

Lens Design I

Lecture 7: Aberrations II

2024-05-30

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Preliminary Schedule - Lens Design I 2024

04.04.	Basics	Zhang	Introduction, Zemax interface, menues, file handling, preferences, Editors, updates, windows, coordinates, System description, 3D geometry, aperture, field, wavelength
18.04.	Properties of optical systems I	Tang	Diameters, stop and pupil, vignetting, layouts, materials, glass catalogs, raytrace, ray fans and sampling, footprints
25.04.	Properties of optical systems II	Tang	Types of surfaces, cardinal elements, lens properties, Imaging, magnification, paraxial approximation and modelling, telecentricity, infinity object distance and afocal image, local/global coordinates
02.05.	Properties of optical systems III	Tang	Component reversal, system insertion, scaling of systems, aspheres, gratings and diffractive surfaces, gradient media, solves
16.05.	Advanced handling I	Tang	Miscellaneous, fold mirror, universal plot, slider, multiconfiguration, lens catalogs
23.05.	Aberrations I	Zhang	Representation of geometrical aberrations, spot diagram, transverse aberration diagrams, aberration expansions, primary aberrations
30.05.	Aberrations II	Zhang	Wave aberrations, Zernike polynomials, measurement of quality
06.06.	Aberrations III	Tang	Point spread function, optical transfer function
13.06.	Optimization I	Tang	Principles of nonlinear optimization, optimization in optical design, general process, optimization in Zemax
20.06.	Optimization II	Zhang	Initial systems, special issues, sensitivity of variables in optical systems, global optimization methods
27.06.	Correction I	Zhang	Symmetry principle, lens bending, correcting spherical aberration, coma, astigmatism, field curvature, chromatical correction
04.07.	Correction II	Zhang	Field lenses, stop position influence, retrofocus and telephoto setup, aspheres and higher orders, freeform systems, miscellaneous
	18.04. 25.04. 02.05. 16.05. 23.05. 30.05. 06.06. 13.06. 20.06.	18.04. systems I 25.04. Properties of optical systems II 02.05. Properties of optical systems III 16.05. Advanced handling I 23.05. Aberrations I 30.05. Aberrations II 06.06. Aberrations III 13.06. Optimization I 20.06. Optimization II	18.04. Properties of optical systems I Tang 25.04. Properties of optical systems II Tang 02.05. Properties of optical systems III Tang 16.05. Advanced handling I Tang 23.05. Aberrations I Zhang 30.05. Aberrations II Tang 13.06. Optimization I Tang 20.06. Optimization II Zhang 27.06. Correction I Zhang

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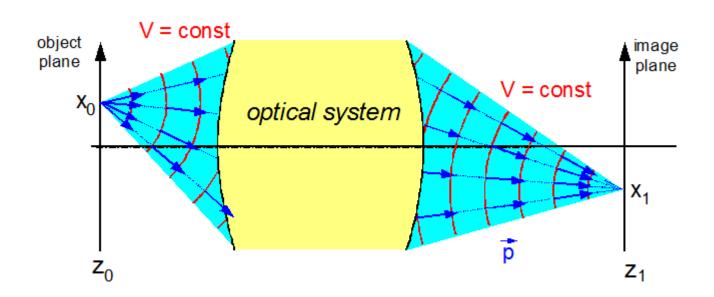


- 1. Optical path difference
- 2. Definition of wave aberrations
- 3. Zernike polynomials
- 4. Measurement of wave aberrations

Rays and Wavefronts



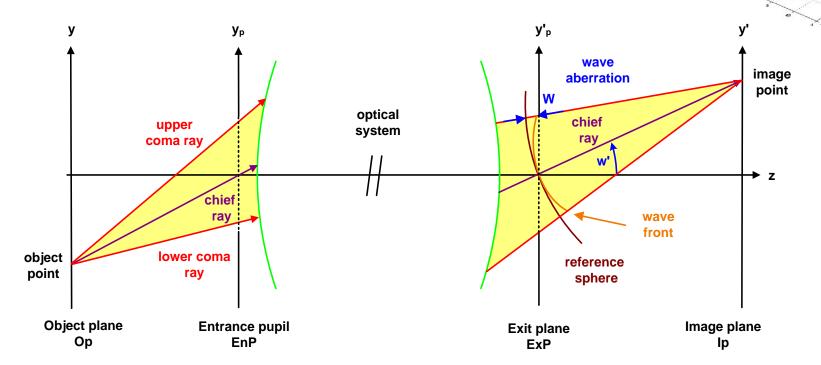
- Rays and Wavefront forms an orthotomic system
- Any closed path integral has zero value
- Corresponds to law of Malus and Fermats principle



