# THE DIFFERENTIAL GEOMETRY DIFFERENTIAL TOPOLOGY DUMP

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37. Lecture 22: Black Holes

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ABSTRACT. Everything about Differential Geometry, Differential Topology

### Part 1. Combinatorics, Probability Theory

**Theorem 1** (4.2. of Feller (1968) [1]). Let  $r_1, \ldots r_k \in \mathbb{Z}$ , s.t.  $r_1 + r_2 + \cdots + r_k = n$ ;,  $r_i \geq 0$ . Let

$$\frac{N!}{r_1!r_2!\dots r_k!} =$$

number of ways in which n elemnts can be divided into k ordered parts (partitioned into k subpopulations). cf. Eq. (4.7) of are to be occupied by bars. Thus  $\binom{r-1}{n-1}$  choices. Feller (1968) [1].

Note that the order of the subpopulations is essential in the sense that  $(r_1 = 2, r_2 = 3)$  and  $(r_1 = 3, r_2 = 2)$  represent different partitions. However, no attention is paid to the order within the groups.

Proof.

(2) 
$$\binom{n}{r_1} \binom{n-r_1}{r_2} \binom{n-r_1-r_2}{r_3} \dots \binom{n-r_1-\dots-r_{k-2}}{r_{k-1}} = \frac{n!}{r_1!r_2!\dots r_k!}$$

i.e. in order to effect the desired partition, we have to select  $r_1$  elements out of n, remaining  $n-r_1$  elements select a second group of size  $r_2$ , etc. After forming the (k-1)st group there remains  $n-r_1-r_2-\cdots-r_{k-1}=r_k$  elements, and these form the last group.

cf. pp. 37 of Feller (1968) [1] Examples. (g) Bridge. 32 cards are partitioned into 4 equal groups  $\rightarrow 52!/(13!)^4$ . Probability each player has an ace (?).

The 4 aces can be ordered in 4! = 24 ways, each order presents 1 possibility of giving 1 ace to each player. Remaining 48 cards distributed  $(48!)/(12!)^4$  ways.

$$\rightarrow p = 24 \frac{48!}{(12!)^4} / \frac{52!}{(13!)^4}$$

(h) A throw of 12 dice  $\rightarrow$  6<sup>12</sup> different outcomes total. Event each face appears twice can occur in as many ways as 12 dice can be arranged in 6 groups of 2 each.

$$\frac{12!}{(2!)^6} / \frac{52!}{(13!)^4}$$

0.0.1. Application to Occupancy Problems; binomial coefficients. cf. Sec. 5 Application to Occupancy Problems of Feller (1968) [1].

Consider randomly placing r balls into n cells.

Let  $r_k =$  occupancy number = number of balls in kth cell.

Every *n*-tuple of integers satisfying  $r_1 + r_2 + \cdots + r_n = r$ ;  $r_k \ge 0$ . describes a possible configuration of occupancy numbers. With indistinguishable balls 2 distributions are distinguishable only if the corresponding *n*-tuples  $(r_1, \dots r_n)$  are not identical.

(i) number of distinguishable distributions is

(3) 
$$A_{r,n} = \binom{n+r-1}{r} = \binom{n+r-1}{n-1}$$

cf. Eq. (5.2) of Feller (1968) [1]

(ii) number of distinguishable distributions in which no cell remains empty is  $\binom{r-1}{1}$ .

72 Proof. Represent balls by stars, indicate n cells by n spaces between n+1 bars. e.g. r=8 balls

n=6 cells

Such a symbol necessarily starts and ends with a bar, but remaining n-1 bars and r starts appear in an arbitrary order. In this way, it becomes apparent that the number of distinguishable distributions equals the number of ways of selecting.

r places out of 
$$n+r-1$$
,  $\frac{(n+r-1)!}{(n-1)!r!} = \binom{n-1+r}{r}$ 

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(4)

$$|||| \dots ||$$
  $n+1$  bars  
 $*** \dots **$   $r$  stars leave  $r-1$  spaces

Condition that no cell be empty imposes the restriction that no 2 bars be adjacent. r stars leave r-1 spaces of which n-1 re to be occupied by bars. Thus  $\binom{r-1}{n-1}$  choices.

Probability to obtain given occupancy numbers  $r_1, \dots r_n = \frac{r!}{r_1! r_2! \dots r_n!} / n^r$ , with r balls given by Thm. 4.2. of Feller (1968)

[1], which is the Maxwell-Boltzmann distribution.

(a) Bose-Einstein and Fermi-Dirac statistics. Consider r indistinguishable particles, n cells, each particle assigned to 1 cell. State of the system - random distribution of r particles in n cells.

If n cells distinguishable,  $n^r$  arrangements equiprobable  $\rightarrow$  Maxwell-Boltzmann statistics.

Bose-Einstein statistics: only distinguishable arrangements are considered, and each assigned probability  $\frac{1}{A_{n,n}}$ 

$$A_{r,n} = \binom{n+r-1}{r} = \binom{n-1+r}{n-1}$$

cf. Eq. 5.2 of Feller (1968) [1]

Fermi-Dirac statistics.

- (1) impossible for 2 or more particles to be in the same cell.  $\rightarrow r < n$ .
- (2) all distinguishable arrangements satisfying the first condition have equal probabilities.  $\rightarrow$  an arrangement is completely described by stating which of the n cells contain a particle r particles  $\rightarrow \binom{n}{r}$  ways r cells chosen.

Fermi-Dirac statistics, there are  $\binom{n}{r}$  possible arrangements, prob.  $1/\binom{n}{r}$ .

pp. 39. Feller (1968) [1]. Consider cells themselves indistinguishable! Disregard order among occupancy numbers. cf. Feller (1968) [1]

### Part 2. Linear Algebra Review

cf. Change of Basis, of Appendix B of John Lee (2012) [3]

**Exercise B.22.** Suppose V, W, X finite-dim. vector spaces

- $S:V \to W, \ T:W \to X$ 
  - (a) rankS < dimV with rankS = dimV iff S injective
  - (b)  $\operatorname{rank} S \leq \operatorname{dim} W$  with  $\operatorname{rank} S = \operatorname{dim} W$  iff S surjective
  - (c) if  $\dim V = \dim W$  and S either injective or surjective, then S isomorphism
  - (d)  $\operatorname{rank} TS < \operatorname{rank} S$   $\operatorname{rank} TS = \operatorname{rank} S$  iff  $\operatorname{im} S \cap \ker T = 0$
  - (e)  $\operatorname{rank} TS \leq \operatorname{rank} T$   $\operatorname{rank} TS = \operatorname{rank} T$  iff  $\operatorname{im} S + \ker T = W$
  - (f) if S isomorphism, then rankTS = rankT
  - (g) if T isomorphism, then rankTS = rankS

EY: Exercise B.22(d) is useful for showing the chart and atlas of a Grassmannian manifold, found in the More examples, for smooth manifolds.

#### Proof. (a) Recall the rank-nullity theorem:

**Theorem 2** (Rank-Nullity Theorem).

(5) 
$$dim(im(S)) + dim(ker(S)) = dimV$$

Now

$$\operatorname{rank}(S) + \dim(\ker(S)) \equiv \dim(\operatorname{im}(S)) + \dim(\ker(S)) = \dim V$$

$$\Longrightarrow \boxed{\operatorname{rank}(S) \le \dim V}$$

If rank(S) = dim V,

then by rank-nullity theorem,  $\dim(\ker(S)) = 0$ , implying that  $\ker S = \{0\}$ 

Suppose  $v_1, v_2 \in V$  and that  $S(v_1) = S(v_2)$ . By linearity of S,  $S(v_1) - S(v_2) = S(v_1 - v_2) = 0$ , which implies, since **Definition 1** (upper bound). Schuller, original, version 1:  $\ker S = \{0\}, \text{ that } v_1 - v_2 = 0.$ 

 $\implies v_1 = v_2$ . Then by definition of injectivity, S injective.

If S injective, then S(v) = 0 implies v = 0. Then  $\ker S = \{0\}$ . Then by rank-nullity theorem,  $\operatorname{rank}(S) = \dim V$ .

(b)  $\forall w \in \text{im}(S), w \in W$ . Clearly rank S < dimW.

If S surjective, im(S) = W. Then dim(im(S)) = rankS = dim W.

If  $\operatorname{rank} S = \dim W = m$ , then  $\operatorname{im}(S)$  has basis  $\{y_i\}_{i=1}^m$ ,  $y_i \in \operatorname{im}(S)$ , so  $\exists x_i \in V, i = 1 \dots m$  s.t.  $S(x_i) = y_i$ , with  $\{S(x_i)\}_{i=1}^m$  linearly independent.

Since  $\{S(x_i)\}_{i=1}^m$  linearly independent and  $\dim W = m$ ,  $\{S(x_i)\}_{i=1}^m$  basis for W.  $\forall w \in W, w = \sum_{i=1}^m w^i S(x_i) = S(\sum_{i=1}^m w^i x_i)$ .  $\sum_{i=1}^m w^i x_i \in V$ . S surjective.

(d) Now

$$\dim V = \operatorname{rank} TS + \operatorname{nullity} TS$$

$$\dim V = \operatorname{rank} S + \operatorname{nullity} S$$

 $\ker S \subseteq \ker TS$ , clearly, so nullity  $S \leq \operatorname{nullity} TS$ 

$$\Longrightarrow \boxed{\operatorname{rank} TS \leq \operatorname{rank} S}$$

If rankTS = rankS,

then nullity S = nullity TS

Suppose  $w \in \text{Im} S \cap \text{ker} T$ ,  $w \neq 0$ 

Then  $\exists v \in S$ , s.t. w = S(v) and T(w) = 0

Then T(w) = TS(v) = 0. So  $v \in \text{ker}TS$ 

 $v \notin \ker S \text{ since } w = S(v) \neq 0$ 

This implies nullity TS > nullity S. Contradiction.

$$\Longrightarrow \operatorname{Im} S \cap \ker T = 0$$

If  $\text{Im}S \cap \text{ker}T = 0$ ,

Consider  $v \in \ker TS$ . Then TS(v) = 0.

. Then  $S(v) \in \ker T$ 

S(v) = 0; otherwise,  $S(v) \in \text{Im}S$ , contradicting given  $\text{Im}S \cap \text{ker}T = 0$  $v \in \ker S$ 

 $\ker TS \subseteq \ker S$ 

 $\Longrightarrow \ker TS = \ker S$ 

So nullityTS = nullityS

 $\implies \operatorname{rank} TS = \operatorname{rank} S$ 

- (f)
- (g)

#### Part 3. Manifolds

Schuller (2015) [28]

- (1) (strikethrough)
- (3) transitivity
- (4) totality:  $a \le b$  or  $b \le a$ ,  $\forall a, b$

 $u \in P$  upper bound to a subset  $T \subseteq P$  if

 $\forall t \in T : t \le u \ (EY: in \ t \le u, \le is \ for \ partial \ ordered \ set)$ 

version 2, possible alternative (bottomline: not right): (of is it  $u \in P$  upper bound to a subset  $T \subseteq P$  if  $\exists t \in T$  s.t. u < t)

#### 1. Inverse Function Theorem

Shastri (2011) had a thorough and lucid and explicit explanation of the Inverse Function Theorem [5]. I will recap it here. The following is also a blend of Wienhard's Handout 4 https://web.math.princeton.edu/~wienhard/teaching/M327/handout4.

**Definition 2.** Let (X, a) metric space.

contraction 
$$\phi: X \to X$$
 if  $\exists$  constant  $0 < c < 1$  s.t.  $\forall x, y \in X$ 

$$d(\phi(x), \phi(y)) \le cd(x, y)$$

**Theorem 3** (Contraction Mapping Principle). Let (X, d) complete metric space.

Then  $\forall$  contraction  $\phi: X \to X$ ,  $\exists ! y \in X$  s.t.  $\phi(y) = y$ , y fixed pt.

*Proof.* Recall def. of complete metric space X, X metric space s.t.  $\forall$  Cauchy sequence in X is convergent in X (i.e. has limit in X).

$$x_1 = \phi(x_0)$$

$$x_2 = \phi(x_1)$$

 $\forall x_0 \in X$ , Define :

$$x_j = \phi(x_{j-1})$$

$$x_n = \phi(x_{n-1})$$

$$d(x_{n+1}, x_n) = d(\phi(x_n), \phi(x_{n-1})) \le cd(x_n, x_{n-1}) \le \dots \le c^n d(x_1, x_0)$$

for some 0 < c < 1.

$$d(x_m, x_n) \le d(x_n, x_{n-1}) + d(x_{n-1}, x_m) \le d(x_n, x_{n-1}) + d(x_{n-1}, x_{n-2}) + \dots + d(x_{m+1}, x_m) \le \sum_{k=n-1}^m c^k d(x_1, x_0)$$

Thus,  $\forall \epsilon > 0$ ,  $\exists n_0 > 0$ ,  $(n_0 \text{ large enough}) \text{ s.t. } \forall m, n \in \mathbb{N} \text{ s.t. } n_0 < n < m$ ,

$$d(x_m, x_n) \le \sum_{k=n-1}^m c^k d(x_1, x_0) < \epsilon d(x_1, x_0)$$

Thus,  $\{x_n\}$  Cauchy sequence. Since X complete,  $\exists$  limit pt.  $y \in X$  of  $\{x_n\}$ .

$$\phi(y) = \phi(\lim_n x_n) = \lim_n \phi(x_n) = \lim_n x_{n+1} = y$$

Since by def. of y limit pt. of  $\{x_n\}, \forall \epsilon > 0$ , then  $\{n | |x_n - y| \le \epsilon, n \in \mathbb{N}\}$  is infinite.

Consider  $\delta > \mathbb{N}$ . Consider  $\{n||x_n - y| \leq \delta, n \in \mathbb{N}\}$ 

 $\exists N_{\delta} \in \mathbb{N} \text{ s.t. } \forall n > N_{\delta}, |x_n - y| < \delta; \text{ otherwise, } \forall N_{\delta}, \exists n > N_{\delta} \text{ s.t. } |x_n - y| \geq \delta. \text{ Then } \{n||x_n - y| \leq \delta, n \in \mathbb{N}\} \text{ finite.}$  Contradiction.

 $\phi$  cont. so by def.  $\forall \epsilon > 0, \exists \delta > 0$  s.t. if  $|x_n - y| < \delta$ , then  $|\phi(x_n) - \phi(y)| < \epsilon$ .

Pick  $N_{\delta}$  s.t.  $\forall n > N_{\delta}$ ,  $|x_n - y| < \delta$ , and so  $|\phi(x_n) - \phi(y)| < \epsilon$ . There are infinitely many  $\phi(x_n)$ 's that satisfy this, and so  $\phi(y)$  is a limit pt.

If  $\exists y_1, y_2 \in X$  s.t.  $\phi(y_1) = y_1$ , then

$$\phi(y_2) = y_2$$

$$d(y_1, y_2) = d(\phi(y_1), \phi(y_2)) \le cd(y_1, y_2)$$
 with  $c < 1$ 

so c = 1

**Theorem 4** (Inverse Function Theorem). Suppose open  $U \subset \mathbb{R}^n$ , let  $C^1 f: U \to \mathbb{R}^n$ ,  $x_0 \in U$  s.t.  $Df(x_0)$  invertible. Then  $\exists$  open neighborhoods  $V \ni x_0, W \ni f(x_0)$  s.t.  $V \subseteq U$  and  $W \subseteq \mathbb{R}^n$ , respectively, and s.t.

- (i)  $f: V \to W$  bijection
- (ii)  $q = f^{-1}: V \to U$  differentiable, i.e.  $q = f^{-1}: W \to V$  is  $C^1$
- (iii)  $D(f^{-1})$  cont. on W.
- (iv)  $Dg(y) = (Df(g(y)))^{-1} \quad \forall y \in W$

Also, notice that  $f(g(y)) = y \forall y \in W$ .

