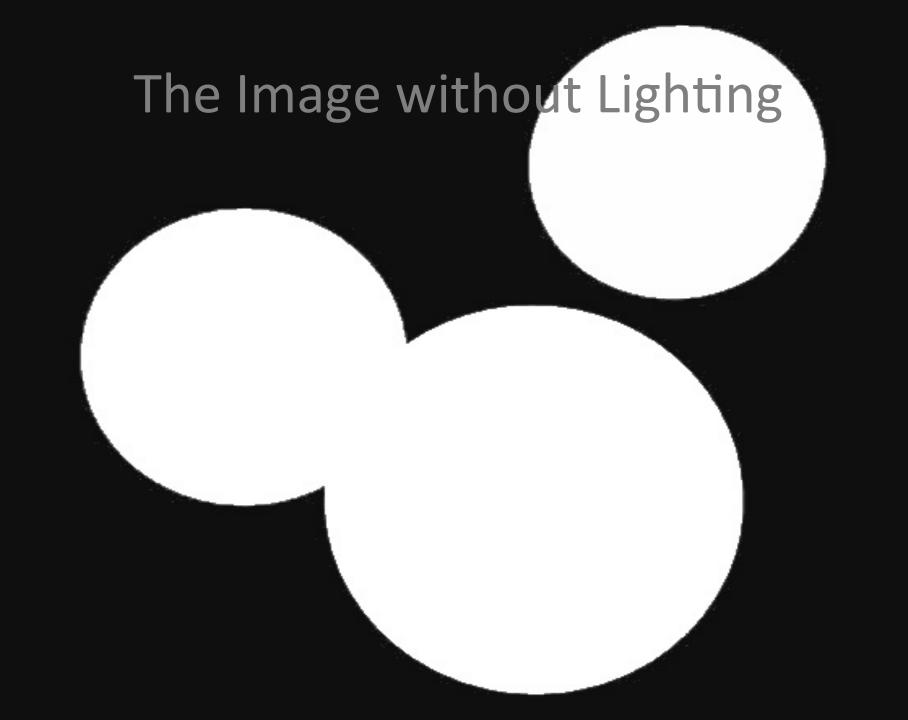
#### **Local Illumination**

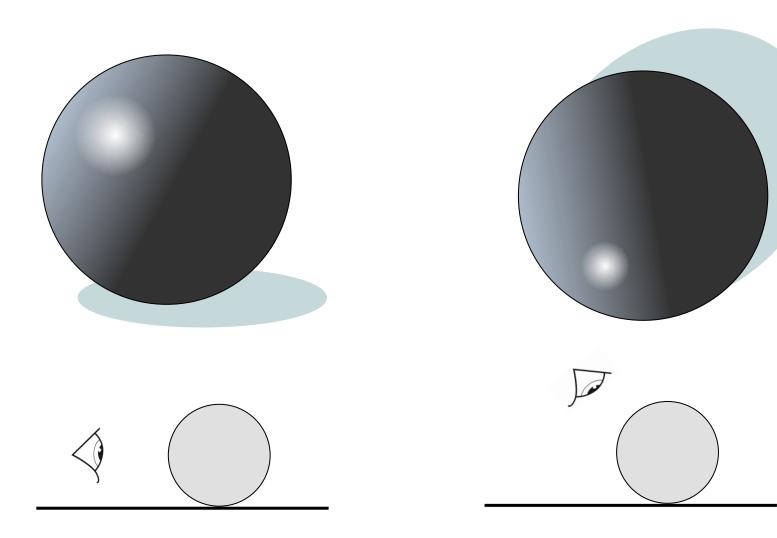
**Tobias Ritschel** 



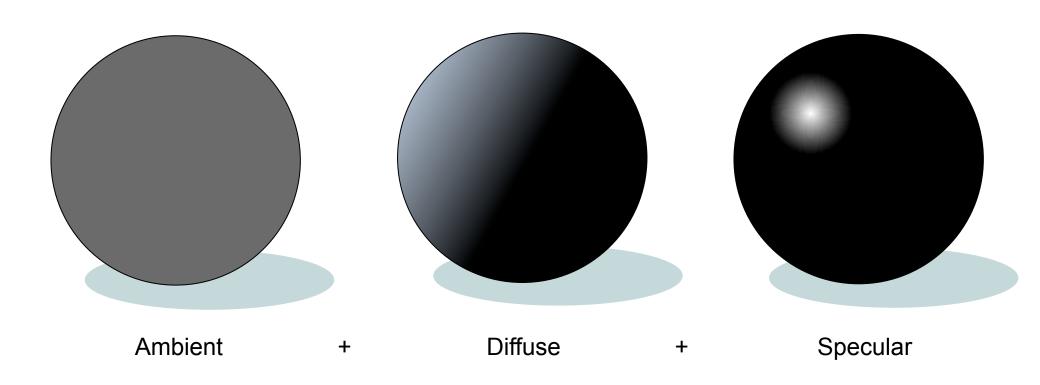
#### Introduction

- Local illumination
  - How a point light and one surface location interact
  - Valid for ray-tracing and for z buffer (projection)
  - Notation
    - $I_r$  Intensity radiating from the object (What we're looking for)
    - $I_i$  Normalized intensity of the light (Characteristic of the light)
    - k proportion of the light reflected rather than absorbed by the material (Characteristic of the surface; varies with light wavelength)

#### Visual features



## Main idea



#### Color

- Light has different wavelengths