Wave Aberration in Optical Systems



- Definition of optical path length in an optical system:
 Reference sphere around the ideal object point through the center of the pupil
- Chief ray serves as referenceDifference of OPL : optical path difference OPD
- Practical calculation: discrete sampling of the pupil area,
 real wave surface represented as matrix



Relationships



- Concrete calculation of wave aberration: addition of discrete optical path lengths (OPL)
- Reference on chief ray and reference sphere (optical path difference)
- Relation to transverse aberrations

- Conversion between longitudinal transverse and wave aberrations
- Scaling of the phase / wave aberration:
 - 1. Phase angle in radiant
 - 2. Light path (OPL) in mm
 - 3. Light path scaled in λ

$$l_{OPL} = \int_{OE}^{AP} n \cdot d\vec{r}$$

$$\Delta_{OPD}(x, y) = l_{OPL}(x, y) - l_{OPL}(0, 0)$$

$$\frac{\partial W}{\partial y_p} = -\frac{\Delta y'}{R - W} \approx -\frac{\Delta y'}{R}$$

$$\Delta s' = \frac{R}{y_p} \cdot \Delta y' = \frac{\Delta y'}{\sin u'} = -\frac{R^2}{y_p} \cdot \frac{\partial W(x_p, y_p)}{\partial y_p}$$

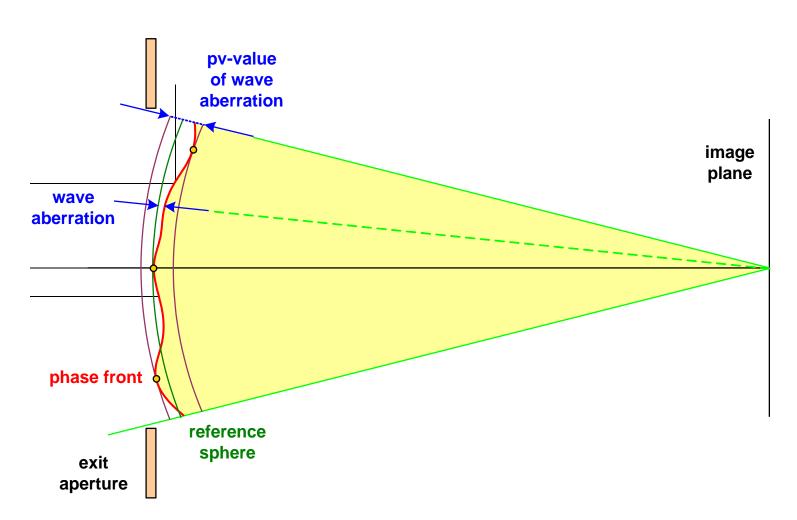
$$E(x) = A(x) \cdot e^{i \cdot \varphi(x)}$$

$$E(x) = A(x) \cdot e^{i \cdot k \Delta_{OPD}(x)}$$

$$E(x) = A(x) \cdot e^{2\pi i \cdot W(x)}$$



Definition of the peak valley value





Mean quadratic wave deviation (W_{Rms} , root mean square)

$$W_{rms} = \sqrt{\left\langle W^2 \right\rangle - \left\langle W \right\rangle^2} = \sqrt{\frac{1}{A_{ExP}}} \iint \left[W \left(x_p, y_p \right) - W_{mean} \left(x_p, y_p \right) \right]^2 dx_p dy_p$$
 with pupil area
$$A_{ExP} = \iint dx dy$$

Peak valley value W_{pv}: largest difference

$$W_{pv} = \max \left[W_{\text{max}} \left(x_p, y_p \right) - W_{\text{min}} \left(x_p, y_p \right) \right]$$

