

5th exercise

1. treat the whole process as a completely inelastic collision under the gravity of Schwarzschild metric.

the outcome (depend on the mass of monkey and coconut and the initial orbital radius)
maybe i. drop into black hole. ii. moving along other circular geodesic orbit. iii. leave off the black hole.

3.

a. according to the equation: in Schwarzschild metric.

$$\begin{aligned}\left(\frac{dr}{d\tau}\right)^2 + \left(1 - \frac{2GM}{r}\right)\epsilon &= E^2 \\ \Rightarrow \left(\frac{dr}{d\tau}\right)^2 &= E^2 + \left(\frac{2GM}{r} - 1\right)\epsilon\end{aligned}$$

since here inside the blackhole, namely $r < 2GM \Rightarrow \frac{2GM}{r} - 1 > 0$

therefore

$$\left|\frac{dr}{d\tau}\right| = \sqrt{E^2 + \left(\frac{2GM}{r} - 1\right)\epsilon}$$

for timelike particle $\epsilon = 1$ then could see

$$\begin{aligned}\left|\frac{dr}{d\tau}\right| &= \sqrt{E^2 + \left(\frac{2GM}{r} - 1\right)\epsilon} \\ &= \sqrt{E^2 + \left(\frac{2GM}{r} - 1\right)} \geq \sqrt{\frac{2GM}{r} - 1}\end{aligned}$$

b.

$$d\tau \leq \frac{dr}{\sqrt{\frac{2GM}{r} - 1}}$$

to calculate the maximum of τ then

$$\begin{aligned}\Delta\tau &= \int_{2GM}^0 \frac{-dr}{\sqrt{\frac{2GM}{r} - 1}} \\ \frac{u = \sqrt{\frac{2GM}{r} - 1}}{r = \frac{2GM}{u^2 + 1}} &= - \int_0^\infty \frac{2GM}{u} \times \left(-\frac{2u}{(u^2 + 1)^2}\right) du \\ &= \\ &= 4GM \int_0^\infty \frac{du}{(u^2 + 1)^2} \\ &= 4GM \times \frac{\pi}{4} \\ &= \pi GM\end{aligned}$$

c. solar mass: $M = 1.989 \times 10^{35} \text{kg}$, $G = 6.674 \times 10^{-11} \text{m}^3 \text{kg}^{-1} \text{s}^{-2}$

$$\begin{aligned} (\Delta\tau)_s &= \frac{\pi G M}{c^3} = \frac{3.1415 \times 1.989 \times 10^{35} \text{kg} \times 6.674 \times 10^{-11} \text{m}^3 \text{kg}^{-1} \text{s}^{-2}}{(2.998 \times 10^8 \text{m s}^{-1})^3} \\ &= 1.548 \text{s} \end{aligned}$$

b'. below may be not right

$$\begin{aligned} d\tau &= \frac{-dr}{\sqrt{E^2 + \left(\frac{2GM}{r} - 1\right)}} = \frac{-dr}{\sqrt{\frac{2GM}{r} + E^2 - 1}} \\ \Delta\tau &= - \int_{2GM}^0 \frac{dr}{\sqrt{\frac{2GM}{r} + E^2 - 1}} \\ &= \frac{u = \sqrt{\frac{2GM}{r} + E^2 - 1}}{r = \frac{2GM}{u^2 + 1 - E^2}} \int_E^\infty \frac{2GM}{u} \times (-1) \frac{2u du}{(u^2 + 1 - E^2)^2} \\ &= 4GM \int_E^\infty \frac{du}{(u^2 + 1 - E^2)^2} \\ &= \frac{v^2 = u^2 - E^2}{u = \sqrt{v^2 + E^2}} - 4GM \int_0^\infty \frac{1}{(v^2 + 1)^2} \times \frac{v}{\sqrt{v^2 + E^2}} dv \\ &\leq 4GM \int_0^\infty \frac{du}{(u^2 + 1)^2} = 4GM \times \frac{\pi}{4} = \pi GM \end{aligned}$$

