POLAR background prediction

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Overview

Gamma Ray Bursts (GRBs)

- What are GRBs?

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  - Bursts of high-energy photons
  - Bursts of gamma rays with an undetermined cause
- Lead to spikes in photon counts/rates

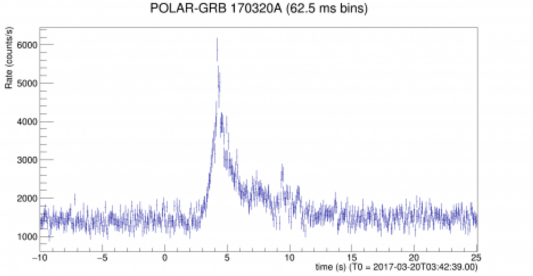

Figure: A light-curve containing a Gamma Ray Burst[1]

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- For this purpose, there's a need for:
  - Data taking → POLAR detector

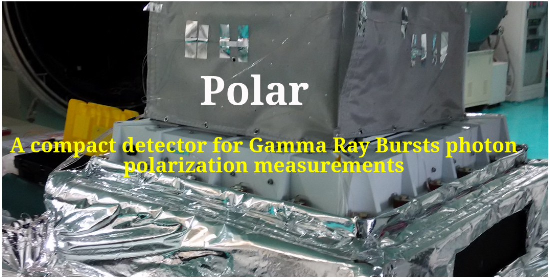

Figure: Picture taken from `https://www.astro.unige.ch/polar/`

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Related work

- Koziol's master thesis [1]: HAGRID[2]

[2] High Accuracy GRB Rapid Inference with Deep Learning in Space

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    - Predict the occurrence of a GRB at time t

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Related work

- Koziol's master thesis [1]: HAGRID[3]
  - Possible data distribution mismatch

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  - But low miss-detection: correctly detected all real GRBs within 10 days of POLAR data.

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  - $\rightarrow$ Gradients w.r.t input variables

Methodology

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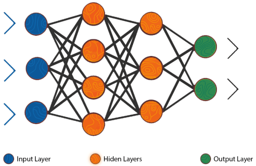
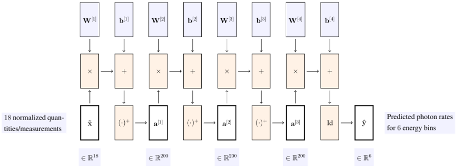
Background model

- Multi-Layer Perceptron (MLP) background model[5]



(a) Multi-Layer Perceptron high-level



(b) Graph of operators describing my Multi-Layer Perceptron

[5]First picture taken from https://medium.com/unpackai/from-anns-artificial-neural-networks-to-rnns-recurrent-neural-networks-93b638772fd1

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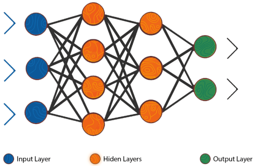

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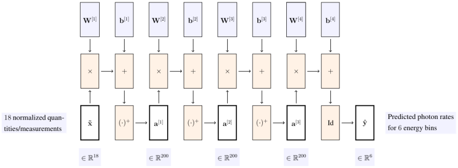

(b) Graph of operators describing my Multi-Layer Perceptron

- Trained to minimize a weighted MSE

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Poor predictions

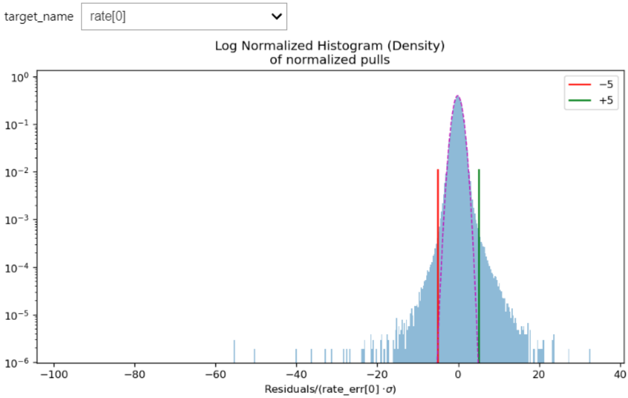

Figure: Normalized pulls for rate[0]

Cluster and cluster intersections

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- GRBs should be visible in different energy bins
- Cluster of data points extracted using different energy bins:
  - Each from the same set of energy bins
  - Or each satisfying same conditions (e.g. number of energy bins)

Results & discussion: Known GRBs

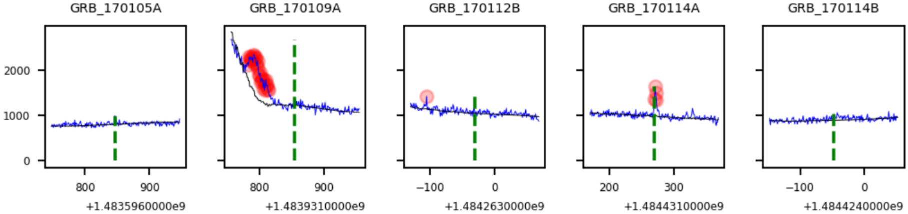

Figure: Energy bin 0, time windows of ±100 [s] around 5 known GRB trigger times [2], based on pulls with $k = 5$. Predictions are in black.

