Influence of energetic particles on MHD modes—application of NOVA-K code

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Background

- In addition to thermal ions and electrons, tokamak plasmas usually contain super-thermal species, which are called fast or energetic particles
- These energetic particles (EP) can be created from various sources, such as ion/electron cyclotron heating, neutral beam injection, and fusion reaction
- The energetic particles can destabilize various MHD waves
- The destabilizing effects (drive effects) of EP can be calculated by the NOVA-K code

NOVA-K code

NOVA-K is a 2D Kinetic-MHD linear stability code for tokamaks with energetic particles.

- Authors: C. Z. Cheng, N. Gorelenkov, G. Y. Fu (PPPL)
- NOVA-K is written in Fortran Language
- Approximate number of code lines: 10,000 lines
- Code was first used in 1988 after the discovery of TAE

Physical model of NOVA-K code

- NOVA-K uses the perturbative method to calculate the destabilizing effects of EP
- NOVA-K are composed of two codes, one of which, NOVA, calculates the ideal MHD modes without the EP
- Then NOVA-K post-processor uses the calculated MHD modes structure to calculate the orbit integration of EP to obtain the contribution of EP to the growth rate of the modes
- Besides the EP drive effects, NOVA-K also includes various damping effects:
 - (1) Thermal electron/ion Landau damping
 - (2) Continuum damping
 - (3) Collisional damping of trapped electrons
 - (4) Radiative damping

Ideal MHD equations

Momentum equation

$$\rho_m \left(\frac{\partial \boldsymbol{u}}{\partial t} + \boldsymbol{u} \cdot \nabla \boldsymbol{u} \right) = -\nabla p + (\nabla \times \boldsymbol{B}) / \mu_0 \times \boldsymbol{B}$$

Faraday's Law

$$\frac{\partial \boldsymbol{B}}{\partial t} = \nabla \times (\boldsymbol{u} \times \boldsymbol{B})$$

Equation of state

$$\frac{\partial p}{\partial t} + \boldsymbol{u} \cdot \nabla p = -\gamma p \nabla \cdot \boldsymbol{u}$$

Linearized ideal MHD equations

- ξ , B_1 , and p_1 are perturbed plasma displacement, magnetic field and plasma pressure, respectively.
- Time dependence of perturbation: $A(\mathbf{r},t) = A(\mathbf{r})\exp(-i\omega t)$
- ω^2 is a real number
- Linearized ideal MHD equations:

$$-\omega^2 \rho_{m0} \boldsymbol{\xi} = -\nabla p_1 + (\nabla \times \boldsymbol{B}_1)/\mu_0 \times \boldsymbol{B}_0 + (\nabla \times \boldsymbol{B}_0)/\mu_0 \times \boldsymbol{B}_1$$
$$\boldsymbol{B}_1 = \nabla \times (\boldsymbol{\xi} \times \boldsymbol{B}_0)$$

 $p_1 + \boldsymbol{\xi} \cdot \nabla p_0 + \gamma p_0 \nabla \cdot \boldsymbol{\xi} = 0$

Kinetic effect of EP on linear MHD stability

 NOVA-K includes kinetic effect through perturbed pressure of energetic/thermal particles in the momentum equation

$$-\omega^{2}\rho_{m0}\boldsymbol{\xi} = -\nabla p_{1} - \nabla \cdot \delta \boldsymbol{p}_{h} + (\nabla \times \boldsymbol{B}_{1})/\mu_{0} \times \boldsymbol{B}_{0} + (\nabla \times \boldsymbol{B}_{0})/\mu_{0} \times \boldsymbol{B}_{1}$$

Quadratic form and perturbative calculation of kinetic effects

$$\delta W_f + \delta W_k - \omega^2 \delta K = 0$$

Zeroth-order equation

$$\delta W_f(\xi_0^*, \xi_0) - \omega_0^2 \delta K(\xi_0^*, \xi_0) = 0$$

The first-order equation

$$\delta W_f(\boldsymbol{\xi}_0^{\star}, \boldsymbol{\xi}_1) - \omega_0^2 \delta K(\boldsymbol{\xi}_0^{\star}, \boldsymbol{\xi}_1) - 2\omega_0 \omega_1 \delta K(\boldsymbol{\xi}_0^{\star}, \boldsymbol{\xi}_0) + \delta W_k(\boldsymbol{\xi}_0^{\star}, \boldsymbol{\xi}_0, \omega_0) = 0$$

$$\Longrightarrow \frac{\omega_1}{\omega_0} = \frac{\delta W_k(\boldsymbol{\xi}_0^{\star}, \boldsymbol{\xi}_0, \omega_0)}{2\omega_0^2 \delta K(\boldsymbol{\xi}_0^{\star}, \boldsymbol{\xi}_0)}$$

where

$$\delta W_k = \int d^3 \boldsymbol{x} \boldsymbol{\xi}^* \cdot \nabla \cdot \delta P_k$$