4.

the equation :

$$\begin{aligned} R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} + \Lambda g_{\mu\nu} &= 0 \\ g^{\mu\nu} R_{\mu\nu} - \frac{1}{2} R g^{\mu\nu} g_{\mu\nu} + \Lambda g^{\mu\nu} g_{\mu\nu} &= 0 \\ R - 2R + 4\Lambda &= 0 \\ R &= 4\Lambda \end{aligned}$$

then

$$\begin{aligned} R_{\mu\nu} - \frac{1}{2} \times 4\Lambda g_{\mu\nu} + \Lambda g_{\mu\nu} &= 0 \\ R_{\mu\nu} - \Lambda g_{\mu\nu} &= 0 \end{aligned}$$

the general spherically symmetric metric

$$ds^2 = -e^{2\alpha(r,t)} dt^2 + e^{2\beta(t,r)} dr^2 + r^2 (d\theta^2 + \sin^2\theta d\phi^2)$$

using the Ricci tensor in textbook and add a term $-\Lambda g_{\mu\nu}$

$$\begin{aligned} R_{tt} &= (\partial_t^2 \beta + (\partial_t \beta)^2 - \partial_t \alpha \partial_t \beta) + e^{2(\alpha - \beta)} \left(\partial_r^2 \alpha + (\partial_r \alpha)^2 - \partial_r \alpha \partial_r \beta + \frac{2}{r} \partial_r \alpha \right) \\ R_{rr} &= - \left(\partial_r^2 \alpha + (\partial_r \alpha)^2 - \partial_r \alpha \partial_r \beta - \frac{2}{r} \partial_r \beta \right) + e^{2(\beta - \alpha)} (\partial_t^2 \beta + (\partial_t \beta)^2 - \partial_t \alpha \partial_t \beta) \\ R_{tr} &= \frac{2}{r} \partial_t \beta \end{aligned}$$

$$\begin{aligned} R_{\theta\theta} &= e^{-2\beta}(r(\partial_r\beta - \partial_r\alpha) - 1) + 1 \\ R_{\phi\phi} &= R_{\theta\theta}\sin^2\theta \end{aligned}$$

the equation $R_{\mu\nu} - \Lambda g_{\mu\nu} = 0$

$$\begin{aligned} &\left\{ \begin{aligned} (\partial_t^2\beta + (\partial_t\beta)^2 - \partial_t\alpha\partial_t\beta) + e^{2(\alpha-\beta)}(\partial_r^2\alpha + (\partial_r\alpha)^2 - \partial_r\alpha\partial_r\beta + \frac{2}{r}\partial_r\alpha) - \Lambda e^{2\alpha} &= 0 \\ -(\partial_r^2\alpha + (\partial_r\alpha)^2 - \partial_r\alpha\partial_r\beta - \frac{2}{r}\partial_r\beta) + e^{2(\beta-\alpha)}(\partial_t^2\beta + (\partial_t\beta)^2 - \partial_t\alpha\partial_t\beta) + \Lambda e^{2\beta} &= 0 \\ \frac{2}{r}\partial_t\beta &= 0 \\ e^{-2\beta}(r(\partial_r\beta - \partial_r\alpha) - 1) + 1 + \Lambda r^2 &= 0 \\ (e^{-2\beta}(r(\partial_r\beta - \partial_r\alpha) - 1) + 1)\sin^2\theta + \Lambda r^2\sin^2\theta &= 0 \end{aligned} \right. \\ &\text{as } \partial_t\beta = 0 \\ \Rightarrow &\left\{ \begin{aligned} (\partial_r^2\alpha + (\partial_r\alpha)^2 - \partial_r\alpha\partial_r\beta + \frac{2}{r}\partial_r\alpha) - \Lambda e^{2\beta} &= 0 \\ -(\partial_r^2\alpha + (\partial_r\alpha)^2 - \partial_r\alpha\partial_r\beta - \frac{2}{r}\partial_r\beta) + \Lambda e^{2\beta} &= 0 \\ e^{-2\beta}(r(\partial_r\beta - \partial_r\alpha) - 1) + 1 + \Lambda r^2 &= 0 \end{aligned} \right. \end{aligned}$$

