

# Microscopic ensemble bootstrap in phase space

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# Abstract

The bootstrap method which has been studied under many quantum mechanical models turns out feasible in microcanonical ensemble as well. While the approach in [Y. Nakayama, Modern Physics Letters A **37**, 10.1142/s0217732322500547 (2022)] produces a sector when Energy is negative, in this paper we report a method that has stronger constraints thus resulting in a smaller region. We also study other models to demonstrate the effectiveness of our method.

## I. INTRODUCTION

Bootstrap, a novel yet promising method has been studied in quantum mechanics recently, is a way to utilize the very general self-consistency condition and solve the system numerically [1][2][3], many have reported the accuracy and high precision of the method. And in some case the bootstrap evolves into Dirac's ladder operator approach and can be solved analytically, suggesting some underlying mechanism of this method [4].

Although widely studied in quantum mechanics, one can also apply the method into microcanonical ensembles as the fundamental relations of bootstrap are still applicable and have their classical correspondence (3)(4)(5) as  $\hbar \rightarrow 0$ .

Nakayama [5] first introduced this method into the classical scenario, despite reporting the feasibility of the approach, when bootstrapping the double well potential he mentioned a peninsula in  $E < 0$  which doesn't converge even for larger bootstrap matrices. In this paper, we use a different approach that incorporates more information in phase space and thus exhibiting a much stronger constraint, we will see that the result of double well bootstrap in phase space cancels the sector region in  $E < 0$  compared to the  $x$  only bootstrap.

We also investigate coulomb potential, a harmonic oscillator and a non relativistic toda model, the first two can be solved analytically(here by analytically we mean that the averages of observables can be written in a function of  $E$ ) via the bootstrap approach which, however, are trivial cases. As for the non relativistic toda model, our approach in phase space once again demonstrates more powerful restriction, maybe over powerful that we can merely see few points in the result if the sample isn't big enough.

Yet a stronger constraint our approach may have, it still cannot converge to the exact solution in some places, like the non relativistic toda model. Not to mention that our approach consumes much more computing power, the size of our bootstrap matrix is  $O(n^4)$ .

But the benefit is that we can easily achieve high precision when the result converges.

## II. MICROCANONICAL ENSEMBLE BOOTSTRAP

### A. Bootstrap Equations and Matrices

Starting with the Hamiltonian we have

$$H = \frac{p^2}{2M} + V(x) \quad (1)$$

for microcanonical ensemble the average of an observable  $\mathcal{O}(x, p)$  is

$$\langle \mathcal{O}(x, p) \rangle = \frac{\int dx dp \mathcal{O}(x, p) \delta(E - H)}{\int dx dp \delta(E - H)} \quad (2)$$

and we can easily find that

$$\langle \{H, \mathcal{O}\} \rangle = 0 \quad (3)$$

$$\langle H\mathcal{O} \rangle = E \langle \mathcal{O} \rangle \quad (4)$$

here  $\{H, \mathcal{O}\}$  is the poisson bracket. As for the positivity constraints, obviously that for any observable  $\mathcal{O}$  we would have

$$\langle \mathcal{O}^* \mathcal{O} \rangle \geq 0 \quad (5)$$

and by writing the observable as a polynomial of certain observable  $o$ ,  $\mathcal{O} = \sum_{i=0}^k a_i o^i$ , one can construct a bootstrap matrix which can be defined as

$$\mathcal{M}_{ij} = \langle (o^*)^i o^j \rangle, \quad i, j = 0, 1, \dots, k \quad (6)$$

we can then rewrite the constraints (5) with matrix and vectors

$$\alpha^\dagger \mathcal{M} \alpha \geq 0 \quad (7)$$

as (7) should hold true for any vector  $\alpha$ , the bootstrap matrix  $\mathcal{M}$  must embody the positive semi definiteness i.e.  $\mathcal{M} \succeq 0$ , which is essentially an eigenvalue problem

$$\forall (\mathcal{M})_{\text{eigenvalue}} \geq 0 \quad (8)$$

When the observable  $\mathcal{O}$  is a coupling of two observables, say,  $A$  and  $B$

