## Semi-Relativistic Coulomb Collision Operators for Current Drive Applications

Y. J.  $Hu^{1,2}$  Y. M.  $Hu^{1,2}$  Y. R. Lin-Liu<sup>3</sup>

<sup>1</sup>Institute of Plasma Physics, Chinese Academy of Sciences, Hefei.

<sup>2</sup>Center for Magnetic Fusion Theory, Chinese Academy of Sciences, Hefei.

<sup>3</sup>Department of Physics and Center for Mathematics and Theoretical Physics, National Central University, Taiwan.

EC-16, Sanya, China, 2010

### Introduction and Motivation

- A semi-relativistic Fokker-Planck operator for electron-electron collision was first used by Karney and Fisch to calculate lower-hybrid and electron cyclotron current drive efficiencies.
- There is evidence from our recent numerical work that this collision operator should be appropriate for current drive applications under ITER conditions.
- We extended Karney and Fisch's work by expressing the semi-relativistic Fokker-Planck coefficients in terms of a pair of Rosenbluth-Trubnikov potential functions, and worked out the general Legendre expansion of these two potentials.
- This general Legendre expansion can be used to implement the semi-relativistic collision term efficiently in Fokker-Planck codes, such as in CQL3D.

# Whether the semi-relativistic collision operator is precise enough under ITER conditions?

- To justify the use of semi-relativistic collision operator under ITER conditions, here we give some numerical examples about the precision of the semi-relativistic collision operator.
- These numerical results indicate that the semi-relativistic collision operator should be precise enough for current drive applications under ITER conditions.

# Whether the semi-relativistic collision operator is precise enough under ITER conditions?

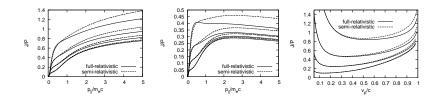


Figure 1: Efficiencies for localized excitation of Landau-damped waves (left figure), cyclotron-damped waves (middle figure) and narrow Landau-damped waves (right figure). The curves from bottom to top show the efficiencies for different temperature  $\Theta = T_e/m_ec^2 = 0.02, 0.05, 0.1, 0.2.$  The solid lines correspond to the full-relativistic model, while the dashed lines correspond to the semi-relativistic model.

### Collision operator in Fokker-Planck form

The collision term for a relativistic plasma of species a colliding off species b may be written in the Fokker-Planck form as[1, 4, 5]

$$C(f_a, f_b) = -\frac{\partial}{\partial \boldsymbol{u}} \cdot \left[ -\boldsymbol{D}^{a/b} \cdot \frac{\partial f_a}{\partial \boldsymbol{u}} + \boldsymbol{F}^{a/b} f_a \right],$$

in which the coefficient  $oldsymbol{D}^{a/b}$  and  $oldsymbol{F}^{a/b}$  are defined by

$$\boldsymbol{D}^{a/b}(\boldsymbol{u}) = \frac{q_a^2 q_b^2}{8\pi \epsilon_0^2 m_a^2} \ln \Lambda^{a/b} \int \boldsymbol{U}(\boldsymbol{u}, \boldsymbol{u}') f_b(\boldsymbol{u}') d\boldsymbol{u}', \qquad (1)$$

$$\boldsymbol{F}^{a/b}(\boldsymbol{u}) = -\frac{q_a^2 q_b^2}{8\pi\epsilon_0^2 m_a m_b} \ln \Lambda^{a/b} \int \frac{\partial}{\partial \boldsymbol{u}'} \cdot \boldsymbol{U}(\boldsymbol{u}, \boldsymbol{u}') f_b(\boldsymbol{u}') d\boldsymbol{u}', \quad (2)$$

Here  $f_a$  and  $f_b$  are the distribution functions for the two species,  $\boldsymbol{u}$  is the ratio of momentum to species mass, and  $\ln \Lambda^{a/b}$  is the Coulomb logarithm.

### **Extend Karney and Fisch's work**

 If either the test or the background species is weakly relativistic, the full relativistic collision kernel can be approximated by the simple Landau semi-relativistic kernel[2, 4],

$$oldsymbol{U}(oldsymbol{v},oldsymbol{v}') = rac{oldsymbol{I}}{|oldsymbol{v}-oldsymbol{v}'|} - rac{(oldsymbol{v}-oldsymbol{v}')\,(oldsymbol{v}-oldsymbol{v}')}{|oldsymbol{v}-oldsymbol{v}'|^3}.$$

 In this work, we extend Karney's work by expressing the semi-relativistic Fokker-Planck coefficients in terms of two potential functions, and worked out the general Legendre decomposition of these two potential functions.

