

Linear (Longitudinal/Seif) Dune Depositional System

Hierarchical Multiscale Structure of Linear Dune Fields

Dune Fields (Ergs): Linear dunes typically occur in extensive dune fields (ergs) as sets of nearly parallel ridges stretching over vast areas ¹. These ridges can be tens to hundreds of kilometers long while only a few hundred meters wide ¹. Intervening **interdune corridors** are often broad and can span hundreds of meters to several kilometers, sometimes consisting of sand sheets, gravel flats, or exposed bedrock surfaces cleared by wind ¹. Within an erg, linear dunes maintain a consistent orientation, giving the field a striped appearance from above. *For example, the Namib Sand Sea in Namibia contains north-south linear dunes (see image below), with long, continuous crests and wide interdune valleys.* ² ³

Satellite view of linear dunes in the Namib Desert. Long, parallel dune ridges are visible, and smaller secondary dunes ride along some crests (lighter lines atop the main ridges). Strong southerly winds produce the north-oriented linear dunes here ².

Individual Seif Dunes: A single linear (or *seif*) dune is an elongated ridge of sand with a sharp, narrow **crestline** running along its length ⁴. Crestlines are often relatively straight or gently sinuous; any curvature or undulation along the crest reflects subtle shifts in wind or variations in sand supply. The crest is typically **knife-edged** and very linear, a result of sand avalanching down both sides over time ⁴. Unlike a barchan dune's single leeward slipface, a seif dune can develop slipfaces on **either flank**, giving it steep, symmetric cross-sectional profiles ⁴. Each flank can act as a lee side under different wind directions, yielding a steep (30–34°) angle of repose surface on whichever side was leeward in the most recent wind event. The flanks themselves are long and relatively planar, separating the crest from the interdune floor.

Slipfaces and Flanks: Because winds from alternating directions drive these dunes, slipfaces tend to form intermittently on alternating sides. At any given time a well-developed slipface may appear on the side opposite the prevailing wind, but a change in wind direction will cause sand to avalanche down the other side, building a slipface there ⁵ ⁴. This process creates a zigzag pattern in cross-sectional layering (with cross-beds dipping in opposite directions on each flank). The result is a dune with a sharp, sometimes slightly *serrated* crest and evidence of avalanching on both sides, although the dune as a whole remains longitudinally oriented. The flanks between slipface events often have lower-angle slopes covered in wind-ripple patterns.

Interdune Corridors: The spaces between linear dunes (interdunes) vary from sand-covered corridors to hardpan flats. In sand-rich fields, interdunes may contain sand sheets or smaller dunes, and they tend to be relatively narrow. In sand-starved fields, interdunes can be broad, exposing coarse lag (gravel) or moist substrate which resists deflation ¹. These corridors often align with the dominant wind flow, effectively acting as wind lanes that funnel airflow. Interdune surfaces might host sparse vegetation or seasonal water pools if the water table is high, which can locally stabilize sand and influence dune spacing.

Superimposed Features: Linear dune systems display a hierarchy of superimposed bedforms. Large linear ridges (primary dunes) often have **secondary dunes or mega-ripples** on their backs and flanks. For instance, the highest seif dunes in the Namib Sand Sea have smaller linear ridges riding along their crests ² – essentially “dunelets” aligned with the same wind regime atop the larger dune. Similarly, meter-scale **ripples** and small transverse dunes (formed by local winds or eddies) commonly appear on the flanks, oriented roughly perpendicular or oblique to the main crestline ⁶. These superimposed features indicate active sand transport at smaller scales, and they can migrate along the main dune or even contribute sand to the main crest. In addition, linear dunes sometimes terminate in or spawn other dune types; for example, a long seif dune may end downwind in a **barchan** or network of barchans if sand supply tapers off ⁷. Conversely, two linear dunes can merge into a single ridge, creating a Y-shaped junction (a **compound dune** crest) where the smaller branch joins a main trunk ⁸. Such **Y-junctions** reflect interactions among dunes and contribute to a multi-scale pattern of branching and convergence within the field.

Physical Drivers and Controls on Linear Dune Formation

Bidirectional Wind Regimes: The defining driver for longitudinal dune formation is a wind regime with **two prevailing directions** that are distinct but fairly balanced. Typically, seasonal or alternating winds blow from two directions that diverge by a large angle (often $>90^\circ$ apart) ⁹. Under these reversing winds, sand is pushed from alternating sides, and dunes tend to elongate **parallel to the resultant sand transport direction** (nearly along the bisector of the two wind directions) ¹. In other words, the dune’s long axis aligns within $\sim 15^\circ$ of the vector sum of sand flux from each wind ¹⁰. The two wind directions do not necessarily have equal strength or frequency, but each must be significant enough to periodically drive sand along the dune. This bidirectional wind regime creates the conditions for *reversing slipfaces* and along-crest transport that characterize seif dunes. If the directional variability is only moderate (e.g. two main directions with one somewhat dominant), linear dunes still form, whereas very low variability (unimodal winds) yields transverse/crescentic dunes and very high variability (multimodal winds) yields star dunes ¹¹. Thus, linear dunes are associated with intermediate wind variability: strong winds from two persistent directions. Wind energy is usually high in these environments (strong, frequent winds) to continuously mobilize sand ¹¹. Notably, the dune orientation is such that each of the two winds blows obliquely *against* a dune flank, moving sand along the ridge rather than fully across it. This helps maintain a longitudinal form by minimizing net sand transport directly across the crest.