$$\mathcal{O} = \sum_{i,j=0}^{k-1} a_i b_j A^i B^j \quad (9)$$

$$\mathcal{O}^* \mathcal{O} = \sum_{i_1, j_1, i_2, j_2=0}^{k-1} a_{i_1}^* b_{j_1}^* (B^*)^{j_1} (A^*)^{i_1} A^{i_2} B^{j_2} a_{i_2} b_{j_2} \quad (10)$$

here we define two auxiliary matrices  $\mathcal{M}_{ij}^0 = (B^*)^j (A^*)^i$  and  $\mathcal{M}_{ij}^1 = A^i B^j$ , the constraints may be written as

$$\mathcal{M}_{ij} = \langle (\mathcal{M}^0 \otimes \mathcal{M}^1)_{ij} \rangle, \quad i, j = 0, 1, 2, \dots, k^2 - 1 \quad (11)$$

$$\mathcal{M} \succeq 0 \quad (12)$$

## B. Recursion Formula

Taking  $\mathcal{O}$  as  $x^m p^n$  and substituting Hamiltonian(1) into (3) and (4), we immediately obtain

$$\begin{aligned} n \langle \frac{dV}{dx} x^m p^{n-1} \rangle &= 2Em \langle x^{m-1} p^{n-1} \rangle - 2m \langle V x^{m-1} p^{n-1} \rangle \\ E \langle x^m p^n \rangle &= \frac{1}{2m} \langle x^m p^{n+2} \rangle + \langle V x^m p^n \rangle \end{aligned} \quad (13)$$

This would do the trick for the potential with polynoial of  $x$ , when the potential is in the form of exponentials, we need to take  $\mathcal{O}$  as  $e^{mx} p^n$

$$\begin{aligned} n \langle \frac{dV}{dx} e^{mx} p^{n-1} \rangle &= 2Em \langle e^{mx} p^{n-1} \rangle - 2m \langle V e^{mx} p^{n-1} \rangle \\ E \langle e^{mx} p^n \rangle &= \frac{1}{2M} \langle e^{mx} p^{n+2} \rangle + \langle V e^{mx} p^n \rangle \end{aligned} \quad (14)$$

## C. Methodology Framework

With the recursion formula(13) or (14) and a few initial values we can construct a whole bootstrap matrix  $\mathcal{M}$ , and by testing the positive semi definiteness of the matrix the validity of the initial values can be determined. Doing so over all the possible initial values will eventually generate the allowed regions restricted by positivity constraint.

## III. NUMERICAL EXAMPLES

### A. Double Well Potential

The Hamiltonian of a double well potential can be written as

$$H = p^2 - x^2 + x^4 \quad (15)$$

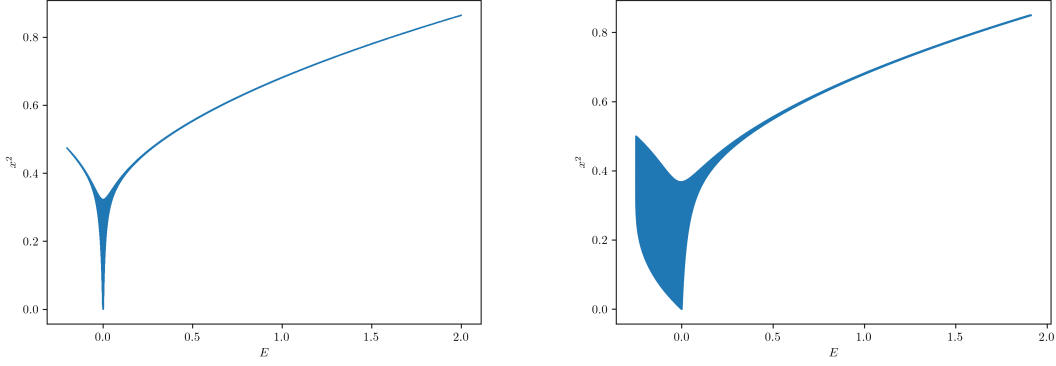


FIG. 1.  $xp$  bootstrap with  $k = 5$ (left),  $x$  only bootstrap with  $k = 9$ (right)

taking  $M = 2$  and with (13) we have

$$2(2n + m) \langle x^{m+3} p^{n-1} \rangle = 2(m + n) \langle x^{m+1} p^{n-1} \rangle + 2mE \langle x^{m-1} p^{n-1} \rangle \quad (16)$$

