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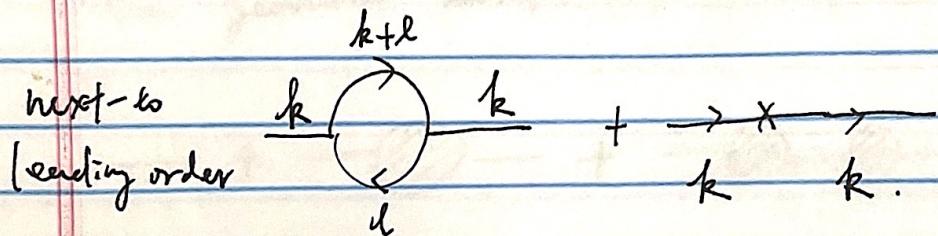
Loop corrections to the propagator (Sec 14)

In path integral, we wrote the exact propagator as

$$\frac{1}{i} \Delta(x_1 - x_2) \equiv \langle 0 | T \psi(x_1) \psi(x_2) | 0 \rangle = \delta_{x_1} \delta_{x_2} iW(J) \Big|_{J=0}$$

$iW(J)$ is the sum of all connected diagrams

(leading order (tree level)) — which is $\Delta(x_1 - x_2)$



using Feynman Rules

$$\frac{1}{i} \tilde{\Delta}(k^2) = \frac{1}{i} \Delta(k^2) + \frac{1}{i} \tilde{\Delta}(k^2) [i\Gamma(k^2)] \frac{1}{i} \tilde{\Delta}(k^2) + O(g^4)$$

where $\tilde{\Delta}(k^2) = \frac{1}{k^2 + m^2 - i\epsilon}$

which is the free-field propagator.

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from the diagram we can read that

$$i\pi(k^2) = \frac{1}{2} (ig)^2 \left(\frac{1}{i}\right)^2 \int \frac{d^4 l}{(2\pi)^4} \tilde{\Delta}((l+k)^2) \tilde{\Delta}(l^2)$$

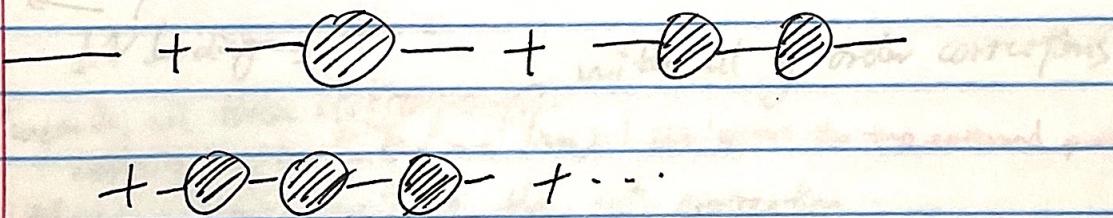
Symmetry factor

$$- i(Ak^2 + Bm^2) + O(g^4)$$

$\uparrow \quad \uparrow$

$$(Z_Q - 1) \quad (Z_m - 1).$$

but we can see these effect can be further summed via geometric series.



$$\begin{aligned} \frac{1}{i} \triangle(k^2) &= \frac{1}{i} \tilde{\Delta}(k^2) + \frac{1}{i} \tilde{\Delta}(k^2) [i\pi(k^2)] \frac{1}{i} \tilde{\Delta}(k^2) \\ &\quad + \frac{1}{i} \tilde{\Delta}(k^2) [i\pi(k^2)] \frac{1}{i} \tilde{\Delta}(k^2) [i\pi(k^2)] \frac{1}{i} \tilde{\Delta}(k^2) \end{aligned}$$

+ ...

$$\begin{aligned} \triangle(k^2) &= \frac{i\tilde{\Delta}(k^2)}{1 - \pi(k^2)\tilde{\Delta}(k^2)} \\ &= \frac{1}{k^2 + m^2 - ie - \pi(k^2)} \end{aligned}$$

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We know that the exact propagator has a pole at $k^2 = -m^2$ with residue of one, this means (only and only if)

$$\Pi(-m^2) = 0$$

$$\Pi'(-m^2) = 0 \quad \left(\frac{d\Pi(k^2)}{dk^2} \Big|_{k^2=-m^2} = 0 \right)$$

This is equivalent to canonical quantization
 $\langle k | k' \rangle$

Note that here

$$-\bullet - \subset -\circ- + \text{---} + -\circ-$$

one-particle irreducible

\nwarrow

IPI diagrams + ...