*Proof.* Consider  $\widetilde{f}(x) = (Df(x_0))^{-1}(f(x+x_0) - f(x_0))$ . Then

$$\widetilde{f}(0) = 0$$
 and

$$D\widetilde{f} = (Df(x_0))^{-1}(Df(x+x_0) - 0)$$

$$D\widetilde{f}(0) = (Df(x_0))^{-1}Df(x_0) = 1$$

So let  $\widetilde{f} \to f$  (notation) and so assume, without loss of generality, that  $U \ni 0$ , f(0) = 0, Df(0) = 1

Choose  $0 < \epsilon \le \frac{1}{2}$ . Let  $0 < \delta < 1$  s.t. open ball  $V = B_{\delta}(0) \subseteq U$ , and  $||Df(x) - 1|| < \epsilon$ .  $\forall x \in U$ , since Df cont. at 0. Let W = f(V).

 $\forall y \in W$ , define  $\phi_y : V \to \mathbb{R}^n$ 

$$\phi_y(x) = x + (y - f(x))$$

$$D(\phi_y)(x) = 1 + -Df(x) \quad \forall x \in V$$

$$||D(\phi_u)(x)|| = ||1 - Df(x)|| \le \epsilon < 1$$

 $\forall x_1, x_2 \in V$ , by mean value Thm. (not the equality that is only valid in 1-dim., but the inequality, that's valid for  $\mathbb{R}^d$ ,

$$\|\phi_y(x_1) - \phi_y(x_2)\| \le \|D(\phi_y)(x')\| \|x_1 - x_2\|$$

for some  $x' = cx_2 + (1 - c)x_1$ ,  $c \in [0, 1]$ . V only needed to be convex set.

$$\Longrightarrow \|\phi_y(x_1) - \phi_y(x_2)\| \le \epsilon \|x_1 - x_2\|$$

Then  $\phi_u$  contraction mapping.

Suppose  $f(x_1) = f(x_2) = y, x_1, x_2 \in V$ .

$$\phi_y(x_1) = x_1$$
 
$$\phi_y(x_2) = x_2$$
 
$$\|\phi_y(x_1) - \phi_y(x_2)\| = \|x_1 - x_2\| \le \epsilon \|x_1 - x_2\| \quad \forall \epsilon > 0 \Longrightarrow x_1 = x_2$$

 $\implies f|_{U}$  injective.

W = f(V), so  $f: V \to W$  surjective. f bijective.

Fix  $y_0 \in W$ ,  $y_0 = f(x_0)$ ,  $x_0 \in V$ .

Let r > 0 s.t.  $B_r(x_0) \subset V$ .

Consider  $B_{r\epsilon}(y_0)$ . If  $y \in B_{r\epsilon}(y_0)$ .

$$r\epsilon > ||y - y_0|| = ||y - f(x_0)|| = ||\phi_y(x_0) - x_0||$$
 with  $\phi_y(x) = x + (y - f(x))$ 

If  $x \in B_r(x_0)$ ,

$$\|\phi_{u}(x) - x_{0}\| \le \|\phi_{u}(x) - \phi_{u}(x_{0})\| + \|\phi_{u}(x_{0}) - x_{0}\| \le \epsilon \|x - x_{0}\| + r\epsilon < 2r\epsilon = r$$

Thus  $\phi(B_r(x_0)) = B_r(x_0)$ .

By contraction mapping principle,  $\exists a \in B_r(x_0)$ , s.t.  $\phi_y(a) = a$ . Then  $\phi_y(a) = a + (y - f(a)) = a \Longrightarrow f(a) = y$ .  $y \in f(V) = W$ .

So  $B_{r\epsilon}(y_0) \subset W$ . W open.

Let  $Mat(n, n) \equiv \text{space of all } n \times n \text{ matrices}; Mat(n, n) = \mathbb{R}^{n^2}$ .

There is a proof of the implicit function theorem and its various forms in Shastri (2011) [5], but I found Wienhard's Handout 4 for Math 327 to be clearer. 

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**Theorem 5** (Implicit Function Theorem). Let open  $U \subset \mathbb{R}^{m+n} \equiv \mathbb{R}^m \times \mathbb{R}^n$ 

$$C^1 f: U \to \mathbb{R}^n$$

 $(a,b) \in U$  s.t. f(a,b) = 0 and  $D_y f|_{(a,b)}$  invertible.

Then  $\exists$  open  $V \ni (a,b), V \subset U$ 

 $\exists open \ neighborhood \ W \ni a, \ W \subseteq \mathbb{R}^m$ 

 $\exists ! \quad C^1 \ q : W \to \mathbb{R}^n \ s.t.$ 

$$\{(x,y) \in V | f(x,y) = 0\} = \{(x,g(x)) | x \in W\}$$

Moreover.

$$dg_x = - (d_y f)^{-1}|_{(x,g(x))} d_x f|_{(x,g(x))}$$

and q smooth if f.

*Proof.* Define  $F: U \to \mathbb{R}^{m+n}$ 

$$F(x,y) = (x, f(x,y))$$

Then F(a,b) = (a,0) (given), and

$$DF = \begin{bmatrix} 1 \\ \frac{\partial f^{i}(x,y)}{\partial x^{j}} & \frac{\partial f^{i}(x,y)}{\partial y^{j}} \end{bmatrix} \equiv \begin{bmatrix} 1 \\ D_{x}f & D_{y}f \end{bmatrix}$$

DF(a,b) invertible.

<sup>&</sup>lt;sup>1</sup>https://web.math.princeton.edu/~wienhard/teaching/M327/handout4.pdf

By inverse function theorem, since DF(a,b) invertible at pt. (a,b),

 $\exists$  open neighborhoods  $V \ni (a,b) \subseteq \mathbb{R}^m \times \mathbb{R}^n$  s.t. F diffeomorphism with  $F^{-1} : \widetilde{W} \to V$ .

$$\widetilde{W} \ni (a,0) \subseteq \mathbb{R}^m \times \mathbb{R}^n$$

Set  $W = \{x \in \mathbb{R}^m | (x,0) \in \widetilde{W}\}$ . Then  $\pi_1(\widetilde{W}) = W$  open in  $\mathbb{R}^m$ .

Define  $q:W\to\mathbb{R}^n$ ,

$$g(x) = \pi_2 \circ F^{-1}(x, 0)$$
 or  $F^{-1}(x, 0) = (h(x), g(x))$ 

Now  $FF^{-1}(x,0) = (x,0) = (h(x), f(h(x), g(x)))$  so  $h(x) = x \,\forall x \in W, 0 = f(x, g(x))$ .

$$\{(x,y) \in V | f(x,y) = 0\} = \{(x,y) \in V | F(x,y) = (x,0)\} = \{(x,g(x)) | x \in W, 0 = f(x,g(x))\}$$

Since  $\pi$  smooth and  $F^{-1}$  is  $C^1$ , q is  $C^1$ .

To reiterate, f(x, g(x)) = 0 on W.

Using chain rule while differentiating f(x, q(x)) = 0,

$$\partial_{x^j} f(x, g(x)) = \frac{\partial f(x, g(x))}{\partial x^k} \frac{\partial x^k}{\partial x^j} + \frac{\partial f(x, g(x))}{\partial y^k} \frac{\partial g^k(x)}{\partial x^j} = D_x f|_{(x, g(x))} + (D_y f)|_{(x, g(x))} \cdot (Dg)_x = 0 \text{ or } (Dg)_x = -(D_y f)|_{(x, g(x))} D_x f|_{(x, g(x))}$$

#### 2. Immersions

**Definition 3** (Immersion). smooth  $f: M \to N$ , s.t.  $Df(p): T_pM \to T_{f(p)}N$  injective. Then f immersion at p.

Absil, Mahony, and Sepulchre [8] pointed out that another definition for a immersion can utilize the theorem that rank of  $Df \equiv DF = \dim T_n M$ . Indeed, recall these facts from linear algebra: for  $T: V \to W$ ,

It's always true that rank $T \leq V$ , and

$$\operatorname{rank} T \leq W$$

rankT = dimV iff T injective. rankT = dimW iff T surjective.

$$T_{x}M \xrightarrow{DF(x)} T_{F(x)}N = T_{y}N$$

$$\uparrow \qquad \qquad \uparrow \qquad \qquad \uparrow$$

$$x \in M \xrightarrow{F} y = F(x) \in N$$

$$M \longrightarrow F \longrightarrow N$$

Now

$$\dim T_x M = \dim M$$
$$\dim T_{F(x)} N = \dim N$$

And

$$rank(DF(x)) \equiv rank \text{ of } F$$

I know that the notation above is confusing, but this is what all Differential Geometry books apparently mean when they say "rank of F".

Now

$$\operatorname{rank}(DF(x)) = \dim(\operatorname{im}(DF(x))) = \dim T_x M \text{ iff } DF(x) \text{ injective}$$

If  $\forall x \in M$ , this is the case, then F an immersion.

Apply the rank-nullity theorem in this case:

$$\operatorname{rank}(DF(x)) + \operatorname{dimker}(DF(x)) = \operatorname{dim}T_x M = \operatorname{dim}M$$

$$\Longrightarrow \operatorname{rank}(DF(x)) = \operatorname{dim}M \le \operatorname{dim}T_{F(x)}N = \operatorname{dim}N \text{ or } \operatorname{dim}M \le \operatorname{dim}N$$

Now

$$\operatorname{rank}(DF(x)) = \dim T_{F(x)}N \text{ iff } DF(x) \text{ surjective}$$

If  $\forall x \in M$ , this is the case, then F an **submersion** 

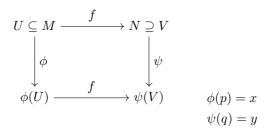
$$\operatorname{rank}(DF(x)) = \dim T_{F(x)}N = \dim N \le \dim M$$

Shastri (2011) has this as the "Injective Form of Implicit Function Theorem", Thm. 1.4.5, pp. 23 and Guillemin and Pollack (2010) has this as the "Local Immersion Theorem" on pp. 15, Section 3 "The Inverse Function Theorem and Immersions" [4].

**Theorem 6** (Local immersion Theorem i.e. Injective Form of Implicit Function Theorem). Suppose  $f: M \to N$  immersion at

Then  $\exists$  local coordinates around p, q, x, y, respectively s.t.  $f(x_1 \dots x_m) = (x_1 \dots x_m, 0 \dots 0)$ .

*Proof.* Choose local parametrizations



 $D(\psi f \varphi^{-1}) \equiv Df$ . Df(p) injective (given f immersion).  $Df(p) \in Mat(n,m)$ 

By change of basis in  $\mathbb{R}^n$ , assume  $Df(p) = \begin{pmatrix} I_m \\ 0 \end{pmatrix}$ 

Now define  $G: \phi(U) \times \mathbb{R}^{n-m} \to \mathbb{R}^n$ 

$$G(x,z) = f(x) + (0,z)$$

Thus, DG(x,z) = 1 and for open  $\phi(U) \times U_2$ ,  $G(\phi(U) \times U_2)$  open.

By inverse function theorem, G local diffeomorphism of  $\mathbb{R}^n$ , at 0.

Now  $f = G \circ i$ , where i is canonical immersion.

$$G(x,0) = f(x)$$

$$\Longrightarrow G^{-1}G(x,0) = (x,0) = G^{-1}f(x)$$

Use  $\psi \circ G$  as the local parametrization of N around pt. q. Shrink U, V so that

**Theorem 7** (Implicit Function Thm.). Let open subset  $U \subseteq \mathbb{R}^n \times \mathbb{R}^d$ ,  $(x,y) = (x^1 \dots x^n, y^1 \dots y^k)$  on U. Suppose smooth  $\Phi: U \to \mathbb{R}^k$ ,  $(a,b) \in U$ ,  $c = \Phi(a,b)$ 

If  $k \times k$  matrix  $\frac{\partial \Phi^i}{\partial y^j}(a, b)$  nonsingular, then  $\exists$  neighborhoods  $V_0 \subseteq \mathbb{R}^n$  of a and smooth  $F: V_0 \to W_0$  s.t.  $W_0 \subseteq \mathbb{R}^k$  of b

$$\Phi^{-1}(c) \bigcap (V_0 \times W_0)$$
 is graph of  $F$ , i.e.  $\Phi(x,y) = c$  for  $(x,y) \in V_0 \times W_0$  iff  $y = F(x)$ .

### 3. Submersions; Rank Theorem

cf. pp. 20, Sec. 4 "Submersions", Ch. 1 of Guillemin and Pollack (2010) [4]. Consider  $X, Y \in \mathbf{Man}$ , s.t.  $\dim X \geq \dim Y$ .

**Definition 4** (submersion). If  $f: X \to Y$ , if  $Df_x \equiv df_x$  is surjective,  $f \equiv submersion$  at x.

Recall that,

$$Df_x: T_xX \to T_{f(x)}Y$$
 
$$\dim T_xX \ge \dim T_{f(x)}Y$$
 
$$\mathrm{rank}Df_x \le \dim T_{f(x)}Y, \text{ in general, while}$$
 
$$\mathrm{rank}Df_x = \dim T_{f(x)}Y \text{ iff } Df_x \text{ surjective}$$

Canonical submersion is standard projection:

If 
$$\dim X = k, k \ge l$$
,  
 $\dim Y = l$ 

$$(a_1 \ldots a_k) \mapsto (a_1 \ldots a_l)$$

**Theorem 8** (Local Submersion Theorem). Suppose  $f: X \to Y$  submersion at x, and y = f(x), Then  $\exists$  local coordinates around x, y s.t.

$$f(x_1 \dots x_k) = (x_1 \dots x_l)$$

i.e. f locally equivalent to canonical submersion near x

*Proof.* I'll have a side-by-side comparison of my notation and the 1 used in Guillemin and Pollack (2010) [4] where I can. For charts  $(U, \phi), (V, \psi)$  for X, Y, respectively, y = f(x) for  $x \in X$ ,

$$U \subseteq X \xrightarrow{f} Y \supseteq V \qquad \qquad \downarrow \phi \qquad \downarrow \psi \circ G \qquad \qquad \downarrow \phi \qquad \qquad \downarrow \psi \qquad$$

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 $Dg_x$  surjective, so assume it's a  $l \times k$  matrix  $\begin{bmatrix} \mathbf{1}_l & 0 \end{bmatrix}$ . Define

(6) 
$$G: U \subset \mathbb{R}^k \to \mathbb{R}^k$$

$$G(a) \equiv G(a^1 \dots a^k) := (q(a), a_{l+1}, \dots, a_k)$$

Now

$$DG(a) = \begin{bmatrix} \mathbf{1}_l & 0 \\ \mathbf{1}_{k-l} \end{bmatrix} = \mathbf{1}_k$$

so G local diffeomorphism (at 0).

So  $\exists G^{-1}$  as local diffeomorphism of some U' of a into  $U \subset \mathbb{R}^k$ . By construction,

$$q = \mathbb{P}_l \circ G$$

where  $\mathbb{P}_l$  is the *canonical submersion*, the projection operator onto  $\mathbb{R}^l$ .

$$g \circ G^{-1} = \mathbb{P}_l$$

(since G diffeomorphism)

$$U \subseteq X \xrightarrow{f} V \subseteq Y$$

$$\phi^{-1} \circ G^{-1} \qquad \psi^{-1} \qquad \qquad \psi^{-1} \qquad \qquad \text{for}$$

$$\phi^{-1} \circ G^{-1}(a) \equiv \phi^{-1} \circ G^{-1}(a^{1} \dots a^{k}) = x \xrightarrow{f} f(x) = y = \psi^{-1}(a^{1} \dots a^{l})$$

$$\phi^{-1} \circ G^{-1} \qquad \qquad \psi^{-1} \qquad \qquad \psi^{-1} \qquad \qquad \downarrow$$

$$(a^{1} \dots a^{k}) \longmapsto \qquad \qquad \downarrow \qquad \downarrow$$

$$(a^{1} \dots a^{l}) \mapsto (a^{1} \dots a^{l})$$

"An obvious corollary worth noting is that if f is a submersion at x, then it is actually a submersion in a whole neighborhood of x." Guillemin and Pollack (2010) [4]

Suppose f submersion at  $x \in f^{-1}(y)$ .

By local submersion theorem

$$f(x_1 \dots x_k) = (x_1 \dots x_l)$$

Choose y = (0, ..., 0).

