## Completo 21/Output file fort.1

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## fort.1: Time evolution

- 1. time step number iloop
- 2. time/years temps/an
- 3. number of layers n\_frontieres
- 4. core mass masse(1)/mearth
- 5. total mass masse(n\_frontieres)/mearth
- 6. core radius / RJ X(1)/RJ
- 7. total radius / RJ X(nvar\*(n\_frontieres-1)+1)/RJ
- 8. core radius / Rearth X(1)/Rearth
- 9. total radius / Rearth at tau=0.01 (usually) X(nvar\*(n\_frontieres-1)+1)/Rearth
- 10. core luminosity / LJ X(2)/LJ
- 11. total luminosity / LJ X(nvar\*(n\_frontieres-1)+2)/LJ
  12. total luminosity / Lsun X(nvar\*(n\_frontieres-1)+2)/Lsun
- 13. pressure at envelope core boundary / bar X(3)/1D6
- 14. surface pressure /bar (at tau=0.01 usually) X(nvar\*(n\_frontieres-1)+3)/1D6
- 15. temperature at core envelope boundary temperature K X(4)
- 16. surface temperature (at tau=0.01 usually) K X(nvar\*(n\_frontieres-1)+4)
- 17. Kelvin Helmholtz timescale TKH/an
- 18. timestep / year dt/an
- 19. total energy Etot
- 20. total energy Etot2
- 21. kenerg
- 22. kenerg2
- 23. lumi correction factor corrLesti
- 24. energy correction factor not used corrEesti
- 25. actual luminosity from energy differences Lactual/LJ from contraction
- 26. mean core density rhocore
- 27. radiogenic core luminosity Lradio/LJ
- 28. entropy at core envelope interface kb / baryon entro(1)
- 29. entropy at top of atmosphere kb / baryon entro(n\_frontieres)
- 30. envelope luminosity from TdS / LJ lumitds(n\_frontieres)/LJ
- 31. total D burning luminosity Lburnoldtot/LJ
- 32. D burning timescale tauburn/an
- 33. Mixing timescale taumix/an
- 34. Number of eddies in planet edNumber
- 35. deuterium mass fraction at bottom xDeu(1)
- 36. X(nvar\*(n\_frontieres-1)+5) tau at outer boundary
- 37. masse23/mearth Mass at tau=2/3
- 38. radius23/RJ Radius at tau=2/3
- 39. radius23/Rearth Radius at tau=2/3
- 40. Mdotevap/Mearth\*an envelope evaporation rate
- 41. envstatus envelope status flag
- 42. X(5) optical depth at core-envelope boundary
- 43. rho(1) gas density at core-env [g/cm\*\*3]
- 44. rho(n\_frontieres) density at outermost layer (usually tau=0.01)
- 45. convect(1) convective or rad at core?