Sand Supply and Sediment Characteristics: Adequate sand supply is crucial for linear dune development. These dunes typically form in **large sand seas or corridors** where sand is plentiful enough to build continuous ridges, but the sand may be more patchy than in transverse dune fields ¹² ¹³. If sand supply is too low, dunes might remain isolated barchans instead of linking into long ridges. If sand supply is very high, dunes can grow taller and wider, but they may also transition to different forms (e.g. if combined with highly variable winds, star dunes might form). Linear dunes often indicate a balance: **moderate to high sand supply** with winds reworking the sand into elongate forms. In many cases, sand supply is fed from an upwind source and transported down the length of the dune. For example, linear dunes commonly nucleate near a sand source like a river delta, a breach in a mountain range, or an upwind erg margin, then extend outward from that source ¹⁴. The grain size is usually uniform fine to medium sand, allowing easy saltation; too much coarse material would armor the surface and inhibit dune growth. Sand on linear dunes can be well-sorted and often forms a **surface armor of coarse grains** in interdune areas (lag deposits), which helps confine the sand to the dunes.

Surface Moisture and Vegetation: Moisture and vegetation play secondary but important roles. **Surface moisture** (from occasional rain or a high water table) can stabilize sand in interdunes or lower dune flanks. Damp sand resists wind erosion, so moist interdunes can act as steady sand supply sources (via deflation of drying surfaces) or impose spacing constraints (dunes form where dry sand is available and stop where interdunes are too wet). In some coastal deserts or fog deserts, moisture may harden interdune crusts, forcing dunes into parallel belts separated by firm flats. **Vegetation** can anchor parts of a dune, reducing sand mobility. In linear dune systems, vegetation (when present) tends to colonize interdune floors and lower slopes where moisture is higher, creating partially stabilized dunes. Over long timescales, increased vegetation can halt dune migration entirely, turning active seif dunes into **fixed linear dunes** (sometimes called *relict* dunes) ¹⁵. Many linear dunes in semi-arid regions (e.g. parts of Australia's Simpson Desert) are vegetated and have not migrated for thousands of years, preserving their form from past climate conditions. However, even in vegetated states, the overall linear morphology (long, narrow ridge) persists. Vegetation thus tends to **limit height and migration rate** of dunes (by binding sand), and can cause dune bifurcation or termination as sand accumulates behind vegetated patches. If vegetation stabilizes the arms of dunes, the system may evolve into parabolic dunes, though for longitudinal ridges this is less common – instead they just become stationary ridges.

Topographic and Boundary Conditions: Linear dunes are often influenced by underlying topography or obstacles. A slight rise or obstacle can trigger sand accumulation and act as the **initiation point** for a linear dune ¹⁶. Many linear dunes begin on the leeward side of a rock outcrop, hill, or even a vegetation clump, where wind flow converges and drops sand. Once started, the dune can extend far beyond the obstacle. The orientation of the dune may also be guided by topography – for instance, valleys can channel one of the wind directions, modifying the local resultant wind. A notable example is in the Namib Sand Sea: linear dunes run N-S under dominant southerly winds, but near the Tsondab valley they are deflected westward because strong easterly winds funnel through the valley and reorient the dunes downwind ³ ¹⁷. Over large areas, subtle changes in basin shape or underlying slope can affect dune spacing and continuity: dunes may thin or terminate at valley margins and cluster in topographic lows where sand is abundant.

In summary, **bidirectional wind** energy is the primary driver setting linear dune orientation and slipface dynamics, while **sand supply** determines if long ridges can form (and how big they get). **Moisture, vegetation, and topography** modulate the dune field pattern – anchoring dunes in place, causing gaps, or initiating new ridges – thereby adding complexity to the ideal pattern predicted by wind alone.

Time-Evolution and Dynamic Processes of Linear Dunes

Linear dunes exhibit dynamic behaviors over time, adjusting their length, shape, and interactions with neighbors under the influence of variable winds and sand supply. Key evolutionary processes include elongation, lateral migration, crest restructuring, and dune-dune interactions:

- **Crestline Elongation:** The foremost growth mode of a seif dune is lengthening of its crest. Under a reversing wind regime, sand transport along the dune axis leads to **endwise extension** – typically at the downwind end (in the resultant transport direction) ¹⁴. With each season of wind from one direction, sand is pushed along the crest toward the opposite end, causing the dune to gradually **prograde** forward. If there is a persistent net transport (one wind direction slightly stronger or more frequent), the dune will elongate in that net direction. Field and remote-sensing observations confirm that active linear dunes can extend their length by a few meters per year. *For example, Martian linear dunes measured over 8 years showed elongation on the order of ~1.3 m per Earth year*

under a bimodal wind regime ¹⁸. Terrestrial dunes in strong wind environments can elongate even faster (several m/year) if sand is abundant. Over centuries, this process can produce huge finger-like dunes extending far from their sand source. Conversely, if sand supply diminishes, elongation may stall and the dune ends become fixed or eroded. Some linear dunes show **age gradients** along their length (with younger sand deposits toward the actively growing end), supporting the idea of progressive forward growth over time ¹⁴.