Then, near x,  $f^{-1}(y) = \{(0, \dots, 0, x_{l+1} \dots x_k)\}$  i.e. let  $V \ni x$  neighborhood of x, define  $(x_1 \dots x_k)$  on V.

Then  $f^{-1}(y) \cap V = \{(0 \dots 0, x_{l+1}, \dots x_k) | x_1 = 0, \dots x_l = 0\}.$ 

Thus  $x_{l+1}, \ldots x_k$  form a coordinate system on open set  $f^{-1}(y) \cap V \subseteq f^{-1}(y)$ .

$$U \subseteq X \xrightarrow{f} V \subseteq Y \qquad \qquad \downarrow \psi \qquad \qquad$$

$$\begin{array}{c}
f^{-1}(y) & \longleftarrow & y \\
\phi^{-1} & & \downarrow \psi \\
\{(0, \dots 0, x^1 \dots x^k)\} & \longleftarrow & (0 \dots 0)
\end{array}$$

3.1. Rank Theorem. Lee (2012) [3] in pp. 85, Ch. 4 Submersions, Immersions, and Embeddings, combines Theorems 6, 8 (local immersion and local submersion theorems, respectively) into the "Rank Theorem" (cf. Thm 4.12 "Rank Theorem" of Lee (2012)):

**Theorem 9** (Rank Theorem). Suppose smooth manifolds  $M, N, dim M = m, dim N = n, smooth map <math>F: M \to N, F$  has constant rank r.

 $\forall p \in M, \exists smooth charts (U, \varphi) for M, centered at p, (V, \psi) for N, centered at F(p), s.t.$ 

$$F(U) \subseteq V$$

in which F has coordinate representation of form

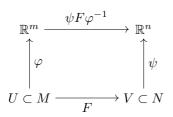
(9) 
$$\widehat{F}(x^1 \dots x^r, x^{r+1} \dots x^m) = (x^1 \dots x^r, 0 \dots 0)$$

Particularly, if F smooth submersion,

$$\widehat{F}(x^1 \dots x^n, x^{n+1} \dots x^m) = (x^1 \dots x^n)$$

and if F smooth immersion

$$\widehat{F}(x^1 \dots x^m) = (x^1 \dots x^m, 0 \dots 0)$$



Also remember that  $DF(p): T_pM \to T_{F(p)}N$ 

Proof. DF(p) has rank r (given). Then DF(p) is some  $r \times r$  submatrix of a  $n \times m$  matrix s.t.  $\det DF(p)$  nonzero. By change of basis in  $\mathbb{R}^n$ , or reordering coordinates, assume DF(p) is upper left submatrix  $\left(\frac{\partial F^i}{\partial x^j}\right) \quad \forall i, j = 1, \dots r$ . Relabel standard coordinate as

$$(x,y) = (x^1 \dots x^r, y^1 \dots y^{m-r}) \in \mathbb{R}^m$$
  
 $(v,w) = (v^1 \dots v^r, w^1 \dots w^{n-r}) \in \mathbb{R}^n$ 

By initial translations of coordinates, assume without loss of generality p = (0,0), F(p) = (0,0)Suppose

$$F(x,y) = (Q(x,y), R(x,y))$$

for some smooth maps  $Q:U\to\mathbb{R}^r,\ R:U\to\mathbb{R}^{n-r}$ Define

$$\varphi: U \to \mathbb{R}^m$$
  
 $\varphi(x, y) = (Q(x, y), y)$ 

SO SO

$$D\varphi(0,0) = \begin{pmatrix} \frac{\partial Q^i}{\partial x^j}(0,0) & \frac{\partial Q^i}{\partial y^j}(0,0) \\ 0 & \delta^i_j \end{pmatrix}$$

 $D\varphi(0,0)$  nonsingular, since  $\det \frac{\partial Q^i}{\partial x^j} \neq 0$  (by hypothesis).

By inverse function thm.,  $\exists$  connected neighborhoods  $U_0$  of (0,0),  $\widetilde{U}_0$  of  $\varphi(0,0) = (0,0)$  s.t.

$$\varphi: U_0 \to \widetilde{U}_0$$

is a diffeomorphism.

By shrinking  $U_0, \widetilde{U}_0$ , assume  $\widetilde{U}_0$  open cube.

Write  $\varphi^{-1}(x,y) = (A(x,y), B(x,y))$ , for some smooth functions  $A: \widetilde{U}_0 \to \mathbb{R}^r$ 

$$B: \widetilde{U}_0 \to \mathbb{R}^{m-r}$$

$$\begin{split} (x,y) &= \varphi(A(x,y),B(x,y)) = (Q(A(x,y),B(x,y)),B(x,y)) \\ \Longrightarrow & \frac{B(x,y) = y}{\varphi^{-1}(x,y) = (A(x,y),y)} \end{split}$$

$$\varphi \varphi^{-1} = 1 \Longrightarrow x = Q(A(x, y), y)$$

Recall that we had hypotehsized that

$$F(x,y) = (Q(x,y), R(x,y))$$

Then

$$F \circ \varphi^{-1}(x,y) = F(A(x,y),y) = (Q(A(x,y),y), R(A(x,y),y)) = (x, R(A(x,y),y))$$

and so

$$F \circ \varphi^{-1}(x,y) = (x, \widetilde{R}(x,y))$$

where  $\widetilde{R}:\widetilde{U}_0\to\mathbb{R}^{n-r}$ 

$$\widetilde{R}(x,y) = R(A(x,y),y)$$

Compute

$$D(F \circ \varphi^{-1})(x,y) = \begin{pmatrix} \delta_j^i & 0\\ \frac{\partial \tilde{R}^i}{\partial x^j}(x,y) & \frac{\partial \tilde{R}^i}{\partial y^j}(x,y) \end{pmatrix}$$

Since composing with a diffeomorphism doesn't change rank of map,  $D(F \circ \varphi^{-1})$  has rank r everywhere in  $\widetilde{U}_0$ .

$$\begin{pmatrix} \delta_j^i \\ \frac{\partial \widetilde{R}^i}{\partial x^j}(x,y) \end{pmatrix} j = 1 \dots r$$
 are linearly independent, so  $\frac{\partial \widetilde{R}^i}{\partial y^j}(x,y) = 0$  on  $\widetilde{U}_0$ , so  $\widetilde{R}^i$  independent of  $y^j$ .

Let  $S(x) = \widetilde{R}(x,0)$ , then

$$F \circ \varphi^{-1}(x,y) = (x, S(x))$$

Let open  $V_0 \subseteq V$ ,  $(0,0) \in V$  be an open subset  $V_0 = \{(v,w) \in V : (v,0) \in \widetilde{U}_0\}$ . Then  $V_0$  is a neighborhood of (0,0).

Because  $\widetilde{U}_0$  is a cube,  $F \circ \varphi^{-1}(x,y) = (x,S(x)),$ 

$$F \circ \varphi^{-1}(\widetilde{U}_0) \subseteq V_0$$

so  $F(U_0) \subseteq V_0$ .

Define 
$$\psi: V_0 \to \mathbb{R}^n$$

$$\psi(v, w) = (v, w - S(v))$$

Because  $\psi^{-1}(s,t) = (s, t + S(s)),$ 

it is a diffeomorphism.

Thus  $(V_0, \psi)$  is a smooth chart.

$$\psi \circ F(\varphi^{-1}(x,y)) = \psi(x,S(x)) = (x,S(x)-S(x)) = (x,0)$$

**Definition 5** (regular value). For smooth  $f: X \to Y$ ,  $X, Y \in Man$ ,  $y \in Y$  is a regular value for f if  $Df_x: T_xX \to T_yY$  surjective  $\forall x \ s.t. \ f(x) = y$ .

 $y \in Y$  critical value if y not a regular value of f.

Absil, Mahony, and Sepulchre [8] pointed out that another definition for a regular value can utilize the theorem that rank of  $Df \equiv DF = \dim T_p N = \dim N$ , iff DF(p) surjective, for  $p \in M$ ,  $F: M \to N$ . Then regular value  $y \in N$ , of F, if rank of  $F \equiv \operatorname{rank}(DF(x)) = \dim N$ ,  $\forall x \in F^{-1}(y)$ , for  $F: M \to N$ .

**Theorem 10** (Preimage theorem). If y regular value of  $f: X \to Y$ ,  $f^{-1}(y)$  is a submanifold of X, with  $dimf^{-1}(y) = dimX - dimY$ 

*Proof.* Given y is a regular value of  $f: X \to Y$ ,

 $\forall x \in f^{-1}(y), Df_x : T_x X \to T_y Y$  is surjective. By local submersion theorem,

$$f(x^1 \dots x^k) = (x^1 \dots x^l) = y$$

Since  $x \in f^{-1}(y)$ ,  $(x^1 \dots x^k) = (y^1 \dots y^l, x^{l+1} \dots x^k)$ .

For this chart for  $(U, \varphi)$ ,  $U \ni x$ , consider  $(U \cap f^{-1}(y), \psi)$  with  $\psi(x) = (x^{l+1} \dots x^k) \quad \forall x \in U \cap f^{-1}(y)$ .  $\forall f^{-1}(y)$  submanifold with  $\dim f^{-1}(y) = k - l = \dim X - \dim Y$ .

Examples for emphasis

If  $\dim X > \dim Y$ ,

if  $y \in Y$ , regular value of  $f: X \to Y$ ,

f submersion,  $\forall x \in f^{-1}(y)$ 

If  $\dim X = \dim Y$ ,

f local diffeomorphism  $\forall x \in f^{-1}(y)$ 

If  $\dim X < \dim Y$ ,  $\forall y \in f(X)$  is a critical value.

Example: O(n) as a submanifold of Mat(n, n)

Given  $\operatorname{Mat}(n,n) \equiv M(n) = \{n \times n \text{ matrices }\}$  is a manifold; in fact  $\operatorname{Mat}(n,n) \cong \mathbb{R}^{n^2}$ ,

Consider  $O(n) = \{A \in \text{Mat}(n, n) | AA^T = 1\}.$ 

(10) 
$$AA^T \in \operatorname{Sym}(n) \equiv S(n) = \{ S \in \operatorname{Mat}(n, n) | S^T = S \} = \{ \text{ symmetric } n \times n \text{ matrices } \}$$

 $\operatorname{Sym}(n)$  submanifold of  $\operatorname{Mat}(n,n)$ ,  $\operatorname{Sym}(n)$  diffeomorphic to  $\mathbb{R}^k$  (i.e.  $\operatorname{Sym}(n) \cong \mathbb{R}^k$ ),  $k := \frac{n(n+1)}{2}$ .

$$f: \operatorname{Mat}(n, n) \to \operatorname{Sym}(n)$$
  
 $f(A) = AA^T$ 

Notice f is smooth,

$$f^{-1}(1) = O(n)$$

$$Df_A(B) = \lim_{s \to 0} \frac{f(A+sB) - f(A)}{s} = \lim_{s \to 0} \frac{(A+sB)(A^T + sB^T) - AA^T}{s} = AB^T + BA^T$$

If  $Df_A: T_A \operatorname{Mat}(n,n) \to T_{f(A)} \operatorname{Sym}(n)$  surjective when  $A \in f^{-1}(1) = O(n)$  (???).

**Proposition 1.** If smooth  $g_1 
ldots g_l 
ldots C^{\infty}(X)$  on X are independent  $\forall x \in X$ , s.t.  $g_i(x) = 0$ ,  $\forall i = 1 
ldots ld$ 

*Proof.* Suppose smooth  $g_1 \dots g_l \in C^{\infty}(X)$  on manifold X s.t.  $\dim X = k \ge l$ .

Consider  $g = (g_1 \dots g_l) : X \to \mathbb{R}^l, Z \equiv g^{-1}(0)$ .

Since  $\forall g_i \text{ smooth, } D(g_i)_x : T_x X \to \mathbb{R} \text{ linear.}$ 

Now for

$$Dg_x = (D(g_1)_x \dots D(g_l)_x) : T_x X \to \mathbb{R}^l$$

By rank-nullity theorem (linear algebra),  $Dg_x$  surjective iff rank $Dg_x = l$  i.e. l functionals  $D(g_1)_x \dots D(g_l)_x$  are linearly independent on  $T_x X$ .

"We express this condition by saying the l functions  $g_1 \dots g_l$  are independent at x." (Guillemin and Pollack (2010) [4])

4. Submanifolds; immersed submanifold, embedded submanifolds, regular submanifolds

**Definition 6** (Embedded Submanifold).

Recall immersion:

 $F: M \to N$  immersion iff DF injective, i.e. iff rankDF = dim M.

Consider manifolds  $M \subseteq N$ .

Consider inclusion map  $i: M \to N$ .

$$i: x \mapsto x$$

If *i* immersion,  $Di(x) = \frac{\partial y^i}{\partial x^j} = \delta_i^{\ i}$  if  $y^i = x^i, \ \forall i = 1, \dots \text{dim} M$ .

**Definition 7** (immersed submanifold). *immersed submanifold*  $M \subseteq N$  if inclusion  $i: M \to N$  is an immersion.

cf. 3.3 Embedded Submanifolds of Absil, Mahony, and Sepulchre [8], also Ch. 5 Submanifolds, pp. 108, **Immersed Submanifolds** of John Lee (2012) [3].

Immersed submanifolds often arise as images of immersions.

**Proposition 2** (Images of Immersions as submanifolds). Suppose smooth manifold M,

smooth manifold with or without boundaries N.

injective, smooth immersion  $F: M \to N$  (F injective itself, not just immersion)

Let S = F(M)

Then S has unique topology and smooth structure of smooth submanifolds of N s.t.  $F: M \to S$  diffeomorphism.

cf. Prop. 5.18 of John Lee (2012) [3].

*Proof.* Define topology of S: set  $U \subseteq S$  open iff  $F^{-1}(U) \subseteq M$  open  $(F^{-1}(U \cap V) = F^{-1}(U) \cap F^{-1}(V), F^{-1}(U \cup V) = F^{-1}(U) \cup F^{-1}(V))$ .

Define smooth structure of  $S: \{F(U), \varphi \circ F^{-1} | (U, \varphi) \in \text{atlas for } M, \text{ i.e. } (U, \varphi) \text{ any smooth chart of } M\}.$  "smooth compatibility condition":

$$(\varphi_2 \circ F^{-1})(\varphi_i F^{-1})^{-1} = \varphi_2 \circ F^{-1} F \varphi_1^{-1} = \varphi_2 \varphi_1^{-1}$$

since  $\varphi_2\varphi_1^{-1}$  diffeomorphism ( $\varphi_2\varphi_1^{-1}$  bijection and it and inverse is differentiable)

F diffeomorphism onto F(M).

and these are the only topology and smooth structure on S with this property:

$$S \xrightarrow{F^{-1}} M \xrightarrow{F} N \qquad = \qquad S \hookrightarrow M$$

and  $F^{-1}$  diffeomorphism, F smooth immersion, so  $i: S \to M$  smooth immersion.

5. Curves, Integral Curves, and Flows

cf. John Lee (2012) [3], Ch. 9, deals with time-dependent vector fields and I don't see other texts or references handling such an important, but overlooked, case.

### 5.1. Curves in Euclidean space. cf. 4.1 Jeffrey Lee (2009) [2]

If C is 1-dim. submanifold of  $\mathbb{R}^n$ ,  $p \in C$ ,  $\exists$  chart (V, y) of C,  $p \in V$  s.t. y(V) is a connected open interval  $I \subset \mathbb{R}$ , inverse map  $y^{-1}: I \to V \subset M$  is a local parametrization.

idea is to extract information that's appropriately independent of parametrization.

If 
$$\gamma: I \to \mathbb{R}^n$$
,  $c: J \to \mathbb{R}^n$  curves with same image,

c is a positive reparametrization of  $\gamma$  if  $\exists$  smooth  $h: J \to I$  with h' > 0 s.t.  $c = \gamma \circ h$ ,

in this case,  $\gamma$ , c have same sense and same orientation

Assume 
$$\gamma: I \to \mathbb{R}^n$$
 has  $||\gamma'|| > 0$ ,

Such a curve is **regular**, i.e. curve is an immersion (Recall, an immersion would be smooth  $\gamma: I \to \mathbb{R}^n$ , s.t.  $D\gamma(p): T_pI \to$  $T_{\gamma(p)}\mathbb{R}^n$  injective).