(meaning all those still any cuts). with all g^2 order corrections
 connected after cutting one line) [We ignore the two external propagators]

Now we can evaluate the g^2 correction

(let's do the integral)

[This avoids double counting of
 the geometric series we did
 earlier]

To evaluate, we need to invoke

"Feynman's formula" that. (to combine
 denominators)

$$\frac{1}{A_1 A_2 \dots A_n} = \int dF_n (x_1 A_1 + x_2 A_2 + \dots + x_n A_n)^{-n}$$

Here the integration is over Feynman parameter x_i (4)

$$\int dF_n = (n-1)! \int_0^1 dx_1 dx_2 dx_3 \dots dx_n \delta(x_1 + x_2 + x_3 + \dots + x_n - 1).$$

Let's prove such an identity in HW.
Hence, knowing one more trick, we now can calculate.

$$\begin{aligned} \Gamma((k+l)^2) \Gamma(l^2) &= \frac{1}{\ell^2 + m^2} \frac{1}{(k+l)^2 + m^2} \\ &= \int_0^1 dx [x((\ell+k)^2 + m^2) + (1-x)(\cancel{\ell^2 + m^2})]^{-2} \\ &= \int_0^1 dx [\ell^2 + 2k\ell x + m^2 + xk^2]^{-2} \\ &= \int_0^1 dx [(l+xk)^2 + x(1-x)k^2 + m^2]^{-2} \\ &= \int_0^1 dx [g^2 + D]^{-2} \\ g &= \ell + xk, \quad D = x(1-x)k^2 + m^2. \end{aligned}$$

The jacobian for changing variable
from $d^d l \rightarrow d^d g$ is trivially one.

Now as g^0 and g^i have different signs in the metric.
let's think about g^0 . as ~~g^0~~ g appears in
squares (why not? the L.I. quantity)

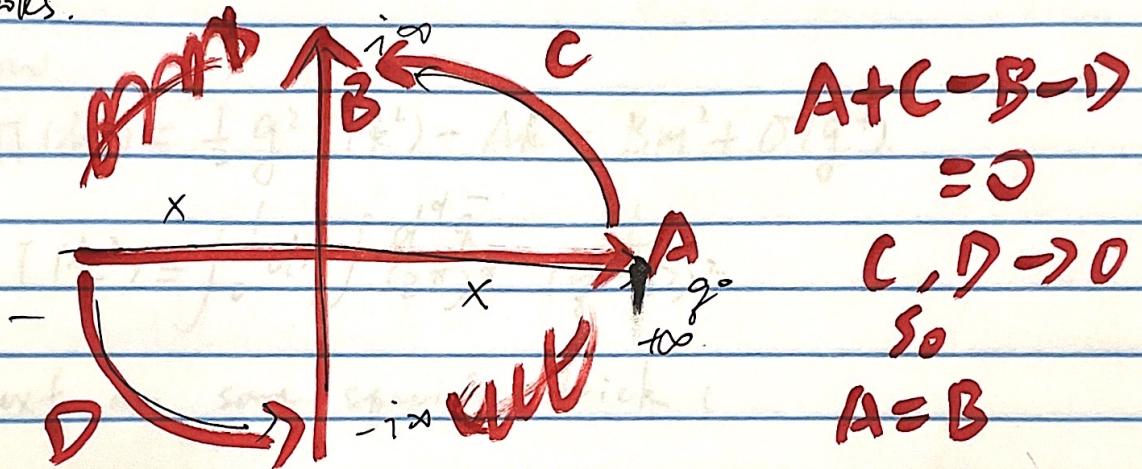
Let's perform the Wick Rotation
 $g^0 \rightarrow ig^0$.

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If the integrand vanishes fast enough in $|q^0| \rightarrow \infty$

This integral won't change value with such rotating
contour

Here, it are super helpful in keeping track of the poles.