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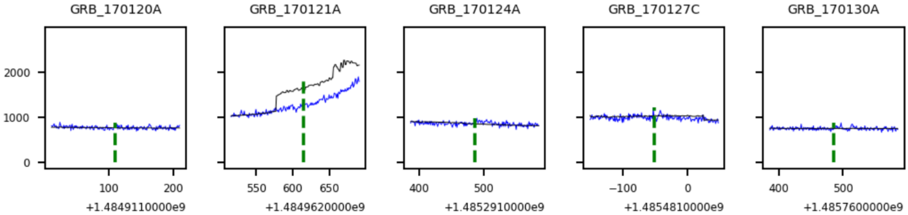
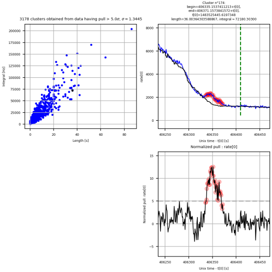
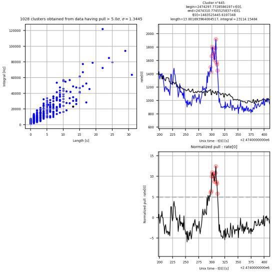


Figure: Energy bin 0, time windows of ± 100 [s] around 5 other known GRB trigger times [2], based on pulls with $k = 5$.

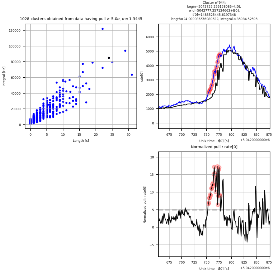
Results & discussion: Cluster intersections



(a) Cluster intersection with more the 3 energy bins; Solar flare confirmed by Prof. Nicolas Produit

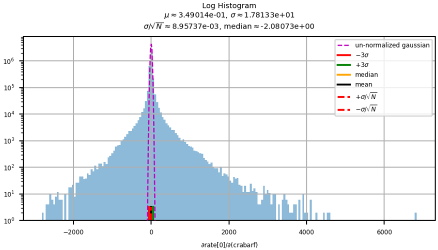


(b) Undetermined cluster intersection with all energy bins and with a 30-second discard window.

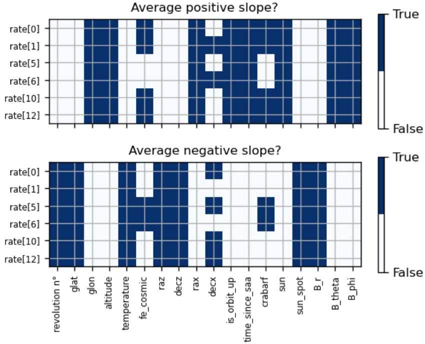


(c) Undetermined cluster intersection with all energy bins and with a 30-second discard window.

Results & discussion: Model interpretability



(a) Histogram of $\frac{\partial \text{rate}[0]}{\partial \text{crabarf}}$ (un-normalized features)



(b) Avg slope μ greater (lesser) than its (negative) standard error $\frac{\sigma}{\sqrt{N}}$?

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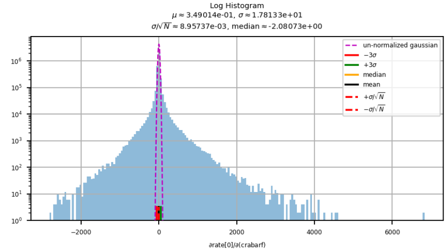

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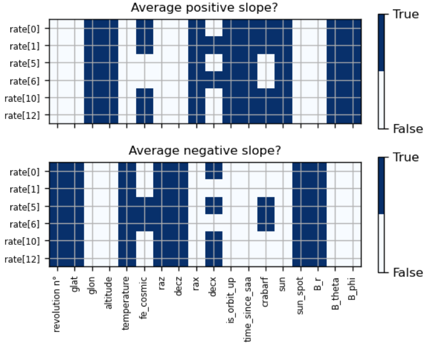

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- Need for further analysis

Results & discussion

Table 6: Mean ± standard deviation (over 10 different seeds) number of positive/negative cluster intersections for different sets of energy bins bs or conditions[29]. This time, there's a discard window of 30 seconds.