**Definition 8** (unit tangent field in Euclidean space, 4.3 Lee (2009) [2]). If  $\gamma: I \to \mathbb{R}^n$  regular curve, then  $\mathbf{T}(t) := \frac{\gamma'(t)}{\|\gamma'(t)\|}$  defines unit tangent field along  $\gamma$  (||T|| = 1)

**length of a curve** defined on closed interval  $\gamma: [t_1, t_2] \to \mathbb{R}^n$ ,

$$L = \int_{t_1}^{t_2} \|\gamma'(t)\| dt$$

Define arc length function for curve  $\gamma: I \to \mathbb{R}^n$  by choosing  $t_0 \in I$ ,

$$s = h(t) := \int_{t_0}^t \|\gamma'(t)\| d\tau$$

If curve is smooth and regular, then  $h' = \|\gamma'(\tau)\| > 0$ , so by inverse function theorem,  $\exists$  smooth  $h^{-1}$  (since h' invertible (with 1/h')).

If  $c(s) = \gamma h^{-1}(s)$ , then ||c'||(s) := ||c'(s)|| = 1.  $\forall s$ 

$$c'(s) = (\gamma(h^{-1}(s)))' = \gamma'(h^{-1}(s)) \cdot \frac{dh^{-1}}{ds}(s)$$
$$\|c'(s)\| = \|\gamma'(t)\| \|\frac{dh^{-1}}{ds}(s)\| = h' \cdot \frac{1}{h'} = 1$$

Curves parametrized by arc length are unit speed curves.

For a unit speed curve,  $\frac{dc}{ds}(s) = \mathbf{T}(s)$ .

**Definition 9** (curvature vector, curvature function, principal normal in Euclidean space, 4.4 Lee (2009) [2]). Let  $c: I \to \mathbb{R}^n$  be a unit speed curve. vector valued function

$$\kappa(s) := \frac{d\mathbf{T}}{ds}(s)$$

is called the **curvature vector**.

$$\kappa(s) := \|\kappa(s)\| = \|\frac{d\mathbf{T}}{ds}(s)\|$$

is called the curvature function.

If  $\kappa(s) > 0$ , then define principal normal

$$\mathbf{N}(s) = \|\frac{d\mathbf{T}}{ds}(s)\|^{-1} \frac{d\mathbf{T}}{ds}(s)$$

s.t. 
$$\frac{d\mathbf{T}}{ds} = \kappa \mathbf{N}$$

5.2. **Integral Curves.** cf. Ch. 9 of Lee (2012) [3]

integral curves - smooth curves whose velocity at each point is equal to the value of the smooth vector field there. Let smooth manifold M with or without boundary.

If smooth curve  $\gamma: J \to M$ , then  $\forall t \in J$ , velocity vector  $\gamma'(t)$  is a vector in  $T_{\gamma(t)}M$ .

**Definition 10** (Integral Curve  $\Gamma$ ). Integral curve:  $\mathbf{v}, \gamma$ .

If vector field  $\mathbf{v}$  on M,

integral curve of v is differentiable curve  $\gamma: J \to M$ , whose velocity  $\forall p \in M, V_{\gamma(t)}$ 

$$\gamma'(t) = V_{\gamma(t)}, \, \forall \, t \in J$$

Finding integral curves boils downs to solving a system of ordinary differential equations in a smooth chart.

Suppose smooth vector field V on M, smooth curve  $\gamma: J \to M$ 

On smooth coordinate domain  $U \subseteq M$ ,

Write  $\gamma$  in local coordinates as

$$\gamma(t) = (\gamma^1(t), \dots, \gamma^n(t))$$
$$\gamma^i(t) = x^i \circ \gamma(t) \text{ where } x^i : U \subseteq M \to \mathbb{R}$$

Then condition for  $\gamma$  to be an integral curve.

$$\gamma'(t) \equiv \dot{\gamma}(t) = V_{\gamma(t)}$$

can be written as

$$\dot{\gamma}^{i}(t) \left. \frac{\partial}{\partial x^{i}} \right|_{\gamma(t)} = V^{i}(\gamma(t)) \left. \frac{\partial}{\partial x^{i}} \right|_{\gamma(t)}$$

which reduces to an autonomous system of ordinary differential equations (ODEs

$$\dot{\gamma}^1(t) = V^1(\gamma^1(t), \dots, \gamma^n(t))$$

(11)

$$\dot{\gamma}^n(t) = V^n(\gamma^1(t), \dots, \gamma^n(t))$$

20200306 EY: I use this notation which is more useful for me, e.g. taking the double derivative of the curve  $\gamma$ :

$$\dot{\gamma}(t) = v(\gamma(t)) \in T_{\gamma(t)}U$$

$$\dot{\gamma}(t) = v^i(\gamma^1(t), \dots, \gamma^n(t))$$

$$\gamma(t) = v^{i}(\gamma^{i}(t), \dots, \gamma^{n}(t))$$

$$\frac{d}{dt}\dot{\gamma}^{i}(t) = \frac{d}{dt}v^{i}(\gamma^{1}(t), \dots, \gamma^{n}(t)) = \frac{d}{dt}v^{i}(x^{1}, \dots, x^{n})$$

$$= \frac{\partial v^{i}}{\partial x^{j}}(x)\dot{x}^{j} = \dot{\gamma}^{j}\frac{\partial \dot{\gamma}^{i}}{\partial x^{j}}$$

### 6. Push-forward, Pull-back

6.1. Push-forward in coordinates,  $f_*: T_pM \to T_{f(p)}N$ . Using Nakahara (2003) [7]'s notation.

Consider a smooth mapping  $f: M \to N$ ,  $\dim M = m, \dim N = n$ .

$$f_*: T_pM \to T_{f(p)}N$$

With

$$(U,\varphi) \subset M \quad \varphi: U \to \mathbb{R}^m \quad \varphi^i \equiv x^i: U \to \mathbb{R}$$

$$(V, \psi) \subset N \quad \psi : V \to \mathbb{R}^n \quad \psi^j \equiv u^j : V \to \mathbb{R}$$

Now  $X \in T_p M$ , so we can write  $X = X^i \frac{\partial}{\partial x^i}$ , and  $f_* X \in T_{f(p)} N$ , so we can write  $f_* X = Y^j \frac{\partial}{\partial y^j}$ .  $Y \in T_{f(p)} N$   $Y = Y^j \frac{\partial}{\partial y^j}$ 

Let  $g \in \mathcal{F}(N)$ 

 $g:N\to\mathbb{R}$ 

Now  $g \circ \psi^{-1} : \mathbb{R}^n \to \mathbb{R}$  or  $g \circ \psi^{-1} : \psi(N) \to \mathbb{R}$ .

For notation, let  $g \circ \psi^{-1}(y) \equiv g(y)$ , in that we treat y to be a point with coordinates in  $\mathbb{R}^n$  notation-wise, and not as  $y^j : V \to \mathbb{R}$ . Theorem 12 (Pullback in coordinates). For smooth  $f : M \to N$ ,

Since  $g \circ f : M \to \mathbb{R}$ , so  $g \circ f \in \mathcal{F}(M)$ .

$$(f_*X)[g] \equiv X[g \circ f]$$

$$(12) \qquad (f_*X)[g] = (Y^j \frac{\partial}{\partial y^j})[g] = Y^j \frac{\partial g}{\partial y^j} = X^i \frac{\partial}{\partial x^i}[g \circ f] = X^i \frac{\partial}{\partial x^i}g(f^j(x)) = X^i \frac{\partial g}{\partial y^j}(y) \frac{\partial f^j}{\partial x^i}(x) = X^i \frac{\partial f^j}{\partial x^i}(x) \frac{\partial}{\partial y^j}g$$

$$Y^j = X^i \frac{\partial f^j}{\partial x^i} = X^i \frac{\partial y^j}{\partial x^i} = F^j_i(x)X^i$$

where  $f: M \to N$ , but  $\psi \circ f \circ \varphi^{-1}: \varphi(U) \subset \mathbb{R}^m \to \mathbb{R}^n$ , but write this notation as  $f^j(x) = y^j$ .

Also, notice that for Eq. 12,

$$g \circ f : M \to \mathbb{R}$$
$$g \circ \psi^{-1} \circ \psi \circ f \circ \varphi^{-1} \text{ and } g \circ f \circ \varphi^{-1} : \varphi(U) \subset \mathbb{R}^m \to \mathbb{R}$$
$$(g \circ \psi^{-1})(\psi \circ f \circ \varphi^1)(x) \equiv g(y(x))$$

6.2. Pullback in coordinates,  $f^*: T^*_{f(p)}N \to T^*_pM$ . For pullback  $f^*: T^*_{f(p)}N \to T^*_pM$ , since  $\omega \in T^*_{f(p)}N$ , then write as  $\omega_j dy^j$  since  $f^*\omega \in T^*_pM$ , so  $f^*\omega = \psi_i dx^i$ 

$$\langle f^*\omega, X \rangle \equiv f^*\omega(X) = \langle \psi_i dx^i, X^j \frac{\partial}{\partial x^j} \rangle \equiv \psi_i dx^i (X^j \frac{\partial}{\partial x^j}) = \psi_i X^i = \langle \omega, f_* X \rangle \equiv \omega(f_* X) =$$

$$= \langle \omega_j dy^j, Y^i \frac{\partial}{\partial y^i} \rangle = \omega_j dy^j (Y^i \frac{\partial}{\partial y^i}) = \omega_j dy^j \left( X^i \frac{\partial y^k}{\partial x^i} \frac{\partial}{\partial y^k} \right) = \omega_j \frac{\partial y^i}{\partial x^i} X^i$$

$$\Longrightarrow \psi_i = \omega_j \frac{\partial y^j}{\partial x^i}$$

Thus

Theorem 11 (Push-forward in coordinates).

For smooth 
$$f: M \to N$$
,  $dim M = m$   
 $dim N = n$   
 $X = X^i \frac{\partial}{\partial x^i} \in T_p M$   
 $Y = Y^j \frac{\partial}{\partial y^j} \in T_{f(p)} N$  then

(13) 
$$f_*: T_pM \to T_{f(p)}N$$

$$f_*(X) = Y$$

$$Y^j = X^i \frac{\partial y^j}{\partial x^i}$$

where  $f_*$  is the **push-forward**, and where colloquially,  $y^j = y^j(x)$ , but actually  $y^j \equiv \psi \circ f \circ \varphi^{-1}$ And

**Theorem 12** (Pullback in coordinates). For smooth  $f: M \to N$ , dimM = m dimN = n  $\omega \in T^*_{f(n)}N$ ,  $\psi = \psi_i dx^i \in T^*_n M$ , then

(14) 
$$f^*: T^*_{f(p)}N \to T^*_pM$$

$$f^*\omega = \psi_i dx^i$$

$$\psi_i = \omega_j \frac{\partial y^j}{\partial x^j}$$

7. Tensors

I'll go through Ch.7 Tensors of Jeffrey Lee (2009) [2].

**Definition 11** (7.1[2]). Let V, W be modules over commutative ring R, with unity. Then, algebraic W-valued tensor on V is multilinear map.

$$\tau: V_1 \times V_2 \times \dots \times V_m \to W$$

where  $V_i = \{V, V^*\}$   $\forall i = 1, 2, ... m$ .

If for r, s s.t. r + s = m, there are r  $V_i = V^*$ ,  $sV_i = V$ , tensor is r-contravariant, s-covariant; also say tensor of total type  $\binom{r}{s}$ .

EY: 20170404 Note that

$$(\tau_{\beta}^{i\alpha} \frac{\partial}{\partial x^{i}} \text{ or } \tau_{\beta}^{i\alpha} e_{i})(\omega_{j} dx^{j} \text{ or } \omega_{j} e^{j} \in V^{*})$$
$$(\tau_{i\alpha}^{\beta} dx^{i} \text{ or } \tau_{i\alpha}^{\beta} e^{i})(X^{j} \frac{\partial}{\partial x^{j}} \text{ or } X^{j} e_{j} \in V)$$

 $\exists$  natural map  $V \to V^{**}$ ,  $\widetilde{v} : \alpha \mapsto \alpha(v)$ . If this map is an isomorphism, V is **reflexive** module, and identify V with  $V^{**}$ .

**Exercise 7.5.** Given vector bundle  $\pi: E \to M$ , open  $U \subset M$ , consider sections of  $\pi$  on U, i.e. cont.  $s: U \to E$ , where  $(\pi \circ s)(u) = u$ ,  $\forall u \in U$ .

Consider  $E^* \ni \omega = \omega_i e^i$ 

 $\forall s \in \Gamma(E), \, \omega(s) = \omega_i(s(x))^i, \ \forall x \in U \subset M. \text{ So define } \widetilde{s} : \omega, x \mapsto \omega(s(x)), \ \forall x \in U.$ 

If  $\widetilde{s} = 0$ ,  $\widetilde{s}(\omega, x) = \omega(s(x)) = 0$   $\forall \omega \in E^*$ ,  $\forall x \in U$ , and so s = 0. (Let  $\omega_i = \delta_{iJ}$  for some J, and so  $s^J(x) = 0$   $\forall J$ ).

s = 0. So  $\ker(s \mapsto \tilde{s}) = \{0\}$  (so condition for injectivity is fulfilled).

Since  $\widetilde{s}: \omega, x \mapsto \omega(s(x)), \forall \omega \in E^*, \forall x \in U, s \mapsto \widetilde{s}$  is surjective.

 $s \mapsto \widetilde{s}$  is an isomorphism so  $\Gamma(E)$  is a reflexive module.

**Proposition 3.** For R a ring (special case),  $\exists$  module homomorphism:

tensor product space  $\rightarrow$  tensor, as a multilinear map, i.e.  $\exists$ 

$$(16) \qquad (\bigotimes_{i=1}^{r} V) \otimes (\bigotimes_{j=1}^{s} V^{*}) \to T_{s}^{r}(V; R) u_{1} \otimes \cdots \otimes u_{r} \otimes \beta^{1} \otimes \cdots \otimes \beta^{s} \in (\bigotimes^{r} V) \otimes (\bigotimes^{s} V^{*}) \mapsto (\alpha^{1} \dots \alpha^{r}, v_{1} \dots v_{s}) \mapsto \alpha^{1}(u_{1}) \dots \alpha^{r}(u_{r}) \beta^{1}(v_{1}) \dots \beta^{s}(v_{s})$$

Indeed, consider

$$(\alpha^1 \dots \alpha^r, v_1 \dots v_s) \in \underbrace{V^* \times \dots \times V^*}_r \times \underbrace{V \times \dots \times V}_s \mapsto \alpha^1(u_1) \dots \alpha^r(u_r) \beta^1(v_1) \dots \beta^s(v_s)$$

and so for

$$\alpha^{i} = \alpha_{\mu}^{i} e^{\mu}, \quad i = 1, 2, \dots r, \ \mu = 1, 2, \dots \dim V^{*} \qquad \alpha^{i}(u_{i}) = \alpha_{\mu}^{i} u_{i}^{\mu}$$

$$v_{i} = v_{i}^{\mu} e_{\mu}, \quad i = 1, 2, \dots s, \ \mu = 1, 2, \dots \dim V \qquad \beta^{i}(v_{i}) = \beta_{\mu}^{i} v_{i}^{\mu}$$

So that

$$\alpha^{1}(u_{1}) \dots \alpha^{r}(u_{r})\beta^{1}(v_{1}) \dots \beta^{s}(v_{s}) = \alpha_{\alpha_{1}}^{1} u_{1}^{\alpha_{1}} \dots \alpha_{\alpha_{r}}^{r} u_{r}^{\alpha_{r}} \beta_{\mu_{1}}^{1} v_{1}^{\mu_{1}} \dots \beta_{\mu_{s}}^{s} v_{s}^{\mu_{s}} = (u_{1}^{\alpha_{1}} \dots u_{r}^{\alpha_{r}} \beta_{\mu_{1}}^{1} \dots \beta_{\mu_{s}}^{s})(\alpha_{\alpha_{1}}^{1} \dots \alpha_{\alpha_{r}}^{r} v_{1}^{\mu_{1}} \dots v_{s}^{\mu_{s}})$$

Identify  $u_1 \otimes \cdots \otimes u_r \otimes \beta^1 \otimes \cdots \otimes \beta^s$  with this multiplinear map.