With such rotation, we change the integration to Euclidean d-dimensional \bar{q}^0 with

$$\bar{q} = (\bar{q}_1, \bar{q}^0) \quad \bar{q}^0 = \sqrt{\bar{q}_j^2} \quad \text{and} \quad \bar{q}_j = \bar{q}_j$$

$$\bar{q}^2 = \bar{q}_1^2 + \bar{q}_2^2 + \dots + \bar{q}_d^2 \quad \text{so} \int_{-\infty}^{+\infty} d\bar{q}^0 f(\bar{q}^0)$$

$$\text{with} \quad \bar{q}^2 = \bar{q}_1^2 + \bar{q}_2^2 + \dots + \bar{q}_d^2$$

$$d^d \bar{q} = i d^d \bar{q}^0$$

$$= \int_{-\infty}^{+\infty} d\bar{q}^0 f(\bar{q}^0)$$

Wick Rotation
avoid pole to make
the calculation simple

(6)

Then

$$\int d^d q f(q^2 - i\epsilon) = i \int d^d \bar{q} f(\bar{q}^2).$$

as long as $f(\bar{q}^2) \rightarrow 0$ faster than $\frac{1}{\bar{q}^2}$ as $\bar{q} \rightarrow \infty$.

Now,

$$\Pi(k^2) = \frac{1}{2} g^2 I(k^2) - Ak^2 - Bm^2 + O(g^4).$$

$$I(k^2) = \int_0^1 dx \int \frac{d^d \bar{q}}{(2\pi)^d} \frac{1}{(\bar{q}^2 + D)^2}.$$

Next are some specific trick ().

$$\Pi''(k^2) = \frac{1}{2} g^2 I''(k^2) + O(g^4)$$

~~skip~~ (this gets rid of A & B terms) (for our expression
for q^3 theory)

$$I''(k^2) = \int_0^1 dx x^2 (1-x)^2 \int \frac{d^d \bar{q}}{(2\pi)^d} \frac{1}{(\bar{q}^2 + D)^4}.$$

We can perform this integral first and then $\int k^2$ twice
(which allows us to match b.c.)

$$\Pi(-m^2) = 0, \quad \Pi'(-m^2) = 0 \quad .$$

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This integral $\int_0^q \frac{1}{(g^2 + D)^4}$

is convergent up to $d < 8$

while the original $\Pi(k^2)$ integration up to $d < 4$

this new range contains $d = 6$.

Around $k^2 = -m^2$

skip

Let's view it again from Taylor expansion point of view.

$$\Pi(k^2) = \left[\frac{1}{2} g^2 I(-m^2) + (A-B)m^2 \right] \quad d \geq 4$$

$$+ \left[\frac{1}{2} g^2 I'(-m^2) - A \right] (k^2 + m^2). \quad d \geq 6$$

$$+ \frac{1}{2!} \left[\frac{1}{2} g^2 I''(-m^2) \right] (k^2 + m^2)^2 + \dots \quad d \geq 8$$

$$+ \tilde{O}(g^4).$$

It's clear that $I^{(n)}(-m^2)$ is divergent for $d \geq 4 + 2n$.

so but A can cancel the divergence at 2nd line from I'

and $A-B$ can cancel the divergence of 1st line from I .

no more free parameters and $\Pi(k^2)$ now is finite for

$$d < 8.$$

For $d > 8$, the procedure breaks down,

the theory is non-renormalizable.

(and we will see this through dim-analysis)

Now it is about How to handle (keep track of) infinities.

Method A. (mostly commonly used in earlier textbooks but also useful for clarifying many physics confusions) (we can have a brief discussion about it in the RGE part).

Pauli-Villars regularization (adding an effective cut-off scale to the theory)

$$\Delta(p^2) \rightarrow \frac{1}{p^2 + m^2 - i\epsilon} - \frac{\Lambda^2}{p^2 + \Lambda^2 - i\epsilon}$$

Λ is the ultra-violet cut off, this will make

$d < 8$ finite as it would be ~~$d \int g^2 \bar{g}^2 \bar{g}^2$~~

$$d \int \frac{1}{(\bar{g}^2 + \lambda)^4} \cdot \cancel{(\bar{g}^2 + \lambda)^4}$$

such a term also ad hoc violates gauge symmetry so one needs to be careful.