bs or condition	negative	positive	bs or condition	negative	positive	bs or condition	negative	positive
# inter > 0	5762.1 ± 965.708	6487.6 ± 826.131	0,2	6.9 ± 2.726	10.7 ± 4.111	1,3	15.6 ± 6.398	51.2 ± 42.695
# inter > 1	4802.1 ± 822.34	5450.6 ± 752.954	0,2,3	6.0 ± 3.432	7.1 ± 1.37	1,3,4	1.0*	–
# inter > 2	3493.8 ± 610.247	4083.5 ± 564.85	0,2,3,4	8.0*	–	1,3,4,5	1.333 ± 0.577	1.5 ± 0.577
# inter > 3	3238.4 ± 602.769	3821.5 ± 538.957	0,2,3,4,5	212.2 ± 42.158	242.7 ± 65.109	1,3,5	–	5.5 ± 6.364
# inter > 4	1150.2 ± 232.379	1595.0 ± 202.499	0,2,3,5	1.0 ± 0.0	3.5 ± 3.536	1,4	1.0*	3.0*
# inter > 5	791.3 ± 165.14	1135.2 ± 149.199	0,2,4	1.5 ± 0.707	1.0*	1,4,5	2.125 ± 1.356	5.5 ± 3.742
0	23.8 ± 8.23	71.6 ± 18.368	0,2,4,5	5.8 ± 4.367	15.8 ± 5.095	1,5	–	2.0 ± 0.0
0,1	16.6 ± 6.059	62.3 ± 45.631	0,2,5	1.0 ± 0.0	2.5 ± 2.38	2	734.8 ± 169.59	939.8 ± 235.236
0,1,2	4.0 ± 2.0	15.2 ± 9.739	0,3	1.0*	3.0 ± 1.414	2,3	2459.9 ± 493.897	3265.5 ± 779.499
0,1,2,3	12.4 ± 3.806	30.5 ± 17.219	0,3,4,5	1.667 ± 1.155	1.75 ± 0.957	2,3,4	114.8 ± 335.669	11.75 ± 12.395
0,1,2,3,4	23.0*	2.5 ± 0.707	0,4	2.0*	–	2,3,4,5	3156.9 ± 657.884	4049.1 ± 676.033
0,1,2,3,4,5	791.3 ± 165.14	1135.2 ± 149.199	0,4,5	2.25 ± 1.753	11.6 ± 3.062	2,3,5	60.444 ± 11.479	225.0 ± 392.74
0,1,2,3,5	1.0 ± 0.0	3.667 ± 4.899	0,5	1.0*	1.75 ± 1.5	2,4	18.333 ± 38.563	2.333 ± 1.751
0,1,2,4	1.0*	3.0*	1	52.6 ± 15.16	104.0 ± 118.135	2,4,5	212.3 ± 59.913	244.9 ± 73.942
0,1,2,4,5	2.8 ± 1.751	11.2 ± 7.315	1,2	3.444 ± 1.424	8.5 ± 4.143	2,5	9.8 ± 6.356	55.5 ± 127.411
0,1,2,5	1.0*	2.667 ± 2.082	1,2,3	56.1 ± 19.221	115.1 ± 41.391	3	1206.3 ± 282.876	1633.0 ± 282.032
0,1,3	1.25 ± 0.5	4.875 ± 7.2	1,2,3,4	14.75 ± 27.5	1.4 ± 0.548	3,4	6.25 ± 9.215	1.0 ± 0.0
0,1,3,4,5	2.5 ± 2.811	2.25 ± 0.5	1,2,3,4,5	413.5 ± 141.535	714.4 ± 132.802	3,4,5	19.7 ± 8.166	16.5 ± 4.503
0,1,3,5	–	1.0*	1,2,3,5	2.167 ± 0.753	9.667 ± 19.268	3,5	1.889 ± 0.782	21.0 ± 58.877
0,1,4	1.0*	6.0*	1,2,4	–	4.0*	4	9.333 ± 20.174	3.833 ± 4.167
0,1,4,5	2.875 ± 1.959	5.9 ± 4.202	1,2,4,5	3.143 ± 2.673	4.889 ± 3.333	4,5	123.5 ± 30.395	124.0 ± 31.383
0,1,5	2.0*	2.0*	1,2,5	–	3.0*	5	7.3 ± 3.302	88.1 ± 234.779

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- Brief peek at model interpretability using partial derivatives of output w.r.t. input

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  - Change the methodology, e.g. use sequential models

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- Other possible improvements:
  - Dive deeper into model interpretability
  - Change the methodology, e.g. use sequential models
  - Improve code quality, GPU usage and decrease memory footprint.

The End

References

Figures without references come from author of the slides.

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