**Proposition 4.** If V is finite-dim. vector space, or if  $V = \Gamma(E)$ , for vector bundle  $E \to M$ , map

$$(17) \qquad (\bigotimes_{i=1}^r V) \otimes (\bigotimes_{i=1}^s V^*) \to T_s^r(V; R)$$

is an isomorphism.

**Definition 12.** tensor that can be written as

$$(18) u_1 \otimes \cdots \otimes u_r \otimes \beta^1 \otimes \cdots \otimes \beta^s \equiv u_1 \otimes \cdots \otimes \beta^s$$

is simple or decomposable.

Now well that not *all* tensors are simple.

**Definition 13** (7.7[2], tensor product).  $\forall S \in T_{s_1}^{r_1}(V), \ \forall T \in T_{s_2}^{r_2}(V),$  define tensor product

(19) 
$$S \otimes T \in T_{s_1+s_2}^{r_1+r_2}(V)$$
$$S \otimes T(\theta^1 \dots \theta^{r_1+r_2}, v_1 \dots v_{s_1+s_2}) := S(\theta^1 \dots \theta^{r_1}, v_1 \dots v_{s_1})T(\theta^{r_1+1} \dots \theta^{r_1+r_2}, v_{s_1+1} \dots v_{s_1+s_2})$$

Proposition 5 (7.8[2]).

$$\tau^{i_1 \dots i_r}_{j_1 \dots j_s} e_{i_1} \otimes \dots \otimes e_{i_r} \otimes e^{j_1} \otimes \dots \otimes e^{j_s} = \tau(e^{i_1} \dots e^{i_r}, e_{j_1} \dots e_{j_s}) e_{i_1} \otimes \dots \otimes e_{i_r} \otimes e^{j_1} \otimes \dots \otimes e^{j_s} = \tau(e^{i_1} \dots e^{i_r}, e_{j_1} \dots e_{j_s}) e_{i_1} \otimes \dots \otimes e_{i_r} \otimes e^{j_1} \otimes \dots \otimes e^{j_s} = \tau(e^{i_1} \dots e^{i_r}, e_{j_1} \dots e_{j_s}) e_{i_1} \otimes \dots \otimes e^{j_s} = \tau(e^{i_1} \dots e^{i_r}, e_{j_1} \dots e^{i_r}) e_{i_1} \otimes \dots \otimes e^{i_r} \otimes e^{j_1} \otimes \dots \otimes e^{j_s} = \tau(e^{i_1} \dots e^{i_r}, e_{j_1} \dots e^{j_s}) e_{i_1} \otimes \dots \otimes e^{j_s} = \tau(e^{i_1} \dots e^{i_r}, e_{j_1} \dots e^{j_s}) e_{i_1} \otimes \dots \otimes e^{j_s} = \tau(e^{i_1} \dots e^{i_r}, e^{i_1} \dots e^{i_r}) e^{i_1} \otimes \dots \otimes e^{i_r} \otimes e^{i_r} \otimes e^{i_1} \otimes \dots \otimes e^{i_r} \otimes e^{i_r} \otimes e^{i_r} \otimes e^{i_1} \otimes \dots \otimes e^{i_r} \otimes e^{i_r}$$

So  $\{e_{i_1} \otimes \cdots \otimes e_{i_r} \otimes e^{j_1} \otimes \cdots \otimes e^{j_s} | i_1 \dots i_r, j_1 \dots j_s \in 1 \dots n\}$  spans  $T_s^r(V; R)$ 

**Exercise 7.11.** Let basis for  $V = e_1 \dots e_n$ , corresponding dual basis for  $V^* = e^1 \dots e^n$ 

Let basis for  $V = \overline{e}_1 \dots \overline{e}_n$ , corresponding dual basis for  $V^* = \overline{e}^1 \dots \overline{e}^n$  s.t.

$$\overline{e}_i = C^k_{\ i} e_k$$
$$\overline{e}^i = (C^{-1})^i_{\ k} e^k$$

EY:20170404, keep in mind that

$$Ax = e_i A^i_{\ k} e^k(x^j e_j) = e_i A^i_{\ j} x^j = A^i_{\ j} x^j e_i$$
$$Ae_i = e_k A^k_i e^i(e_i) = A^k_i e_k = \overline{e}_i$$

$$\overline{\tau}^{i}{}_{jk}\overline{e}_{i} \otimes \overline{e}^{j} \otimes \overline{e}^{k} = \overline{\tau}^{i}{}_{jk}C^{l}{}_{i}e_{l}(C^{-1})^{j}{}_{m}e^{m}(C^{-1})^{k}{}_{n}e^{n} = \overline{\tau}^{i}{}_{jk}C^{l}{}_{i}(C^{-1})^{j}{}_{m}(C^{-1})^{k}{}_{n} = \tau^{l}{}_{mn}$$

$$\overline{\tau}^{i}{}_{jk} = C^{c}{}_{k}C^{b}{}_{j}(C^{-1})^{i}{}_{a}\tau^{a}{}_{bc}$$

On Remark 7.13 of Jeffrey Lee (2009) [2]: first, egregious typo for L(V, V); it should be L(V, W). Onward, for L(V, W),

consider  $W \otimes V^* \ni w \otimes \alpha$  s.t.

$$(w \otimes \alpha)(v) = \alpha(v)w \in W, \forall v \in V, \text{ so } w \otimes \alpha \in L(V, W)$$

Now consider (category of) left R-module,

$${}_{R}\mathbf{Mod} \ni {}_{\mathrm{Mat}_{\mathbb{K}}(N,M)}\mathbb{K}^{N}$$

where

$$V = \mathbb{K}^N$$
$$W = \mathbb{K}^M$$

For  $A \in \operatorname{Mat}_{\mathbb{K}}(N, M)$ ,  $x \in \mathbb{K}^N$ ,

$$e_i A^i_{,\mu} e^{\mu}(x^{\nu} e_{\nu}) = Ax = e_i A^i_{\mu} x^{\mu}, \quad i = 1, 2, \dots M, \, \mu = 1, 2, \dots N$$

$$A \in \operatorname{Mat}_{\mathbb{K}}(N, M) \cong W \otimes V^* \cong L(V, W)$$

Consider

$$\alpha \in (\mathbb{K}^N)^* = V^* \qquad \alpha = \alpha_\mu e^\mu$$

$$w \in \mathbb{K}^M = W \qquad w = w^i e_i$$

$$\alpha \otimes w = w \otimes \alpha = w^i \alpha_\mu e_i \otimes e^\mu$$

(remember, isomoprhism between  $\operatorname{Mat}_{\mathbb{K}}(N, M)$  and  $W \otimes V^*$  guaranteed, if V, W are free R-modules,  $R = \mathbb{K}$ ). Let V, W be left R-modules, i.e.  $V, W \in {}_{R}\mathbf{Mod}$ .

$$V^* \in \mathbf{Mod}_R$$

For  $V^* \otimes W \in \mathbf{Mod}_R \otimes {}_R\mathbf{Mod}$ 

$$\alpha \in V^*, w \in W$$
 
$$(\alpha \otimes w)(v) = \alpha(v)w, \text{ for } v \in V \in {}_R\mathbf{Mod}$$

But  $(w \otimes \alpha)(v) = w\alpha(v)$ .

Note  $\alpha(v) \in R$ .

Let V, W be right R-modules, i.e.  $V, W \in \mathbf{Mod}_R$ .

$$V^* \in {}_{R}\mathbf{Mod}$$

For  $W \otimes V^* \in \mathbf{Mod}_R \otimes {}_R\mathbf{Mod}$ .

$$\alpha \in V^*, \ w \in W$$
 
$$(v)(w \otimes \alpha) = w\alpha(v), \ \text{with} \ \alpha(v) \in R, \ v \in V$$

So  $W \otimes V^* \cong L(V, W)$ , for  $V, W \in \mathbf{Mod}_R$ 

**Definition 14** (7.20[2], **contraction**). Let  $(e_1, \ldots e_n)$  basis for V,  $(e^1 \ldots e^n)$  dual basis. If  $\tau \in T^r_s(V)$ , then for k < r, l < s, define

(21) 
$$C_l^k \tau \in T_{s-1}^{r-1}(V)$$

$$C_l^k \tau(\theta^1 \dots \theta^{r-1}, w_1 \dots w_{s-1}) :=$$

$$\sum_{a=1}^n \tau(\theta^1 \dots \underbrace{e^a}_{kth \ position} \dots \theta^{r-1}, w_1 \dots \underbrace{e_a}_{ith \ position} \dots w_{s-1})$$

 $C_l^k$  is called **contraction**, for some single  $1 \le k \le r$ , some single  $1 \le l \le s$ ,

$$C_l^k: T_s^r(V) \to T_{s-1}^{r-1}(V)$$

s.t.

$$(C_l^k \tau)^{i_1 \dots \widehat{i_k} \dots i_r}_{j_1 \dots \widehat{j_l} \dots j_s} := \tau^{i_1 \dots a \dots i_r}_{j_1 \dots a \dots j_s}$$

Universal mapping properties can be invoked to give a basis free definition of contraction (EY: 20170405???) IN general,

$$\forall v_1 \dots v_s \in V, \forall \alpha^1 \dots \alpha^r \in V^*$$

so that

$$v_j = v_j^{\mu} e_{\mu}$$
  $j = 1 \dots s$ ,  $\mu = 1, \dots \dim V$   
 $\alpha^i = \alpha_{\mu}^i e^{\mu}$   $i = 1 \dots r$ ,  $\mu = 1 \dots \dim V^*$ 

then  $\forall \tau \in T^r_s(V)$ ,

$$\tau(\alpha^{1} \dots \alpha^{r}, v_{1} \dots v_{s}) = \tau(\alpha_{\mu_{1}}^{1} e^{\mu_{1}} \dots \alpha_{\mu_{r}}^{r} e^{\mu_{r}}, v_{1}^{\nu_{1}} e_{\nu_{1}} \dots v_{s}^{\nu_{s}} e_{\nu_{s}}) =$$

$$= \alpha_{\mu_{1}}^{1} \dots \alpha_{\mu_{r}}^{r} v_{1}^{\nu_{1}} \dots v_{s}^{\nu_{s}} \tau(e^{\mu_{1}} \dots e^{\mu_{r}}, e_{\nu_{1}} \dots e_{\nu_{s}}) = \alpha_{\mu_{1}}^{1} \dots \alpha_{\mu_{r}}^{r} v_{1}^{\nu_{1}} \dots v_{s}^{\nu_{s}} \tau^{\mu_{1} \dots \mu_{r}}_{\nu_{1} \dots \nu_{s}}$$

which is equivalent to

$$\tau \in T_s^r(V) \xrightarrow{\alpha^1 \otimes \cdots \otimes \alpha^r \otimes v_1 \otimes \cdots \otimes v_s \otimes \alpha^1} \alpha^1 \otimes \cdots \otimes \alpha^r \otimes v_1 \otimes \cdots \otimes v_s \otimes \tau$$

$$C_{s+1}^1 C_{s+2}^2 \dots C_{r+s}^r C_1^r C_2^{r+1} \dots C_s^{r+s}$$

$$\tau(\alpha^1 \dots \alpha^r, v_1 \dots v_s) \in R$$

where I've tried to express the right-R-module, "right action" on  $\alpha^1 \otimes \cdots \otimes \alpha^r \otimes v_1 \otimes \cdots \otimes v_s \in V^* \otimes \cdots \otimes V$ . Conlon (2008) [17]

### Part 4. Lie Groups, Lie Algebra

8. Lie Groups

- : Lie Groups
- : Groups
- : Ring
- : group algebra
- : Group Ring
- : Representation Theory
- : Modules
- : kG-modules

From Sec. 8.1 "Noncommutative Rings" of Rotman (2010) [10]:

**Definition 15.** ring R - additive abelian group equipped with multiplication  $R \times R \to R$  s.t.  $\forall a, b \in R$   $(a,b) \mapsto ab$ 

- (i) a(bc) = (ab)c
- (ii) a(b+c) = ab + ac, (b+c)a = ba + ca
- (iii)  $\exists 1 \in R \text{ s.t. } \forall a \in R, 1a = a = a1$

Example 8.1[10]

(ii) group algebra kG, k commutative ring, G group, "its additive abelian group is free k-module having basis labeled by elements of G,

i.e.  $\forall a \in kG, a = \sum_{g \in G} a_g g, a_g \in k, \forall g \in G, a_g \neq 0$  for only finitely many  $g \in G$ .

define (ring) multiplication 
$$kG \times kG \to kG$$
  $\forall a,b \in kG$ ,  $a = \sum_{g \in G} a_g g$  to be 
$$ab = ab$$
 
$$b = \sum_{h \in G} b_h g$$

$$\left(\sum_{g \in G} a_g g\right) \left(\sum_{h \in G} b_h h\right) = \sum_{z \in G} \left(\sum_{gh=z} a_g b_h\right) z$$

**Definition 16.** Given R ring, left R-module is (additive) abelian group M equipped with

scalar multiplication  $R \times M \to M$  s.t.  $\forall m, m' \in M, \forall r, r', 1 \in R$ 

$$(r,m)\mapsto rm$$

- (i) r(m+m') = rm + rm'
- (ii) (r + r')m = rm + r'm
- (iii) (rr')m = r(r'm)
- (iv) 1m = m

EY: 20150922 Example: for kG-module  $V^{\sigma}$ , for  $r \in kG$ , so  $r = \sum_{g \in G} a_g g$ 

$$\begin{array}{c} R \times M \to M \\ (r,m) \mapsto rm \end{array} \Longrightarrow \begin{array}{c} kG \times V \to V \\ (r,v) \mapsto tv \end{array}$$

For some representation  $\sigma: G \to GL(V)$ ,

$$rv = \sum_{g \in G} a_g g \cdot v = \sum_{g \in G} a_g \sigma_g(v)$$

So a kG-module needs to be associated with some chosen representation.

Note for V as an additive abelian group,  $\forall u, v, w \in V$ ,

$$v + w = w + v$$
,  $(u + v) + w = u + (v + w)$   
 $v + 0 = v \quad \forall v \in V \text{ for } 0 \in V$   
 $v + (-v) = 0 \quad \forall v \in V$ 

So a vector space can be an additive abelian group.

Note that

 $1v = \sigma(1)v = 1v = v$ 

$$r(v+w) = \left(\sum_{g \in G} a_g g\right)(v+w) = \left(\sum_{g \in G} a_g \sigma_g\right)(v+w) = \sum_{g \in G} a_g \sigma_g(v) + \sum_{g \in G} a_g \sigma_g(w) = rv + rw$$

$$(r+r')v = \left(\sum_{g \in G} a_g g + b_g g\right)v = \sum_{g \in G} (a_g \sigma_g + b_g \sigma_g)v = \sum_{g \in G} a_g \sigma_g(v) + \sum_{g \in G} b_g \sigma_g(v) = rv + r'v$$

$$(rr')v = \left(\sum_{g \in G} a_g g \sum_{h \in G} b_h h\right)v = \left(\sum_{z \in G} \sum_{gh = z} a_g b_h z\right)v = \sum_{z \in G} \sum_{gh \in z} a_g b_h \sigma_z(v) = \sum_{g \in G} \sum_{h \in G} a_g b_h \sigma_g \sigma_h(v)$$
since  $\sigma(gh) = \sigma(g)\sigma(h) = \sigma_g \sigma_h = \sigma_{gh}$  ( $\sigma$  homomorphism)

From Sec. 8.3 "Semisimple Ring" of Rotman (2010) [10]:

**Definition 17.** k-representation of group G is homomorphism

$$\sigma: G \to GL(V)$$

where V is vector field over field k

**Proposition 6** (8.37 Rotman (2010)[10]).  $\forall k$ -representation  $\sigma : G \to GL(V)$  equips V with structure of left kG-module, denote module by  $V^{\sigma}$ .