- **Lateral Migration and Adjustment:** Although linear dunes primarily grow longitudinally, they can **shift laterally** over time if the two wind directions are not perfectly balanced. If one wind direction dominates slightly (in strength or duration), it will produce a greater sand flux on one flank, causing the dune to gradually **drift sideways** in the direction of the lesser opposing transport. This lateral migration is usually slow – larger dunes might shift only a few meters over many years. Evidence of lateral motion can be seen in asymmetric cross-sections or offsets in older versus younger dune positions (e.g., older stabilized dune remnant slightly to one side of the active crest). In Australia's Strzelecki Desert, for instance, linear dunes migrated a few meters eastward during the Holocene due to stronger west winds, as indicated by preferential sand accumulation on their east flanks ¹⁹. Importantly, the lateral motion of linear dunes is **much slower than that of barchans**, since the longitudinal alignment means much of the sand transport is along the ridge, not across it. The **migration rate is inversely related to dune size**: larger (taller) dunes shift more slowly because they require more sand movement to translate the whole mass ²⁰. Thus, over time linear dunes tend to maintain spacing and orientation, with only minor lateral adjustments unless one wind direction decisively changes.
- **Crest Morphology Changes:** As winds alternate, the **crestline itself can oscillate and meander** slightly in planform. Each time the wind from one side blows, it may push segments of the crest leeward (laterally) until the next reversal pushes them back. Over long periods, this creates a gentle *sine-wave undulation* or scalloped pattern along the otherwise linear crest. The amplitude of this meandering is controlled by dune dynamics: **dune height and wind regime balance** are key factors ²¹. Higher dunes and more evenly split winds produce more pronounced back-and-forth shifting of the crest, leading to a noticeably wavy dune form ²¹. If one wind dominates heavily, the crest may instead take on a steady curve or remain straight. Shorter-term changes include the formation of *reversal bedforms*: e.g. small temporary transverse dunes can appear on a flank during one wind regime and then be absorbed back into the main ridge when winds reverse. Crest height can also change over time: linear dunes grow taller as they accumulate sand, but there is often an upper height limit imposed by wind flow (once tall enough, wind may start stripping sand off the crest as fast as it is added). Dune height and width typically **increase until a steady-state aspect ratio** is reached, after which the dune elongates rather than growing taller ²² ²³. If sand supply wanes, dunes might lower and flatten (with broader, more tabular crests, especially toward their downwind terminus) ²⁴.
- **Slipface Development (Alternating Sides):** The creation and removal of slipfaces is an ongoing cyclical process on linear dunes. During a period of wind from direction A, sand moves up the windward side and avalanches down the opposite side (leeward relative to A), forming a fresh slipface there. When the wind switches to direction B (from the opposite side), the previously windward side becomes the new lee, and a slipface develops on that side ⁵. As a result, **slipfaces alternate sides with changing winds**, and each side of the dune experiences episodes of avalanche deposition. Over time, this yields cross-bedded deposits dipping in opposite directions on

alternating flanks, often creating an interleaved herringbone pattern in the stratigraphy. Notably, a slipface may not continuously cover the entire flank length at any given moment – it can be patchy or confined to sections – because wind events (like storms) may dump sand unevenly. High dunes sometimes go through periods with **no active slipfaces** along their sides (especially if winds are gentle or nearly equal from both sides, causing sand to just creep along the crest). But once a sufficiently strong wind blows from one side, a slipface will appear on the other side, maintaining the dune's sharp crest. In summary, active linear dunes will show evidence of recent avalanches on whichever flank was leeward last, and older stabilized dunes might preserve sets of cross-beds from multiple wind directions.

- **Dune Merging and Splitting:** Linear dunes interact with neighbors in ways that can change their continuity. A common interaction is the formation of **Y-junctions** where two dunes merge. This can happen when the end of one dune collides with the side of another due to slight differences in orientation or length. The smaller ridge might attach and feed sand into the larger, creating a Y-shaped compound dune pattern ⁸. Such **bifurcations** are frequently observed and indicate a form of pattern coarsening – smaller dunes combining to form fewer, larger dunes over time. Conversely, linear dunes can also **split or branch**. Split can occur if a dune grows very long and sections of its crest become misaligned or separated by wind eddies; a segment of the crest might detach and form a new parallel ridge (similar to a defect spawning a new dune). Splitting is sometimes observed where dunes encounter variability in wind or topography, leading to a forked crestline. In active fields, high sand influx can cause the *tip* of a linear dune to **eject barchans or new dunes**: sand that overruns the linear ridge's capacity breaks off as isolated barchan dunes at the end of the ridge ⁷. These barchans can then be reabsorbed into linear ridges further downwind or remain as separate features if conditions favor barchan shape. This process effectively limits the length of any single ridge by shedding excess sand into new dunes. Over long timescales, dune interactions lead to a self-organizing pattern: some ridges link and grow, others fragment, maintaining an optimal spacing. Dune-field pattern modeling studies have noted that **defects and terminations** in crestlines often lead to Y-junctions or spur dunes, helping dunes avoid overly close spacing by redistributing sand between neighboring ridges. The net effect is that linear dune fields tend to **regularize** over time – erratic smaller dunes join major ridges, and overly long ridges shed sand – resulting in a quasi-stable set of equally spaced, parallel dunes.
- **Climatic and Temporal Changes:** Because linear dunes often have great longevity, they can record changes in climate and wind regimes. If the wind regime shifts (e.g. a new prevailing direction emerges), linear dunes may **reorient** gradually. This can produce overprinting patterns: for instance, older linear dunes aligned to an ancient wind direction can be partially overrun by newer dunes in a slightly different orientation, creating a lattice or rhomboidal pattern where two sets of linear ridges intersect ¹⁷. In the Namib Desert, some linear dunes oriented north-south have been displaced or superimposed by younger dunes oriented west-east when wind patterns changed in the past ¹⁷. Additionally, if climate becomes wetter and vegetation grows, dunes can stabilize (pause migration) – later, if climate dries again, the dunes can reactivate along the same orientation. Thus, the current morphology of a linear dune may be a palimpsest of multiple phases of activity and quiescence. Time-evolution models must consider these temporal layers – in a generative context, one might simulate initial dune growth, then impose a slight orientation change to mimic an older set of crestlines beneath the current ones.

In summary, linear dunes evolve mainly by **growing lengthwise and adjusting laterally** at slow rates, with their form maintained by alternating winds building slipfaces on both sides. They interact by merging and splitting, which over time leads to an organized pattern of ridges. Capturing these dynamics is important for any model: one must include rules for elongation, occasional lateral migration, alternating deposition on flanks, and crest segmentation or merging events to mimic the natural evolution of seif dune fields.