Still, we can calculate and then take the $\Lambda \rightarrow \infty$ limit, that makes the $\propto \Lambda^2, \propto \Lambda^4$ parts cancel.

Λ tracks the divergencies.

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Method B : Dimensional Regularization

(Dim - Reg).

Very popular & beautiful, (+ Hoff & Veltman).

One can directly evaluate from a dimension & analytically continue.

The angular part of $d\bar{g}$ over whole range

yields the area (angular volume) of a unit sphere.

$$S_d = \frac{2\pi}{\Gamma(\frac{d+1}{2})} \quad \text{for } d > 1$$

$$\left(\text{one can verify } \int d\bar{g} e^{-\bar{g}^2} \right) \boxed{\Gamma(z) = \int_0^\infty t^{z-1} e^{-t} dt.}$$

$\Gamma(x)$ is the Euler Gamma function.

$$\Gamma(n+1) = n!$$

$$\Gamma(n+\frac{1}{2}) = \frac{(2n)!}{n! 2^{(2n)}} \sqrt{\pi}$$

$$\Gamma(-n+\gamma) = \frac{(-1)^n}{n!} \left[\frac{1}{x} - \gamma + \sum_{k=1}^n k^{-1} + O(x) \right]$$

↑

$$\text{Setting } x=6-6, \text{ we get } \gamma = 0.5772 \dots$$

Euler Mascheroni Constant.

$$\gamma = \lim_{n \rightarrow \infty} \left(\log n + \sum_{k=1}^n \frac{1}{k} \right)$$

$$= \int_1^\infty \left(-\frac{1}{x} + \frac{1}{\lfloor x \rfloor} \right) dx, \quad \lfloor \cdot \rfloor \text{ is the floor function.}$$

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The radial part is

$$\int \frac{d^d \bar{g}}{(2\pi)^d} \frac{(\bar{g})^a}{(\bar{g}^{\frac{1}{2}} + D)^3} = \frac{\Gamma(b-a-\frac{1}{2}d) \Gamma(a+\frac{1}{2}d)}{(4\pi)^{d/2} \Gamma(3) \Gamma(\frac{1}{2}d)} D^{-(b-a-\frac{1}{2}d)}$$

Very useful formula (Eq. 14.27).

For the theory we are evaluating, $a=0, b=2$

(General case) For general d , the coupling g has a mass dimension $\epsilon/2$

$$\epsilon = 6-d.$$

To account for this, we introduce a new parameter μ .

$$g \rightarrow g \mu^{\epsilon/2}$$

μ is not a physical parameter. (hence g remains dimensionless for general theory)

μ is unphysical so no observable can depend on it (later we revisit and derived RGIE).

Setting $d=6-\epsilon$, we get

$$I(k^2) = \frac{\Gamma(-1+\frac{\epsilon}{2})}{4\pi^3} \int_0^1 dx Q \left(\frac{4\pi}{D}\right)^{\epsilon/2}.$$

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With the substitution of $g \rightarrow g e^{-\epsilon/2}$

[we'll see this convenience later]

$$\text{and defining } \alpha = \frac{g^2}{(4\pi)^3}$$

In normal IEP

counting it would be

$\alpha = \frac{g^2}{4\pi}$, and order-by-order

$$\frac{\alpha}{(4\pi)^2}$$

(not the fine-structure constant, which is $\frac{e^2}{4\pi}$)

$$\Pi(k^2) = \frac{1}{2} \alpha P(-1 + \frac{\epsilon}{2}) \int_0^1 dx D \left(\frac{4\pi \tilde{\mu}^2}{D} \right)^{\epsilon/2}$$

$$- A k^2 - B m^2 + O(\alpha^2) \quad \text{dimless.}$$

Taking the $\epsilon \rightarrow 0$ limit.