Conversely,  $\forall$  left kG-module V determines k-representation  $\sigma: G \to GL(V)$ 

Proof. Given 
$$\sigma: G \to GL(V),$$
  
 $\sigma_q =: \sigma(g): V \to V$ 

define

 $v, w \in V$ 

$$kG \times V \to V$$

$$\left(\sum_{g \in G} a_g g\right) v = \sum_{g \in G} a_g \sigma_g(v)$$

$$\begin{aligned} \operatorname{Let} r, r', 1 &\in kG \\ r &= \sum_{g \in G} a_g g \\ r(v+w) &= \left(\sum_{g \in G} a_g g\right) (v+w) = \left(\sum_{g \in G} a_g \sigma_g\right) (v+w) = \sum_{g \in G} a_g \sigma_g(v) + \sum_{g \in G} a_g \sigma_g(w) = rv + rw \\ (r+r')v &= \left(\sum_{g \in G} a_g g + b_g g\right) v = \sum_{g \in G} (a_g \sigma_g + b_g \sigma_g) v = \sum_{g \in G} a_g \sigma_g(v) + \sum_{g \in G} b_g \sigma_g(v) = rv + r'v \\ (rr')v &= \left(\sum_{g \in G} a_g g \sum_{h \in G} b_h h\right) v = \left(\sum_{z \in G} \sum_{gh = z} a_g b_h z\right) v = \sum_{z \in G} \sum_{gh \in z} a_g b_h \sigma_z(v) = \sum_{g \in G} \sum_{h \in G} a_g b_h \sigma_g \sigma_h(v) \end{aligned}$$

since  $\sigma(gh) = \sigma(g)\sigma(h) = \sigma_a\sigma_h = \sigma_{ah}$  ( $\sigma$  homomorphism)

$$1v = \sigma(1)v = 1v = v$$

Conversely, assume V left kG-module.

If  $g \in G$ , then  $v \mapsto gv$  defines  $T_g: V \to V$ .  $T_g$  nonsingular since  $\exists T_g^{-1} = T_{g^{-1}}$ 

Define 
$$\sigma: G \to GL(V)$$
  
 $\sigma: g \mapsto T_g$ 

 $\sigma$  k-representation

$$\sigma(gh) = T_{gh} = T_g T_h = \sigma(g)\sigma(h)$$
  
$$\sigma(gh)(v) = T_{gh}v = ghv = T_g T_h v = \sigma(g)\sigma(h)v \quad \forall v \in V$$

**Proposition 7.** Let group G, let  $\sigma, \tau : G \to GL(V)$  be k-representations, field k. If  $V^{\sigma}, V^{\tau}$  corresponding kG-modules in Prop. 6 (Prop. 8.37 in Rotman (2010) [10]), then  $V^{\sigma} \simeq V^{\tau}$  as kG-modules iff  $\exists$  nonsingular  $\varphi : V \to V$  s.t.

$$\varphi \tau(g) = \sigma(g) \varphi \quad \forall g \in G$$

*Proof.* If  $\varphi: V^{\tau} \to V^{\sigma}$  kG-isomorphism, then  $\varphi: V \to V$  isomorphism s.t.

$$\varphi(\sum a_g gv) = (\sum a_g g)\varphi(v) \quad \forall v \in V, \forall g \in G$$

in 
$$V^{\tau}$$
,  $kG \times V \to V$  in  $V^{\sigma}$ ,  $kG \times V \to V$  scalar multiplication 
$$gv = \tau(g)(v) \qquad gv = \sigma(g)(v)$$
  $\Longrightarrow \forall g \in G, v \in V, \quad \varphi(\tau(g)(v)) = \sigma(g)(\varphi(v))$ 

I think

$$\varphi(gv) = \varphi(\tau(g)(v)) = g\varphi(v) = \sigma(g)\varphi(v)$$
$$\Longrightarrow \varphi\tau(g) = \sigma(g)\varphi \quad \forall g \in G$$

Conversely, if  $\exists$  nonsingular  $\varphi: V \to V$  s.t.  $\varphi \tau(g) = \sigma(g) \varphi \quad \forall g \in G$ 

$$\varphi\tau(g)v = \varphi(\tau(g)v) = \sigma(g)\varphi(v) \quad \forall g \in G, \forall v \in V$$

Consider scalar multiplication

$$\begin{split} kG \times V \to V \\ \sum_{g \in G} a_g g(v) &= \sum_{g \in G} a_g \tau_g(v) \\ \varphi\left(\sum_{g \in G} a_g \tau_g(v)\right) &= \varphi\left(\sum_{g \in G} a_g \tau(g)v\right) = \sum_{g \in G} a_g \sigma(g)\sigma(g)\varphi(v) = \left(\sum_{g \in G} a_g g\right)\varphi(v) \end{split}$$

Admittedly, after this exposition from Rotman (2010) [10], I still didn't understand how kG-modules relate to representation theory and group rings. I turned to Baker (2011) [11], which we'll do right now. Note that I found a lot of links to online resources on representation theory from Khovanov's webpage http://www.math.columbia.edu/~khovanov/resource/.

Note,

**Definition 18.** vector subspace  $W \subseteq V$  is called a

G-submodule, G-subspace, EY: 20150922 "invariant" subspace?

if  $\forall g \in G$ , for representation  $\rho: G \to GL_k(V)$ ,  $\rho_g(w) \in W$ ,  $\forall w \in W$ ,  $\forall g \in G$  i.e. closed under "action of elements of G" with  $\rho_g =: \rho(g): V \to V$ 

Given basis  $\mathbf{v} = \{v_1 \dots v_n\}$  for V,  $\dim_k V = n$ ,  $\forall g \in G$ ,

$$\rho_g v_j = \rho(g) v_j = r_{kj}(g) v_k$$

for, indeed,

$$\rho_g x^j v_j = \rho(g) x^j v_j = x^j \rho(g) v_j = x^j r_{kj}(g) v_k = r_{kj} x^j v_k$$

so that

$$\rho: G \to GL_k(V)$$
$$\rho(g) = [r_{ij}(g)]$$

Example 2.1 (Baker (2011) [11]): Let  $\rho: G \to GL_k(V)$  where  $\dim_k V = 1$ 

$$\forall v \in V, v \neq 0, \forall g \in G, \lambda_g \in k \text{ s.t. } g \cdot v = \rho_g(v) = \lambda_g v$$

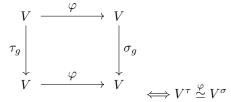
$$\rho(hg)v = \rho_h \rho_g v = \lambda_{hg} v = \lambda_h \lambda_g v \Longrightarrow \lambda_{hg} = \lambda_h \lambda_g$$

 $\Longrightarrow \exists \text{ homomorphism } \Lambda: G \to k^{\times}$ 

$$\Lambda(g) = \lambda_q$$

From Sec. 2.2 "G-homomorphisms and irreducible representations" of Baker (2011) [11], suppose  $\sigma: G \to GL_k(V)$  are representations

Many names for the same thing: G-equivalent, G-linear, G-homomorphism, EY: 20150922 kG-isomorphic? If  $\forall a \in G$ .



Indeed, define

$$\begin{split} \varphi: V^{\tau} &\to V^{\sigma} \\ \varphi(v+w) &= \varphi(v) + \varphi(w) \\ \varphi(rv) &= \varphi(\sum_{g \in G} a_g g \cdot v) = \varphi(\sum_{g \in G} a_g \tau_g(v)) = \sum_{g \in G} a_g \varphi(\tau_g(v)) = \sum_{g \in G} a_g \sigma_g \cdot \varphi(v) = r \varphi(v) \end{split}$$

EY: 20150922 So  $\varphi$  is a kG-isomorphism between left kG modules  $V^{\tau}$  and  $V^{\sigma}$  if it's bijective and is "linear" in "scalars"  $r \in kG$ , i.e.  $\varphi(rv) = r\varphi(v)$ .

Define action of G on  $\operatorname{Hom}_k(V,W)$  ( $\operatorname{Hom}_k(V,W)$  is the vector space of k-linear transformations  $V \to W$ )

$$V \xrightarrow{f} W$$

$$v \longmapsto f(v)$$

Consider

 $\forall f \in \operatorname{Hom}_k(V, W), f : V \to W$ 

 $f(v) \in W$ 

$$G \times \operatorname{Hom}_{k}(V, W) \to \operatorname{Hom}_{k}(V, W)$$
  
$$(g \cdot f) \mapsto (\sigma_{q} f) \circ \rho_{q^{-1}} \text{ i.e. } (g \cdot f)(v) = \sigma_{q} f(\rho_{q^{-1}} v) \quad (f \in \operatorname{Hom}_{k}(V, W))$$

Let  $g, h \in G$ ,

$$(gh \cdot f)(v) = g \cdot \sigma_h f(\rho_{h^{-1}} v) = \sigma_q \sigma_h f(\rho_{h^{-1}} \rho_{g^{-1}} v) = (\sigma_{gh} f(\rho_{gh})^{-1})(v)$$

Thus,  $G \times \operatorname{Hom}_k(V, W) \to \operatorname{Hom}_k(V, W)$  is thus another G-representation of G.

$$(g \cdot f) \mapsto (\sigma_g f) \circ \rho_{g^{-1}}$$

For k-representation  $\rho$ , if the only G-subspaces of V are  $\{0\}$ , V,  $\rho$  irreducible or simple.

$$\rho_g(\{0\}) = \{0\}$$

$$\rho_g(V) = V$$

given subrepresentation  $W \subseteq V$ , V/W admits linear action of G,  $\overline{\rho}_W : G \to GL_k(V/W)$  quotient representation

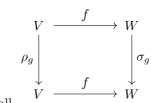
$$\overline{\rho}_W(g)(v+W) = \rho(g)(v) + W$$

if  $v' - v \in W$ 

$$\rho(g)(v') + W = \rho(g)(v + (v' - v)) + W = (\rho(g)(v) + \rho(g)(v' - v)) + W = \rho(g)(v) + W$$

**Proposition 8** (2.7 Baker (2011)[11]). if  $f: V \to W$  G-homomorphism, then

- (a) kerf is G-subspace of V
- (b) imf is G-subspace of W



Proof. Recall

(a) Let  $v \in \ker f$ . Then  $\forall g \in G$ ,

$$f(\rho_q v) = \sigma_q f(v) = 0$$

so  $\rho_g v \in \ker f$ ,  $\forall g \in G$ . So  $\ker f$  is G-subspace of V

(b) Let  $w \in \text{im } f$ . So w = f(u) for some  $u \in V$ 

$$\sigma_a w = \sigma_a f(u) = f(\rho_a u) \in \operatorname{im} f$$

So im f is G-subspace of W

**Theorem 13** (Schur's Lemma). Let  $\rho: G \to GL_{\mathbb{C}}(V)$  be irreducible representations of G over field  $k = \mathbb{C}$ ; let  $f: V \to W$  be  $\sigma: G \to GL_{\mathbb{C}}(W)$ 

G-linear map.

- (a) if  $f \neq 0$ , f isomorphism. True  $\forall k$  field, not just  $\mathbb{C}$
- (b) if V = W,  $\rho = \sigma$ , then for some  $\lambda \in \mathbb{C}$ , f given by  $f(v) = \lambda v$  ( $v \in V$ ) (true for algebraically closed fields)

*Proof.* (a) By Prop. 8,  $\ker f \subseteq V$ ,  $\operatorname{im} f \subseteq W$  are G-subspaces.

For  $\rho$ , only G-subspaces are 0 or V, so if  $\ker f = V$ , f = 0. If  $\ker f = 0$ , f injective.

For  $\sigma$ , only G-subspaces are 0 or V, so  $\operatorname{im} f = 0$ , f = 0. If  $\operatorname{im} f = V$ , f surjective.  $\Longrightarrow f$  isomorphism.

(b) Let  $\lambda \in \mathbb{C}$  be an eigenvalue of f,  $f(v_0) = \lambda v_0$  eigenvector,  $v_0 \neq 0$ .

) Let  $\lambda \in \mathbb{C}$  be an eigenvalue of f,  $f(v_0) = \lambda v_0$  eigenvector,  $v_0 \neq 0$ Let linear  $f_{\lambda}: V \to V$  s.t.

$$f_{\lambda}(v) = f(v) - \lambda v \quad (v \in V)$$

 $\forall g \in G$ 

$$\rho_g f_{\lambda}(v) = \rho_g f(v) - \rho_g \lambda v = f(\rho_g v) - \lambda \rho_g v = f_{\lambda}(\rho_g v)$$

So  $f_{\lambda}$  is G-linear, for

Since  $f_{\lambda}(v_0) = 0$ , by Prop. 8,  $\ker f_{\lambda} = V$ , (for  $\ker f_{\lambda} \neq 0$  and so  $\ker f_{\lambda} = V$ )

By rank-nullity theorem,  $\dim V = \dim \ker f_{\lambda} + \dim \inf f_{\lambda}$ .

So  $\operatorname{im} f_{\lambda} = 0$ , and so  $f_{\lambda}(v) = 0 \ (\forall v \in V) \Longrightarrow f(v) = \lambda v$ 

Schur's lemma, at least the first part, implies that the left kG-modules associated with representations  $\rho, \sigma$  are kG-isomorphic, Proof. i.e.

$$\begin{array}{cccc}
V & \xrightarrow{f} & W \\
\rho_g \downarrow & & \downarrow \sigma_g \\
\downarrow & & \downarrow \sigma_g \\
V & \xrightarrow{f} & W & \iff V^{\rho} \stackrel{f}{\simeq} V^{\sigma}
\end{array}$$

with f being an isomorphism between  $V^{\rho}$  and  $V^{\sigma}$  s.t.

$$f(v+w) = f(v) + f(w) \quad \forall v, w \in (V^{\sigma}, +)$$
$$f(rv) = rf(v) \quad \forall r = \sum_{g \in G} a_g g \in kG$$

Kosmann-Schwarzbach's **Groups and Symmetries**[12] is a very lucid text that's mathematically rigorous enough and practical for physicists. It's really good and very clear. Let's follow its development for SU(2), SO(3),  $SL(2,\mathbb{C})$  and corresponding Lie algebras  $\mathfrak{su}(2)$ ,  $\mathfrak{so}(3)$ ,  $\mathfrak{sl}(2,\mathbb{C})$ .

From Chapter 2 "Representations of Finite Groups" of Kosmann-Schwarzbach (2010) [12]

**Definition 19** (2.1 Kosmann-Schwarzbach (2010)[12]). On  $L^2(G)$ , scalar product defined by

$$\langle f_1|f_2\rangle = \frac{1}{|G|} \sum_{g \in G} \overline{f_1(g)} f_2(g)$$

 $f_1, f_2 \in \mathcal{F}(G) \equiv \mathbb{C}[G]$  vector space of functions on G taking values on  $\mathbb{C}$ 

**Definition 20** (2.3 Kosmann-Schwarzbach (2010)[12]). Let  $(E, \rho)$  be representation of G

character of 
$$\rho \equiv \chi_{\rho} : G \to \mathbb{C}$$
$$\chi_{\rho}(g) = tr(\rho(g)) = \sum_{i=1}^{n} (\rho(g))_{ii}$$

Note: equivalent representations have same character each conjugacy class of G, function  $\chi_n$  is constant

Looking at Def. 19

$$\langle \chi_{\rho_1} | \chi_{\rho_2} \rangle = \frac{1}{|G|} \sum_{g \in G} \chi_{\rho_1}(g^{-1}) \chi_{\rho_2(g)}$$

since  $\overline{\chi_{\rho_1(g)}} = \chi_{\rho_1}(g^{-1})$  by unitarity of representation with respect to scalar product  $\langle , \rangle$ 

**Proposition 9** (2.7 Kosmann-Schwarzbach (2010)[12]). Let  $(E_1, \rho_1)$  be representations of G, let linear  $u: E_1 \to E_2$ .  $(E_2, \rho_2)$ 

Then  $\exists$  linear  $T_u$  s.t.