General case with apodization:
 weighting of local phase errors with intensity, relevance for psf formation

$$W_{rms} = \sqrt{\frac{1}{A_{ExP}^{(w)}} \iint I_{ExP}(x_p, y_p) \cdot [W(x_p, y_p) - W_{mean}^{(w)}(x_p, y_p)]^2 dx_p dy_p}$$

Criteria of Rayleigh and Marechal



- Rayleigh criterion:
 - 1. maximum of wave aberration: $W_{pv} < \lambda/4$
 - 2. beginning of destructive interference of partial waves
 - 3. limit for being diffraction limited (definition)
 - 4. as a PV-criterion rather conservative: maximum value only in 1 point of the pupil
 - 5. different limiting values for aberration shapes and definitions (Seidel, Zernike,...)
- Marechal criterion:
 - 1. Rayleigh crierion corresponds to $W_{rms} < \lambda/14$ in case of defocus

$$W_{rms}^{Rayleigh} \le \frac{\lambda}{\sqrt{192}} = \frac{\lambda}{13.856} \approx \frac{\lambda}{14}$$

- 2. generalization of $W_{rms} < \lambda/14$ for all shapes of wave fronts
- 3. corresponds to Strehl ratio $D_s > 0.80$ (in case of defocus)
- 4. more useful as PV-criterion of Rayleigh

Wave Aberrations



- Wave aberration: relative to reference sphere
- Choice of offset value: vanishing mean

$$\langle W(x,y)\rangle = \frac{1}{F_{ExP}} \iint W(x,y) \, dx \, dy = 0$$

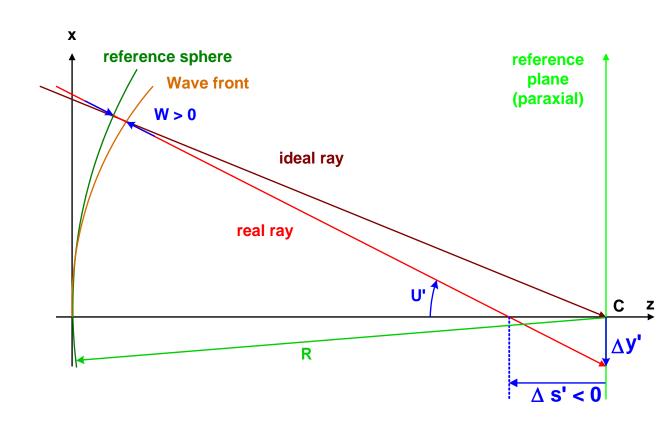
■ Sign of W:

W > 0 : stronger convergence intersection : s < 0

- W < 0 : stronger

divergence

intersection: s < 0

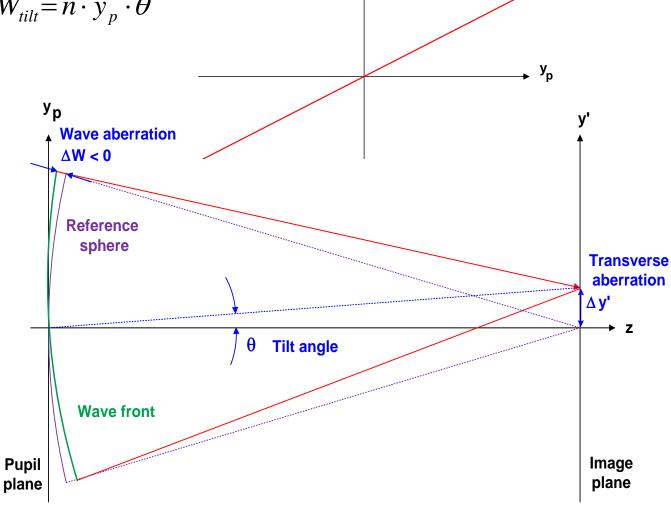


Change of reference sphere: tilt by angle θ linear in y_p

$$\Delta W_{tilt} = n \cdot y_p \cdot \theta$$

Wave aberration due to transverse aberration ∆y'

