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  - Spontaneous fission.

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  - Change of multiplicities.

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- Spontaneous fission.
- Induced fission.
- Change of multiplicities.
- Multiplicities and coincidences can give isotopic composition and mass of the sample.

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### NDA

- Spontaneous fission.
- Induced fission.
- Change of multiplicities.
- Multiplicities and coincidences can give isotopic composition and mass of the sample.
- A full number distribution offers deeper insight than a few factorial moments.

#### **Number Distributions**

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- Factorial moments and probability distributions
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Neutrons in a fissile sample can undergo different processes.

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- Neutrons in a fissile sample can undergo different processes.
  - o Induce fission.
  - o Capture.
  - $\circ$  (n, xn)-reactions.
  - Escape the sample.

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- The event of inducing fission will be expressed with a probability p, while failing to do so with probability 1 p.

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- In each event of induced fission new neutrons are born which will be treated independently of their origin.

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- When including absorption one obtains the statistics of the escaped particles.
- In each event of induced fission new neutrons are born which will be treated independently of their origin.
- These events will lead to a certain number distribution for a sample that will vary with mass and composition.

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To find the probability distributions we used the mathematical tool Probability Generating Functions (PGFs)

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To find the probability distributions we used the mathematical tool Probability Generating Functions (PGFs)

$$h(z) = \sum_{n} p_1(n)z^n \quad \text{and} \quad H(z) = \sum_{n} P(n)z^n. \tag{1}$$

Starting with an initial neutron or a source event (spontaneous fission)

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Master equations

$$h(z) = (1 - p)z + pq_f[h(z)]$$
 (2)

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$$H(z) = q_s[h(z)]. (3)$$

with PGFs

$$q_s(z) = \sum_n p_s(n) z^n \quad , \quad q_f(z) = \sum_n p_f(n) z^n. \tag{4}$$

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The multiplicities (factorial moments) of a function f are obtained as derivatives of its PGF g(z) evaluated at z=1:

$$\langle n \rangle = \left. \frac{\partial g(z)}{\partial z} \right|_{z=1} = \sum_{n} n f(n)$$
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• Calculated at z=0, more terms than for multiplicities.

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- Nested functions, lower order derivatives recurring.

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- Calculated at z=0, more terms than for multiplicities.
- Nested functions, lower order derivatives recurring.
- Longer expressions, which on the other hand can be expressed recursively

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To find the terms expressed in a recursive manner we need to calculate initial terms such as  $p_1(0)$ 

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To find the terms expressed in a recursive manner we need to calculate initial terms such as  $p_1(0)$ 

$$p_1(0) = (1 - \mathbf{p})z + \mathbf{p}q_f[h(z)]\Big|_{z=0} = \mathbf{p}\sum_{n=0}^{N} p_f(n)[p_1(0)]^n.$$
 (7)

8-th degree polynomial to be solved for  $p_1(0)$ , note the p-dependence for  $p_1(0)$ .

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The event of absorption can be included into the fission distribution:

$$\widetilde{p}_f(n) = \frac{p' - p}{p'} \delta_{n,0} + \frac{p}{p'} p_f(n). \tag{8}$$

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$$\widetilde{p}_f(n) = \frac{p' - p}{p'} \delta_{n,0} + \frac{p}{p'} p_f(n). \tag{8}$$

The first master equation will then read as:

$$h(z) = (1 - p')z + p'\widetilde{q}_f[h(z)].$$
 (9)

where  $\widetilde{q}_f(z)$  is the generating function of the  $\widetilde{p}_f(n)$  of Eq. (8).

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 The process of detection can be added by considering the neutrons that have been emitted by the sample.

$$\varepsilon(z) = \epsilon z + (1 - \epsilon). \tag{10}$$

Here,  $\epsilon$  is the detector efficiency for neutrons.

# **Including detection**

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the new master equations will be:

$$h_d(z) = h[\varepsilon(z)]$$
 ,  $H_d(z) = H[\varepsilon(z)]$ . (11)

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 ,  $H_d(z) = H[\varepsilon(z)]$ . (11)

 The derivatives needed for finding factorial moments as well as the statistics change in a simple way:

$$\frac{d^n h_d(z)}{dz^n} = \frac{d^n h(z)}{dz^n} \cdot (\epsilon)^n \quad , \quad \frac{d^n H_d(z)}{dz^n} = \frac{d^n H(z)}{dz^n} \cdot (\epsilon)^n \tag{12}$$

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For the factorial moments the full change is

$$\widetilde{\nu}_{d,n} = (\epsilon)^n \cdot \widetilde{\nu}_n. \tag{13}$$

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The creation of photons is a more complicated process connected to neutrons. No self-multiplication.

