■ EEB

Data analysis of targets in TESS

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- quadruple system with two close eclipsing binaries

1.Observation data ^



Download data

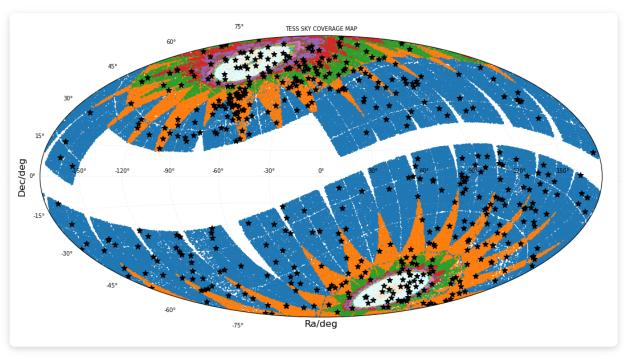


Figure 1. TESS observation targets of 26 sectors in the whole sky area. The black dots represent the large eccentricity binary samples we found.

Code

/home/life/notes/testpy/search.py

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2.Data analysis of eccentric eclipsing binaries in TESS ^

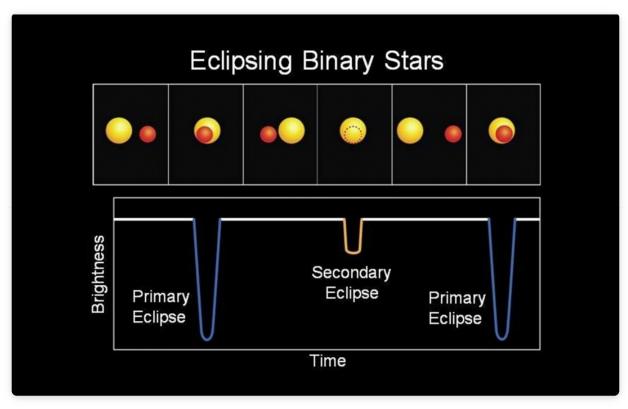


Figure 2. When the smaller star partially blocks the larger star, a primary eclipse occurs, and a secondary eclips occurs when the smaller star is occulted.

The eclipsing binary candidates are first selected by **visually checking** whether more than three eclipses are present in the LightCurve.

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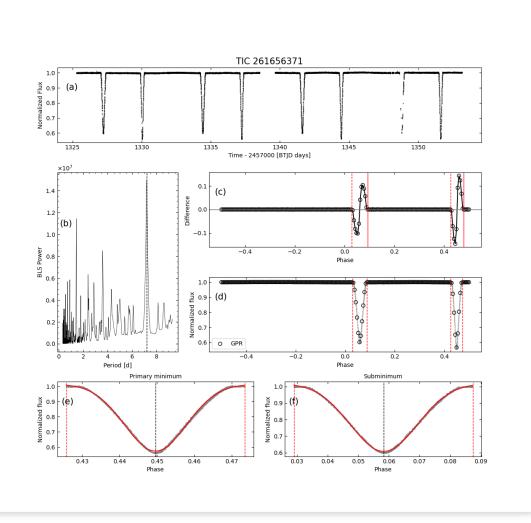
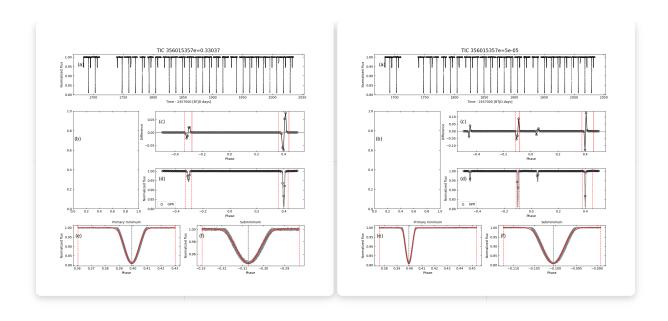


