

# Thermal Mathematical Models Correlation.

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Inés Arauzo Andrés

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MUSE IDR-UPM-ETSIAE



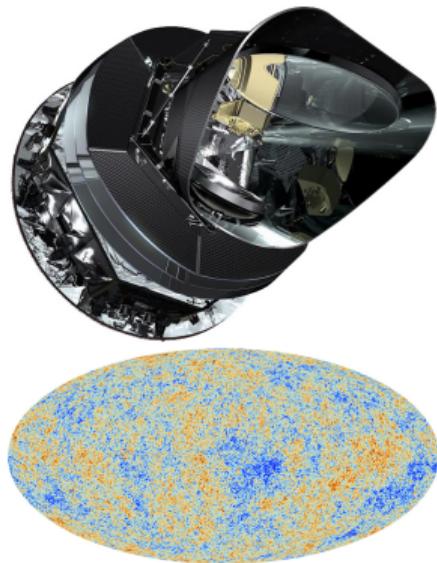
# Outline

# Introduction

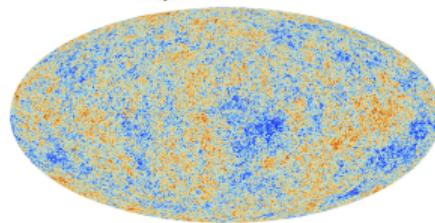
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# State of the art

- Thermal mathematical model of the UPMSat 3 in ESATAN.
- $\approx 1150$  nodes
- Two steady cases defined by the mission requirements:
  1. Steady hot with all the payloads active (PL1 Case).
  2. Steady Cold, latency case, minimum power consumption.



# Objectives



- Develop thermal TVAC test cases in ESATAN.
- Run the models on pycanha and compare the results with the ones from ESATAN.
- Determine and compare the most influent parameters in the thermal model on flight and TVAC.
- Optimize sensor placement for the TVAC test and the Flight cases.
- Correlate the model with the test cases.

# Pycanha model validation

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# ESATAN vs Pycanha

- HFI environment temperature  $\downarrow$  2K.
- LFI environment temperature  $\downarrow$  60K.
- 0.5 W heatlift in the LFI.

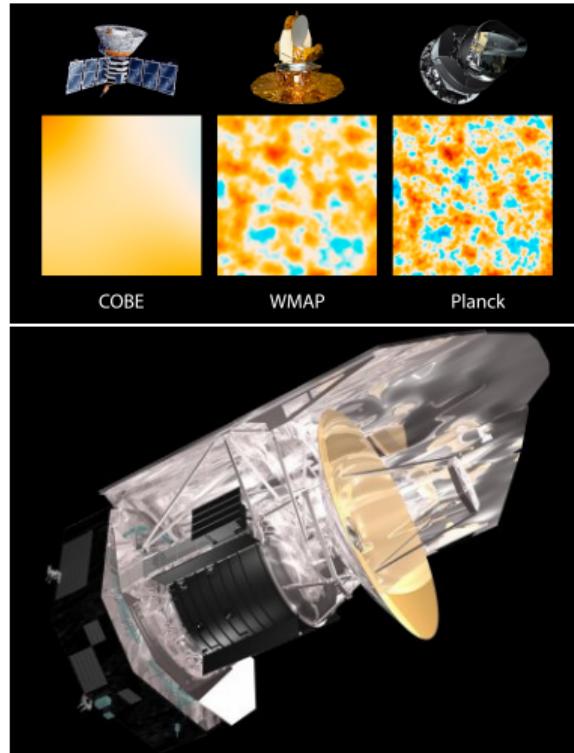
## Other requirements

- No deployables.
- No optical elements such as windows with warm edges between the feed horns and telescope
- 1.5 yr minimum total lifetime.
- A spinning spacecraft.

# Precedents

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# Background



- COBE ( $T = 2.725\text{K}$ ): used a 650 l He cryostat to cool the IR spectro-photometer to 2 K .
- WMAP ( $T = 90\text{K}$ ): used passive cooling. A lot simpler, but unable to reach sub-Kelvin operating temperatures.
- Herschel ( $T = 1.65\text{K}$ ): used 2 He cryostats of 2300 l each to cool the HIFI.

