

# Relativistic Quantum Theory of Atoms and Molecules

## 1 Relativity in atomic and molecular physics

### 1.1 Elementary ideas

nuclei  $\longrightarrow$  point mass

a-th nucleus  $\longrightarrow Z_a e$

The moving particles interact according to Coulomb's law:

$$-\frac{Z_a e^2}{4\pi\epsilon_0 r^2} \quad \text{nucleus a - electrons}$$

$$\frac{e^2}{4\pi\epsilon_0 r^2} \quad \text{electron - electron}$$

$$\frac{Z_a Z_b e^2}{4\pi\epsilon_0 r^2} \quad \text{nuclei a-b repel each other}$$

The electronic intrinsic angular momentum – spin  $s$

$$s = \frac{1}{2} \hbar \sigma \quad \sigma = (\sigma_x, \sigma_y, \sigma_z)$$

Note:  $s^2$  and  $s_z$  have eigenfunction.  $\sigma$  – *spinlabel*

N indistinguishable electrons system wavefunction  $\Psi(q_1, q_2, \dots, q_N, t)$

Spin-Orbit Coupling  $\longrightarrow j = l + s$

### 1.2 The one-electron atom

#### 1.2.1 Classical Kepler orbits

$$\frac{1}{r} = \frac{mk}{|l|^2} \{1 + \varepsilon \cos(\theta + \alpha)\} \quad k = \frac{Ze^2}{4\pi\epsilon_0} \quad \varepsilon = \sqrt{1 + \frac{2E|L|^2}{mk^2}}$$

$$\bullet -\frac{mk}{2|l|^2} \leq E < 0 \implies 0 \leq \varepsilon < 1$$

$$\implies \begin{cases} r = \frac{|l|^2}{mk(1 + \varepsilon)} & \text{closest approach} \\ r = \frac{|l|^2}{mk(1 - \varepsilon)} & \text{maximum distance} \end{cases}$$

$$\text{When } \varepsilon = 0 \implies r = \frac{l^2}{mk} \implies E = -\frac{mk}{2|l|^2}.$$

$$\bullet E = 0 \implies \varepsilon = 1 \implies \text{orbit is a parabola.}$$

$$\bullet E > 0 \implies \varepsilon > 1 \implies \text{orbit is hyperbola.}$$

$$r_{min} = \frac{|l|^2}{mk(1 + \varepsilon)} \quad v_{max} = \frac{|l|}{mr_{min}}$$

#### 1.2.2 The Bohr atom

$$E = \frac{1}{2} \langle V \rangle = -\langle T \rangle$$

Where  $E$  is the energy of particle,  $\langle T \rangle$  is the orbital average of the kinetic energy and  $\langle V \rangle$  is the potential energy.

The frequencies of the spectral lines could be fitted to Rydberg's formula:

$$\nu = R \left( \frac{1}{n^2} - \frac{1}{m^2} \right)$$

The transition energy between two states:  $E_n = -\frac{R}{n^2}$

### 1.2.3 X-ray spectra and Moseley's Law

The square root of the frequency of each corresponding X-ray line was approximately proportional to  $Z$ . Relativistic effects modify the  $Z$ -dependence as  $Z$  increases.

### 1.2.4 Transition to quantum mechanics

A particle wavefunction:  $\psi(\mathbf{r}, t)$

Schrödinger equation:  $i\hbar \frac{\partial \psi}{\partial t} \psi(\mathbf{r}, t) = H\psi(\mathbf{r}, t)$

Hamiltonian:  $\hat{H}(\mathbf{p}, \mathbf{r}) = \frac{1}{2m} \mathbf{p}^2 + V(\mathbf{r})$ ,  $\mathbf{p} \rightarrow -i\hbar \Delta$ ,  $\mathbf{r} \rightarrow \mathbf{r}$

