Tillering ___ 1st April 2019

- . Common aims of filtering: x finding a signal in the duta
 - * separating natic and signal
 - * compressing the data
 - # fillers can also be optimized
 - * Deural Dets essentially have a fillers" stared in their layers.
- · Linear fillers: the general fillering equation
 - * dota stream x(E)

* Fillening function FCE)

Tillering function
$$F(E)$$
 $\times_{\overline{T}}(E) = \int \overline{F(E')} \times (t-E') dt'$

The fillered data the filler

of filler output"

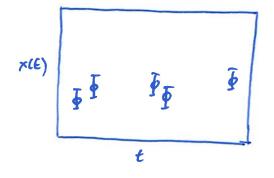
. Integrals are linear operators; and most data are discrete => one often sees filters uniten in terms of linear algebra: rould also be a matrix

Filered data not a scalar product (else the output would not be a cector)

•
$$S_N = \langle \vec{F} \cdot \vec{x} \rangle^2$$
 is an often found, signal to noise "

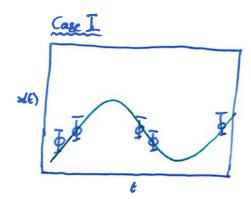
we continue to discuss linear fillers.

- · Fillers can improve (or facilitate) an analysis, if they are sensibly constructed and used.
 - F Wiener filler: is a maximum-a-posteriori solution to
 Gaussian fields under Gaussian shot naise (see script)
 - * Spoursify fillering: adds prior in formation into the analysis, of what sensible signal shapes" are.
- · Sponsity and filtering: Imagine 5 data points

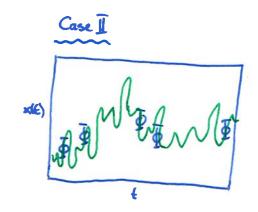


>> How much do these 5 paints tell as about a potential signal?

>> It depends on how smooth the signal is?



no the green symul is very smooth, 5 darks points measure it reliably



whe green signal is wildly variable and the 5 dates points are insufficient to constrain this signal.

If we have a priori a sensible reason to believe that there are no high frequencies in our signal, i.e. it looks more like Case I, then puthing in this prior belief will improve the analysis.

As wavelet fillers allow to do this.

Example: Wave let fillering as done in gravitational wave research

The usual Fourier transform:
S_F(E) =
$$\int_{-\infty}^{100} S(\omega) e^{i\omega t} d\omega$$

wo a bit entreme: either in w-space or in t-space

. The short-time Fourier teams form:

$$S(\mathcal{E},\omega) = \int_{\mathcal{E}} S(\varepsilon) g(\varepsilon-\varepsilon) e^{-i\omega t} dt$$

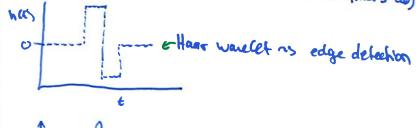
3(1-17) In both t-and w-space.

~> Before etiut was our Besis: the Fourier basis. ~> Now declare g(t-2) e-int our new "basis".

"Basis": a set of fundhers which projects out interesting elements from the data. * " over complete basis": one that is not authornarmal, and strictly speciting Contains too many basis elements coften adapted to real situations).

· Wavelet transform:

is there is a multitude of famous were lets (Goldon, Marcet 1985; Hany wavelets...) is all have in common that they filter patterns from the data that look like the warlet warlet



Mexican hat files

W Gabor-Marlet warrlets

The matched filler

- "Typically astronomers and physicists already know which signal that are searching for in the dota. It is then unnecessary to use any ready-monde filter sach as warelet filters, since we can as well create our own gotimal filter. The most famous of those is the matched filter.
- " applical": if the sought signal is in the data, then this filler is the most sensitive with respect to it (apicles it up the fastest")
- * fix = apply filler f to data z"

Derivation, matched filter:

- · 2 Prerequisites: 1) f
 - 1) for optimization: the Couchy-Schwaz inequality: $|\vec{x}^T\vec{y}|^2 \leq (\vec{x}^T\vec{x})(\vec{y}^T\vec{y})$
 - 2) Symmetric positive matrices have a carrigue matrix root:

 M=M12 M12 and M12 invertible, such that M12 M-1/2 = 1
 - . Aim, maximize signal to notice (SKU) for additive notice,

$$\vec{x} = \vec{s} + \vec{n}$$
, $\langle \vec{n} \rangle = 0$, $\langle \vec{n} \vec{n} \rangle = C$
dotte shear noise

· filking: ftx = fts + fth

correct fulse detections (interpreting noise patterns as signal)

no the mass of the filler (which we want to be as low as possible) is then $\langle \vec{f}^{\dagger} \vec{n} \ \vec{n} \vec{\tau} \vec{f} \rangle = \vec{f}^{\dagger} C \vec{f} =: 0$ (expected false detections)

as the corresponding signal S to this 10 is then 1ft &12 (mind the square).

Our signal to noise is then expected to be:

(SM) =
$$\frac{|\vec{f}| \vec{s}|^2}{|\vec{f}|^2 |\vec{f}|^2}$$
 top and bottom measure in different units, adept those weether from the \vec{f} = $\frac{|\vec{f}| \vec{f} \cdot \vec{f}|^2}{|\vec{f}|^2 |\vec{f}|^2}$ (see that \vec{f} = $\frac{|\vec{f}| \vec{f} \cdot \vec{f}|^2}{|\vec{f}|^2 |\vec{f}|^2}$ = $\frac{|\vec{f}| \vec{f}|^2}{|\vec{f}|^2 |\vec{f}|^2}$

Now we are greedy, and demand that the signal to note rather shall satisfy the Cauchythwarz inequality. This is the optimization step.

(cust) 1 (cust)

= ST CTS => (SKU) = ST CTS after optimization

so now both use the same units

Now we have to find a filler which substies our demands:

$$|\langle S_{NOY} = |\hat{f}^{T}\hat{s}|^{2} \stackrel{!}{=} |\hat{s}^{T}c^{-1}\hat{s}|^{2}$$

$$|\hat{f}^{T}c\hat{f}| \stackrel{!}{=} |\hat{s}^{T}c^{-1}\hat{s}|^{2}$$

$$\Rightarrow$$
 $A = C^1$ solves this

"> So our sought matched filter is $f = C^{-1}\vec{S}$, the inverse-variance very hed signal. ~> also known as a correlating with a inverse variance weighted template"

Matched fillering : in the

X = 0+1 € 3 9 X=0+1 33 } x= (ĥ(ω), ĥ(ω), ... ĥ(ω))

in the tubulal sheet luse are x already. variable x: M= x Mo, here I wisk Myo, Since

fr= (h(2, 1/2, 1/2, 1/2, 1/2, 1/2, 1/2, 1/2), ...

\$ Sea-ching for Gurech collapse him

「= (h(で, 光, 鬼), h(で, 水の))…) {T= (h(2, M'mo, E'), ... seasing for

correct mass 14

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