$$\Delta W_{tilt} = -\frac{y_p}{R_{\text{Re }f}} \cdot \Delta y'$$

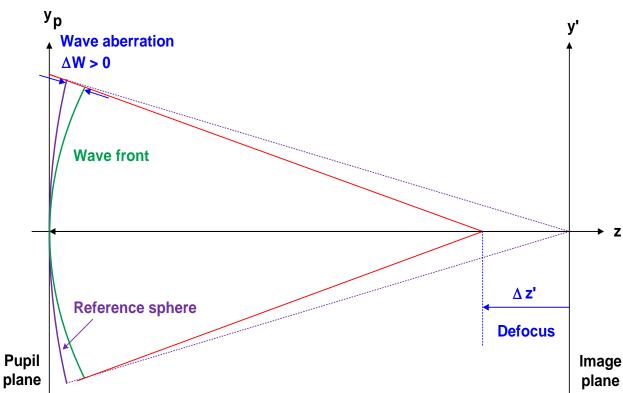


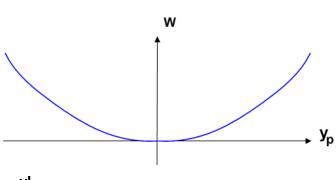
W

Defocussing of Wavefront

Paraxial defocussing by Δz : Change of wavefront

$$\Delta W_{Def} = -\frac{n \cdot r_p^2}{2R_{ref}^2} \cdot \Delta z' = -\frac{1}{2} n \cdot \Delta z' \cdot \sin^2 u$$





Zernike Polynomials

Expansion of the wave aberration on a circular area

$$W(r,\varphi) = \sum_{n} \sum_{m=-n}^{n} c_{nm} Z_{n}^{m}(r,\varphi)$$

$$c_{nm} = \frac{2(n+1)}{\pi(1+\delta_{m0})} \cdot \int_{0}^{1} \int_{0}^{2\pi} W(r,\varphi) Z_{n}^{m*}(r,\varphi) d\varphi r dr$$

Zernike polynomials in cylindrical coordinates:
 Radial function R(r), index n
 Azimuthal function φ, index m

$$Z_n^m(r,\varphi) = R_n^m(r) \cdot \begin{cases} \sin m\varphi \text{ für } m>0\\ \cos m\varphi \text{ für } m<0\\ 1 \qquad \text{ für } m=0 \end{cases}$$

Orthonormality

$$\int_{0}^{1} \int_{0}^{2\pi} Z_{n}^{m}(r,\varphi) Z_{n'}^{m'*}(r,\varphi) d\varphi r dr = \frac{\pi \cdot (1 + \delta_{m0})}{2(n+1)} \cdot \delta_{nn'} \delta_{mm'}$$

- Advantages:
 - 1. Minimal properties due to W_{rms}
 - 2. Decoupling, fast computation
 - 3. Direct relation to primary aberrations for low orders
- Problems:
 - 1. Computation on discrete grids
 - 2. Non circular pupils
 - 3. Different conventions concerning indeces, scaling, coordinate system

Zernike Polynomials: Different Nomenclatures

- 1. Fringe representation
 - CodeV, Zemax, interferometric test of surfaces
 - Standardization of the boundary to ±1
 - no additional prefactors in the polynomial
 - Indexing accordint to m (Azimuth), quadratic number terms have circular symmetry
 - coordinate system invariant in azimuth
- 2. Standard representation
 - CodeV, Zemax, Born / Wolf
 - Standardization of rms-value on ± 1 (with prefactors), easy to calculate Strehl ratio
 - coordinate system invariant in azimuth
- 3. Original Nijboer representation
 - Expansion:

$$W(r,\varphi) = a_{00} + \frac{1}{\sqrt{2}} \sum_{n=0}^{k} a_{0n} R_n^0 + \sum_{n=0}^{k} \sum_{\substack{m=1\\ n-m\\ gerade}}^{n} a_{nm} R_n^m \cos(m\varphi) + \sum_{n=0}^{k} \sum_{\substack{m=1\\ n-m\\ gerade}}^{n} b_{nm} R_n^m \sin(m\varphi)$$

- Standardization of rms-value on ±1
- coordinate system rotates in azimuth according to field point

