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Section #1. Class 1

Attempts in Relativistic QM.

Axioms: < Schrödinger Equation.

States are represented by vectors in Hilbert space

Observables by Hermitian operators

Measurements yield one of the eigenvalues

Identical particles, etc.

$$i\hbar \frac{\partial}{\partial t} |\psi\rangle = H|\psi\rangle$$

Let's consider the simplest

$|\psi, t\rangle$ in Griffith was the general vector in Hilbert space

$$H = \frac{1}{2m} \vec{p}^2$$

$$i\hbar \frac{\partial}{\partial t} \psi(\vec{x}, t) = -\frac{\hbar^2}{2m} \nabla^2 \psi(\vec{x}, t)$$

$|S(t)\rangle$ and

$\psi(x, t) = \langle x | S(t) \rangle$ $\psi(x, t) = \langle x | \psi, t \rangle$ is the position space

$= \int d^3y \delta(x-y) \psi(y, t)$ wave function.

$= \psi(x, t)$ We want to generalize into relativistic version

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$$\mathcal{H} \sim E \rightarrow \mathcal{H} = \sqrt{\phi_c^2 + m_c^2 c^4}$$

$$\mathcal{H} = \underbrace{mc^2}_{\text{constant}} + \underbrace{\frac{1}{2m} p^2}_{\text{QM}} + \underbrace{\dots}_{\text{corrections}}$$

Now SE becomes.

$$i\hbar \frac{\partial}{\partial t} \psi(x, t) = \underbrace{t \sqrt{m^2 c^4 + \cancel{\phi^2 c^2} - \hbar^2 c^2 \nabla^2}}_{\text{still treats space-time differently.}} \psi(x, t)$$

non-linear so solving it is problematic in general.

How about we "square" it?

$$-\hbar^2 \frac{\partial^2}{\partial t^2} \psi(x, t) = (m^2 c^4 + -\hbar^2 c^2 \nabla^2) \psi(x, t)$$

~~$\mathcal{X}^\mu = (x, t)$~~

$$\mathcal{X}^\mu = (ct, \vec{x})$$

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and a contravariant

$$\chi_\mu = (-ct, \vec{x}') \quad . \quad \mu = 0, 1, 2, 3.$$

Minkowski metric.

$$g_{\mu\nu} = \begin{pmatrix} -1 & & & \\ & +1 & & \\ & & +1 & \\ & & & +1 \end{pmatrix}$$

$$\chi_\mu = g_{\mu\nu} \chi^\nu$$

the inverse

$$\chi^\mu = g^{\mu\nu} \chi_\nu \quad \text{require that}$$

$$\text{requires } g^{\mu\nu} = \begin{pmatrix} -1 & & & \\ & +1 & & \\ & & +1 & \\ & & & +1 \end{pmatrix}$$

so we have

$$g^{\mu\nu} g_{\nu\rho} = \delta_\rho^\mu \quad \text{the Kronecker Delta.}$$

$$= \begin{pmatrix} 1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$$

Special Relativity tells us that physics looks the same in all inertial frames.

x^μ specify a inertial frame.

another coordinates \bar{x}^μ can be related to it by

$$\bar{x}^\mu = \underbrace{\Lambda^\mu_\nu}_{\text{const}} x^\nu + \underbrace{a^\mu}_{\text{const.}}$$

↓
Lorentz transformation → translation vector.

Furthermore, we need to require that.

$$g_{\mu\nu} \Lambda^\mu_\rho \Lambda^\nu_\sigma = g_{\rho\sigma} \quad (\text{recall proper time}),$$

which keeps the length unchanged.

"interval" between two points

as

$$\begin{aligned} (x - x')^2 &= g_{\mu\nu} (x - x')^\mu (x - x')^\nu \\ &= (x - x')^2 - c^2 (t - t')^2 \end{aligned}$$

$$(\bar{x} - \bar{x}')^2 = g_{\mu\nu} (\bar{x} - \bar{x}')^\mu (\bar{x} - \bar{x})^\nu$$

$$= g_{\mu\nu} \Lambda^\mu_\rho \Lambda^\nu_\sigma (x - x')^\rho (x - x')^\sigma$$

$$= g_{\rho\sigma} (x - x')(x - x)^\sigma$$

$$= (x - x')^2. \text{ as desired.}$$

Now special relativity Equivalence demands

$$\gamma(x) = \gamma(\bar{x})$$

and obey same equations in their own frame.

lets define

$$\partial_\mu = \frac{\partial}{\partial x^\mu} = \left(\frac{1}{c} \frac{\partial}{\partial t}, \vec{\nabla} \right)$$

$$\partial^\mu = \frac{\partial}{\partial x_\mu} = \left(-\frac{1}{c} \frac{\partial}{\partial t}, \vec{\nabla} \right).$$

$$\underbrace{\partial^\mu x^\nu}_{g_{\mu\nu}} = g_{\mu\nu}.$$

Since $\bar{x}^\mu = \Lambda^\mu_\nu x^\nu + a^\mu$

$$\bar{\partial}^\mu = \Lambda^\mu_\nu \partial^\nu.$$

we can easily check that.