# Thermal control instruments

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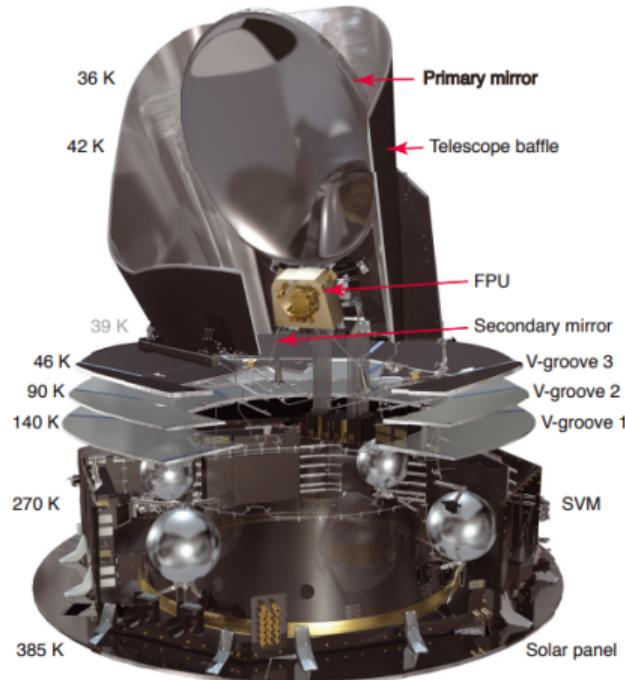
# Thermal control instruments

## 1. Passive Elements

- Telescopic Baffle: provided radiative shield.
- 3 V-groove radiators.

## 2. Active Elements

- Hydrogen sorption cooler.
- Helium Joule Thompson expansion cooler.
- He-He dilution cooler.

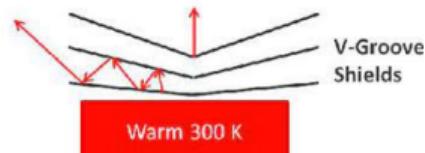


# Passive elements



- Telescopic Baffle: provided radiative shield.
- 3 V-groove radiators: provided thermal isolation and radiative cooling

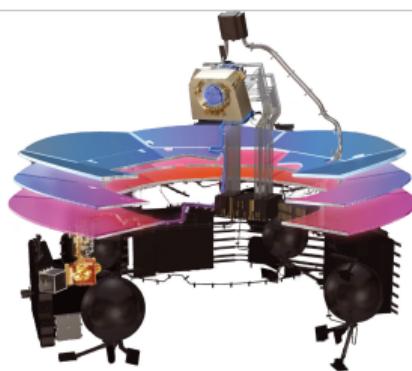
Allows to reach the precooling temperature  $\leq 60\text{K}$ .