Zernike Polynomials

Zernike polynomials orders by indices:

n: radial

m: azimuthal, sin/cos

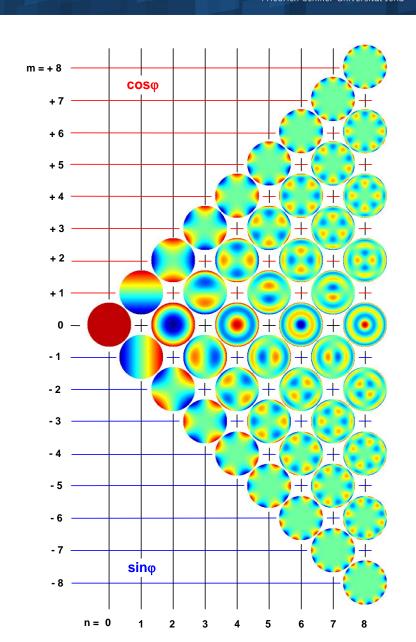
Orthonormal function on unit circle

$$Z_n^m(r,\varphi) = R_n^m(r) \cdot \begin{cases} \sin m\varphi \text{ für } m > 0 \\ \cos m\varphi \text{ für } m < 0 \\ 1 \qquad \text{für } m = 0 \end{cases}$$

Expansion of wave aberration surface

$$W(r,\varphi) = \sum_{n} \sum_{m=-n}^{n} c_{nm} Z_{n}^{m}(r,\varphi)$$

- Direct relation to primary aberration types
- Direct measurement by interferometry
- Orthogonality perturbed:
 - 1. apodization
 - 2. discretization
 - 3. real non-circular boundary



Calculation of Zernike Polynomials

- Assumptions:
 - 1. Pupil circular
- 2. Illumination homogeneous
- 3. Neglectible discretization effects /sampling, boundary)
- Direct computation by double integral:
 - 1. Time consuming
 - 2. Errors due to discrete boundary shape
 - 3. Wrong for non circular areas
 - 4. Independent coefficients
- LSQ-fit computation:
 - 1. Fast, all coefficients c_i simultaneously
- 2. Better total approximation
- 3. Non stable for different numbers of coefficients, if number too low
- Stable for non circular shape of pupil

$$c_{j} = \frac{1}{\pi} \int_{0}^{1} \int_{0}^{2\pi} W(r, \varphi) Z_{j} *(r, \varphi) d\varphi r dr$$

$$\sum_{i=1}^{N} \left[W_i - \sum_{j=1}^{N} c_j Z_j(r_i) \right]^2 = \min$$

$$\vec{c} = \left(\underline{Z}^T \underline{Z}\right)^{-1} \underline{Z}^T \vec{W}$$