$$y^{\frac{\epsilon}{2}} = 1 + \frac{\epsilon}{2} \ln \tilde{\mu} + O(\epsilon^2)$$

and $P(-u+x)$

$$\Pi(k^2) = -\frac{1}{2} \alpha \left[\left(\frac{2}{\epsilon} + 1 \right) \left(\frac{1}{6} k^2 + m^2 \right) \right.$$

$$\left. + \int_0^1 dx D \ln \left(\frac{4\pi \tilde{\mu}^2}{e^x D} \right) \right] - A k^2 - B m^2 + O(\alpha^2)$$

(noting that $D = x(1-x)k^2 + m^2$)

$$\int_0^1 dx D = \frac{1}{6} k^2 + m^2$$

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It is convenient to define

$$\mu = \sqrt{4\pi} e^{-\nu/2} \tilde{\mu}$$

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so we have now

$$\begin{aligned}\Pi(k^2) &= \frac{1}{2} \times \int_0^1 dx D \ln(D/m^2) \\ &\quad - \left\{ \frac{1}{6} \alpha \left[\frac{1}{6} + \ln(\mu/m) + \frac{1}{2} \right] + A \right\} k^2 \\ &\quad - \left\{ \alpha \left[\frac{1}{6} + \ln(\mu/m) + \frac{1}{2} \right] + B \right\} m^2 + O(\alpha^2)\end{aligned}$$

If we demand

$$A = -\frac{1}{6} \alpha \left[\frac{1}{6} + \ln(\mu/m) + \frac{1}{2} + K_A \right] + O(\alpha^2)$$

$$B = -\alpha \left[\frac{1}{6} + \ln(\mu/m) + \frac{1}{2} + K_B \right] + O(\alpha^2)$$

$$\begin{aligned}\Pi(k^2) &= \frac{1}{2} \times \int_0^1 dx D \ln(D/m^2) \\ &\quad + \alpha \left(\frac{1}{6} K_A k^2 + K_B m^2 \right) + O(\alpha)\end{aligned}$$

Now $\Pi(k^2)$ is finite! and independent of μ .To fix K_A and K_B , we must still impose the condition:

$$\Pi(-m^2) = 0 \quad \& \quad \Pi'(-m^2) = 0.$$

(13)

Noting that

$$\text{Eq. 14.41} \quad \Pi(k^2) = \frac{1}{2} \times \int_0^1 dx D \ln(D) + \text{linear in } k^2 \text{ and } m^2 + O(\alpha^2)$$

we can impose

$$\Pi(1/k^2) = \frac{1}{2} \times \int_0^1 dx D \ln(D/D_0) + \text{linear in } (k^2 + m^2) + O(\alpha^2)$$

$$\text{here } D_0 \equiv D|_{k^2 = -m^2} = [1 - x(1-x)] m^2$$

We can then differentiate Eq 14.41 w.r.t. k^2

and find ~~$\Pi'(1/m^2)$~~ $\Pi'(1/m^2)$ vanishes for

$$\Pi(k^2) = \underbrace{\frac{1}{2} \times \int_0^1 dx D \ln(D/D_0)}_{\text{this integral can be done "normally"}}, -\frac{1}{2} \times (k^2 + m^2) + O(\alpha^2)$$

and there is a "branch point" at $k^2 = -4m^2$.

The exact propagator can now be written as

$$\tilde{D}(k^2) = \frac{1}{1 - \Pi(k^2)/(k^2 + m^2)} \frac{1}{k^2 + m^2 - i\epsilon}$$

(14)

Let's see the logic here,

we are essentially try to find solutions that matches the boundary conditions.

(Ask for yourself: is the assignment unique?)

In any case, we found

$$\Pi(k^2) = \frac{1}{2} \frac{g^2}{(4\pi)^3} \underbrace{\int_0^1 dx D \ln(D/D_0)}_{\downarrow} - \frac{1}{12} \frac{g^2}{(4\pi)^3} (k^2 + m^2) + O(g^4)$$

this can be done in a closed form.