Expression of perturbation

- Due to toroidal symmetry, perturbation can be represented by a single toroidal mode, $\exp(-in\zeta)$.
- However, due to the poloidal nonuniform, different poloidal harmonics are coupled, perturbation should include all Fourier series in θ

$$\boldsymbol{\xi}(\psi,\theta,\zeta,t) = \sum_{m=-\infty}^{m=+\infty} \boldsymbol{\xi}_m(\psi) \exp[i(m\theta - n\zeta - \omega t)]$$

ullet ψ dependence of the perturbation is treated in NOVA-K using finite-element expansion

Decomposition of ideal MHD perturbation

NOVA-K decomposes the displacement vector and perturbed magnetic field as

$$\boldsymbol{\xi} = \frac{\xi_{\psi}}{|\nabla \psi|^2} \nabla \psi + \frac{\xi_s}{B^2} (\boldsymbol{B}_0 \times \nabla \psi) + \frac{\xi_b}{B^2} \boldsymbol{B}_0,$$

$$\boldsymbol{B}^{(1)} = \frac{Q_{\psi}}{|\nabla \psi|^2} \nabla \psi + \frac{Q_s}{|\nabla \psi|^2} (\boldsymbol{B}_0 \times \nabla \psi) + \frac{Q_b}{B_0^2} \boldsymbol{B}_0.$$

Ideal MHD eigen-mode equations

$$\begin{pmatrix} A_{11} & A_{12} & A_{13} & A_{14} \\ A_{21} & A_{22} & A_{23} & A_{24} \\ A_{31} & A_{32} & A_{33} & A_{34} \\ A_{41} & A_{42} & A_{43} & A_{44} \end{pmatrix} \begin{pmatrix} \boldsymbol{\xi}_{\psi} \\ \delta p_{1} \\ \boldsymbol{\xi}_{s} \\ \nabla \cdot \boldsymbol{\xi} \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix},$$

where A_{ij} can be

$$(\nabla \psi) \cdot \nabla,$$
 $m{B} \cdot \nabla,$ $(m{B} imes
abla \psi) \cdot
abla,$

Perturbed EP distribution due to ideal MHD perturbation

- Equilibrium distribution: $F = F(P_{\varphi 0}, \varepsilon_0, \mu_0)$
- Ideal MHD perturbation: $\mathbf{B}_1 = \nabla \times (\boldsymbol{\xi} \times \boldsymbol{B}_0)$
- Perturbed distribution:

$$\delta f = -\frac{Ze}{c} \boldsymbol{\xi}_{\perp} \nabla \psi \frac{\partial F}{\partial P_{\varphi 0}} - \mu_0 \frac{B_{1\parallel}}{B_0} \frac{\partial F}{\partial \mu_0} - \delta g,$$

where

$$\delta g = -i(\omega - n\omega_{\star}) \frac{\partial F}{\partial \varepsilon_{0}} \int_{0}^{t} \mathcal{L}^{(1)} d\tau,$$

$$\mathcal{L}^{(1)} = -\left(mv_{\parallel}^{2} - \mu_{0}B_{0}\right) \boldsymbol{\xi} \cdot \boldsymbol{\kappa} + \mu_{0}B_{0}\nabla \cdot \boldsymbol{\xi}_{\perp},$$

$$\omega_{\star} = \frac{\partial F}{\partial P_{c0}} / \frac{\partial F}{\partial \varepsilon_{0}}.$$

Unperturbed guiding-center orbits

Unperturbed guiding-center orbits

$$\dot{\boldsymbol{R}} = v_{\parallel} \boldsymbol{b} + \frac{1}{m\Omega} \boldsymbol{b} \times \left(\mu \nabla B + m v_{\parallel}^2 \boldsymbol{\kappa} \right),$$
$$\dot{v}_{\parallel} = \frac{1}{m} \left(-\mu \boldsymbol{b} \cdot \nabla B + m v_{\parallel} \boldsymbol{\kappa} \cdot \dot{\boldsymbol{R}} \right).$$

Finite-Orbit-Width (FOW) effect is included in NOVA-K.

Toroidicity-Induced Alfvén Eigen-mode

- Toroidicity-Induced Alfvén Eigen-mode (TAE) is a type of global shear Alfvén eigenmode.
- Exists only in toroidal geometry: created due to toroidal coupling.
- Frequency is discrete, lies within the "gaps" in the shear Alfvén continuum.
- Can be strongly destabilized by α -particles in burning tokamak plasmas.