- Illumination is independent
- Red-in-green-out odes not exist (exception: fluorescencence)
- We do all computation independently on RGB 3vectors



#### **Ambient Light**

- Approximation to global illumination
  - Each object is illuminated to a certain extent by "stray" light
  - Constant across a whole object
- Often used simply to make sure everything is lit, just in case it isn't struck by light direct from a light source

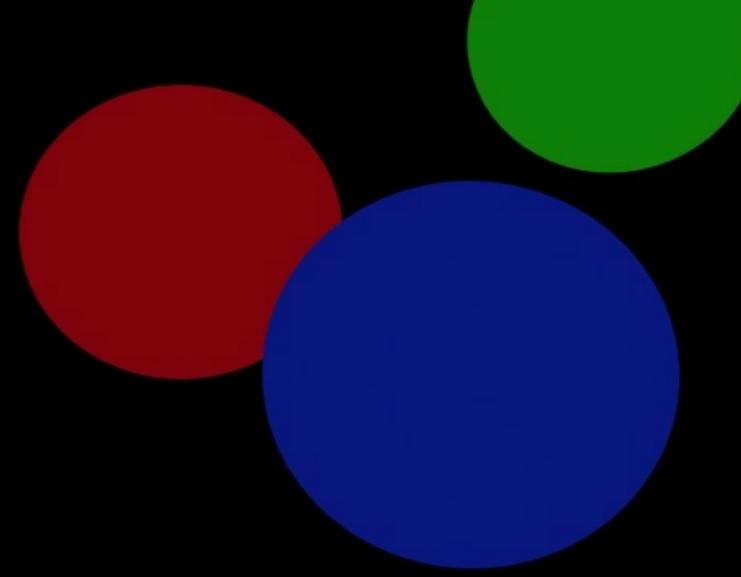
### **Ambient Light**

• Ambient light usually set for whole scene  $(I_a)$ 

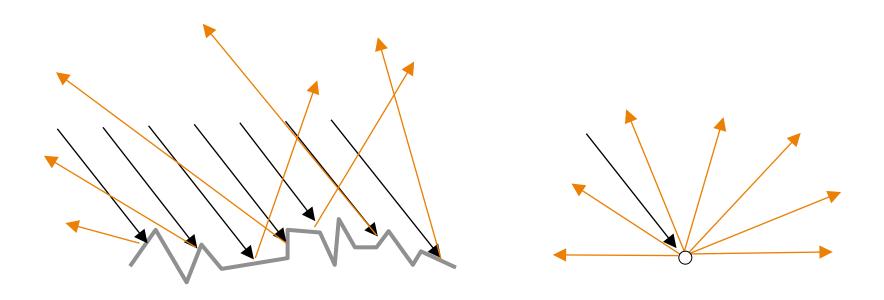
• Each object reflects only a proportion of that  $(k_a)$ 