$$\bar{\partial}^\rho \bar{x}^\sigma = g^{\rho\sigma}$$

as

$$\begin{aligned}\bar{\partial}^\rho \bar{x}^\sigma &= \Lambda^\rho_\nu \partial^\nu (\Lambda^\sigma_\mu x^\mu + a^\sigma) \\ &= \Lambda^\rho_\nu \partial^\nu x^\mu \Lambda^\sigma_\mu \\ &= \Lambda^\rho_\nu \Lambda^\sigma_\mu g^{\mu\nu} = g^{\rho\sigma}.\end{aligned}$$

Now $\partial^2 = \partial^\mu \partial_\mu = -\frac{1}{c^2} \frac{\partial}{\partial t^2} + \vec{\nabla}^2$.

K.G. $(-\partial^2 + m^2 c^2 / h^2) \psi(x, \theta) = 0$.

And in \bar{x} inertial frame

$$(-\bar{\partial}^2 + m^2 c^2 / h^2) \psi(\bar{x}, \theta) = 0$$

Now

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we need to show that the two K.G
in two frames are the same,

$$\bar{\partial}^2 = g_{\mu\nu} \bar{\partial}^\mu \bar{\partial}_\nu = g_{\mu\nu} N_\mu^a N_\nu^b \partial^c \partial^d$$

$$= g_{\rho\sigma} \partial^\rho \partial^\sigma = \partial^2$$

so these are two identical equations.

But K.G Equation doesn't obey SE.

and the equation changes from first order to
second order.

then the normalization of states

$$\langle \psi, t | \psi, t \rangle = \int d^3x \langle (\psi, t) | \psi, t \rangle \langle \psi, t | \psi, t \rangle$$

$$= \int d^3x \psi^*(x) \psi(x)$$

is not time independent.

thus probability is not conserved

one can see
more details by working out

$\frac{d}{dt} \langle \psi | \psi \rangle$ (K.G doesn't help
here since it is 2nd order)

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$$H_{ab} = C P_j (\alpha^j)_{ab} + M C^2 (\rho)_{ab}$$

Skipping Dirac's effort, as we'll be reworking it
 with a fresh eye of symmetries
 linear & conserves probability. & Representations
 (later this of course.)

found anti-particle states and Dirac sea.
 negative Energy.]

And it's already about
 infinite # of particles.

but still we can't describe bosons.

we need a pause: and now we realise that
 realise.

Time is not yet an ~~op~~ Hermitian operator.
 so we never treated it properly.

shopped
Jan. 22. 25

Two operations

} denote \bar{X}' to a label

} promote t to an operator.

When promote $t \rightarrow T$ operator.

one can use the proper time, τ in SE.

coordinate time, in Heisenberg Picture.

$$\hat{X}^{\mu}(\tau) \text{ has } \hat{X}^0 = T.$$

many complications, any monotonic function of τ
is good here so infinite redundancy.

Also one can consider more parameters,

$\hat{X}^{\mu}(s, \tau)$ now classically this gives
a family of world lines that we call world sheets.

$\hat{X}^{\mu}(s, \tau)$ is a propagating string
beyond my scope here.

Another option is to demote X^{μ} as labels on operators.

consider an operator $\phi(\vec{x})$.

A set of such operators are call quantum fields.

$\phi(\vec{x}) \rightarrow$
Heisenberg Picture

$$\phi(\vec{x}, t) = e^{iHt/\hbar} \phi(\vec{x}, 0) e^{-iHt/\hbar}.$$

so now \tilde{x}, t are labels on operators $\varphi(\tilde{x}, t)$.

none is an eigenvalue of an operator.

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Quantum Field.

These two options are equivalent.

but $\varphi(\tilde{x}, t)$ corresponds to Quantum Field Theory

is much more convenient for most problems, which
is the subject of this course.

For example of classical field theory, see, e.g. Tiny, Srednicki, etc.

There is another desired property,

Normal fixed #, QM can be written as
non-relativistic QFT.

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Let's begin with a position space SE. for n particles.

~~$$i\hbar \frac{\partial}{\partial t} \Psi = \left[\sum_{j=1}^n \left(-\frac{\hbar^2}{2m} \vec{\nabla}_j^2 + U(\vec{x}_j) \right) + \sum_{j=1}^n \sum_{k=1}^{j-1} V(x_j - x_k) \right] \Psi$$~~

Here ~~$\Psi = \Psi$~~ $\Psi = \Psi(\vec{x}_1, \vec{x}_2, \vec{x}_3, \dots, \vec{x}_n; t)$

The QM can be rewritten by introducing a quan-

field $a(\vec{x})$ and $a^\dagger(\vec{x})$ to have

$$[a(\vec{x}), a(\vec{x}')] = 0$$

$$[a^\dagger(\vec{x}), a^\dagger(\vec{x}')] = 0$$

$$[a(\vec{x}), a^\dagger(\vec{x}')] = \delta^3(\vec{x} - \vec{x}')$$

$a^\dagger(\vec{x})$ and $a(\vec{x})$ is like SHO's creation & annihilation operators.