$V(\mathbf{r}) \rightarrow$  potential energy of an electron at a distance  $r = |\mathbf{r}|$

$$V(\mathbf{r}) = -\frac{Ze^2}{4\pi\epsilon r}$$

which could deduce the Hamiltonian

$$\hat{H} = -\frac{\hbar^2}{2m} \Delta^2 - \frac{Ze^2}{4\pi\epsilon r}$$

and whose energies are given by the formula

$$\epsilon_{nl} = -\frac{mZ^2e^4}{32\pi^2\epsilon_0^2\hbar^2n^2}$$

The orbital angular momentum vector  $\mathbf{l}$ , and  $l^2$  takes the values  $l(l+1)\hbar^2$ ,  $l_z$  takes the  $2l+1$  values  $m\hbar$ . Due to Rydberg's formula  $R = mZ^2e^4/32\pi^2\epsilon_0^2\hbar^2$ , from the energy relation could deduce that  $\langle T_n \rangle = -E_n$ , and  $T_n = mv^2/2$ , get the relation

$$\frac{v_n}{c} = \frac{\alpha Z}{n}$$

where  $\alpha = e^2/4\pi\epsilon_0\hbar c$  is the dimensionless fine structure constant.

In spherical polar coordinates

$$\psi_{nlm}(\mathbf{r}, t) = \text{const.} \frac{P_{nl}(r)}{r} Y_l^m(\theta, \phi)$$

### 1.2.5 Sommerfeld's relativistic orbits and Dirac's wave equation

In the Kepler problem, the particle speed attains its maximum at closest approach to the centre of force

$$\frac{v_{max}}{c} = \frac{k}{c|l|}(1 + \epsilon), \quad \epsilon > 0$$

$v_{max}$  is inversely proportional to  $|l|$ , so the largest effects in states with the lowest angular momentum.

A particle moving in some reference frame with velocity  $\mathbf{u}$  has four-momentum  $p^u$

$$p^0 = \frac{E}{c} = mc\gamma(u), \quad p^i = mu^i\gamma(u), \quad i = 1, 2, 3.$$

where  $\gamma(u) = 1/\sqrt{1-u^2/c^2}$  and  $u^i$  are the Cartesian components of  $\mathbf{u}$ .

Dirac' relativistic wave equation for hydrogenic atoms:

$$i\hbar \frac{\partial \psi}{\partial t} = \hat{H}\psi, \quad \hat{H} = c\boldsymbol{\alpha} \cdot \mathbf{p} + \beta mc^2 - \frac{Ze^2}{4\pi\epsilon_0 r}$$

The solution in spherical polar coordinates

$$\psi_{E\kappa m}(\mathbf{r}) = \text{const.} \frac{1}{r} \begin{pmatrix} P_{E\kappa}(r)\chi_{\kappa m}(\theta, \varphi) \\ iQ_{E\kappa}(r)\chi_{-\kappa m}(\theta, \varphi) \end{pmatrix}$$

$\kappa = \pm 1, \pm 2, \dots$  is the angular number and,

$$\chi_{\pm\kappa m}(\theta, \varphi) = \sum_{\sigma} (l, m - \sigma, 1/2, \sigma | l, 1/2, j, m) Y_l^{m-\sigma}(\theta, \varphi) \phi_{\sigma}$$

where

$$\phi_{1/2} = \begin{pmatrix} 1 \\ 0 \end{pmatrix}, \quad \phi_{-1/2} = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

The two possible Coupling modes are define by

$$\kappa = \eta(j + 1/2), \quad l = j + \eta/2, \quad \eta = \pm 1$$

The principal quantum number:  $n = n_r + |\kappa|$

### 1.2.6 Dirac and Schrödinger charge distributions

- Dirac *radial density distributions* are more compact than their nonrelativistic counterparts.  $\rho_{n\kappa} < \rho_{nl}$ ,  $\rho_{n\kappa}$  is the Dirac value,  $\rho_{nl}$  is the Schrödinger value.
- Dirac electrons are more tightly bound than their Schrödinger counterparts:  $\varepsilon_{n\kappa} < \varepsilon_{nl}$
- *Spin-orbit splitting*.  $\varepsilon_{n,\kappa=l} < \varepsilon_{n,\kappa=-l-1} < \varepsilon_{nl}$
- The Schrödinger radial wavefunction in the nonrelativistic limit  $\alpha \rightarrow 0, c \rightarrow \infty$ , the Pauli formula

$$Q_{n,\kappa} \approx \frac{1}{2c} \left( \frac{dP_{n,\kappa}}{dr} + \frac{\kappa P_{n,\kappa}}{r} \right) \left\{ 1 + O\left(\frac{1}{c^2}\right) \right\}$$

### 1.2.7 The Dirac hydrogenic spectrum at high Z

The properties of the Dirac hydrogenic states depart more and more from those of the corresponding nonrelativistic model as Z increases.

## 1.3 Many-electron atoms