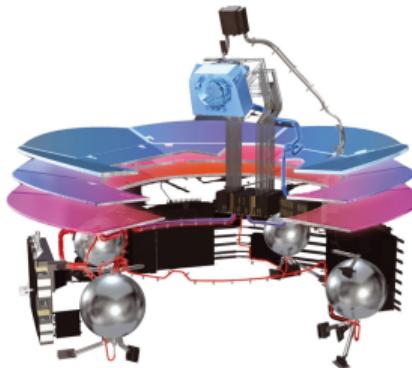
# Active elements



H-Sorption cooler

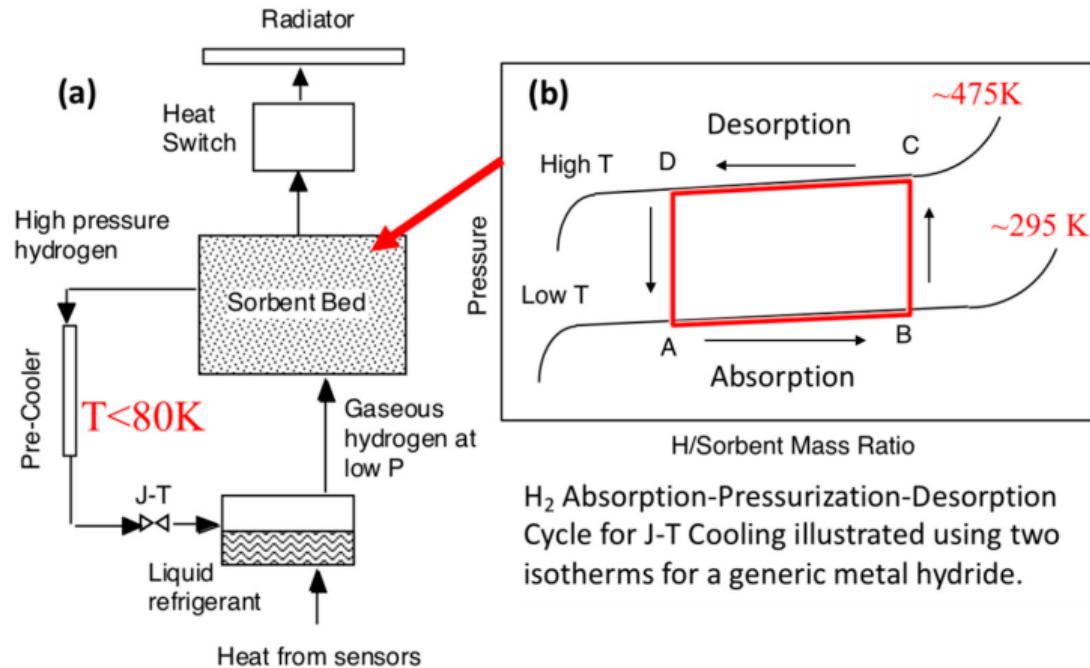


He-JT Expansion cooler



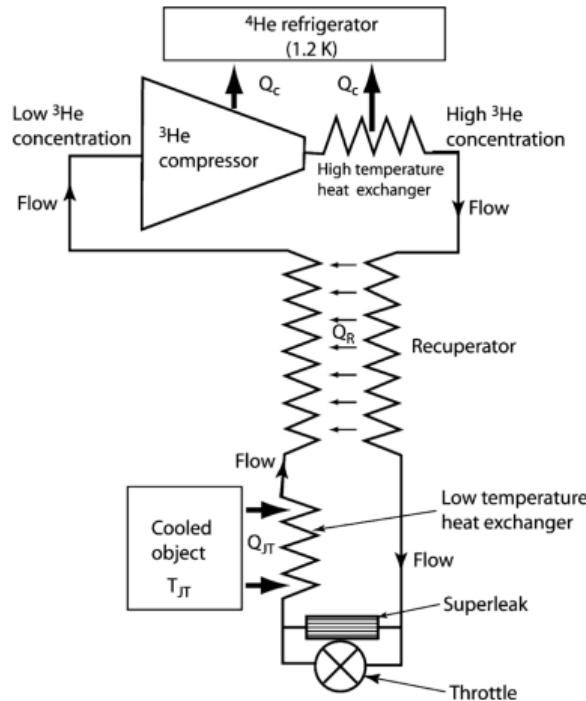
He-He dilution cooler

### **Active elements: H-Sorption Cooler**



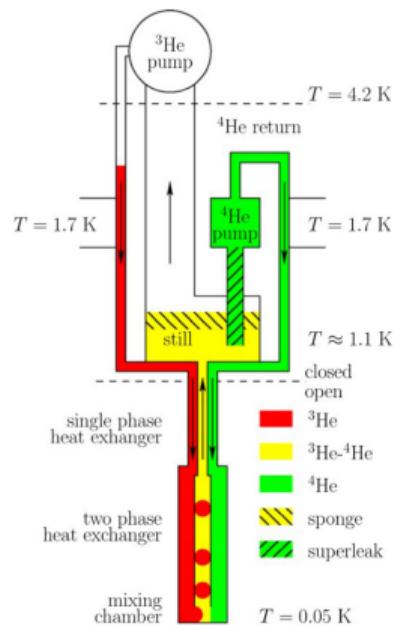
## Functioning of an sorption cooler

# Active elements: He-JT cooler



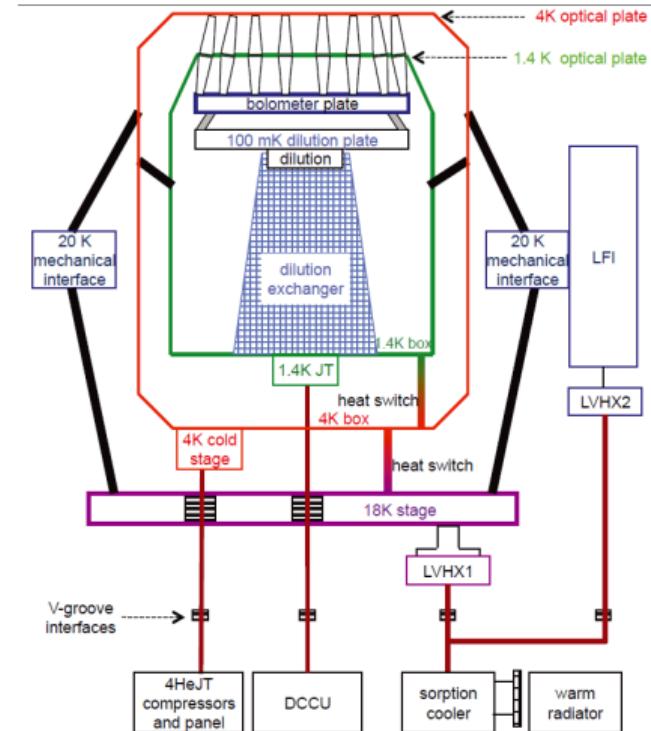
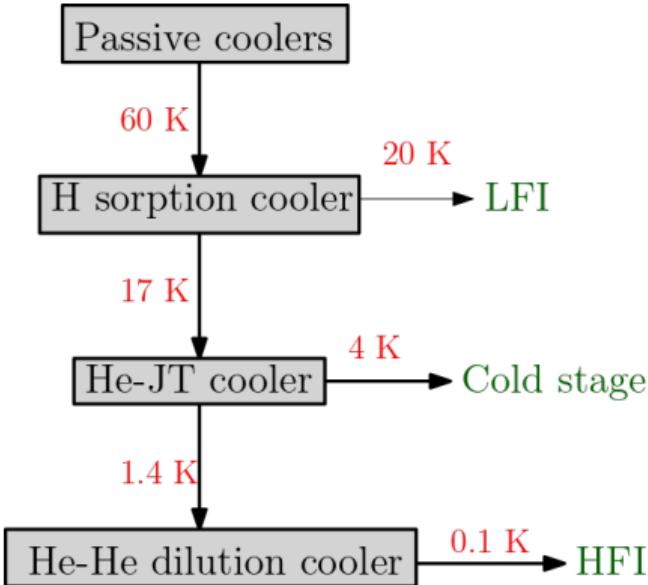
- JT He cooler works by expanding helium gas through a small orifice, resulting in a temperature drop.
- Cold He recirculates through a heat exchanger to cool down a target object.
- He is compressed back to its original pressure to start the cooling cycle again.

