R= 384 × 103 km = 3,8 × 108 M

X= 200 = 2.4x109 m

What is momentum?

Orbit is ~ 27 days = T

V= 27R = WR

= 2,4×109m

27 days x 24k x 60 mm x 60s

= 1029m/s

Mmoon = 7.34 ×1022 Kg

 $\lambda = \frac{h}{p} = \frac{6.63 \times 10^{-34} \text{ J-s}}{7.3 \times 10^{22} \text{ kg} \cdot 10^{3} \text{ m/s}}$

= (0 m

× ~ 1069 Very classical

$$\frac{1}{2}mv^2 = 13.6eV = \frac{p^2}{2m}$$

$$\lambda = \frac{h}{p} = \frac{2\pi 6.58 \times 10^{-16} \text{ eV} = 5}{3700 \text{ eV}}$$

$$X = 2\pi \times 0_0 \approx 3\mathring{A}$$

c) The de Broghe wovelongth is about the same as the hydrogen case above since KE 13, the some and

X 15 also close.

* ~ 1 Quanting.

d) Now the particle is relativistic

$$P = \frac{1}{c} = \frac{1}{6eV}$$

$$\lambda = \frac{1}{P} = \frac{2\pi \times 6.88 \times 10^{-16} \text{ eV-S}}{10^{9} \text{ eV-S}}$$

$$= \frac{1}{10^{9}} \times \frac{3 \times 10^{8} \text{ mg}}{3 \times 10^{8} \text{ mg}}$$

$$= \frac{1}{2} \times 10^{-15} \text{ m}$$

$$= \frac{1}{2} \times 10^{5} \Rightarrow \text{Chossical}$$
e) Moeutron = 940 MeV/c²

$$P = \sqrt{2} \times 940 \text{ MeV/c}^{2} \times 4 \text{ MeV}$$

$$= 87 \text{ MeV}$$

$$= 87 \text{ MeV}$$

$$\Rightarrow 14 \text{ fm}$$

$$\Rightarrow \boxed{X} \times \frac{1}{2} \Rightarrow \text{Quantum}$$

$$\Rightarrow \sqrt{X} \times \frac{1}{2} \Rightarrow \text{Quantum}$$

$$\Rightarrow \sqrt{X} \times \frac{1}{2} \Rightarrow \sqrt{X} \times \frac{1}{2} = \frac{1}{2} \times \frac{1}{2} \frac{1}{2}$$

Q/c wowes

$$p = MV = 1.7 \times 10^{-27} \times 1500 \text{m/s}$$

$$= 2.6 \times 10^{-24} \text{ kg-m/s}$$

λ = L = 6 × 10 - 34 J - 5 2 × 10 - 10 m

= 2 Å

X ~ 1 > Quantum effects
are important, although In practice classical molecular dynamics or somrelassical methods (e.g. path-margral MD) can approach this problem!

9) Harmonte oscMator X = 0.1Åros(211×100THz×t)

Xmax = (0,1A) (2TX100 THZ)

= 6300 m/s

This is a sanitar speed as f) (and the particle will go slower sometimes). So we again expect $\frac{x}{\lambda} = 1$ Quantum

Note: more of proton and mass of H-com are almost the same some (e)'s are light!

42-381 GO SHETTS EYK-EASE" - 5 SOUNRES 42-382 TO SHEETS EYK-EASE" - 5 SOUNRES 82-389 200 SHEETS EYE-EASE" - 5 SOUNRES h) Sample pendulum

$$\omega = \sqrt{\frac{g}{L}} = \sqrt{\frac{9.8 \text{m/s}^2}{50 \text{ cm}}}$$

$$\omega = 4.43 \text{ rad/s}$$

$$V_{max} = \omega \times_{max}$$

= $4.4 \text{ rad/s} \times 10^{-19} = 4.4 \times 10^{-19} \text{ m/s}$

$$\lambda = \frac{h}{p} = \frac{6 \times 10^{-34} \text{ J-s}}{1.8 \times 10^{-17} \text{ kg-m/s}}$$

 $\left[\begin{array}{cccc} \lambda \approx 3 \times 10^{-17} \text{m} \end{array}\right] \frac{\chi}{\lambda} < 2$

Interesting that this comes out loss than one. The real resonant frequency could be higher due to radiation pressure. I think the motion of the test masses in LIGO is still considered to be classical. Certainly the (frequency integrated) excussions are much larger than 1019 m.

i) Agam

Vmax = WXmax

= 2TX105H2X 109m

~ 10 m/s

X = 1 = 6x10-34 J-5 MVmn = 8x10-12 kg x 10-3 m/s

= 10 m << 10 nm

 $\frac{x}{\lambda} \approx 10^{\prime\prime}$ $\frac{x}{\lambda} > 7.1$ classizal

i) M Hompwork & 2 02. 2 0.05 kg

 $\lambda = \frac{6 \times 10^{-34} \text{ J-s}}{0.05 \text{ kg} \times 1 \text{ m/s}} = 1 \times 10^{-32} \text{ m}$

1 x = 3×1031 Very Classical