Static Generation Rules for Geologically Plausible Linear Dune Fields

When constructing a **snapshot** of a linear dune field (for instance, to synthesize a realistic geologic analog image), one should enforce rules that capture the characteristic layout and morphology at multiple scales. The goal is to produce a static scene that *looks* like a believable stage in a linear dune field's development, consistent with known physical constraints. Key rules and steps include:

- **Overall Alignment:** Choose a primary dune orientation corresponding to the resultant wind direction. In practice, determine two main wind directions (A and B) and let the dune crests align roughly with their vector resultant. For example, if wind A is from the south and wind B from the east (two dominant directions ~90° apart), orient dunes along a direction between them (south-east in this case). As a rule of thumb, linear dune crest orientation can be set within ~10–15° of the resultant transport direction of the wind regime ¹. This ensures dunes are neither perpendicular nor parallel to any single wind, but in an intermediate orientation consistent with alternating oblique sand transport. All dunes in the field should follow the same general orientation, with only minor local deviations.
- **Dune Spacing:** Impose a relatively uniform spacing between adjacent dune crests, reflecting the natural **wavelength** of the pattern. Linear dunes tend to be separated by distances on the order of tens to a few hundreds of meters, depending on their size. Empirical relationships show that **dune spacing correlates with dune height** ²⁵. A common guideline is to space dunes about *5–15 times the dune height* apart (crest-to-crest). For instance, if the target dune height is ~20 m, spacing might be ~200 m (with variation). In the Namib linear dunes, Lancaster (1982) documented a positive correlation where taller dunes had larger spacings ²⁶. In generation, one can set an average spacing and add some randomness (e.g. ±20%) to avoid perfect regularity. Also consider that very large dunes (>30 m high) generally require wider interdunes for wind flow and sand availability. Ensure that interdune width is sufficient so that one dune's slipface does not overlap the base of its neighbor – interdunes should appear as distinct corridors, possibly hosting thin sand sheets or gravel flats. If modeling limited sand supply, allow some interdune patches to be sand-free (exposing substrate), especially in a **compound pattern** where dunes might sit on a non-sandy surface.
- **Parallel Crest Consistency:** Arrange dunes in a *sub-parallel array*. Real linear dunes are remarkably parallel over long distances, so the model should reflect that consistency. Minor convergence or divergence can be introduced (dunes may slowly converge into Y-junctions or diverge around an obstacle), but large random orientation swings should be avoided. Dunes can be slightly sinuous: incorporate gentle bends or undulations in the crestline with long wavelengths (on the order of several dune-spacing units). These undulations can be modeled as low-amplitude sine waves superimposed on the straight crest axis. The wavelength of crest meandering could be tied to dune length or wind oscillation scale (e.g. one full sine wave over a dune's length, or a wavelength of a few

kilometers). The amplitude should be small relative to spacing (perhaps 5–10% of spacing) to keep dunes from intersecting or diverging too much. This gives each ridge a realistic slight *wiggle* rather than a mathematically perfect line.

- **Dune Length and Termination:** Vary the lengths of dunes across the field to mimic natural limits and sources. Not all dunes should span the entire field; some may start or end within the domain. Typically, dunes might **initiate** near a sand source margin (one end anchored) and **terminate** downwind. To implement this, you can randomly assign some dunes a “start” within the field (e.g. a dune might begin in the middle where a local source is present) or let all dunes begin at the upwind boundary. Likewise, allow dunes to end before the far downwind edge – possibly by thinning out into a barchan or fading into a sand sheet. Include a few **terminations and junctions**: e.g., have two shorter dunes merge into one longer dune (Y-junction), or one long dune split into two shorter branches. A Y-junction can be constructed by having a secondary crest branch off at a shallow angle (~30° or less) from a main crest and run parallel at a short distance before fully diverging. These should be infrequent and placed logically (often where dunes might collide if one was slightly misaligned or where sand availability changes). The presence of an occasional Y-shaped merge will greatly increase realism ⁸.
- **Elevation Profile & Cross-section:** Assign each dune a height profile along its length. Linear dunes often have their maximum height somewhere near the middle or upwind half, and may taper toward the ends. You can model dune height vs. length as a gentle rise from the upwind foot to a peak, then a slight decrease toward the downwind terminus. For simplicity, assume each dune has a roughly uniform height along most of its length (Type A dunes in some studies maintain fairly constant height for kilometers ²⁷). Alternatively, incorporate Type B behavior for some dunes – where height and width diminish progressively along the dune (particularly for shorter dunes <1 km) ²⁸. This means the dune might be highest near its start and gradually lower toward its tip, reflecting depletion of sand. Relate **dune width to height**: use a typical aspect ratio so that width (base-to-base, across the dune) is about 10–20 times the dune height (implying a cross-sectional slope on each side roughly 10%, consistent with about 5°–10° average flank slope, steepening to 30° at the slipface near the crest). For example, a 20 m tall dune might have ~200–300 m between its toes (from the base of one flank to the base of the other). Ensure flanks are symmetric or only mildly asymmetric in a static snapshot. Each dune should be given a **crest width** (the flat or knife-edge top); in practice this is very narrow for active dunes (a meter or two of rounded top at most), but for modeling, one can treat the crest as a line or a very thin ridge.
- **Slipface Allocation:** In a static image, decide which flank of each dune has the active slipface (steep lee) and which is the gentler windward slope. This could be done by assigning a “current prevailing wind” for the snapshot. For instance, if we assume the latest wind was from direction A, then **flank B** (the side opposite that wind) should show the slipface. Visually, the slipface side will have a sharper, shadowed drop-off in an image, whereas the windward side is more gradual. However, not every dune needs an active slipface along its entire length – you can vary it so that some dunes show patchy slipfaces (perhaps due to local gusts or partially stabilized sections). A simple rule: for each dune, choose one flank to be lee at the moment and make its slope ~33° and mark its base with small talus aprons; the opposite flank make shallower (~10–15°). To add realism, alternate this choice for neighboring dunes or along a single dune’s length if appropriate (since a gust could cause a portion of a dune to avalanche). If modeling reversing dunes, you might even show minor slipface scars on the “windward” side from previous events. In an image, texturing the lee side with

smoother, avalanche sand and the windward with more ripples or erosion features conveys this difference.

