = -\frac{mv^2}{r}$$

$$\rho_0 = n\hbar = mvr \Rightarrow mv^2 = \frac{m^2v^2}{m} = \frac{n^2\hbar^2}{mr^2}$$

$$\frac{q^2}{kr^2} = \frac{1}{mr} \frac{n^2\hbar^2}{r^2} \Rightarrow r_n = \frac{kn^2\hbar^2}{mq^2}$$

 $|r_n|$  is the radius of the nth orbit

$$-\frac{q^2}{kr^2} = -\frac{mv^2}{r} \Rightarrow \frac{n\hbar}{rm} = \frac{q^2}{kn\hbar}$$

by subbing r from above.

$$\Rightarrow$$
 K.E. of  $e^-=$  
$$\frac{1}{2}mv^2=\frac{mq^4}{2k^2n^2\hbar^2}$$

P.E. of 
$$e^-=$$
 
$$-\frac{q^2}{kr_n}=-\frac{mq^4}{k^2n^2\hbar^2}$$

by subbing r

Total Energy of  $e^- =$ 

$$E_n = KE = PE = -\frac{mq^4}{2k^2n^2\hbar^2} = -KE$$

• Energy difference between orbits:

$$E_{n2} - E_{n1} = \frac{mq^4}{2k^2\hbar^2} \left(\frac{1}{n_1^2} - \frac{1}{n_2^2}\right)$$

• Frequency of light given by a transition between orbits:

$$V_{21} = \left[\frac{mq^4}{2k^2\hbar^2h}\right] \left(\frac{1}{n_1^2} - \frac{1}{n_2^2}\right)$$

• Heisenberg uncertainty principle:  $(\Delta x)(\Delta \rho_x) \geq \frac{\hbar}{2} \\ (\Delta E)(\Delta t) \geq \frac{\hbar}{2}$ 

$$(\Delta x)(\Delta \rho_x) \ge \frac{\hbar}{2}$$
$$(\Delta E)(\Delta t) \ge \frac{\hbar}{2}$$

## Quantum Mechanics 2.1

Classical Variable	$\rightarrow$	Quantum Operator
$\overline{x}$	$\rightarrow$	x
f(x)	$\longrightarrow$	f(x)
$\rho(x)$	$\longrightarrow$	$\frac{\hbar}{i}\frac{\partial}{\partial x}$
E	$\rightarrow$	$-rac{\hbar}{j}rac{\ddot{\partial}}{\partial t}$

## Chapter 3 3

 $\bullet\,$  Equilibrium number of EHP's in pure Si at room temp:

$$10^{10}\frac{EHP}{cm^3}$$

• Si atom density in pure Si at room temp:

$$5*10^{22}\frac{atoms}{cm^3}$$