$\Rightarrow -e^{-2\beta}r\partial_r\partial_t\alpha = 0 \Rightarrow$ set $\alpha(t, r) = \alpha_0(t) + \alpha_1(r)$, and could see the equation below have no relation with $\alpha_0(t)$, that's to say for arbitrary $\alpha_0(t)$ the equations will all satisfy.

therefore set $\alpha_0 = 0$ then $\alpha = \alpha_1$

$$\Rightarrow \frac{4}{r}\partial_r(\alpha + \beta) = 0 \Rightarrow \alpha + \beta = c(t) \text{ could set } c(t) = 0 \text{ here, therefore } \alpha = -\beta$$

$$\begin{aligned} \Rightarrow &\left\{ \begin{aligned} (\partial_r^2\alpha + (\partial_r\alpha)^2 - \partial_r\alpha(-\partial_r\alpha) + \frac{2}{r}\partial_r\alpha) - \Lambda e^{-2\alpha} &= 0 \\ -(\partial_r^2\alpha + (\partial_r\alpha)^2 - \partial_r\alpha(-\partial_r\alpha) + \frac{2}{r}\partial_r\alpha) + \Lambda e^{-2\alpha} &= 0 \\ e^{2\alpha}(r(-\partial_r\alpha - \partial_r\alpha) - 1) + 1 + \Lambda r^2 &= 0 \end{aligned} \right. \\ \Rightarrow &\left\{ \begin{aligned} (\partial_r^2\alpha + 2(\partial_r\alpha)^2 + \frac{2}{r}\partial_r\alpha) - \Lambda e^{-2\alpha} &= 0 \\ e^{2\alpha}(-2r\partial_r\alpha - 1) + 1 + \Lambda r^2 &= 0 \end{aligned} \right. \\ \Rightarrow &\left\{ \begin{aligned} (\partial_r^2\alpha + 2(\partial_r\alpha)^2 + \frac{2}{r}\partial_r\alpha) - \Lambda e^{-2\alpha} &= 0 \\ \partial_r\alpha + \frac{1}{2r} + \frac{1}{2r}(1 + \Lambda r^2)e^{-2\alpha} &= 0 \end{aligned} \right. \end{aligned}$$

$$\begin{aligned} (2r\partial_r\alpha + 1)e^{2\alpha} &= 1 + \Lambda r^2 \\ \partial_r(re^{2\alpha}) &= 1 + \Lambda r^2 \\ \frac{d(re^{2\alpha})}{dr} &= 1 + \Lambda r^2 \\ re^{2\alpha} &= \int dr\{1 + \Lambda r^2\} \\ &= r + \frac{\Lambda}{3}r^3 + C \\ e^{2\alpha} &= 1 + \frac{\Lambda}{3}r^2 + \frac{C}{r} \end{aligned}$$

since $\Lambda \rightarrow 0$ the result will be the same as Schwarzschild solution. therefore $C = -2GM$

the metric :

$$ds^2 = -\left(1 - \frac{2GM}{r} + \frac{\Lambda}{3}r^2\right)dt^2 + \left(1 - \frac{2GM}{r} + \frac{\Lambda}{3}r^2\right)^{-1}dr + r^2d\Omega^2$$

b.