- **Surface Texture and Minor Features:** Populate the dune surfaces with small-scale patterns. Apply **sand ripples** along the flanks and interdune surfaces aligned perpendicular to the local wind flow. On the windward flank (current windward), ripples should be oriented roughly upslope (since wind drives them upward), and on the lee flank, any visible older ripples might be oriented downslope from the previous wind. Optionally, add **small secondary dunes**: e.g., low-amplitude secondary linear ridges along the crest (especially if the main dune is very large – these would be compound dunes ²). These could be modeled as slight undulations or small parallel ridges attached to the main crestline. In wide interdunes, one might place occasional **barchan or dome dunes** if sand has accumulated there (especially in sand-rich scenarios), or scatter some sand sheets. If vegetation is part of the scenario, include patches of vegetation in interdunes and maybe a thin cover on some dune flanks (with corresponding stabilization: dunes might be slightly more irregular or have a flatter top where vegetated). Ensure the **color/tonal variation** in the snapshot distinguishes slipface (usually smoother, shadowed, maybe darker) vs. windward (brighter, textured by ripples). Real satellite images often show linear dune crests with alternating bright and dark streaks corresponding to recent slipfaces and interdune exposures.
- **Boundary Conditions and Edge Effects:** At the edges of the dune field, dunes might change form. Incorporate that by, for example, making the upwind boundary of the field have smaller or more isolated dunes that join into the main linear ridges (since sand is just starting to accumulate). On the downwind edge, perhaps fade out the dunes into smaller barchans or hummocks. If the field abuts a mountain or non-sandy area, dunes near that edge might be shorter or divert around the obstacle (curving slightly). These edge adjustments help the scene feel part of a larger system rather than abruptly cut off.

By following these static generation rules, one creates a synthetic dune field that respects geological realism: **parallel, elongate ridges** of sand with consistent spacing, subtle variability in crest lines, evidence of alternating wind influence, and appropriate scaling of dune size to spacing. Such a snapshot can then be used as a starting point or compared against real-world dune fields for validation.

Quantitative Relationships and Formulas for Linear Dune Modeling

In developing a physically accurate model, it's useful to incorporate quantitative relations gleaned from empirical studies and theory. Below are some key formulas and relationships relevant to linear (seif) dunes:

- **Resultant Transport Direction:** Given two prevailing wind vectors (with directions θ_1 and θ_2 , and transport magnitudes $Q_{\{1\}}$ and $Q_{\{2\}}$, e.g. proportional to drift potential), the **resultant sand transport vector** \mathbf{R} can be computed as $\mathbf{R} = Q_{\{1\}}\mathbf{d}_{\{1\}} + Q_{\{2\}}\mathbf{d}_{\{2\}}$ (where \mathbf{d} is a unit vector in direction θ). The angle θ_r dictates the dune alignment. For example, if wind 1 (θ_1) has 60% of total transport and wind 2 (θ_2) 40%, the dune's trend will be closer to θ_1 but somewhat toward θ_2 . The orientation of \mathbf{R} aligns roughly with θ_r , i.e., along \mathbf{R} . Another way to express dune orientation is through the *gross bedform-normal transport* concept: dunes orient such that net transport perpendicular to the crest is minimized. This can be framed as: find crest angle φ such that $Q_{\{1\}}\sin(\theta_1 - \varphi) + Q_{\{2\}}\sin(\theta_2 - \varphi) \approx 0$. Solving that yields φ close to θ_r . In sum, a **formula**

for dune orientation could be written as: $\phi \approx \arctan\frac{Q_1 \sin \theta_1 + Q_2 \sin \theta_2}{Q_1 \cos \theta_1 + Q_2 \cos \theta_2}$, which yields $\phi = \theta_r$ (resultant direction). In practice, for equal-strength winds blowing at an angle $>90^\circ$, ϕ will lie about midway between them ⁹.