$$\langle H x^m p^n \rangle = E \langle x^m p^n \rangle = \langle x^m p^{n+2} \rangle - \langle x^{m+2} p^n \rangle + \langle x^{m+4} p^n \rangle \quad (17)$$

plus  $\langle \{H, x^m\} \rangle = 0$ , we get

$$\langle x^{m-1} p \rangle = 0 \quad (18)$$

and for simplicity, we here assume that the average of  $x$  to the odd powers is 0

$$\langle x^m \rangle = 0, \quad \text{for all odd } m \quad (19)$$

with (16) (17) (18) (19) plus the initial parameters  $E$  and  $\langle x^2 \rangle$  we can construct the bootstrap matrix  $\mathbf{M}$ . The result when  $k = 5$  is shown in Fig. 1. We also reproduced the result in [5] to make a comparison.

In contrast with [5], our result shows no peninsula in where  $E < 0$ , as we include the information of the momentum. When  $E$  turns negative, the curve manifests a similar behavior as the positive part. In our program the scale of the bootstrap matrix is  $k^4$ , which is much larger compared to the single observable bootstrap program with the scale of  $k^2$ . So here the size of our bootstrap matrix is 625, about seven times to the  $x$  bootstrap. But the region still doesn't shrink even when we set  $k = 25$  for the  $x$  only bootstrap program, so we conclude that this bootstrap in phase space does have a much stronger constraints than the single observable one.

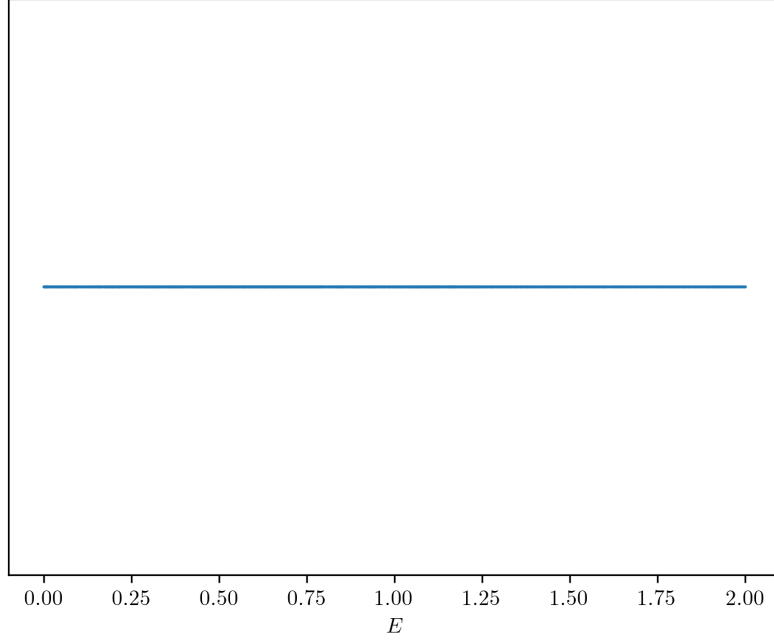


FIG. 2.  $\langle x^m p^n \rangle$  bootstrap ( $k = 4$ ) for harmonic oscillator

### B. Harmonic Oscillator and Coulomb Potential

Consider a harmonic oscillator, its hamiltonian is

$$H = p^2 + x^2 \quad (20)$$

and again, we can obtain the recursion formula

$$\begin{aligned} 2n \langle x^{m+1} p^{n-1} \rangle &= 2Em \langle x^{m-1} p^{n-1} \rangle - 2m \langle x^{m+1} p^{n-1} \rangle \\ E \langle x^m p^n \rangle &= \langle x^m p^{n+2} \rangle + \langle x^{m+2} p^n \rangle \end{aligned} \quad (21)$$

since  $\langle x^0 p^0 \rangle = 1$  and  $\langle x \rangle = 0$ , and (18) also holds true, we only need the initial energy  $E$  to bootstrap. The result of bootstrap is shown in Fig.2, as we sample  $E$  from negative to positive, we can see that the negative energies are rejected by the bootstrap program. While the  $x$  only bootstrap can also generate the same plot, our approach reaches an accuracy 100 times higher than the  $x$  only bootstrap, because to obtain the same result, our method requires sampling more than 100 times as many data points. As for the coulomb potential, assume the Hamiltonian