NOVA-K input

- The input file for the equilibrium is "inequ", which is a file formatted as a table. The entries of the table set various equilibrium parameters
- Most important are (1) pressure and safety factor profiles. (2) shape of LCFS
- The energetic particles equilibrium distribution can be set in NOVA_param. For example, the argument tip is related to the distribution function type, (refer to taem.f)
 - "s" for slowing down distribution, $F_h(\varepsilon, \Lambda, \psi) = n_h(\psi) \delta(\Lambda \Lambda_0) / \varepsilon^{3/2}$
 - "m" for Maxwellian with standard input

NOVA-K output

- Shear Alfvén continuum.
- Frequency and mode structure of ideal MHD modes
 - External/Internal kink modes
 - Toroidal Alfvén Eigenmodes(TAE)
- Drive/damping rate
 - Energetic particles drive/damping
 - Thermal electron/ion Landau damping
 - Continuum damping
 - Collisional damping of trapped electrons
 - Radiative damping

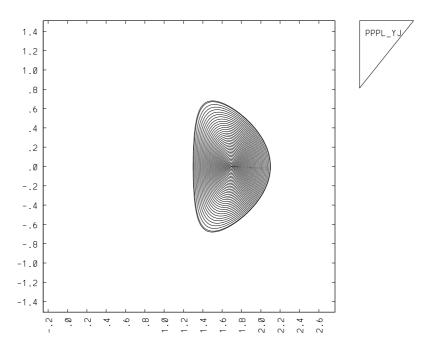


Figure 1. EAST shape parameter: $R_0=1.7m,~a=0.4m,$ elongation $\kappa=1.7,$ triangularity $\delta=0.5.$

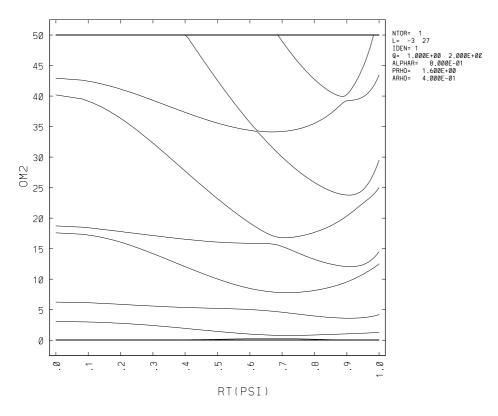


Figure 2. n=1 continuum spectrum

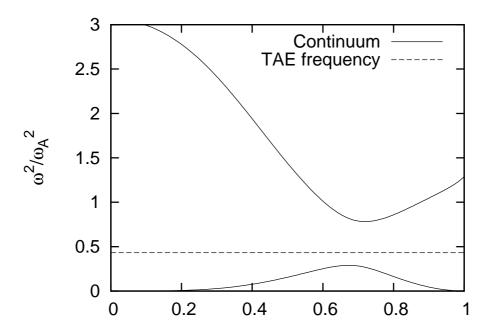


Figure 3. n=1 continuum spectrum and TAE frequency

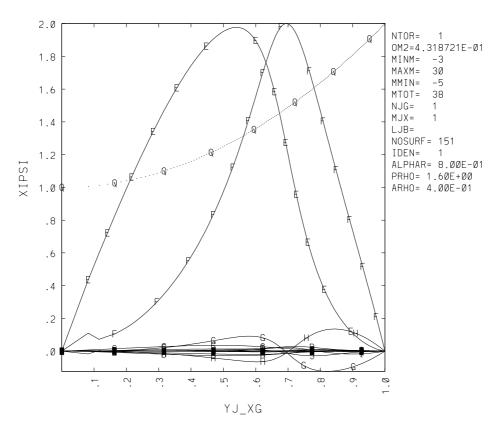


Figure 4. Mode structure of n=1 TAE. Dominant poloidal modes are m=1 and m=2

Example of NOVA-K output

```
gam ecoll -9.784951544844977E-005
gam eLandau -2.373366177356202E-003
gam DLandau -8.925676315388741E-004
gam TLandau -1.885105163939624E-010
gam HLandau -5.120318917926147E-009
gam CLandau 0.000000000000000E+000
radiative damping 0.00000000000000E+000
Fast ion beta, beta h 8.00000000000000E-002
Fast ion growth rate without FLR, gam h -1.319028439767093E-002
Fast ion growth rate with FLR, gam hFLR -1.950236033379125E-003
beta hcrit=-beta h*(all dampings above)/gam h
or beta_hcrit=-beta_h*(all dampings above)/gam hFLR
```

Thanks!