(22) 
$$T_{u}: E_{1} \to E_{2}$$

$$T_{u} = \frac{1}{|G|} \sum_{g \in G} \rho_{2}(g) u \rho_{1}(g)^{-1}$$

so that  $\rho_2(g)T_u = T_u\rho_1(g) \quad \forall g \in G$ 

$$\rho_2(g)T_u = \frac{1}{|G|} \sum_{h \in G} \rho_2(gh)u\rho_1(h^{-1}) = \frac{1}{|G|} \sum_{k \in G} \rho_2(k)u\rho_1(k^{-1}g) = T_u\rho_1(g)$$

Thus, diagrammatically, we have that

$$E_{1} \xrightarrow{T_{u}} E_{2}$$

$$\downarrow \rho_{1}(g) \qquad \downarrow \rho_{2}(g)$$

$$E_{1} \xrightarrow{u} E_{2} \implies E_{1} \xrightarrow{T_{u}} E_{2}$$

From Definition 1.12 of Kosmann-Schwarzbach [12], "representations  $\rho_1$  and  $\rho_2$  are called **equivalent** if there is a bijective intertwining operator for  $\rho_1$  and  $\rho_2$ ." So I will interpret this as if an intertwining operator is not bijective, then the representations  $\rho_1$ ,  $\rho_2$  are not equivalent.

**Proposition 10** (2.8 Kosmann-Schwarzbach (2010)[12]). Let  $(E_1, \rho_1)$  be irreducible representations of G, let linear  $u: E_1 \to E_2$ ,  $(E_2, \rho_2)$ 

define  $T_u$  by  $T_u = \frac{1}{|G|} \sum_{g \in G} \rho_2(g) u \rho_1(g)^{-1}$  by Eq. 22.

- (i) If  $\rho_1$ ,  $\rho_2$  inequivalent, then  $T_u = 0$
- (ii) If  $E_1 = E_2 = E$  and  $\rho_1 = \rho_2 = \rho$ , then

$$T_u = \frac{tr(u)}{dimE} 1_E$$

Proof. (i) if  $\rho_1, \rho_2$  are inequivalent, by definition,  $T_u$  is not isomorphic. Then by Schur's lemma (first part),  $T_u = 0$  (ii) By Schur's lemma,  $T_u(v) = \lambda v \quad \forall v \in E = E_1 = E_2$ . So  $T_u = \lambda 1_E$ .  $\operatorname{tr} T_u = \lambda \operatorname{dim} E$  or  $\lambda = \frac{\operatorname{tr} T_u}{\operatorname{dim} E}$ . Thus,  $T_u = \frac{\operatorname{tr} T_u}{\operatorname{dim} E} 1_E$ 

Let  $(e_1 \dots e_n)$  basis of E $(f_1 \dots f_p)$  basis of F

$$\forall u \in \mathcal{L}(E, F), \begin{array}{l} u : E \to F \\ u(x) = u(x^j e_j) = x^j u(e_j) = x^j u^i_{\ j} f_i \end{array} \text{ for } x = x^j e_j \in E \\ u = u^i_{\ i} e^j \otimes f_i \qquad \qquad y = y^i f_i \in F \end{array}$$

For

$$T: E^* \otimes F \to \mathcal{L}(E, F)$$

$$T(\xi \otimes y) = u^i{}_j e^j \otimes f_i \text{ i.e. set } T(\xi \otimes y) \text{ to this } u$$

$$T(\xi \otimes y) = T(\xi_l e^l \otimes y^k f_k) = \xi_l y^k T(e^l \otimes f_k) = (\xi_l y^k T^{li}_{kj}) e^j \otimes f_i \Longrightarrow \xi_l y^k T^{li}_{kj} = u^i{}_j$$

Exercises. Exercises of Ch. 2 Representations of Finite Groups [12]

Exercise 2.6. [12] The dual representation.

Let  $(E, \pi)$  representation of group G.

 $\forall g \in G, \xi \in E^*, x \in E, \text{ set } \langle \pi^*(g)(\xi), x \rangle = \langle \xi, \pi(g^{-1})(x) \rangle$ 

(a) dual (or contragredient) of  $\pi$ ,  $\pi^*: G \to \text{End}(E^*)$ ,  $\pi^*$  is a representation, since

$$\langle \pi^*(gh)(\xi), x \rangle = \langle \xi, \pi((gh)^{-1})(x) \rangle = \langle \xi, \pi(h^{-1}g^{-1})(x) \rangle = \langle \xi, \pi(h^{-1})\pi(g^{-1})(x) \rangle = \langle \xi, \pi(h^{-1})(\pi(g^{-1})(x)) \rangle = \langle \pi^*(h)(\xi), \pi(g^{-1})(x) \rangle = \langle \pi^*(g)\pi^*(h)(\xi), x \rangle$$

since this is true,  $\forall x \in E, \forall \xi \in E^*, \pi^*(gh) = \pi^*(g)\pi^*(h)$ . dual  $\pi^*$  of  $\pi$  is a representation.

(b) Consider  $G \times \mathcal{L}(E, F) \to \mathcal{L}(E, F)$ .

$$g \cdot u = \rho(g) \circ u \circ \pi(g^{-1})$$

Define

$$\sigma: G \to \operatorname{End}(\mathcal{L}(E, F))$$
  
$$\sigma(g): \mathcal{L}(E, F) \to \mathcal{L}(E, F)$$
  
$$\sigma(g)(u) = \rho(g) \circ u \circ \pi(g^{-1})$$

Let  $(e_1 \dots e_n)$  be a basis of E. Let  $\xi = \xi_i e^i \in E^*$ ,  $x = x^j e_i \in E$ . Consider the isomorphism  $T: E^* \otimes F \to \mathcal{L}(E,F)$  defined as<sup>2</sup>

$$T: E^* \otimes F \to \mathcal{L}(E, F) = \operatorname{Hom}(E, F)$$
  
$$\xi \otimes y \mapsto (x \mapsto \xi(x)y)$$

Choose bases 
$$(e_1 \dots e_n)$$
 of  $E$   
 $(e^1 \dots e^n)$  of  $E^*$ . Then
$$(f_1 \dots f_p) \text{ of } F$$

$$T(e^j \otimes f_i)(x) = T(e^j \otimes f_i)(x^k e_k) = \delta^j_{\ k} x^k f_i = x^j f_i$$

$$T(e^j \otimes f_i)(e_k) = \delta^j_{\ k} f_i$$

Consider

$$u \in \mathcal{L}(E, F)$$

$$u : E \to F$$

$$u(x) = u(x^{j}e_{j}) = x^{j}u(e_{j}) = x^{j}u_{j}^{i}f_{i}$$

$$u(e_{j}) = u_{i}^{i}f_{i} \text{ i.e. } u : e_{j} \to u_{i}^{i}f_{i}$$

Then  $\forall u \in \mathcal{L}(E, F)$ ,

$$T(u^i_{i}e^j \otimes f_i)(e_k) = u^i_{i}\delta^j_{k}f_i = u^i_{k}f_i = u(e_k) \Longrightarrow u = T(u^i_{i}e^j \otimes f_i)$$

so T is surjective.

With  $T(\xi \otimes y) = T(\xi' \otimes y')$ ,

$$T(\xi \otimes y)(x) = T(\xi' \otimes y')(x)$$
  
$$\xi(x)y = \xi'(x)y' \Longrightarrow \xi(x)y - \xi'(x)y' = 0$$

For  $f \in P^{(k)}(V)$ , the general form is

which implies that  $\xi \otimes y = \xi' \otimes y'$ . So T is injective. Or, one could consider that  $T^{-1}: \mathcal{L}(E,F) \to E^* \otimes F$ ,  $T^{-1}: u \mapsto u^i_{\ i}e^j \otimes f_i$ , which is the inverse of T.

### Remark 1.

$$E^* \otimes F \stackrel{T}{\simeq} \mathcal{L}(E, F) = Hom(E, F)$$
  
 $(\xi, y) \mapsto (x \mapsto \xi(x)y)$ 

and so  $(e^j \otimes f_i) \mapsto (x \mapsto e^j(x) f_i = x^j f_i)$ So  $E^* \otimes F$  is isomorphic to  $\mathcal{L}(E,F) = Hom(E,F)$ 

For representation  $\pi$ ,

$$\pi: G \to \operatorname{End}(E)$$

$$\pi(g): E \to E$$

$$\pi(g)(x) = \pi(g)(x^j e_j) = x^j \pi(g)(e_j) = x^j \pi(g)^i{}_i e_i = (\pi(g)^i{}_i x^j e_i$$

Consider this matrix formulation:

$$\pi^*(g)(\xi) = \pi^*(g)(\xi_i e^i) = \xi_i \pi^*(g)(e^i) = \xi_i (\pi^*(g))^i{}_j e^j$$
$$\Longrightarrow \langle \pi^*(g)(\xi), x \rangle = \xi_i (\pi^*(g))^i{}_j x^j$$

and

$$\langle \xi, \pi(g^{-1})(x) \rangle = \xi_i \pi(g^{-1})^i{}_i x^j$$

so that

$$\langle \pi^*(g)(\xi), x \rangle = \langle \xi, \pi(g^{-1})(x) \rangle \Longrightarrow \pi(g^{-1})^i_{\ j} = (\pi^*(g))^i_{\ j}$$

Thus, given a choice of basis for E, the dual of  $\pi$ ,  $\pi^*(g)^i{}_i$ , and  $\pi(g^{-1})^i{}_i$  are formally equal. So for a choice of basis of E and of F,

 $(\pi^* \otimes \rho)(g)(\xi, y) = (\pi^*(g) \otimes \rho(g))(\xi, y) = \pi^*(g)\xi \otimes \rho(g)y = \xi_l \pi(g^{-1})^l{}_i e^j \otimes \rho(g)^i{}_k y^k f_i = \rho(g)^i{}_k y^k \xi_l \pi(g^{-1})^l{}_i e^j \otimes f_i$ Applying T,

$$T(\pi^* \otimes \rho)(g)(\xi, \rho) = \rho(g)^i{}_k y^k \xi_l \pi(g^{-1})^l{}_i = \rho(g) T(\xi, y) \pi(g^{-1})$$

$$E^* \otimes F \xrightarrow{T} \mathcal{L}(E, F)$$

$$(\pi^* \otimes \rho)(g) \downarrow \qquad \qquad \downarrow \sigma(g)$$

$$E^* \otimes F \xrightarrow{T} \mathcal{L}(E, F)$$

$$(\xi, y) \longmapsto \xrightarrow{T} (x \mapsto \xi(x)y) = y^i \xi_j$$

$$(\pi^* \otimes \rho)(g) \downarrow \qquad \qquad \downarrow \sigma(g)$$

$$\pi^*(g)(\xi) \otimes \rho(g)y \longmapsto_{T} \rho(g)y^i \xi_j \pi(g^{-1}) = \rho(g)T(\xi, y)\pi(g^{-1})$$

Thus

Thus, representation  $\sigma(g)$  is equivalent to representation  $(\pi^* \otimes \rho)$ , a tensor product of representations.

**Exercise 2.15.** Representation of  $GL(2,\mathbb{C})$  on the polynomials of degree 2

Let group G, let representation  $\rho$  of G on  $V = \mathbb{C}^n$ , i.e.  $\rho: G \to \operatorname{End}(V)$ 

Let  $P^{(k)}(V)$  vector space of complex polynomials on V that are homogeneous of degree k.

$$f = \sum_{\substack{i_1 + i_2 + \dots + i_n = k \\ 0 \le i_j \le k}} a_{i_1 i_2 \dots i_d} x_1^{i_1} x_2^{i_2} \dots x_n^{i_n}$$

http://math.stackexchange.com/questions/57189/understanding-isomorphic-equivalences-of-tensor-product

<sup>&</sup>lt;sup>2</sup>Mathematics stackexchage Isomorphism between Hom and tensor product [duplicate] http://math.stackexchange.com/questions/428185/isomorphism-between-hom-and-tensor-product

Given

$$\binom{n+k}{k} = \binom{k-1}{k-1} + \binom{k}{k-1} + \dots + \binom{n+k-1}{k-1} = \sum_{i=0}^{n} \binom{k-1+i}{k-1}$$

 $\binom{k+n-1}{n-1}$  is number of monomials of degree k.

So  $\dim P^{(k)}(V) = \binom{k+n-1}{n-1}$ . This is a very lucid and elementary exposition on the basics of polynomials which I found was useful for the basic facts I forgot<sup>3</sup>.

So we have the graded algebra

$$P(V) = \bigoplus_{k=0}^{\infty} P^{(k)}(V)$$

$$\rho^{(k)} : G \to \operatorname{End}(P^{(k)}(V))$$

$$\rho^{(k)}(g) : P^{(k)}(V) \to P^{(k)}(V)$$

$$\rho^{(k)}(g)(f) = f \circ \rho(g^{-1})$$

This is a representation of G since

(a)

$$\rho^{(k)}(gh)(f) = f \circ \rho((gh)^{-1}) = f \circ \rho(h^{-1}g^{-1}) = f \circ \rho(h^{-1}\rho(g^{-1})) = \rho^{(k)}(gh)(f) = \rho^{(k)$$

(b) Choose basis  $(e_1 
ldots e_n)$  of V,  $x = x^j e_j \in V$ ,  $\rho : G \to \text{End}(V)$ , and so  $\rho(g)(x) = \rho(g)(x^j e_j) = x^j \rho(g)(e_j) = x^j (\rho(g))^i{}_j e_i$ . With  $\xi(e_i) = \xi_i \Longrightarrow \langle \xi, \rho(g^{-1})x \rangle = \xi_i x^j (\rho(g^{-1}))^i{}_j$  $\forall \xi \in V^*, \ \xi = \xi_i e^i$ .

$$\rho^{*}(g)(\xi) = \rho^{*}(g)(\xi_{i}e^{i}) = \xi_{i}\rho^{*}(g)^{i}{}_{j}e^{j}$$

$$\Longrightarrow \langle \rho^{*}(g)(\xi), x \rangle = \xi_{i}x^{j}(\rho^{*}(g))^{i}{}_{j} \Longrightarrow (\rho^{*}(g))^{i}{}_{j} = (\rho(g^{-1}))^{i}{}_{j}$$
So  $\forall f \in P^{(1)}(V), x \in V, \rho(g^{-1})x = x^{j}(\rho(g^{-1}))^{i}{}_{j}e_{i}$ . So  $f \circ \rho(g^{-1})(x) = \sum_{i=1}^{n} a_{i}(\rho(g^{-1}))^{i}{}_{j}x^{j} = \sum_{i=1}^{n} a_{i}(\rho^{*}(g))^{i}{}_{j}x^{j}$ 

$$\Longrightarrow \rho^{(1)}(g)(f) = f \circ \rho^{*}(g)$$

(c) Suppose  $G = GL(2, \mathbb{C})$ ,  $V = \mathbb{C}^2$ ,  $\rho$  fundamental representation  $g = \begin{pmatrix} a & b \\ c & d \end{pmatrix}$ ,  $g^{-1} = \frac{1}{\det g} \begin{pmatrix} d & -b \\ -c & a \end{pmatrix}$  for  $\det g = ad - bc$ . Let k = 2,  $\dim P^{(2)}(\mathbb{C}^2) = \binom{2+2-1}{2-1} = \binom{3}{1} = 3$   $\forall f \in P^{(2)}(\mathbb{C}^2)$ ,  $f(x,y) = Ax^2 + 2Bxy + Cy^2$ 

$$P^{(2)}(\mathbb{C}^2) \to \mathbb{C}^3$$
$$f(x,y) = Ax^2 + 2Bxy + Cy^2 \mapsto \begin{pmatrix} A \\ B \\ C \end{pmatrix} \in \mathbb{C}^3$$

Call this transformation  $T, T: P^{(2)}(\mathbb{C}^2) \to \mathbb{C}^3$ .

$$\forall \begin{pmatrix} A \\ B \\ C \end{pmatrix} \in \mathbb{C}^3, \ f(x,y) = Ax^2 + 2Bxy + Cy^2 \text{ and } Tf(x,y) = \begin{pmatrix} A \\ B \\ C \end{pmatrix}. \ T \text{ surjective.}$$

Suppose Tf(x,y) = Tf'(x,y),

$$\implies Ax^2 + 2Bxy + Cy^2 = A'x^2 + 2B'xy + C'y^2 \implies (A - A')x^2 + 2(B - B')xy + (C - C')y^2 = 0$$

Then since the monomials form a basis, and its basis elements are independent (by definition), then A = A', B = B', C = C'. T injective. So T is bijective, an isomorphism.