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Higher multiplicities can be favourable from a detection view-point.

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$$g(z) = (1 - p) + pr_f(z)q_f[g(z)]$$
 (14)

and

$$G(z) = r_s(z)q_s[g(z)]. \tag{15}$$

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With PGFs for gammas produced in spontaneous and induced fission

$$g(z) = \sum_{n} f_1(n)z^n$$
 ,  $G(z) = \sum_{n} F(n)z^n$ . (16)

# **Absorption and detection**

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Photon absorption will be accounted for by the probability  $l_{\gamma}$  that describes the leakage probability for one single photon:

$$l(z) = l_{\gamma}z + (1 - l_{\gamma}). \tag{17}$$

Note that for photons the absorption of neutrons also play a fundamental role, while the opposite is not true.

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Note that for photons the absorption of neutrons also play a fundamental role, while the opposite is not true.

The next step in the simulation of the statistics obtained from measurements is to incorporate the process of detection:

$$\varepsilon_{\gamma}(z) = \epsilon_{\gamma}z + (1 - \epsilon_{\gamma}).$$
 (18)

Using this equation we can obtain the detection statistics as:

$$g_d(z) = g[l\{\varepsilon(z)\}]$$
 ,  $G_d(z) = G[l\{\varepsilon(z)\}]$ . (19)

## **Photon-neutron correlation**

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Initial source photons are uncorrelated to source neutrons.

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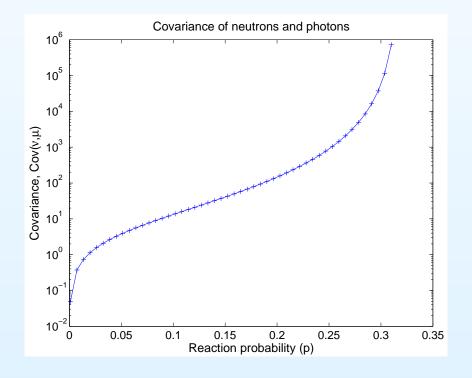
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- Initial source photons are uncorrelated to source neutrons.
- Multiplication of neutrons on the other hand is the reason for additional photon multiplication as well:

$$\mathbf{Cov} \{ \widetilde{\boldsymbol{\nu}}, \widetilde{\boldsymbol{\mu}} \} = \frac{p}{1-p} \left( \nu_{s,1} \nu_{r,1} + \nu_{s,2} - \nu_{s,1}^2 \right) \mu_{r,1} \mathbf{M}^2 + \left( \frac{p}{1-p} \right)^2 \left( \nu_{s,1} + \mu_{r,1} \nu_{r,2} \right) \mathbf{M}^3.$$
(20)



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The symbolic code Mathematica have been used to do the derivations and find formulae for higher order terms which grow rapidly in size.

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The symbolic code Mathematica have been used to do the derivations and find formulae for higher order terms which grow rapidly in size.

$$\begin{cases} \mathsf{h_d}\text{'}(\mathsf{z}) \to \frac{\mathsf{p'-1}}{\mathsf{p'}\nu_{d,s}(1)\text{-1}}, \mathsf{h_d''}(\mathsf{z}) \to -\frac{\mathsf{p'}\nu_{d,s}(2)\mathsf{h_d'}(\mathsf{z})^2}{\mathsf{p}\nu_{d,s}(1)\text{-1}}, \\ \mathsf{h_d}^{(3)}(\mathsf{z}) \to \frac{-\mathsf{p'}\nu_{d,s}(3)\mathsf{h_d'}(\mathsf{z})^3\text{-3p'}\nu_{d,s}(2)\mathsf{h_d''}(\mathsf{z})}{\mathsf{p'}\nu_{d,s}(1)\text{-1}}, \dots \end{cases}$$

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$$\begin{cases} h_{\text{d}}\text{'}(z) \to \frac{p\text{'-1}}{p\text{'}\nu_{d,s}(1)\text{-1}}, h_{\text{d}}\text{''}(z) \to -\frac{p\text{'}\nu_{d,s}(2)h_{\text{d}}\text{'}(z)^2}{p\nu_{d,s}(1)\text{-1}}, \\ h_{\text{d}}^{(3)}(z) \to \frac{-p\text{'}\nu_{d,s}(3)h_{\text{d}}\text{'}(z)^3\text{-3p'}\nu_{d,s}(2)h_{\text{d}}\text{''}(z) \ h_{\text{d}}\text{'}(z)}{p\text{'}\nu_{d,s}(1)\text{-1}}, \\ \end{cases}$$

As a final step the parameters are replaced with values to get numerical expressions.