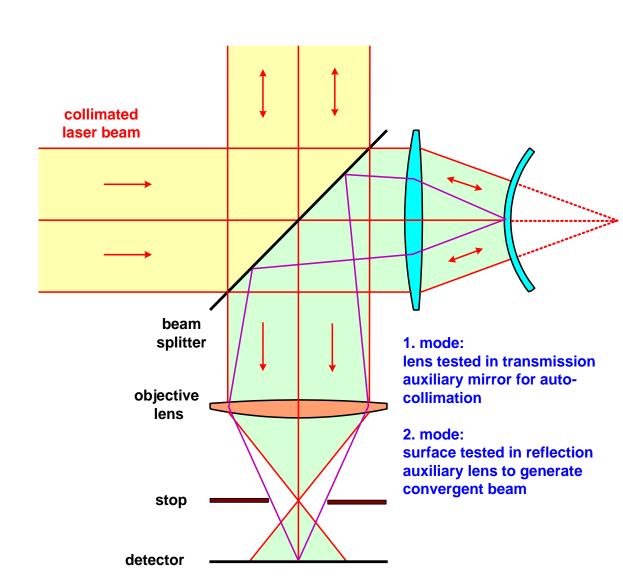
Zernike Polynomials: Explicite Formulas

n	m	Polar coordinates	Cartesian coordinates	Interpretation
0	0	1	1	piston
1	1	$r\sin\varphi$	x	tilt in x
1	-1	$r\cos\varphi$	У	tilt in y
2	2	$r^2 \sin 2\varphi$	2xy	Astigmatism 45°
2	0	$2r^2-1$	$2x^2 + 2y^2 - 1$	defocussing
2	-2	$r^2\cos 2\varphi$	y^2-x^2	Astigmatism 0°
3	3	$r^3 \sin 3\varphi$	$3xy^2 - x^3$	trefoil 30°
3	1	$(3r^3-2r)\sin\varphi$	$3x^3 - 2x + 3xy^2$	coma x
3	-1	$(3r^3-2r)\cos\varphi$	$3y^3 - 2y + 3x^2y$	coma y
3	-3	$r^3\cos 3\varphi$	$y^3 - 3x^2y$	trefoil 0°
4	4	$r^4 \sin 4\varphi$	$4xy^3 - 4x^3y$	Four sheet 22.5°
4	2	$\left(4r^4 - 3r^2\right)\sin 2\varphi$	$8xy^3 + 8x^3y - 6xy$	Secondary astigmatism
4	0	$6r^4 - 6r^2 + 1$	$6x^4 + 6y^4 + 12x^2y^2 - 6x^2 - 6y^2 + 1$	Spherical aberration
4	-2	$(4r^4 - 3r^2)\cos 2\varphi$	$4y^4 - 4x^4 + 3x^2 - 3y^2 - 4x^2y^2$	Secondary astigmatism
4	-4	$r^4\cos 4\varphi$	$y^4 + x^4 - 6x^2y^2$	Four sheet 0°

Testing with Twyman-Green Interferometer

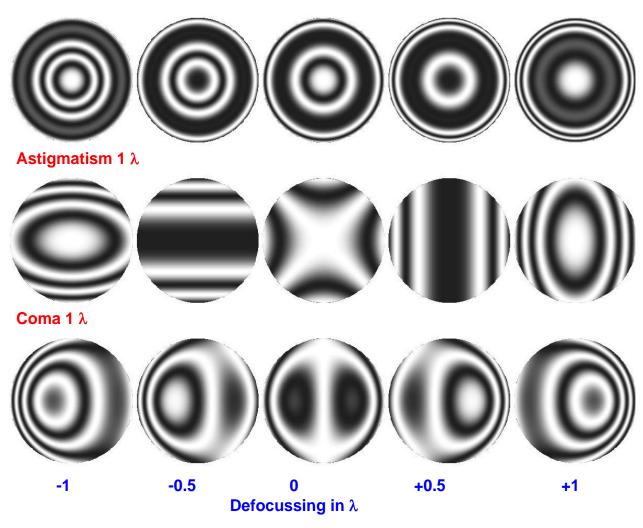


- Short common path, sensible setup
- Two different operation modes for reflection or transmission
- Always factor of 2 between detected wave and component under test



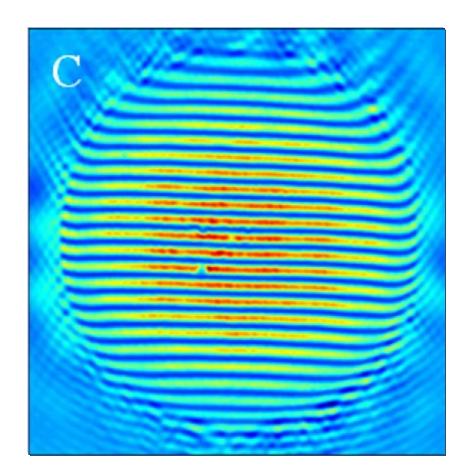
Interferograms of Primary Aberrations

Spherical aberration 1 λ



Interferogram - Definition of Boundary

 Critical definition of the interferogram boundary and the Zernike normalization radius in reality



Exercise I: Wave Aberrations and Zernike Coefficients



Load the system data from Moodle. It is a microscopic lens with high NA = 1.28 from the book of Laikin.

- a) Show the rms wave aberrations as a function of the defocussing. Discuss the results
- b) Show the rms wave aberration as a function of the field for all wavelengths. Is the system diffraction limited?
- c) Calculate the Zernike coefficients for the primary wavelength on axis and for the maximum field size. What kind of aberration limits the performance in the field?
- d) Calculate the Zernikes on axis behind the first three components and in the image. What can be seen for the changes and the compensation effects in the spherical aberration coefficients?