- **Dune Height-Width Scaling:** Linear dunes maintain a roughly constant aspect ratio once mature. Field measurements show dune height (H) scales nearly linearly with dune width (W). A simple approximation is $H \approx 0.05W$ (meaning dune height is ~5% of the total base width). For instance, a dune 100 m wide would be ~5 m tall; a 200 m wide dune ~10 m tall. Observations in deserts and simulations suggest aspect ratios on the order of 1:10 to 1:20 (height:width) are common for transverse cross-sections of linear dunes ²⁹. Some studies define aspect ratio as height divided by half-width, in which case linear dunes often have values ~0.3–0.5 ³⁰ (this corresponds to height ~30–50% of half-width, or again ~6–10% of full width). **Crest sharpness** (width of crest vs. height) is also important: active dunes have very narrow crests (essentially knife-edged), but vegetated or stabilized linear dunes can develop broader, flattened crests; this can be quantified by a shape factor, though qualitative rules usually suffice (e.g. treat active crest as <5 m wide flat area for a dune tens of meters high). If modeling sub-grid topography, one might implement a **parabolic cross-section** for dunes (per simulations that show a parabola fits a dune's profile when no slipface is present ³¹), and then truncate at the angle of repose to form the slipface.
- **Dune Height-Spacing Relationship:** Numerous surveys have found that **dune spacing (λ)** is positively correlated with dune height (H) across different dune fields ²⁵. An empirical relationship from Sahara and Namib linear dunes is roughly linear: $\lambda = a + bH$. Lancaster (1982) reported a correlation ($R^2 \approx 0.59$) in the Namib Sand Sea where spacing increased with height ²⁶. A simplified rule is to take **spacing as proportional to height**, e.g., $\lambda \approx 10H$ (with a range from $5H$ to $15H$ depending on local factors). This can be refined: Wilson (1972) for Saharan dunes found a similar trend ²⁶. This relationship likely arises because larger dunes accumulate more sand and require more "breathing room" for winds to travel between them without immediately forming another dune. In modeling, one could set a base spacing and then adjust it if dunes grow: as a dune's height increases, slightly increase the spacing to its neighbor (or vice versa, if two dunes start crowding, one might eventually merge or one gets taller forcing spacing increase). This can also be related to wind regime: more variable winds (closer to star-dune conditions) could allow **closer spacing** because dunes don't grow as large vertically, whereas strongly bidirectional winds (with high sand flux) produce bigger dunes spaced further apart.
- **Migration and Elongation Rates:** For linear dunes, *elongation* along the crest is the dominant movement. The elongation rate v_e can be estimated from the along-crest sand flux. If Q_{\parallel} is the volume sand flux along the dune (per unit crest length) and A_{cs} is the cross-sectional area of the dune, then $v_e = \frac{Q_{\parallel}}{A_{\text{cs}}}$. This says the speed at which the dune end advances is the volume of sand arriving per year divided by the dune's cross-section that must be filled. For example, if $Q_{\parallel} = 30 \text{ m}^3/\text{m/year}$ (a reasonable sand flux along a dune crest ¹⁸) and the dune cross-section is 300 m^2 , then $v_e = 0.1 \text{ m/year}$. In a real case, Parteli et al. (2009) simulated seif dunes elongating ~1–3 m/year under strong bimodal winds ¹⁸. The **lateral migration** (sideways drift) velocity v_{lat} is much smaller; it can be related to any imbalance in transport from each side. One could estimate $v_{\text{lat}} \approx \frac{Q_A - Q_B}{A_{\text{cs}}}$ (projected normal to crest), where Q_A and Q_B are sand flux contributions from winds on either side. If winds are balanced, $Q_A \approx Q_B$ and $v_{\text{lat}} \approx 0$. If not, the dune will creep laterally. Observationally, migration rates for large

linear dunes are low: e.g., a 20 m high linear dune might shift lateral position by <1 m/year if one wind is moderately stronger. This aligns with general dune physics where migration speed inversely scales with height ²⁰. A known approximate formula for transverse dune migration is $v \sim \frac{1}{H}$ for a given wind climate ²⁰; linear dunes, being typically larger and bidirectionally fed, move even more slowly laterally.

- **Wind Variability Index (RDP/DP):** Fryberger & Dean's drift potential metrics are useful to quantify the wind regime for dune type. The *Drift Potential* (DP) is the total sand transport potential (sum of contributions from all wind directions), and the *Resultant Drift Potential* (RDP) is the magnitude of the vector sum of those contributions ¹¹. The ratio RDP/DP indicates how unidirectional the wind is. For linear dunes, RDP/DP tends to be intermediate (often around 0.4–0.8). In fact, Fryberger and Dean (1979) noted that **crescentic dunes and linear dunes** both had relatively high wind energy and moderately high RDP/DP, whereas star dunes had lower RDP/DP (more multidirectional winds) ¹¹. As a rule, one could set a threshold: require $RDP/DP \geq 0.3$ and ≤ 0.8 for linear dunes. If $RDP/DP > 0.9$, expect transverse dunes; if < 0.3 , expect star dunes. This ratio could also be tied to dune spacing: a higher RDP/DP (winds more consistently from two directions) might produce more aligned and possibly more widely spaced dunes, whereas a lower RDP/DP (winds more variable) could lead to irregular, closer-spaced or slightly net-curved linear dunes.
- **Crestline Sinuosity and Meander Wavelength:** The sinuosity S_c of a linear dune's crest (ratio of actual crestline length to straight-line length) is usually low (close to 1) but not exactly 1. If one wants to quantify it, natural seif dunes might have S_c values like 1.05–1.2 for mildly sinuous forms. This could be correlated with the wind directional imbalance: more alternation (and equal winds) yields larger oscillations (higher sinuosity). A **meander wavelength** L_m along the crest can be on the order of the distance the dune extends during one full cycle of wind regime oscillation. If seasonal winds alternate yearly, one might imagine a full S-shape forming over many cycles – but in practice dunes reach a statistically steady meander. No simple formula exists for L_m in nature, but one could impose $L_m \sim \frac{2\pi}{|f_A - f_B|} v_e$ where f_A, f_B are fractions of time of wind A and B, and v_e elongation speed, implying the crest might complete a full lateral shift cycle after enough length is added that each wind has dominated in turn. This is speculative; a simpler modeling choice is to just use a fixed wavelength or random long perturbation for crest shape.
- **Sand Flux and Slipface Angle:** The presence of slipfaces on alternating sides can be modeled with a threshold sand flux. For example, whenever the instantaneous sand transport from a given wind exceeds some value (enough to build a critical slope), a slipface forms. One could use the repose angle ($\sim 33^\circ$) to check dune flank slopes and enforce avalanching if exceeded. If sand accumulates such that local slope $> 33^\circ$, that segment of the model dune "fails" and sand is moved down to the base of that flank (creating the slipface deposit). This can be turned into a rule: *if local flank slope $> 33^\circ$, redistribute sand to lower layers until slope = 33°* . This rule ensures realistic steep faces develop. The avalanching process can be thought of in discrete steps or continuous slumping in a simulation.

These formulas and relationships provide a quantitative backbone to the qualitative rules. Incorporating them helps ensure that the generated dune field is not just visually plausible but also scales and behaves consistently. For instance, if one increases wind strength in the model, the dune spacing might increase (dunes grow taller, so they space out more), or if one changes the wind balance, the crest orientation formula can update the dune alignment. Using the above relationships, one can calibrate a generative

model to produce a range of linear dune scenarios that match real-world metrics (heights, spacings, etc.) observed in places like the Sahara, Namib, or Arabian deserts.