# Active elements: He-He Dilution cooler



- Cooling medium is a mixture of Helium-3 and Helium-4 gradually cooled down to temperatures approaching 0 K.
- Uses the superfluid helium-4 to carrying heat away from the target object, as it flows without any resistance.
- Can achieve extremely low temperatures, down to a few millikelvin.

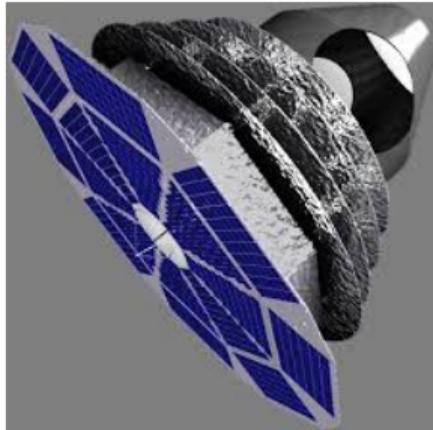
# Complete system



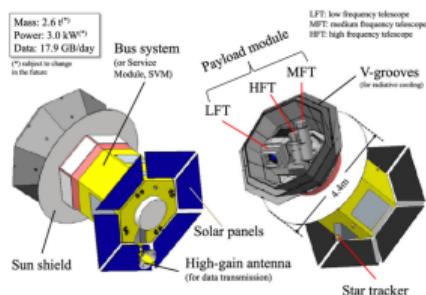
## Future missions

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# Active Elements



COrE+  
Cosmic Origins Explorer,  
ESA



LiteBird  
Primordial B modes mission,  
JAXA



Pixie Absolute  
Spectrophotometer, NASA

## References

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## References

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# Thank you for attending

## Questions?