$$\alpha = \frac{g^2}{(4\pi)^3}$$

$$\Pi(k^2) = \frac{1}{12} \alpha [C_1 k^2 + C_2 m^2 + 2 k^2 f(r)] + O(\alpha^2)$$

$$f(r) = r^3 \tanh^{-1}(1/r), \quad C_1 = 3 - \pi i \sqrt{3}$$

$$r = (1 + 4m^2/k^2)^{1/2}, \quad C_2 = 3 - 2\pi i \sqrt{3}.$$

$$\underbrace{\qquad}_{\downarrow}$$

there is a square root branch point at $k^2 < -4m^2$

Recall that

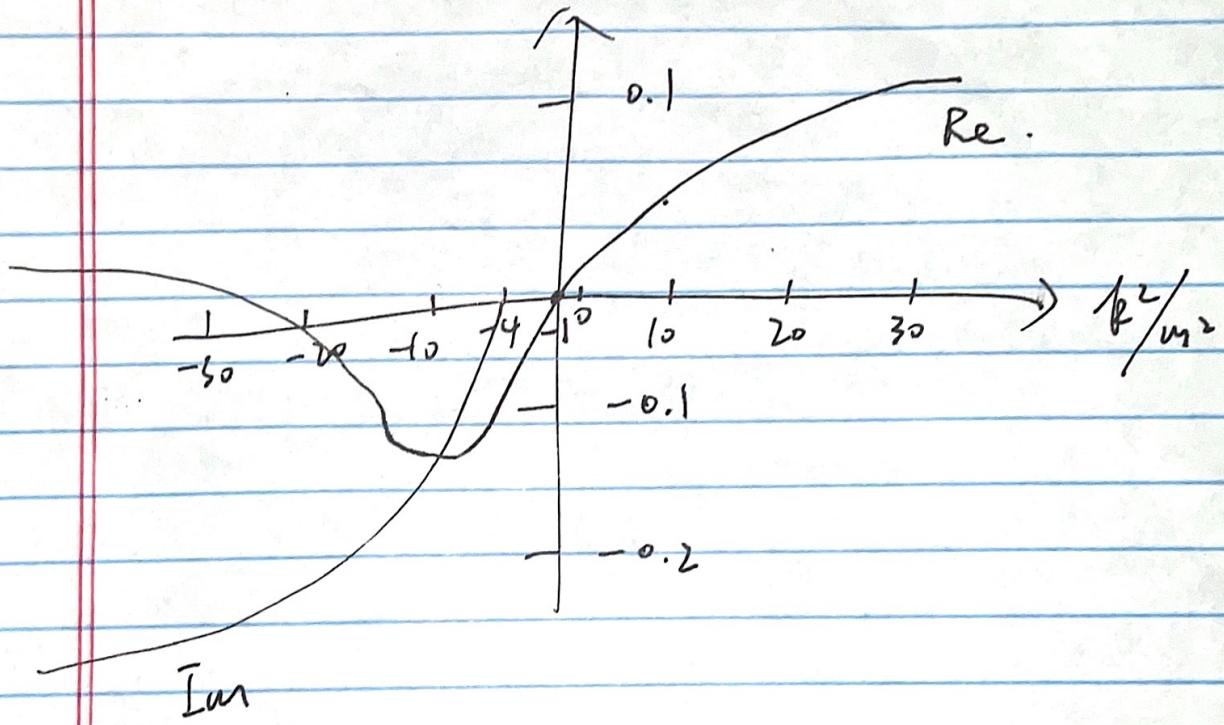
$$\tilde{\Delta}(k^2) = \tilde{\Delta}(k^2) \frac{1}{1 - \Pi(k^2) \tilde{\Delta}(k^2)}$$

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For large values of $|k^2|$

$$\frac{\text{Tr}(k^2)}{k^2 + m^2} \approx \frac{1}{2} \alpha \left[C_1 + \ln \frac{k^2}{m^2} \right]$$

k^2 being negative & $\ln(k^2 - i\epsilon) = \ln|k^2| - i\pi$



Without much extra work, one can make connection to the Lehmann - Källén form.

~~Seeds~~

$$\frac{1}{\text{Tr}P(s)} = \frac{\text{Im}\text{Tr}(-s)}{(-s + m^2 - \text{Re}\text{Tr}(-s))^2 + (\text{Im}\text{Tr}(-s))^2}$$

spectral density, which starts to grow and negative $k^2 \leq -4m^2$