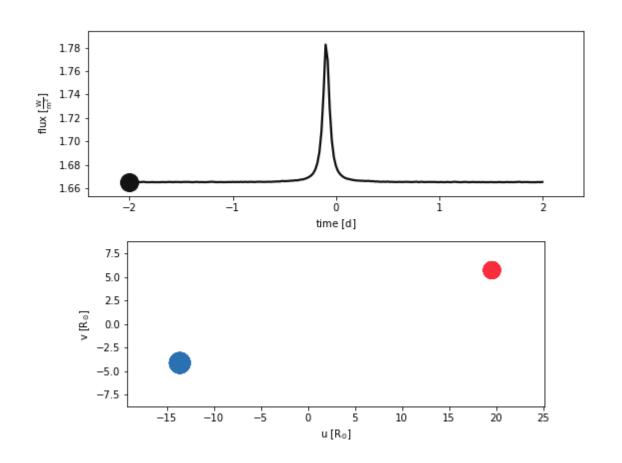
Figure 3. LC analysis and eclipses determination for EEB TIC 261656371 as an example.(a)Light curve.(b)The highest peak of the periodogram is determined as the inital period.(c,d)The original LC is folded and smoothed to obtain the differential folded LC and folded LC.The two eclipses are denoted by red dashed lines.(e,f)Two eclipses(black dotes)are fitted by the Gaussian function(red solid curve) to determine phases(black dashed lines) of the primary and secondary minimum.

code

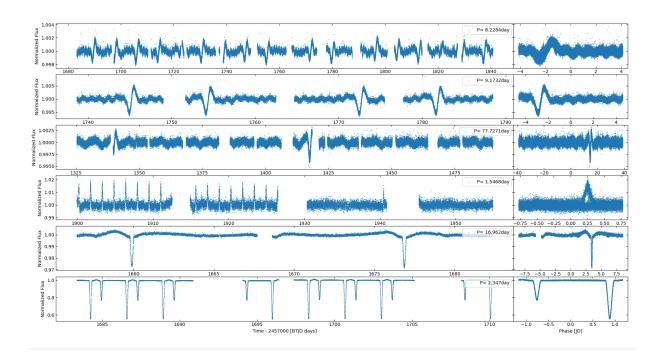
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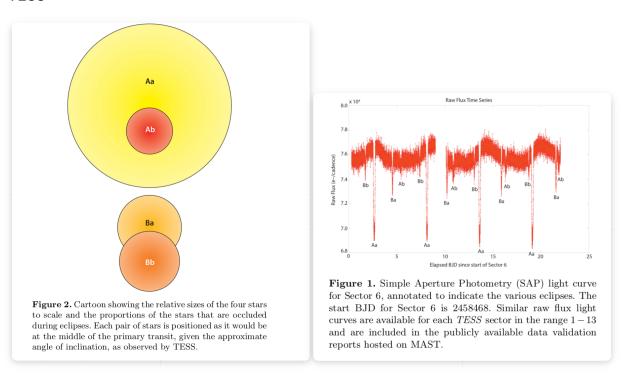
3.Other intersting things in TESS ^



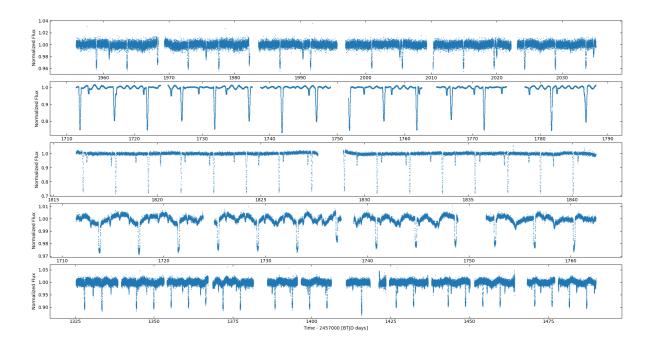
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TIC 278956474:Two close binaries in one young quadruple system, identified by TESS



Guo et al.



The Astrophysical Journal, 888:95 (15pp), 2020 January 10 11.608 11.610 11.610 11.612 11.614 11.616 11.616 11.618 1.0 Orbital Phase 0.6 8.0 1.2 1.4 0.0008
0.0006
0.0004
0.0002
0.0000 0.0008 26.00f_{orb} 29.00f_{orb} 0.0000

2

0

Figure 2. Phase-folded Kepler light curve of KIC 8719324 and its Fourier spectrum. The two dominant TEOs at 26 and 29 times of orbital frequency (f_{orb}) are labeled.

4 6 Frequency (d⁻¹) 8

10

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