here choose $\theta = \frac{\pi}{2}$, too

$$\begin{aligned}
\epsilon &= -g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \\
&= \left(1 - \frac{2GM}{r} + \frac{\Lambda}{3}r^2\right) \left(\frac{dt}{d\tau}\right)^2 - \left(1 - \frac{2GM}{r} + \frac{\Lambda}{3}r^2\right)^{-1} \left(\frac{dr}{d\tau}\right)^2 - r^2 \left(\frac{d\phi}{d\tau}\right)^2 \\
&\Rightarrow \\
\left(\frac{dr}{d\tau}\right)^2 + \left(1 - \frac{2GM}{r} + \frac{\Lambda}{3}r^2\right) \left(r^2 \left(\frac{d\phi}{d\tau}\right)^2 + \epsilon\right) &= \left(\left(1 - \frac{2GM}{r} + \frac{\Lambda}{3}r^2\right) \frac{dt}{d\tau}\right)^2 \\
\left(\frac{dr}{d\tau}\right)^2 + \left(1 - \frac{2GM}{r} + \frac{\Lambda}{3}r^2\right) \left(\frac{L^2}{r^2} + \epsilon\right) &= E^2
\end{aligned}$$

for massive particle $\epsilon = 1$, $V(r) = \frac{1}{2} \left(1 - \frac{2GM}{r} + \frac{\Lambda}{3}r^2\right) \left(\frac{L^2}{r^2} + 1\right)$

therefore the radical geodesic equation of motion:

$$\left(\frac{dr}{d\tau}\right)^2 + 2V(r) = E^2$$

5.

as $g_{\mu\nu} U^\mu U^\nu = -1$, $U_\mu = (U_0, 0, 0, 0)$

$$\begin{aligned}
g^{\mu\nu} U_\mu U_\nu &= -\left(1 - \frac{2GM}{r_*}\right)^{-1} U_0^2 = -1 \\
\Rightarrow U_0^2 &= \left(1 - \frac{2GM}{r_*}\right) \\
\Rightarrow U_0 &= -\sqrt{1 - \frac{2GM}{r_*}}
\end{aligned}$$

a. bacon move along the geodesic, and angular momentum $L = 0$

$$\begin{aligned}
\left\{ \begin{array}{l} \left(1 - \frac{2GM}{r}\right) \frac{dt}{d\tau} = E \\ \left(\frac{dr}{d\tau}\right)^2 + \left(1 - \frac{2GM}{r}\right) = E^2 \end{array} \right. &\Rightarrow \frac{\left(\frac{dr}{d\tau}\right)^2}{\left(\frac{dt}{d\tau}\right)^2} = \frac{E^2 - \left(1 - \frac{2GM}{r}\right)}{\frac{E^2}{\left(1 - \frac{2GM}{r}\right)^2}} \\
&\Rightarrow \left(\frac{dr}{dt}\right)^2 = \left(1 - \frac{1}{E^2} \left(1 - \frac{2GM}{r}\right)\right) \left(1 - \frac{2GM}{r}\right)^2 \\
&\Rightarrow \left|\frac{dr}{dt}\right| = \sqrt{\left(1 - \frac{1}{E^2} \left(1 - \frac{2GM}{r}\right)\right) \left(1 - \frac{2GM}{r}\right)^2}
\end{aligned}$$

b.

$$\begin{aligned}\left(\frac{dr}{d\tau}\right)^2 &= E^2 - \left(1 - \frac{2GM}{r}\right) \\ \left|\frac{dr}{d\tau}\right| &= \sqrt{E^2 - \left(1 - \frac{2GM}{r}\right)}\end{aligned}$$

the comoving coordinate observe speed

$$\frac{dr_*}{dt} = \left|\frac{dr}{d\tau}\right| = \sqrt{E^2 - \left(1 - \frac{2GM}{r}\right)}$$

when $r = 2GM$, the $\left|\frac{dr}{d\tau}\right| = E$

c.

bacon emit a photo ,moving along the geodesic.