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As a final step the parameters are replaced with values to get numerical expressions.

Saves time and makes multiple evaluations easy.

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$$\begin{cases} \mathsf{h_d}\text{'}(\mathsf{z}) \to \frac{\mathsf{p'}\text{-}1}{\mathsf{p'}\nu_{d,s}(1)\text{-}1}, \mathsf{h_d}\text{''}(\mathsf{z}) \to -\frac{\mathsf{p'}\nu_{d,s}(2)\mathsf{h_d}\text{'}(\mathsf{z})^2}{\mathsf{p}\nu_{d,s}(1)\text{-}1}, \\ \mathsf{h_d}^{(3)}(\mathsf{z}) \to \frac{-\mathsf{p'}\nu_{d,s}(3)\mathsf{h_d}\text{'}(\mathsf{z})^3\text{-}3\mathsf{p'}\nu_{d,s}(2)\mathsf{h_d}\text{''}(\mathsf{z}) \ \mathsf{h_d}\text{'}(\mathsf{z})}{\mathsf{p'}\nu_{d,s}(1)\text{-}1}, \dots \end{cases}$$

As a final step the parameters are replaced with values to get numerical expressions.

- Saves time and makes multiple evaluations easy.
- Comparisons with Monte Carlo simulations done with the MCNP-PoliMi code, from which we have taken the numerical value of p and the leakage and detection probabilities for neutrons and photons.

# **Effect of absorption**

**Number Distributions** 

**Neutron Distribution** 

Gamma Distribution

### Results

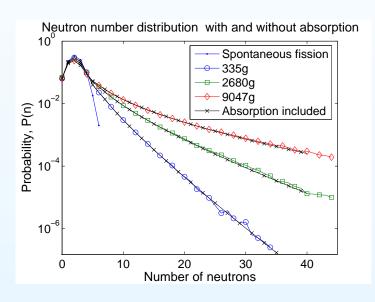
- Tools
- Effect of absorption
- Comparison with

Monte Carlo

- Detection
- Simulated scintillation detector results

Conclusions

• Dependence on mass shown in the parameter p, and also an increased probability of absorption. 20 wt%  $^{240}$ Pu and 80 wt%  $^{239}$ Pu.



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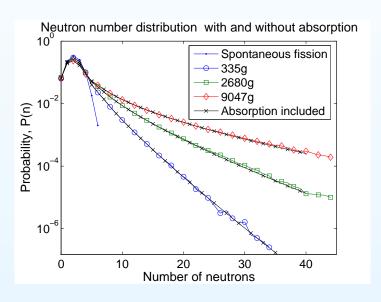
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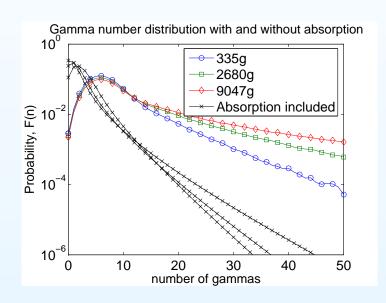
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- Change compared to non-multiplying case.
- Small effect of absorption on neutrons for such a heavy element.
- Large self-shielding for photons, which are still generated with high multiplicity, but few leak out.

# **Comparison with Monte Carlo**

### **Number Distributions**

### **Neutron Distribution**

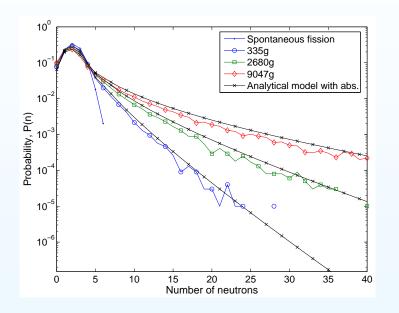
### Gamma Distribution

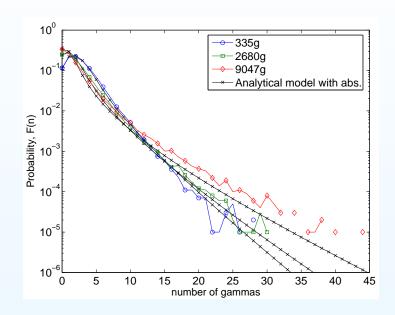
#### Results

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- Good agreement with MCNP-PoliMi.
- For photons one have higher multiplicities per source event for smaller samples.