Visual References and Spatial Layout Guidance

Understanding real examples is key to modeling, so a few visual references can guide the spatial organization:

- **Satellite Images of Linear Dune Fields:** Overhead imagery (like from Google Earth or NASA satellites) reveals the striking parallelism and extent of linear dunes. For instance, satellite views of the **Rub' al Khali (Empty Quarter) in Arabia** or the **western Egyptian deserts** show nearly parallel seif dunes running for tens of kilometers, separated by flat interdunes. These images highlight features like *Y-junctions* (where two dunes join into one), consistent spacing, and gentle curvature responding to geography. In the Namib Desert image provided earlier, note how dunes are straight in uniform conditions but **bend near a valley** due to secondary winds ³. Such references suggest that when implementing a model, one might introduce controlled curvature in dunes near known wind corridors or edges.
- **Aerial and Ground Photographs:** Oblique aerial photos can illustrate dune cross-sections and surface details. For example, an oblique view of vegetated linear dunes in Australia's Simpson Desert shows razorback crests and the vast length of ridges with green interdune flats. Ground-level photos, while less common for linear dunes due to their remote locations, depict the dual slipfaces: one can stand atop a linear dune and see steep drops on both sides. They also show the crest often has a series of small scallops or crenulations from the alternating avalanches. These details remind us to include fine-scale irregularity in the crest line in a model (not perfectly smooth, but with small-scale undulations).
- **Schematic Diagrams:** Geologic schematics (from textbooks or papers) often label the parts of a linear dune: showing a cross-section with two slipfaces, the angle of repose, and the internal cross-bedding structure (alternating sets dipping in opposite directions). Such diagrams help in conceptualizing how to layer sediment in a synthetic model: one could generate internal strata by sweeping a cross-sectional template along the dune axis, alternating the dip direction per wind regime. Spatial layout schematics also illustrate dune interactions – e.g., a plan-view diagram might show several linear dunes with arrows indicating wind directions and where they merge. Incorporating those patterns (like a *bouquet* of dunes merging downwind, or a dune splitting around a hard obstacle) can make the generated output more true to nature.
- **Analogs in Other Planetary Settings:** Interestingly, linear dunes are not unique to Earth; they appear on Mars and Titan as well ³² ³³. For example, Cassini radar images of Titan show massive linear dunes hundreds of kilometers long, and HiRISE images on Mars show sinuous linear dunes in some craters. These alien examples reaffirm that the fundamental geometry – long parallel ridges shaped by two-directional winds – is a robust outcome of physics. They also sometimes make patterns more obvious (Titan's dunes, for instance, are extremely regular in orientation and spacing due to presumably consistent winds). Such patterns can inspire the **simplified base state** for modeling Earth dunes, before adding Earth's complexity (like vegetation or topographic interruptions).

In practice, it's beneficial to use a real dune field as a blueprint: e.g., trace a portion of a known linear dune field and analyze its crest spacing, sinuosity, and junction frequency, then use those statistics in generation. If no direct data, these visual guides at least ensure that the synthetic field *looks right* – with dunes mostly parallel, similar lengths, occasional junctions, and realistic variability.

Procedural Generation Algorithm (Pseudocode)

To integrate all the above rules and simulate a linear dune field, one can outline a step-by-step algorithm. Below is pseudocode that captures the multiscale structure and key physical rules for generating a static linear dune field, with notes for dynamic evolution:

```

Initialize wind regime parameters:
wind_dir1 = θ1 (e.g., 0° from north), wind_dir2 = θ2 (e.g., 120° from north)
wind_strength1 = W1 (relative drift potential from θ1)
wind_strength2 = W2 (relative drift potential from θ2)
resultant_dir = atan2(W2*sin θ2 + W1*sin θ1, W2*cos θ2 + W1*cos θ1) // dune orientation

Initialize domain dimensions (e.g., width X by length Y in meters)
Set average dune spacing S (e.g., S = k * H_avg, choose H_avg)
Set number of primary dunes N ≈ domain_width / S

// Generate base dune ridges
for i from 1 to N:
    // Starting position for dune i (jitter around a line so they are not perfectly aligned at edges)
    x_start = 0 + random_uniform(-0.1*S, 0.1*S) // slight offset at upwind boundary
    y_start = (i - 0.5)*S + random_uniform(-0.2*S, 0.2*S) // stagger initial lateral positions

    // Determine dune length L for this dune
    if using_sand_source_model:
        // Dune might start later (not at boundary) if source is patchy
        x_start = random_choice([0, small_positive]) // some start at boundary, some a bit in
        L = domain_length * random_uniform(0.8, 1.0) // many dunes span full length, some shorter
        if random_event(prob=0.2):
            L *= random_uniform(0.3, 0.7) // 20% of dunes are significantly shorter (gap or terminate early)

    // Create crest polyline for dune i
    crest = Polyline()
    crest.angle = resultant_dir + random_normal(0, 2°) // small random orientation variance

