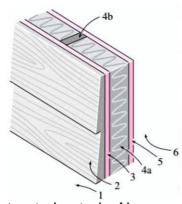
Determine the overall unit thermal resistance (the R  $_{value}$ ) of a wood frame wall that is built around 38-mm 90-mm wood studs with a center-to-center distance of 400 mm. The 90-mm-wide cavity between the study is filled with ure than erigidin sulation. The inside is finished with 13-mm gypsum wall board and the outside with 13-mm plywood and 13-mm 200-mm wood bevel lapped siding.



Theinsulated cavity constitutes 75 percent of the heat transmission are awhile the study, plates, and sills constitute

21 percent. The headers constitute 4 percent of the area, and they can be treated as studs. Also, determine the rate of heat loss through the walls of a house whose perimeter is 50 mand wall height is 2.5 min Las Vegas, Nevada, whose winter design temperature is -2 C. Take the indoor design temperature to be 22 Candas sume 20 percent of the wall area is occupied by glazing (glass) (Wall without glass 80%).

Unit thermal resistance (the R-value) of common components used in buildings

	R-Value			R-Value	
Component	$m^2 \cdot {^\circ\!C/\!W}$	$ft^2 \cdot h \cdot {^o}F/Btu$	Component	m² · °C/W	ft² · h · °F/Bt
Outside surface (winter)	0.030	0.17	Wood stud, nominal 2 in × 6 in		
Outside surface (summer)	0.044	0.25	(5.5 in or 140 mm wide)	0.98	5.56
Inside surface, still air	0.12	0.68	Clay tile, 100 mm (4 in)	0.18	1.01
Plane air space, vertical,			Acoustic tile	0.32	1.79
ordinary surfaces			Asphalt shingle roofing	0.077	0.44
$(\varepsilon_{\text{eff}} = 0.82)$ :			Building paper	0.011	0.06
13 mm ( $\frac{1}{2}$ in)	0.16	0.90	Concrete block, 100 mm (4 in):		
20 mm ( $\frac{3}{4}$ in)	0.17	0.94	Lightweight	0.27	1.51
40 mm (1.5 in)	0.16	0.90	Heavyweight	0.13	0.71
90 mm (3.5 in)	0.16	0.91	Plaster or gypsum board,		
Insulation, 25 mm (1 in)			13 mm (½ in)	0.079	0.45
Glass fiber	0.70	4.00	Wood fiberboard, 13 mm ( $\frac{1}{2}$ in)	0.23	1.31
Mineral fiber batt	0.66	3.73	Plywood, 13 mm ( $\frac{1}{2}$ in)	0.11	0.62
Urethane rigid foam	0.98	5.56	Concrete, 200 mm (8 in)		
Stucco, 25 mm (1 in)	0.037	0.21	Lightweight	1.17	6.67
Face brick, 100 mm (4 in)	0.075	0.43	Heavyweight	0.12	0.67
Common brick, 100 mm (4 in)	0.12	0.79	Cement mortar, 13 mm (1/2 in)	0.018	0.10
Steel siding	0.00	0.00	Wood bevel lapped siding,		
Slag, 13 mm ( $\frac{1}{2}$ in)	0.067	0.38	13 mm × 200 mm		
Wood, 25 mm (1 in)	0.22	1.25	(1/2 in × 8 in)	0.14	0.81
Wood stud, nominal 2 in ×					
4 in (3.5 in or 90 mm wide)	0.63	3.58			

	m² °C	m² °C
	W	W
	Insulation	Wood Studs
1. Outside air surface	0.03	0.3
2. Wood bevel (13*200) mm	0.14	0.14
3. Plywood (13 mm)	0.11	0.11
4.a Urethane rigid foam	90	/
	0.98* = 3.528	
4.b Wood Studs (90 mm)	/	0.63
5. Gypsum Wallboard (13 mm)	0.079	0.079
6. Inside air surface	0.12	0.12

The R\_unit values are respectively:

R' with insulation =	4,007 <sup>m² °C</sup> / <sub>W</sub>	
R' with wood studs =		1,379 <sup>m²°C</sup> / <sub>W</sub>

$$R'_{tot} = R*A = R = \frac{R'_{tot}}{A}$$
 (R' is R<sub>unit</sub>)

$$R_{unit} = \frac{1}{U} = \sum_{R'} \frac{1}{R'} = U$$

$$\frac{1}{R \text{ tot}} = \frac{1}{R \text{ insul}} + \frac{1}{R \text{ wood}} = > \frac{A \text{ tot}}{R' \text{ tot}} = \frac{A \text{ insul}}{R' \text{ insul}} + \frac{A \text{ wood } R'}{\text{wood}}$$

