GR notes

Yucun Xie

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Conventions

- 1. Greek index (e.g. α, β, μ, ν) take value from $\{0, 1, 2, 3\}$.
- 2. Events denoted by cursive capitals (e.g. $\mathscr{A}, \mathscr{B}, \mathscr{E}$).
- 3. $(x^0, x^1, x^2, x^3) \equiv (t, x, y, z) \equiv x^{\alpha}$
- 4. Latin index (e.g.i, j, k) take value from $\{1, 2, 3\}$.
- 5. New unit that speed of light c = 1
- 6. Einstein summation convention $ds^2 = g_{\mu\nu}x^{\mu}x^{\nu} = \sum_{\mu=0}^{3} \sum_{\nu=0}^{3} g_{\mu\nu}x^{\mu}x^{\nu}$

1 Special Relativity

1.1 Spacetime

Definition 1.1. Inertial coordinate

The coordinate system must satisfy three property to be consider inertial coordinat:

- 1. The distance between two points are independent of time.
- 2. The clocks at every points ticking off time coordinate t at same rate.
- 3. The geometry of space is always Euclidean (flat).

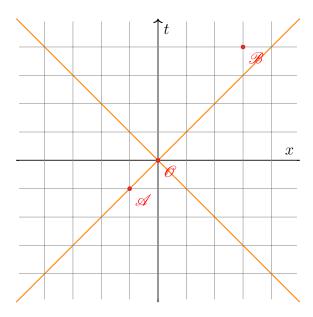


Figure 1: two events with coordinate (-1, -1, 0, 0) and (4, 3, 0, 0). Orange line is light's worldline.

The event in 4-D spacetime is defined by a set of coordinate (t, x, y, z). For simplicity, we assume those events have y = 0, z = 0 so that we can draw a 2D graph to represent them.

Analog to Euclidean geometry, just like the euclidean distance $\Delta l^2 = \Delta x^2 + \Delta y^2 + \Delta z^2$, we define the spacetime interval $\Delta s^2 = -\Delta t^2 + \Delta x^2 + \Delta y^2 + \Delta z^2$.

Remark. There are a lot different conventions to define the sign of interval, here we just use the popular one (-,+,+,+).

Example.

Interval for the two events in Figure 1 is $\Delta s^2 = -\Delta t^2 + \Delta x^2 + \Delta y^2 + \Delta z^2 = -9$.

The universality speed of light means that $\frac{\Delta r}{\Delta t} = \frac{\sqrt{\Delta x^2 + \Delta y^2 + \Delta z^2}}{\Delta t} = 1$ are always hold, then we can then write the interval $\Delta s^2 = -\Delta t^2 + \Delta x^2 + \Delta y^2 + \Delta z^2 = 0$. This experimental fact yield all law of special relativity.

When the interval Δs^2 is less than 0, we call the separation bewteen events is **timelike**; When the interval Δs^2 is equal to 0, we call it **lightlike** or null; When the interval Δs^2 is greater than 0, we call it **spacelike**.

- 1.2 Tensor
- **Energy and Momentum**

Differential Geometry

- **Geodesics** 2.1
- **Riemann Tensor**
- 2.3 Connection

Proof. Here is a proof shows that connection not a tensor by show connection does not obey tensor transformation law.

$$\begin{split} \nabla_{\beta'} e_{\alpha'} &= \Gamma_{\alpha'\beta'}^{\gamma'} e_{\gamma'} \\ &= \frac{\partial x^{\lambda}}{\partial x^{\beta'}} \nabla_{\lambda} (\frac{\partial x^{\mu}}{\partial x^{\alpha'}} e_{\mu}) \\ &= \frac{\partial x^{\lambda}}{\partial x^{\beta'}} (\frac{\partial}{\partial x^{\lambda}} \frac{\partial x^{\mu}}{\partial x^{\alpha'}} e_{\mu} + \frac{\partial x^{\mu}}{\partial x^{\alpha'}} \Gamma_{\mu\lambda}^{\gamma} e_{\gamma}) \\ &= \frac{\partial x^{\lambda}}{\partial x^{\beta'}} \frac{\partial}{\partial x^{\lambda}} \frac{\partial x^{\mu}}{\partial x^{\alpha'}} e_{\mu} + \frac{\partial x^{\lambda}}{\partial x^{\beta'}} \frac{\partial x^{\mu}}{\partial x^{\alpha'}} \Gamma_{\mu\lambda}^{\gamma} e_{\gamma} \\ &= \frac{\partial x^{\lambda}}{\partial x^{\beta'}} \frac{\partial}{\partial x^{\lambda}} \frac{\partial x^{\mu}}{\partial x^{\alpha'}} \frac{\partial x^{\gamma'}}{\partial x^{\mu}} e_{\gamma'} + \frac{\partial x^{\lambda}}{\partial x^{\beta'}} \frac{\partial x^{\mu}}{\partial x^{\alpha'}} \frac{\partial x^{\gamma'}}{\partial x^{\gamma'}} \Gamma_{\mu\lambda}^{\gamma} e_{\gamma'} \end{split}$$

which yield

$$\Gamma^{\gamma'}_{\alpha'\beta'} = \frac{\partial x^{\lambda}}{\partial x^{\beta'}} \frac{\partial}{\partial x^{\lambda}} \frac{\partial x^{\mu}}{\partial x^{\alpha'}} \frac{\partial x^{\gamma'}}{\partial x^{\mu}} + \frac{\partial x^{\lambda}}{\partial x^{\beta'}} \frac{\partial x^{\mu}}{\partial x^{\alpha'}} \frac{\partial x^{\gamma'}}{\partial x^{\gamma'}} \Gamma^{\gamma}_{\mu\lambda}$$
 There is an extra term in transformation of connection, so connection is not a tensor.

A Topology Space

B Manifolds