We can introduce a new H for our QFT.

$$H = \int d^3x \ a^\dagger(\vec{x}) \left(-\frac{\hbar^2}{2m} \vec{\nabla}^2 + U(\vec{x}) \right) a(\vec{x})$$

$$+ \frac{1}{2} \int d^3x d^3y \ V(\vec{x} - \vec{y}) a^\dagger(\vec{x}) a^\dagger(\vec{y}) a(\vec{y}) a(\vec{x})$$

Now consider a state

$$|\Psi, t\rangle = \int d^3x_1 d^3x_2 \dots d^3x_n \ \Psi(\vec{x}_1, \vec{x}_2, \dots, \vec{x}_n, t) \\ a^\dagger(\vec{x}_1) a^\dagger(\vec{x}_2) \dots a^\dagger(\vec{x}_n) |0\rangle$$

here $|0\rangle$ is the vacuum state that

$$a(\vec{x})|0\rangle = 0.$$

We can show that SE of \mathcal{H}^{QFT}
is satisfied if and only if

$$\# \psi(\vec{x}_1, \vec{x}_2, \dots, \vec{x}_{n+1})$$

satisfies SE of \mathcal{H} .

so they are equivalent

and $a^+(\vec{x}_i)|0\rangle$ creates a particle of at position \vec{x}_i
can be interpreted as

the operator $N = \int d^3x a^+(\vec{x})a(\vec{x})$ counts the total # of
particles.

$[N, H] = 0$ so particle # is conserved.

but for a general theory, e.g.

$$\mathcal{H} > \int d^3x [a^+(\vec{x})a(\vec{x}) + \text{h.c.}]$$

$$\begin{aligned} [a(x), a^+(y)a(y)] \\ \stackrel{\text{e.g.}}{=} \delta^3(x-y) \end{aligned}$$

$$\begin{aligned} [a(x), a^+(y)a(y)] \\ = -\delta^3(x-y) \end{aligned}$$

so even powers
cancel & commutes

And it acts on Ψ

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$\mathcal{Q}M$ will fail
and $\mathcal{Q}FT$ prevail.

There are also requirements for bosons & fermions,
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different
that we will study in near future classes.

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falls unter numerischer in $\langle 0 | \dots$

$[V, N]$

$$= \left[\int d^3x d^3y V(\vec{x} - \vec{y}) a^\dagger(x) a(y) a(y) a(x), \int d^3z a^\dagger(z) a(z) a(z) \right]$$

$\xrightarrow{\text{to } \mathcal{H}}$ falls numerisch in $\langle 0 | \dots$

$$\int d^3x d^3y d^3z \left(V(\vec{x} - \vec{y}) a^\dagger(x) a^\dagger(y) a(y) a(x) a^\dagger(z) a(z) \right)$$

$\xrightarrow{\text{faktor } \dots, \int d^3x \dots}$

$$- \underbrace{a^\dagger(z) a(z) \left(V(\vec{x} - \vec{y}) a^\dagger(x) a^\dagger(y) a(y) a(x) \right)}$$

$\xrightarrow{\text{faktor } \dots}$

numerisch in $\langle 0 | \dots$

$\xrightarrow{\text{Faktor } \dots}$ abheben \rightarrow andere $\langle 0 | \dots \rangle$. Nun

$$\left(\int d^3x \dots + a^\dagger(x) a(x) \right) a^\dagger(y) \xrightarrow{\text{zu konservativeren und aus}}$$

$$\xrightarrow{\text{faktor } \dots}$$

$$\left(\int d^3x \dots a^\dagger(y) + a^\dagger(x) \left(\int d^3y \dots + a^\dagger(y) a(y) a^\dagger(z) \right) \right)$$

$\xrightarrow{\text{faktor } \dots}$

$$\left(\int d^3x \dots + a^\dagger(x) \left(\int d^3y \dots + a^\dagger(y) a(y) a^\dagger(z) \right) \right) = \mathcal{V} \quad \text{retarriert mit}$$

abheben

numerisch in $\langle 0 | \dots \rangle$ oder $\langle 0 | \dots \rangle$

$$[\psi^\dagger(x) \psi^\dagger(y), \psi(x) \psi(y)]$$

$$(\psi - \psi^\dagger)^2 =$$

$$[\psi^\dagger(x) \psi^\dagger(y), \psi(x) \psi(y)]$$

$$(\psi - \psi^\dagger)^2 =$$

numerisch in $\langle 0 | \dots \rangle$

oder numerisch in $\langle 0 | \dots \rangle$

p.v. - joint numerisch nicht funden

$$[\psi^\dagger(x) \psi^\dagger(y), \psi(x) \psi(y)] \propto \delta^3(p) < \infty$$