observers

$$\begin{aligned}p^\mu U_\mu &= -E_{\text{photo}} \\ \frac{dt}{d\tau} U_t &= -\frac{2\pi}{\lambda} \\ \frac{\frac{2\pi}{\lambda_{\text{em}}}}{\left(1 - \frac{2GM}{r_{\text{em}}}\right)} \times \left(-\sqrt{\left(1 - \frac{2GM}{r_*}\right)}\right) &= -\frac{2\pi}{\lambda_{\text{ob}}} \\ \lambda_{\text{ob}} &= \lambda_{\text{em}} \left(\frac{1 - \frac{2GM}{r_{\text{em}}}}{\sqrt{1 - \frac{2GM}{r_*}}} \right)\end{aligned}$$

d. for null curve

$$\begin{aligned}\begin{cases} \left(1 - \frac{2GM}{r}\right) \frac{dt}{d\tau} = E \\ \left(\frac{dr}{d\tau}\right)^2 = E^2 \end{cases} &\Rightarrow \frac{\left(1 - \frac{2GM}{r}\right)^2 \left(\frac{dt}{d\tau}\right)^2}{\left(\frac{dr}{d\tau}\right)^2} = 1 \\ &\Rightarrow \left(\frac{dt}{dr}\right)^2 = \left(\frac{1}{1 - \frac{2GM}{r}}\right)^2\end{aligned}$$

as $r \rightarrow r_*$ and $r_* > r$ therefore:

$$\begin{aligned}\frac{dt}{dr} &= \frac{1}{1 - \frac{2GM}{r}} \\ dt &= \frac{1}{1 - \frac{2GM}{r}} dr \\ \int_0^{t_{\text{ob}}} dt &= \int_{r_{\text{em}}}^{r_*} \frac{dr}{1 - \frac{2GM}{r}} \\ t_{\text{ob}} &= \int_{r_{\text{em}}}^{r_*} \frac{r dr}{r - 2GM} \\ &= \int_{r_{\text{em}}}^{r_*} \frac{(r - 2GM) + 2GM}{r - 2GM} dr \\ &= \int_{r_{\text{em}}}^{r_*} 1 + \frac{2GM}{r - 2GM} dr \\ t_{\text{ob}} &= r_* - r_{\text{em}} + 2GM \ln\left(\frac{r_* - 2GM}{r_{\text{em}} - 2GM}\right)\end{aligned}$$

e. using the result (d)

$$\begin{aligned}
e^{\frac{t_{\text{ob}} - r_* + r_{\text{em}}}{2GM}} &= \frac{r_* - 2GM}{r_{\text{em}} - 2GM} \\
\frac{\lambda_{\text{ob}}}{\lambda_{\text{em}}} &= \frac{1 - \frac{2GM}{r}}{\sqrt{1 - \frac{2GM}{r_*}}} = \frac{1 - \frac{2GM}{r}}{1 - \frac{2GM}{r_*}} \times \sqrt{1 - \frac{2GM}{r_*}} \\
&= \frac{r - 2GM}{r_* - 2GM} \times \frac{r_*}{r} \sqrt{1 - \frac{GM}{r_*}} \\
&= e^{-\frac{t_{\text{ob}}}{2GM}} \times \frac{e^{\frac{r - r_*}{2GM}}}{r} \times r_* \sqrt{1 - \frac{GM}{r_*}} \\
&\stackrel{r \rightarrow 2GM}{=} e^{-\frac{t_{\text{ob}}}{2GM}} \frac{e^{\frac{r_*}{2GM} + 1}}{2GM} r_* \sqrt{1 - \frac{GM}{r_*}}
\end{aligned}$$

when $r \rightarrow 2GM \Rightarrow \frac{\lambda_{\text{ob}}}{\lambda_{\text{em}}} \rightarrow 0$, could see that require $t_{\text{ob}} \rightarrow \infty$ consisting with imagination, therefore the approximation is not bad.

$$\begin{aligned}
\frac{\lambda_{\text{ob}}}{\lambda_{\text{em}}} &\propto e^{-\frac{t_{\text{ob}}}{2GM}} \\
&\Rightarrow T = -2GM
\end{aligned}$$