- 46. convect(n\_frontieres) convective or rad at top?
- 47. nconvlayer nb of convective layers
- 48. P23/1D6 pressure at tau=2/3 in bars
- 49. T23 temperature at tau=2/3 in K
- 50. Teq equilibrium temp [K]
- 51. aplanete/AU
- 52. fract\_ice ice mass fraction in core
- 53. menve/mearth envelope mass
- 54. menve/menveinit current envelope mass / initial envelope mass
- 55. Tint intrinsic temperature K
- 56. waterstate
- 57. thab/an
- 58. rho mean planet density [g/cm\*\*3]
- 59. Eintcore [erg]
- 60. Egravcore [erg]
- 61. Lintcore/LJ core lumiosity due to core cooling
- 62. Lgravcore/LJ core luminosity due to core contraction
- 63. (Lint+Labs)/LJ
- 64. lambda max Wien mu maximum wavelength of radiation
- 65. rroche/Rearth Hill sphere radius
- 66. ABS(mdotover)/Mearth\*an
- 67. H23 Scale height at tau=2/3
- 68. G23 gravitational acceleration at tau=2/3
- 69. Rtrans/Rearth Transit radius in Rearth
- 70. Rtrans/RJ Transit radius in Rjupiter
- 71. R1bar/Rearth 1 Bar radius
- 72. T1bar Temperature at 1 Bar radius
- 73. tau1bar Tau at 1 Bar radius
- 74. M1bar/Mearth Mass at 1 Bar radius
- 75. Lbloating/LJ Bloating luminosity
- 76. Lintenve/LJ Luminosity of envelope from cooling77. Lgravenve/LJ Luminosity of envelope from contraction
- 78. Eintenve [erg]
- 79. Egravenve [erg]
- 80. kv2kth ratio opacity in visual to thermal for atmo model 81. Rstar/Rsun stellar radius
- 82. Tstar stellar effective temperaure
- 83. Lstar/Lsun stellar luminosity
- 84. FhalfDeg(1) Fermi integral for degeneracy
- 85. thetaEDeg(1) Degeneracy parameter 1=fully nondegenerate 0=fully degenerate
- 86. psiDeg(1) degeneracy parameter
- 87. Rrcb/Rearth Radius innermost radiative convective boundary
- 88. Lrcb/LJ Radius innermost radiative convective boundary
- 89. Prcb/Bar Pressure at innermost radiative convective boundary 90. Trcb Temperature at innermost radiative convective boundary
- 91. Mrcb/Mearth Mass at innermost radiative convective boundary
- 92. taurcb Optical depth at innermost radiative convective boundary
- 93. Kapparcb Kappa at innermost radiative convective boundary
- 94. Rhorcb density at innermost radiative convective boundary
- 95. dlnTdlnPradrcb radiative gradient at innermost radiative convective boundary
- 96. dlnTdlnPconvrcb convective gradient at innermost radiative convective boundary
- 97. totally emitted energy (integral L dt)
- 98. contrast black body 1 mu
- 99. contrast black body 2 mu

100. contrast black body 3 mu 101. contrast black body 4 mu 102. contrast black body 5 mu 103. contrast black body 6 mu 104. contrast black body 8 mu 105. contrast black body 10 mu 106. contrast black body 12 mu 107. contrast black body 16 mu 108. contrast black body 20 mu 109. contrast black body 40 mu 110. contrast black body 60 mu 111. contrast black body 100 mu 112. Abs Magnitudes: AMES Cond 2MASS J 113. H 114. K 115. AMES Cond NACO J 116. H 117. Ks 118. Lp 119. Mp 120. AMES Cond SPHERE Y 121. J 122. H 123. Ks 124. AMES Dusty 2MASS J 125. H 126. K 127. AMES Dusty NACO J 128. H 129. Ks 130. Lp 131. Mp 132. AMES Dusty SPHERE Y 133. J 134. H 135. Ks 136. BT Settl CFIST 2011 bc 2MASS J 137. H 138. K 139. BT Settl CFIST 2011 bc NACO J 140. H 141. Ks 142. Lp 143. Mp 144. BT Settl CFIST 2011 bc SPHERE Y 145. J 146. H 147. Ks 148. Abs Mags for black body, Vega normalized U 149. B 150. V 151. R 152. I

153. J

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154. H
155. Ks
156. L
157. M
158. N
159. Q
160. W1
161. W2
162. W3
163. W4
164. Abs Fluxes in Jansky for black body
165. B
166. V
167. R
168. I
169. J
170. H
171. Ks
172. L
173. M
174. N
175. Q
176. W1
177. W2
178. W3
179. W4
180. Abs magnitudes from BT settl grid 2MASS J
181. 2MASS H
182. 2Mass Ks
183. VLT NACO CONICA J
184. VLT NACO CONICA H
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185. VLT NACO CONICA Ks

186. VLT NACO LPRIME

187. VLT NACO MPRIME

188. HICIAO H

189. Pcent [BAR] (estimated central pressure)

190. Tcent [K] estimated central temperature for adiabatic profile

To check for energy conservation at one timestep (without D burning): it must hold that 11=25+27 and that 11=30+61+62+27

To check for energy conservation across time: 19+97 should be constant (without Dburning, Lradio, and Lbloating, and non-irradiated planets).

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