```

```

    crest.base_curve = straight_line(x_start, y_start, angle=crest.angle,
length=L)
    // Add gentle sinusoidal perturbation to crest line for realism
    crest = apply_sine_wave(crest.base_curve, amplitude=0.05*S,
wavelength=random_uniform(5*S, 10*S))
    // Note: ensure crest stays within domain bounds, clip if necessary

    // Assign dune height profile along crest
    H_max = H_avg * random_uniform(0.8, 1.2)
    if L < 0.5*domain_length:
        // shorter dune - perhaps type B that tapers
        H_max *= 0.8 // shorter dunes a bit smaller
    // Height gradually increases from start, then maybe decreases near end
    crest.height(z) = H_max * sin( π * distance_along(crest)/L )^p
        // using a sine-shaped longitudinal profile (p ~1 for
symmetric rise/fall, or skew it if needed)
    // Optionally, for type A vs B differentiation:
    if L > threshold_length:
        // Type A: maintain height or increase slightly toward end
        crest.height(z) = H_max (nearly flat profile except small undulations)
    else:
        // Type B: decrease height towards downwind end
        crest.height(z) = H_max * (1 - distance_along(crest)/L)^0.5 // taper
height

    // Assign dune width from height (simple linear scaling)
    crest.width(z) = 10 * crest.height(z) // e.g., width ~10*height (can refine
or randomize factor)

    // Store dune crest geometry and attributes
    dunes[i].crest = crest

// Introduce dune interactions (mergers, splits) after initial placement
for each adjacent pair (dune_i, dune_j):
    if distance_between_endpoints(dune_i, dune_j) < merge_threshold:
        // Create Y-junction merge
        merge_point = midpoint(dune_i.end, dune_j.end)
        // Connect j's end to i's crest near merge_point
        dunes[j].crest.attach_to(dunes[i].crest, at=merge_point, angle_diff <
30°)
        mark dunes[j] as merged_into i
    if random_event(prob=0.1) and dune_i.length > split_threshold:
        // Split dune_i into two branches at some point
        split_pt = random_point_on( dune_i.crest, between=0.3L and 0.7L )
        create new crest branch starting at split_pt with same orientation +
small offset
        length_new = (dune_i.total_length - split_pt.distance_to_start) *
random_uniform(0.5,0.8)

```

```

branch = extend_line_from(split_pt, angle=crest.angle + 5°, length_new)
// treat branch as new dune or secondary crest attached to dune_i
dunes.add(branch as new dune or sub-dune)

remove dunes marked merged // those become part of longer dune

// Assign slipfaces based on assumed last wind
current_wind = choose(wind_dir1 or wind_dir2) // e.g., pick the dominant or a
random one for snapshot
lee_side_angle = current_wind == θ1 ? θ1 + 180° : θ2 + 180°
for each dune in dunes:
    // Determine which flank is lee in this snapshot
    crest_orient = dune.crest.angle
    // Angle between current wind and crest orientation
    diff = angle_between(current_wind, crest_orient)
    if |diff| < 90°:
        // wind blows from one side of crest
        lee_flank = downwind side relative to wind (could determine via cross
product or so)
    else:
        // If current wind almost parallel to crest (rare, since usually
oblique), choose a side arbitrarily or none
        lee_flank = None
    // Set slope on lee_flank to 33° and windward flank to ~10°
    if lee_flank is not None:
        dune.profile.cross_section.setSlope(lee_flank, 33°)
        dune.profile.cross_section.setSlope(other_flank, 10°)
    else:
        // no pronounced slipface in this case, keep both flanks ~15°
        dune.profile.cross_section.setSlope(both_flanks, 15°)
    // Optionally, if lee_flank spans entire dune? Could vary slipface
continuity:
    if random_event(prob=0.3):
        introduce a gap in the slipface (no slipface for a segment to simulate
patchiness)

// Smaller scale surface features
for each dune:
    generate ripples on flanks:
        ripple_wavelength = 1-5 m (based on grain size/wind)
        ripple_orientation: perpendicular to crest on windward side, somewhat
radial on lee side
        texture dune flanks with these ripples (e.g., sinusoidal micro
undulations)
    if dune.height > large_threshold and random_event(prob=0.5):
        // add secondary dune on crest (compound dune)
        sec_length = dune.width * random_uniform(1,2) // small dune maybe 1-2

```

```

dune widths long
    sec_crest = small ridge on top of main crest (offset slightly to one
side) running parallel
        // no need to be detailed, just visual hint of smaller dune on big one

// Vegetation or moisture effects if any
if veg_cover_fraction > 0:
    for each dune:
        apply vegetation to lower flanks/interdune randomly (e.g., cluster
vegetation in interdunes)
            where vegetated, reduce local dune height slightly and smooth the top
(simulate stabilization)
            maybe remove slipface if heavily vegetated (stabilized slope)
if water_table_shallow:
    mark some interdune areas as wet (no loose sand) and ensure dunes don't
occupy those fully (they stop or divert)

// Finalize the elevation model
Initialize a grid or mesh for the domain
for each dune:
    for each point along crest (discretize crest polyline):
        generate cross-sectional dune shape (e.g., triangular with given height,
width, slipface etc.)
            add to elevation grid (raise elevation)
        ensure superposition: if dunes overlap, take the highest elevation (dunes
merging) or sum? (though merges handled above)
    Optionally add subtle noise to surface to mimic small-scale irregularities

Output: Digital elevation model (or rendered image) of linear dune field

```

This pseudocode first establishes wind and orientation, then places parallel dune ridges with some randomness in position and length. It then handles special cases of merging and splitting to create compound structures. Next, it assigns slipfaces according to a chosen current wind and adds detailed features like ripples and secondary dunes. Finally, it can output a topographic model. During dynamic time-stepping, one could loop through winds and incrementally adjust crest positions (elongating ends, shifting laterally) and update slipfaces, but for a single snapshot the above procedure suffices.

Note: This algorithm can be adjusted for different scenarios – for instance, increasing the probability of splits for a more complex network, or introducing a second set of dunes at a different orientation to simulate an older generation of dunes beneath the current one (as seen in regions with shifting paleowind regimes). The key is to maintain the **physical consistency**: dunes aligned with resultant winds, alternating slipfaces, correlated height/spacing, and logical interactions. By following these rules and steps, the generated analog images of linear dune fields should convincingly mirror real-world geology, capturing the multiscale structure from erg down to ripple.

Sources: The above rules and parameters are informed by field measurements and studies of linear dunes in places like the Namib and Sahara [2](#) [1](#), theoretical and modeling work on bimodal wind dune

formation 9 21, and empirical relationships noted by researchers (e.g. Lancaster 1982 on dune spacing 26, Fryberger & Dean 1979 on wind regimes 11).

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