Exercise II: Aplanatic Lens



Consider a collimated incoming beam with wavelength 500 nm and diameter 10 mm. This bundle should be focussed by a perfect lens of focal length f = 50 mm.

- a) Place an aplanatic-concentric lens shortly behind the ideal lens with the material SF57. What is the resulting numerical aperture in the image space? Show at least two different methods to find the best image position.
- b) Show that the spherical aberration of this setup is exactly zero for all orders.
- c) Aplanatic means, that the linear coma vanishes and the imaging is free of coma for a small but finite field size. Show this property by using a small field of 2° for the current system. What is the largest present aberration?

Delano's Representation of Spherical Aberration

- Paraxial optics: Delano relation
- Real ray comparison:
 Delano surface contribution

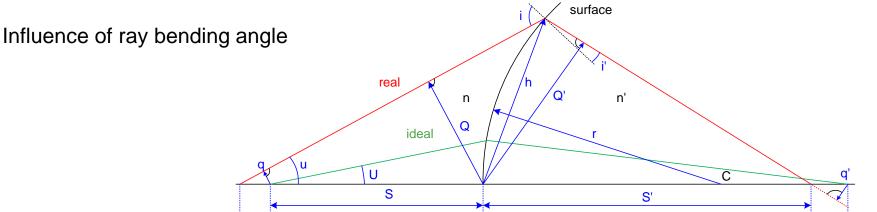
$$n' \cdot q' \cdot U' = n \cdot q \cdot U + n \cdot i \cdot (Q' - Q)$$

$$\Delta s'_{SPH} = \Delta s_{SPH} \cdot \frac{n_1 U_1 \sin u_1}{n'_k U'_k \sin u'_k} + \sum_{j} \frac{(Q - Q') \cdot i \cdot n_j}{n'_j U'_j \sin u'_j}$$

$$\Delta s'_{SPH} = \Delta s_{SPH} \cdot \frac{n_1 U_1 \sin u_1}{n'_k U'_k \sin u'_k} + \sum_j \frac{n_j}{n'_j} \cdot h \cdot \sin \frac{i' - i}{2} \cdot \frac{2i \cdot \sin \frac{i' - u}{2}}{U'_j \sin u'_j}$$

Surface contribution grows with

- 1. ratio of refractive indices
- 2. height of the marginal ray
- 3. Influence of ray bending angle



Aplanatic Surfaces



- Aplanatic surfaces: zero spherical aberration:
 - 1. Ray through vertex

$$s' = s = 0$$

2. concentric

$$s' = s$$
 und $u = u'$

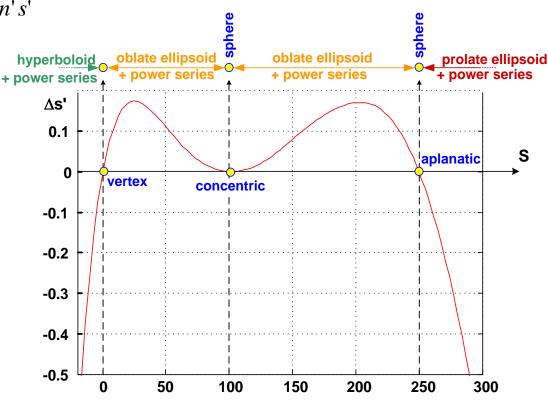
3. Aplanatic

ns = n's'

Condition for aplanatic surface:

$$r = \frac{ns}{n+n'} = \frac{n's'}{n+n'} = \frac{ss'}{s+s'}$$

- Virtual image location
- Applications:
 - 1. Microscopic objective lens
 - 2. Interferometer objective lens



Aplanatic Lenses



- Aplanatic lenses
- Combination of one concentric and one aplanatic surface: zero contribution of the whole lens to spherical aberration
- Not useful:
 - 1. aplanatic-aplanatic
 - 2. concentric-concentric bended plane parallel plate, nearly vanishing effect on rays