• So far then  $I_r = k_a I_a$ 

## The Image - Ambient

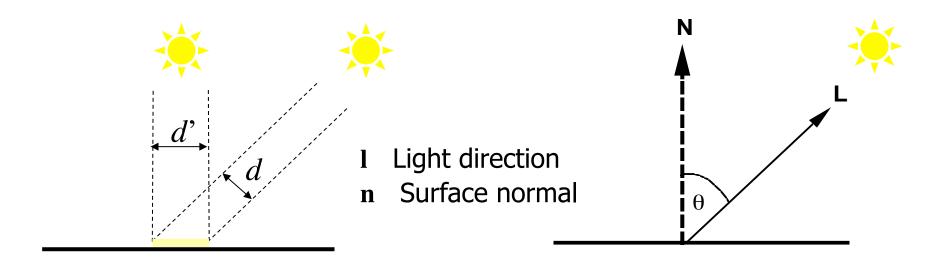


#### Lambert's Law

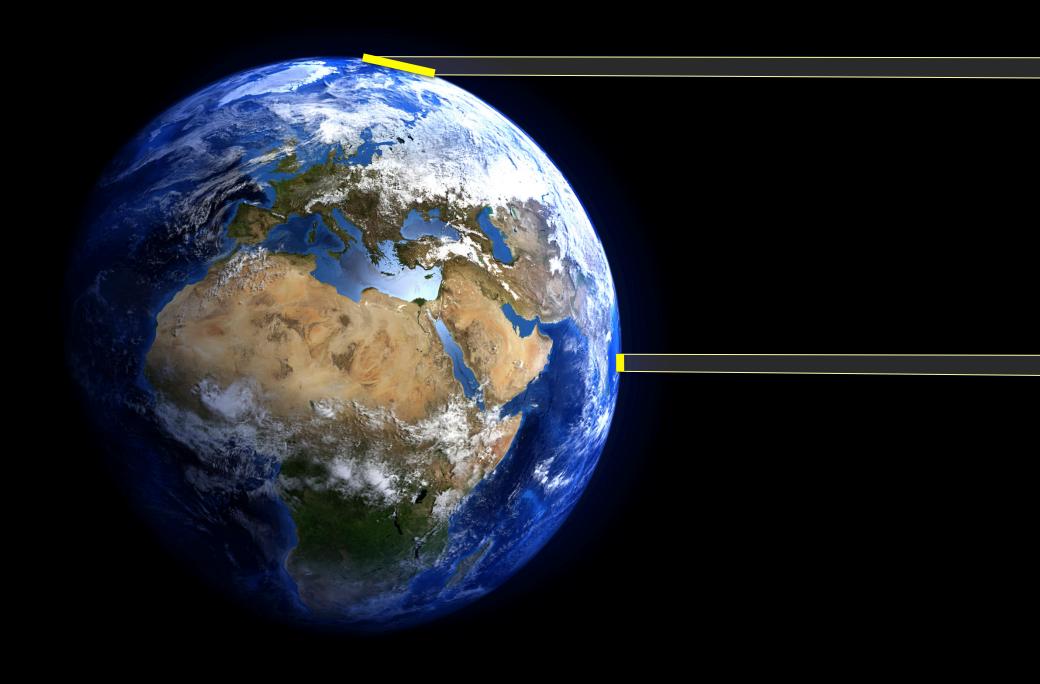


- Diffuse reflector scatters light
- Assume equality in all directions
- Called Lambertian surface
- Angle of incoming light is still critical

#### Lambert's Law



- Incoming intensity of light is proportional to d
- d is proportional to  $\cos \theta = \langle \mathbf{n}, \mathbf{l} \rangle^+ = \max(0, \langle \mathbf{n}, \mathbf{l} \rangle)$
- No negative length or light
- Reflected intensity proportional to  $\cos \theta$



#### Diffuse Light

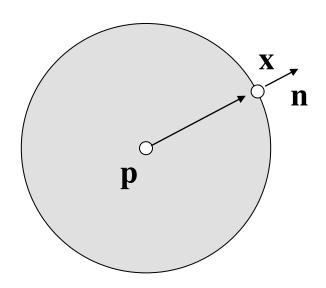
 The normalised intensity of the light incident on the surface due to a ray from a light source

The light reflected due to Lambert's law

• Proportion of light reflected rather than absorbed ( $k_{
m d}$ )

#### **Normals**

To do Lambertian shading, we need the normal n
 of a sphere at p at the intersection point x