$$H = p^2 - \frac{1}{r} + \frac{1}{r^2} \quad (22)$$

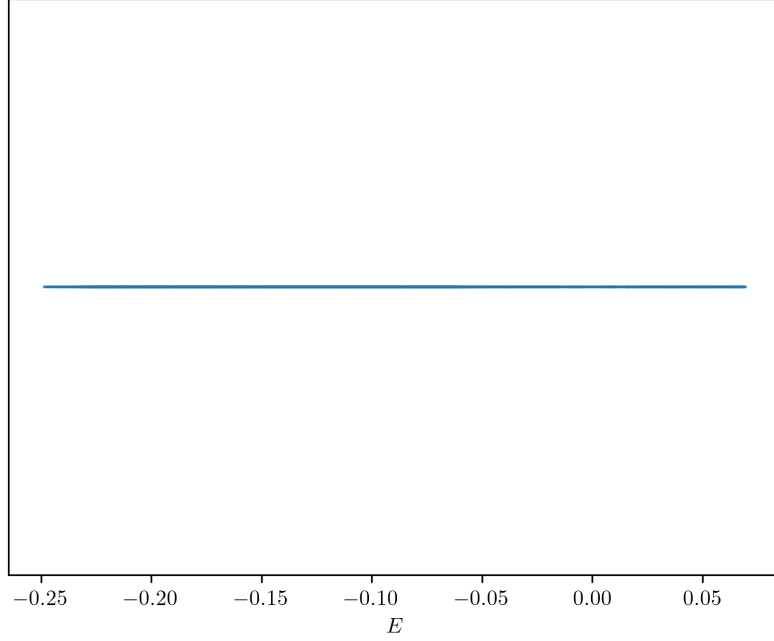


FIG. 3.  $\langle x^m \rangle$  bootstrap ( $k = 8$ ) for coulomb potential

here  $-\frac{1}{r} + \frac{1}{r^2}$  is the effective potential. The recursion equations are

$$\begin{aligned} 2mE \langle r^{m-1} p^{n-1} \rangle &= 2(m-n) \langle r^{m-3} p^{n-1} \rangle + (n-2m) \langle r^{m-2} p^{n-1} \rangle \\ E \langle r^m p^n \rangle &= \langle r^m p^{n+2} \rangle - \langle r^{m-1} p^n \rangle + \langle r^{m-2} p^n \rangle \end{aligned} \quad (23)$$

Substituting  $m = n = 1$  into the first equation we can get

$$2E = -\langle r^{-1} \rangle \quad (24)$$

which is Virial theorem. With the equation (18), again we only need  $E$  to bootstrap the coulomb potential. The result is shown in Fig3. However, this time we failed to find any points using  $x^m p^n$  bootstrap, we speculate that this is due to insufficient numerical precision. But the  $x^m$  bootstrap is doing great, at least we successfully find the allowed energy.