## Potential form of the semi-relativistic Fokker-Planck coefficients

Define two potential functions,

$$h_b(\boldsymbol{v}) = -\frac{1}{8\pi} \int |\boldsymbol{v} - \boldsymbol{v}'| \gamma'^5 f_b(\gamma' \boldsymbol{v}') d^3 \boldsymbol{v}',$$

$$g_b(\boldsymbol{v}) \equiv -\frac{1}{4\pi} \int \left| \frac{1}{|\boldsymbol{v} - \boldsymbol{v}'|} \left( \frac{1}{\gamma'} + \frac{1}{\gamma'^3} \right) + \frac{(\boldsymbol{v} \cdot \boldsymbol{v}' - \boldsymbol{v}'^2)^2}{|\boldsymbol{v} - \boldsymbol{v}'|^3 \gamma'} \right| \gamma'^5 f_b(\gamma' \boldsymbol{v}') d^3 \boldsymbol{v}'.$$

In terms of these two potential functions, the diffusion and friction coefficients can be written as.

$$\boldsymbol{D}^{a/b}(\boldsymbol{u}) = -\frac{4\pi c_{ab}}{m_a^2} \frac{\partial^2 h_b(\boldsymbol{v})}{\partial \boldsymbol{v} \partial \boldsymbol{v}},$$

$$\mathbf{F}^{a/b}(\mathbf{u}) = -\frac{4\pi c_{ab}}{m_{rr}m_{b}}\frac{\partial}{\partial \mathbf{v}}g_{b}(\mathbf{v}).$$

### Legendre expansion of Fokker-Planck coefficients I

If the background distribution  $f_b$  and the potential functions are expanded in terms of Legendre polynomials ( assuming axial symmetry ),

$$f_b(u,\theta) = \sum_{l=0}^{\infty} f_b^l(u) P_l(\cos \theta),$$

$$h_b(v,\theta) = \sum_{l=0}^{\infty} h_b^l(v) P_l(\cos \theta),$$

$$g_b(v,\theta) = \sum_{l=0}^{\infty} g_b^l(v) P_l(\cos \theta),$$

one can get a relation between  $\boldsymbol{h}_b^l$  and  $\boldsymbol{f}_b^l$ ,

$$h_b^l(v) = \frac{1}{2(4l^2 - 1)} \left[ \int_0^v \frac{(v')^{l+2}}{v^{l-1}} \left( 1 - \frac{2l - 1}{2l + 3} \times \frac{(v')^2}{v^2} \right) \gamma'^5 f_b^l(\gamma' v') dv' \right] + \int_v^c \frac{v^l}{(v')^{l-3}} \left( 1 - \frac{2l - 1}{2l + 3} \times \frac{v^2}{(v')^2} \right) \gamma'^5 f_b^l(\gamma' v') dv' \right].$$

Я

## Legendre expansion of Fokker-Planck coefficients II

And a relation between  $g_b^l$  and  $f_b^l$ ,

$$\begin{split} g_b^l(v) &= -\frac{1}{2l+1} \int_0^v \left[ \frac{1}{c^2} \frac{v'^{l+2}}{v^{l-1}} \gamma'^4 \frac{l(l-1)}{2l-1} - \frac{1}{c^2} \frac{v'^{l+4}}{v^{l+1}} \gamma'^4 \left( \frac{l^2+l-1}{2l+3} \right) \right. \\ &+ \left. \frac{(v')^{l+2}}{v^{l+1}} \left( \gamma'^2+1 \right) \gamma'^2 \right] f_b^l(\gamma'v') dv' \\ &- \left. \frac{1}{2l+1} \int_v^c \left[ -\frac{1}{c^2} \frac{v^{l+2}}{v'^{l-1}} \gamma'^4 \frac{(l+1)(l+2)}{2l+3} + \frac{1}{c^2} \frac{v^l}{v'^{l-3}} \gamma'^4 \left( \frac{l^2+l-1}{2l-1} \right) \right. \\ &+ \left. \frac{v^l}{(v')^{l-1}} \left( \gamma'^2+1 \right) \gamma'^2 \right] f_b^l(\gamma'v') dv'. \end{split}$$

Q

### Summary

- We have derived a general Legendre expansion for the Fokker-Planck coefficients of the semi-relativistic collision operator.
- This general Legendre decomposition can be used to implement the semi-relativistic collision term efficiently in Fokker-Planck codes, such as in CQL3D[9].

#### References



[1] S. T. Beliaev and G. I. Budker. Sov. Phys. Dokl, 1:218, 1956.



[2] Charles F. F. Karney and Nathaniel J. Fisch. Physics of Fluids, 28(1):116-126, 1985.



[3] Charles F. F. Karney. Comp. Phys. Rep., 4:183-244, 1986.



[4] L. D. Landau. Phys. Z. Sowjetunion, 10:154, 1936.



[5] Bastiaan J. Braams and Charles F. F. Karney. Phys. Fluids, 1B(7):1355-1368, July 1989.



[6] Y. M. Hu, Y. J. Hu, and Y. R. Lin-Liu. In Poster at 9th Joint Workshop on Electron Cyclotron Emission and Electron Cyclotron Heating. Sanva. China. 2010.



[7] Marshall N. Rosenbluth, William M. MacDonald, and David L. Judd. Phys. Rev., 107(1):1-6, Jul 1957.



[8] B. A. Trubnikov. Consultants Bureau, New York, 1965.



[9] R.W. Harvey and M.G. McCoy. In *IAEA Conf. Proc.*, Technical Committee Meeting on Advances in Simulation and Modeling of Thermonuclear Plasmas, page 527, Montreal, Canada, 1992. International Atomic Energy Agency.



[10] Bastiaan J. Braams and Charles F. F. Karney. Phys. Rev. Lett, 59(16):1817, 1987.



[11] Jr. T.M. Antonsen and K.R. Chu. Physics of Fluids, 25(8):1295–1296, 1982.



[12] Ronald H. Cohen. Physics of Fluids, 30(8):2442-2449, 1987.



[13] Y. R. Lin-Liu, V. S. Chan, and R. Prater. Physics of Plasmas, 10(10):4064-4071, 2003.