## Lighting Equation #2

$$I_{\rm r} = k_{\rm a}I_{\rm a} + k_{\rm d}I_{\rm i} < {\bf n}, {\bf l} > +$$

• Ambient and diffuse components  $k_{\rm a}$  and  $k_{\rm a}$ 

## Multiple Lights?

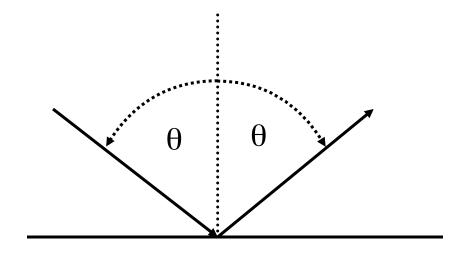
- Light adds linear
- Just add

$$I_{\rm r} = k_{\rm a}I_{\rm a} + k_{\rm d}I_{\rm 1} < {\bf n}, {\bf l} > + k_{\rm d}I_{\rm 2} < {\bf n}, {\bf l} > + \dots$$

 We see importance of clamping: Adding without clamping, lights would cancel! Not in this universe

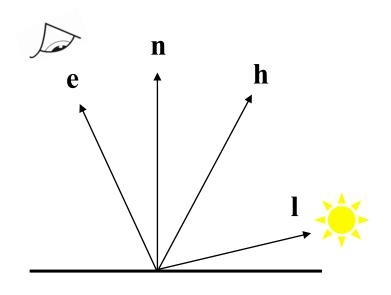
# The Image - Diffuse

## **Perfect Specularity**



Would almost never see the specular highlight

## Imperfect Specularity (Phong)

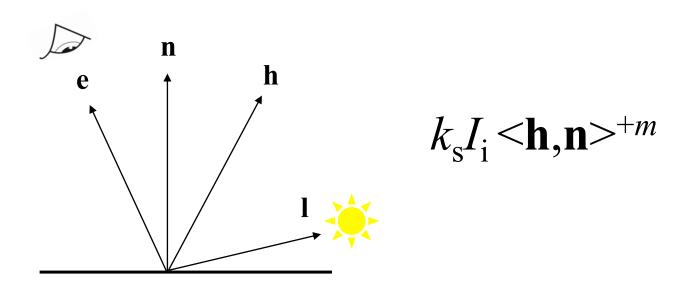


- e is the direction to the eye
- n is the normal

• I is the direction to the light

h bisects e and l

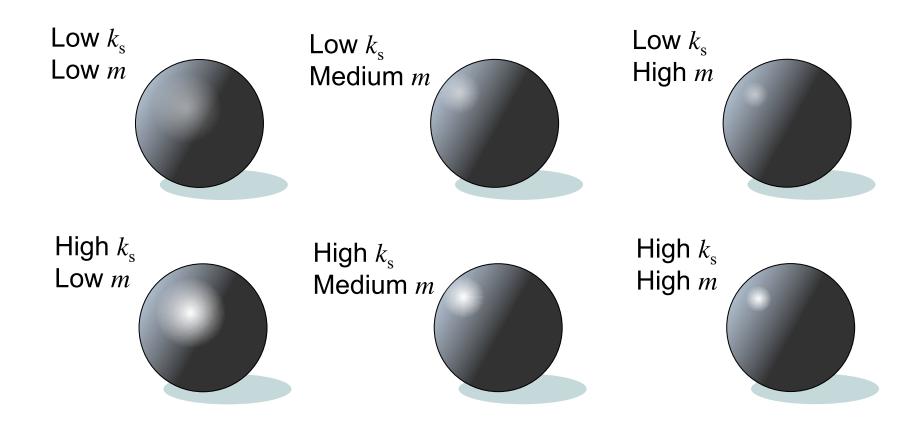
#### Specular Component



- *m* is the power of the light (shininess)
  - High m imply sharp, small highlights
  - Low m imply blurred, large highlights

## Specular phenomenology

$$k_{\rm s}I_{\rm i}<{\bf h},{\bf n}>^m$$



### Lighting Equation #3

$$I_{\rm r} = k_{\rm a}I_{\rm a} + I_{\rm i} (k_{\rm d} < {\bf n}, {\bf l})^{+} + k_{\rm s} (< {\bf h}, {\bf n})^{+})^{m}$$

- Ambient, diffuse & specular components
- Again if there are multiple lights there is a sum of the specular and diffuse components for each light

# The Image – Specular

#### Web Page

- Web page for exercises (soon)
- Web page for demos (now)

cg.cs.ucl.ac.uk

#### Conclusions

- We can now colour the pixels by combining
  - Ambient light
  - Diffuse reflections
  - Specular reflections
  - Summed over several light sources
- We need
  - Shadows
  - Better model for light reflection of the object: BRDF
  - Global illumination