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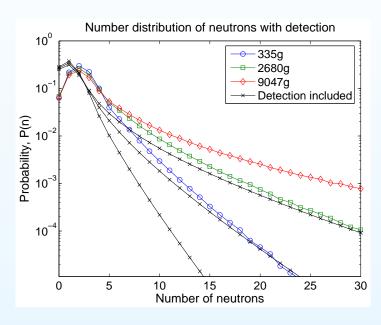
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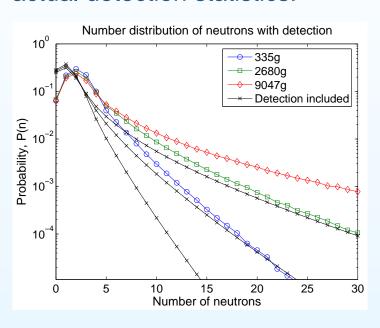
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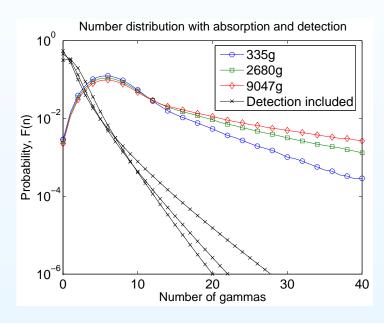
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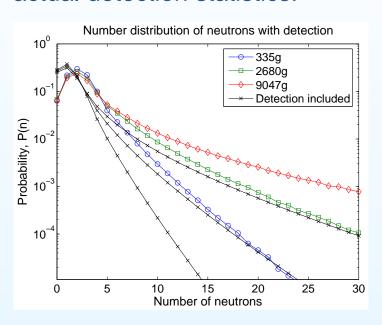
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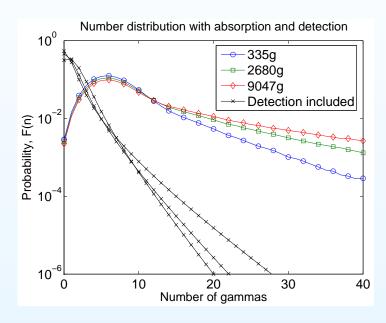
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- Detection efficiencies of 50% were used here.
- The chance to detect many particles from the same source event decreases for both neutrons and photons.
- The photon multiplicities decrease so much that neutrons might be more favourable to observe even though they have lower source multiplicities.

## Simulated scintillation detector results

**Number Distributions** 

**Neutron Distribution** 

Gamma Distribution

### Results

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- Effect of absorption
- Comparison with

### Monte Carlo

- Detection
- Simulated scintillation detector results

- Organic scintillators have the advantage to be able to detect both neutrons and photons, and to discriminate between them.
- MCNP-PoliMi was also used to simulate a scintillator detector setup with six detectors:

### Simulated scintillation detector results

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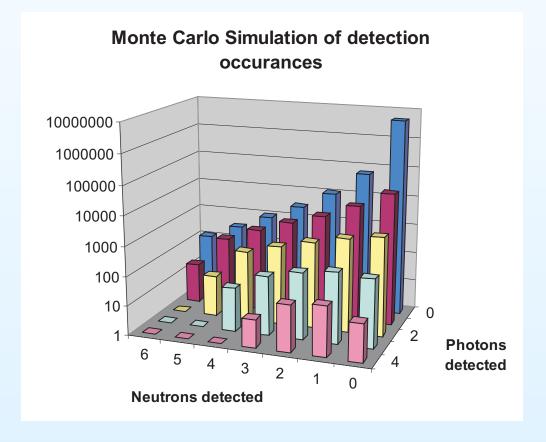
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**Number Distributions Neutron Distribution** Gamma Distribution Results Conclusions Conclusions **Conclusions** 

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 Master equations, earlier used for finding factorial moments, can be used to find probability distributions.

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- Symbolic computation makes it possible to find higher-order terms in a recursive manner.

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- Extensions
  - Compare to experiments
  - Simulate realistic multiplicity counters