$$U_{tot} * A_{tot} = U_{ins} * A_{ins} + U_{wood} * A_{wood}$$

$$U_{tot} = U_{ins}^* \frac{Ainsul}{Atot} + U_{wood}^* \frac{A wood A}{tot}$$

 $U_{tot} = U_{ins} * 0.75 + U_{wood} * 0.25$  (insulated cavity -> 75% = 0.75 of the heat transmission area studs, plates, sills + headers (/studs) -> 21+4% = 0.25)

$$U_{\text{insul}} = \frac{1}{R'_{\text{insul}}} = \frac{1}{4.007} = 0.2496 \frac{W}{m^{2} \circ C}$$

$$U_{\text{wood}} = \frac{1}{R' \text{wood}} = \frac{1}{1.379} = 0,7252 \frac{W}{m^2 \, ^{\circ} \text{C}}$$

$$U_{\text{tot}} = 0.2496 * 0.75 + 0.7252 * 0.25 = 0.1872 + 0.1813 = 0.3685 \frac{W}{\text{m}^2 \circ \text{C}}$$

 $A_{tot} = 50 * 2.5 * 0.8 = 100 \text{ m}^2$  (0.8 is the 80% of the wall without glass)

$$\Delta T = 22 - (-2) = 24 \, ^{\circ}C$$

$$Q_{tot} = U_{tot} * A_{tot} * \Delta T = 884.4 W$$

## THE RADIATION AND THE RADIATIVE HEAT TRANSFER

Inphysics, the term radiation is generally used to indicate the set of phenomena characterized by the transport of energy in space, where there are different surroundings conditions (in which there is n't thermal equilibrium). Typical examples of radiation are light and heat.

So, radiation differs from conduction and convection because it doesn't require the presence of a material medium to take place.

A particular type of radiation is constituted by electromagnetic waves: they represent the energy emitted by matter as a result of the changes in the electronic configurations of the molecules.

Electromagnetic wavestrans portenergy, and this energy, once in contact with another medium, produces a change in heat. An increase in temperature is caused by an increase in energy in the movement of the particles that compose it.

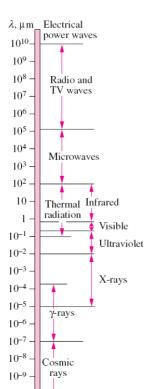
Electromagnetic waves that transport energy is characterized by a constant, finite vacuum speed of approximately 299792 km/s (rounded to 3.0 x 108 m/s), indicated by c, from a particular frequency v and its wavelength, indicated with the symbol  $\lambda$ .

The frequency vis defined as the number of wavelengths that pass in a second for a given point in space:

$$v = \frac{c}{\lambda}$$

Higher is the wave frequency value, smaller its wavelength is, so the speed of light is constant for every electromagnetic radiation in a vacuum.

Visible light (or better, visible light for the human eye) is the electromagnetic radiation with a wavelength ranging from about 380 nm of violet to about 760 nm for red.



The electromagnetic wave spectrum

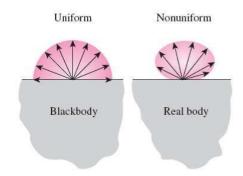
## **BLACK BODY:**

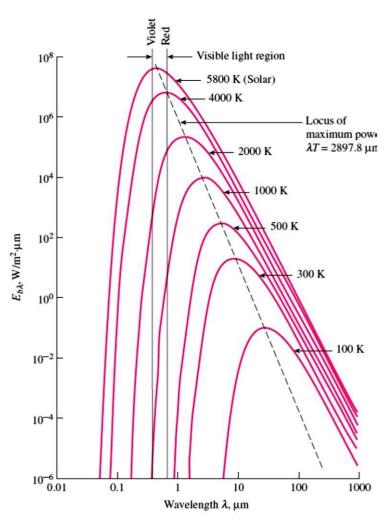
Ablackbody is an *idealized body* that *emits the maximum* amount of radiation: it is a *perfect emitter* and absorber of all incident radiation.

The radiation energy emitted by a black body: E (T) =  $\sigma$ T<sup>4</sup> [ Black body emissive power  $\sigma$  b

$$\sigma = 5.670*10^{-8}$$
  $\frac{W}{\text{m} * \text{K}^4}$ 

Stefan-Boltzmann constant





The graphic represents the variety of the blackbody emissive power with wavelength for several temperatures.

The axis of wavelength is on a logarithmic scale, while the axis of the emissive power is on a linear scale.

The amount of <u>emitted radiation</u> increases with increasing temperature.

The radiation emitted by the sun, which is considered to be a blackbody at 5800 K, reaches its *peak* in the visible light region of the spectrum.

Surfaces at T < 800 K emit almost entirely in the infrared region and thus are not visible to the eye unless they reflectlightcomingfromothersources.