### C. Non-Relativistic Toda model

For a non-relativistic Toda model, the Hamiltonian may be written as

$$H = p^2 + e^x + e^{-x} \quad (25)$$

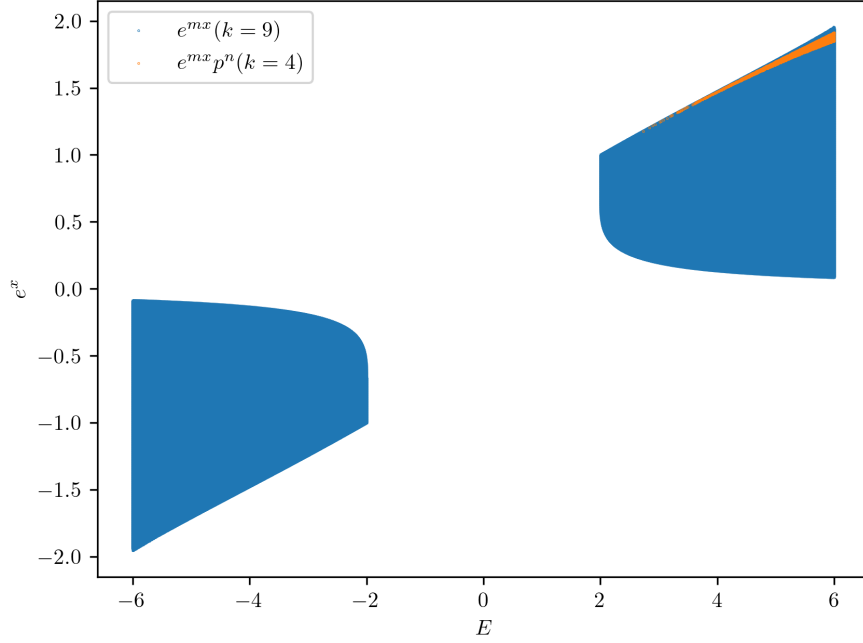


FIG. 4. Bootstrap for non-relativistic Toda model

the recursion equations are

$$\begin{aligned}
 (n + 2m) \langle e^{m+1} p^{n-1} \rangle &= 2E \langle e^m p^{n-1} \rangle + (n - 2m) \langle e^{m-1} p^{n-1} \rangle \\
 E \langle e^{mx} p^n \rangle &= \langle e^{mx} p^{n+2} \rangle + \langle e^{(m+1)x} p^n \rangle + \langle e^{(m-1)x} p^n \rangle
 \end{aligned}
 \tag{26}$$

and studying  $\langle \{H, e^{mx}\} \rangle$  we have

$$\langle e^{(m-1)x} p \rangle = 0
 \tag{27}$$

using the initial parameters  $E$  and  $\langle e^x \rangle$ , the bootstrap result is shown in Fig.4

Using the two observables bootstrap shows a much stronger constraint than the single one, in this non-relativistic Toda mode the  $e^{mx}$  bootstrap even fails to reject the negative  $E$ , and resulting in a much larger region. Our approach may have better precision, yet the allowed region still tends to diverge when  $E$  goes large5.

#### IV. CONCLUSIONS

In this paper, we study two different bootstrap approaches in four different models, our approach outperforms the single observable bootstrap program in some models e.g. double



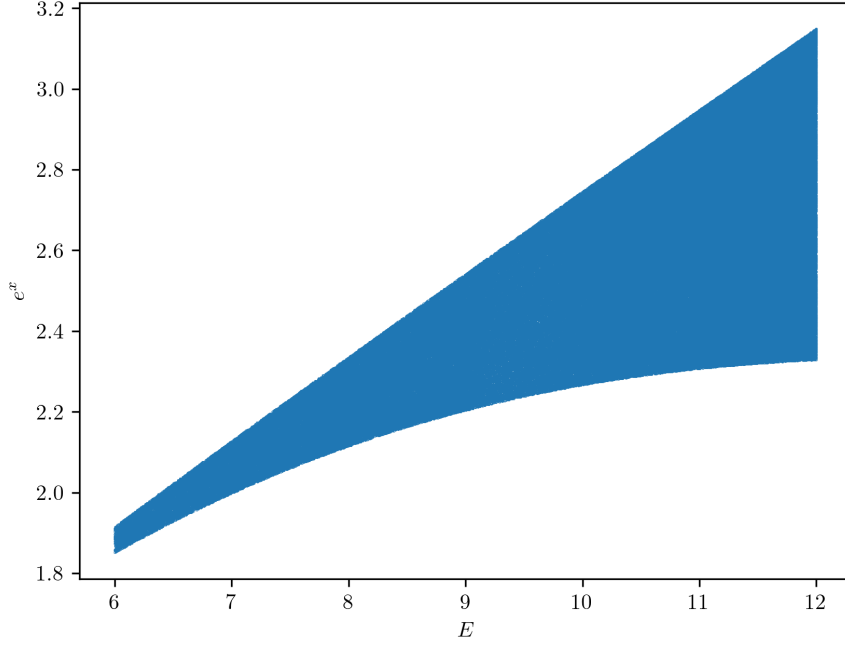


FIG. 5. Allowed region diverges as  $E$  grows

well, non-relativistic Toda model, while in other two models our bootstrap in phase space fails to achieve a satisfactory result, the single observable one already restricts the model pretty well.

Although our approach performs better and achieves a much higher precision, it still fails to converge to the exact solution, and even tends to diverge. Like  $E \rightarrow 0$  in the double well potential,  $E \rightarrow \infty$  in the non-relativistic Toda model, we may still be missing some information to pin point the final answer.

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