(This is all in groups.sage)

```
sage: P2CC.<x,y> = PolynomialRing(CC,2) # this declares a PolynomialRing of field of complex numbers,
# of order 2 (i.e. only 2 variables for a polynomial, such as x, y)
sage: A = var('A')
sage: assume(A, ''complex'')
sage: B = var('B')
sage: assume(B, ''complex'')
sage: C = var('C')
     assume (C, ''complex'')
sage: f(x,y) = A*x**2 + 2*B*x*y + C*y**2
sage: a = var('a')
sage: assume(a, ''complex'')
     b = var('b')
      assume(b, ''complex'')
      assume(c.''complex
     d = var('d')
     assume(d, ''complex'')
     g = Matrix([[a,b],[c,d]]
sage: X = Matrix([[x],[y]])
     f((g.inverse()*X)[0,0], (g.inverse()*X)[1,0]).expand()
sage: f((g.inverse()*X)[0,0], (g.inverse()*X)[1,0]).expand().coefficient(x^2).full_simplify()
(C*c^2 - 2*B*c*d + A*d^2)/(b^2*c^2 - 2*a*b*c*d + a^2*d^2)
sage: f((g.inverse()*X)[0,0], (g.inverse()*X)[1,0]).expand().coefficient(x*y).full_simplify()
-2*(C*a*c + A*b*d - (b*c + a*d)*B)/(b^2*c^2 - 2*a*b*c*d + a^2*d^2)
sage: f( (g.inverse()*X)[0,0], (g.inverse()*X)[1,0] ).expand().coefficient(v^2).full_simplify()
(C*a^2 - 2*B*a*b + A*b^2)/(b^2*c^2 - 2*a*b*c*d + a^2*d^2)
```

So

$$\rho^{(2)}(g)(f)(x,y) = f \circ \rho(g^{-1})(x,y) =$$

$$= \frac{Cc^2 - 2Bcd + Ad^2}{(ad - bc)^2}x^2 + -2\frac{(Cac + Abd - (bc + ad)B)}{(ad - bc)^2}xy + \frac{Ca^2 - 2Bab + Ab^2}{(ad - bc)^2}y^2$$

So define  $\widetilde{\rho}: G \to \operatorname{End}(\mathbb{C}^3)$ .  $\widetilde{\rho}$  is a representation, for

$$\forall v = \begin{pmatrix} A \\ B \\ C \end{pmatrix} \in \mathbb{C}^3, \quad \widetilde{\rho}(gh)(v) = T \circ f \circ \rho((gh)^{-1}) = T \circ f \circ \rho(h^{-1}g^{-1}) = T \circ f \circ \rho(h^{-1})\rho(g^{-1})$$

$$\text{Now } \widetilde{\rho}(h)(v) = T \circ f \circ \rho(h^{-1})$$

$$\Longrightarrow \widetilde{\rho}(g)\widetilde{\rho}(h)(v) = T \circ (f \circ \rho(h^{-1})) \circ \rho(g^{-1}) = T \circ f \circ \rho(h^{-1})\rho(g^{-1}) \text{ and so}$$

$$\widetilde{\rho}(gh) = \widetilde{\rho}(g)\widetilde{\rho}(h)$$

And so

$$\widetilde{\rho}^*(g)(v) = Tf\rho(g^{-1})$$

and consider this commutation diagram, that (helped me at least and) clarifies the relationships:

<sup>&</sup>lt;sup>3</sup>Polynomials. Math 4800/6080 Project Course http://www.math.utah.edu/~bertram/4800/PolyIntroduction.pdf

$$P^{(2)}(\mathbb{C}^2) \xrightarrow{T} \mathbb{C}^3$$

$$\rho^{(2)}(g) \downarrow \qquad \qquad \downarrow \widetilde{\rho}(g)$$

$$P^{(2)}(\mathbb{C}^2) \xrightarrow{T} \mathbb{C}^3$$

$$\begin{array}{ccc}
f & & T & \longrightarrow \begin{pmatrix} A \\ B \\ C \end{pmatrix} \\
\downarrow & & \downarrow & \downarrow \\
f \circ \rho(g^{-1}) & & \longrightarrow & \begin{pmatrix} D \\ E \\ F \end{pmatrix}
\end{array}$$

with

$$\begin{pmatrix} D \\ E \\ F \end{pmatrix} = \begin{pmatrix} \frac{Cc^2 - 2Bcd + Ad^2}{(ad - bc)^2} \\ -2\frac{(Cac + Abd - (bc + ad)B)}{(ad - bc)^2} \\ \frac{Ca^2 - 2Bab + Ab^2}{(ad - bc)^2} \end{pmatrix}$$

Now define the dual  $\tilde{\rho}^*$  as such:

$$\begin{split} \widetilde{\rho}^*(g): (\mathbb{C}^3)^* &\to (\mathbb{C}^3)^* \\ \widetilde{\rho}^*(g) &= \widetilde{\rho}(g^{-1}) \\ &\forall \, \xi \in (\mathbb{C}^3)^* \\ \widetilde{\rho}^*(g) \xi &= \xi_i (\widetilde{\rho}^*(g))^i{}_j e^j = \xi_i (\widetilde{\rho}(g^{-1}))^i{}_j e^j \end{split}$$

So for 
$$v = \begin{pmatrix} A \\ B \\ C \end{pmatrix} \in \mathbb{C}^3$$
,  $f = T^{-1}v = Ax^2 + 2Bxy + Cy^2 \in P^2(\mathbb{C}^2)$ ,

$$\widetilde{\rho}(g^{-1})(v) = T \circ (f\rho(g)) = \begin{bmatrix} Aa^2 + 2Bac + Cc^2 \\ Aab + Bbc + Bad + Ccd \\ Ab^2 + 2Bbd + Cd^2 \end{bmatrix}$$

which was found using Sage Math:

```
sage: f((g*X)[0,0],(g*X)[1,0])
(a*x + b*y)^2*A + 2*(a*x + b*y)*(c*x + d*y)*B + (c*x + d*y)^2*C
sage: f((g*X)[0,0],(g*X)[1,0]).expand()
sage: f((g*X)[0,0],(g*X)[1,0]). expand().coefficient(x<sup>2</sup>)
A*a^2 + 2*B*a*c + C*c^2
sage: f((g*X)[0,0],(g*X)[1,0]).expand().coefficient(x*y)
2*A*a*b + 2*B*b*c + 2*B*a*d + 2*C*c*d
sage: f((g*X)[0,0],(g*X)[1,0]).expand().coefficient(y^2)
A*b^2 + 2*B*b*d + C*d^2
```

or

sage: T( f((g\*X)[0,0],(g\*X)[1,0]).expand() )
[A\*a^2 + 2\*B\*a\*c + C\*c^2,
2\*A\*a\*b + 2\*B\*b\*c + 2\*B\*a\*d + 2\*C\*c\*d,
$$A*b^2 + 2*B*b*d + C*d^2$$
]

So then

$$\widetilde{\rho}(g^{-1}) = \begin{bmatrix} a^2 & 2ac & c^2 \\ 2ab & 2(ad+bc) & 2cd \\ b^2 & 2bd & d^2 \end{bmatrix}$$

So then

$$\widetilde{\rho}^*(g) = \begin{bmatrix} a^2 & 2ac & c^2 \\ 2ab & 2(ad+bc) & 2cd \\ b^2 & 2bd & d^2 \end{bmatrix}$$

and operate on row vectors  $\xi \in (\mathbb{C}^3)^*$  with  $\widetilde{\rho}^*(g)$  from the row vector's right.

More: Let 
$$G = SU(2)$$
. Then  $U = e^{i\phi} \begin{bmatrix} a & b \\ -\overline{b} & \overline{a} \end{bmatrix}$ 

$$\widetilde{\rho}: SU(2) \to \operatorname{End}(\mathbb{C}^3)$$

$$\widetilde{\rho}(U): \mathbb{C}^3 \to \mathbb{C}^3$$

$$\widetilde{\rho}(U)(v) = e^{-2i\varphi} \begin{bmatrix} A\overline{a}^2 + 2B\overline{a}\overline{b} + C\overline{b}^2 \\ -A\overline{a}b + B + Ca\overline{b} \\ Ab^2 - 2Bab + Ca^2 \end{bmatrix}$$

$$\Longrightarrow \widetilde{\rho}(U) = e^{-2i\varphi} \begin{bmatrix} -\overline{a}^2 & 2\overline{a}\overline{b} & \overline{b}^2 \\ -\overline{a}b & 1 & a\overline{b} \\ b^2 & -2ab & a^2 \end{bmatrix}$$

cf. Ch. 5 Lie Groups of Jeffrey Lee (2009) [2]

**Definition 21** (Lie Group). Lie Group  $G := smooth \ manifold \ G$  is a Lie Group if G is a group (abstract group), s.t.

multiplication map  $\mu: G \times G \to G$  $\mu(g,h) = gh$ 

inverse map  $inv: G \to G$ 

$$inv(q) = q^{-1}$$

are  $C^{\infty}$  maps.

If group is abelian, use additive notation q + h for group operation.

**Definition 22**  $(GL(n,\mathbb{R}))$ .  $GL(n,\mathbb{R}) := group \ of \ all \ invertible \ real \ n \times n \ matrices.$ global chart on  $GL(n,\mathbb{R}) = \{x_i^i\}$ ,  $n^2$  functions  $x_i^i$ , where if  $A \in GL(n,\mathbb{R})$ , then  $x_i^i(A)$  is if the entry of A.

Claim:  $GL(n,\mathbb{R})$  is a Lie group.

$$\frac{\partial}{\partial x_{-}^{l}}(x_{k}^{i}(A)x_{j}^{k}(B)) = \delta_{l}^{i}\delta_{k}^{m}x_{j}^{k}(B) + x_{k}^{i}(A)\delta_{l}^{k}\delta_{j}^{m}$$

inversion map; appeal to formula for  $A^{-1}$ ,  $A^{-1} = \text{adj}(A)/\text{det}(A)$ ,  $\text{adj}(A) \equiv \text{adjoint matrix}$  (whose entries are cofactors)  $\implies A^{-1}$  depends smoothly on entries of A.

Similarly,  $GL(n,\mathbb{C})$ , group of invertible  $n \times n$  complex matrices, is a Lie group.

**Exercise 5.5.** Let subgroup H of G, consider cosets gH,  $g \in G$ .

Recall G is disjoint union of cosets of H.

Claim: if H open, so are all its cosets. And H closed.

*Proof.* cf. stackexchange: Open subgroups of a topological group are closed

 $gH = \{gh|h \in H\}$  is an open neighborhood of g (since  $1 \in H$ , and mapping  $h \mapsto gh$  sends open sets to open sets, since its inverse,  $qh \mapsto h$ , is  $C^{\infty}$  (so continuous)).

$$gH \to H$$
  $H \to gH$  
$$gh \xrightarrow{g^{-1}} h = \mu(g^{-1}, gh) \qquad h \xrightarrow{g} gh = \mu(g, h)$$

Then  $\forall$  coset qH, qH is open.

Suppose  $g' \in H^c \equiv G - H \equiv G \backslash H$ .

Consider  $h \in H$ , if  $g'h \in H$ , then  $g' = (g'h)h^{-1} \in H$  (recall  $h^{-1} \in H$ , and H is a subgroup).

Contradiction.

 $\Longrightarrow \forall g' \in H^c$ ,  $\exists$  open neighborhood  $g'H \subset H^c$ , so  $H^c$  open (by definition). Then H closed.

cf. Thm. 5.6 in Jeffrey Lee (2009) [2].

**Theorem 14.** If G connected Lie group, U neighborhood of identity element e, then U generates the group, i.e.  $\forall g \in G$ , g is a product of elements of U.

*Proof.* Note  $V = \text{inv}(U) \cap U$  is an open neighborhood of e. Note inv(V) = V.  $\text{inv}(V) \equiv V^{-1} = \{V^{-1} | v \in V\}$ . We say that V is symmetric.

Claim: V generates G.

 $\forall$  open  $W_1$ , open  $W_2 \subset G$ ,

 $W_1W_2 = \{w_1w_2 | w_1 \in W_1, w_2 \in W_2\}$  is an open set being a union of open sets  $\bigcup_{g \in W_1} gW_2$ .

Thus, inductively defined sets

$$V^n = VV^{n-1}, \quad n = 1, 2, 3, \dots$$

are open.

$$e \in V \subset V^2 \subset \dots V^n \subset \dots$$

It's easy to check that each  $V^n$  is symmetric.

$$\operatorname{inv}(V) = V$$

$$\operatorname{inv}(V^2) = \operatorname{inv}(\bigcup_{v \in V} vV) = V \operatorname{inv}(V) = V = V^2$$

$$\operatorname{inv}(V^{n+1}) = \operatorname{inv}\left(\bigcup_{v \in V} vV^n\right) = V \operatorname{inv}(V^n) = VV^n = V^{n+1}$$

so  $V^{\infty} := \bigcup_{n=1}^{\infty} V^n$  is symmetric.

 $V^{\infty}$  closed under inversion, also multiplication. Thus  $V^{\infty}$  is an open subgroup.

From Exercise 5.5, Jeffrey Lee (2009) [2], i.e. Exercise 8,  $V^{\infty}$  also closed, since G is connected,  $V^{\infty} = G$ . (a topological space X is **connected** iff the only open and closed (clopen) sets are  $\emptyset$  and X).

**Definition 23.** Identity component of  $G, G_0$ .

 $G_0 := connected component of Lie group G that contains identity;$ 

 $G_0$  is a Lie group, and is generated by any open neighborhood of the identity.

**Definition 24.** For Lie group G, fixed element  $g \in G$ ,

left translation (by g)  $L_g: G \to G$ ,  $L_g x = g x$ ,  $\forall x \in G$ right translation (by g)  $R_g: G \to G$ ,  $R_g x = x g$ ,  $\forall x \in G$ 

 $L_a$ ,  $R_a$  are diffeomorphisms with  $L_a^{-1} = L_{a^{-1}}$ ,  $R_a^{-1} = R_{a^{-1}}$ .

**Definition 25** (Product Lie group). If G, H are Lie groups, then product manifold  $G \times H$  is a Lie group, where multiplication

$$(g_1, h_1) \cdot (g_2, h_2) = (g_1g_2, h_1h_2)$$

Lie group  $G \times H$  is called **product Lie group** 

e.g. product group  $S^1 \times S^1 \equiv 2$ -torus group.

Generally, higher torus groups  $T^n = S^1 \times \cdots \times S^1$  (n factors).

**Definition 26** (Lie subgroup of G, H). Let H be an abstract subgroup of Lie group G.

If H is a Lie group s.t. inclusion map  $i: H \to G \equiv H \hookleftarrow G$  is an immersion, then H is a **Lie subgroup** of G.

Recall  $i: H \to G$  immersion iff Di injective, i.e. iff  $\operatorname{rank} Di = \dim H$  cf. Prop. 5.9 in Jeffrey Lee (2009) [2].

□ **Proposition 11.** If H abstract subgroup of Lie group G, that's also a regular submanifold  $\equiv$  embedded submanifold, then H closed Lie subgroup.

Recall that

 $embedded\ submanifold\ \equiv\ regular\ submanifold$ 

Each name is used frequently and we shouldn't be biased against one or the other; we'll have to refer to both, to emphasize they're exactly the same.

embedded submanifold  $\equiv$  regular submanifold is an immersed submanifold s.t. inclusion map i is a topological embedding, i.e. embedded submanifold  $\equiv$  regular submanifold  $S \subset M$ ,

immersed submanifold S if  $i: S \to M \equiv S \hookrightarrow M$  is an immersion, i.e. Di injective, i.e. rank $Di \equiv \dim S$ .

topological embedding := homeomorphism onto its image, i.e.

injective cont. map  $f: X \to Y, X, Y$  topological spaces, is a **topological embedding** 

if f is a homeomorphism between X and f(X).

f homeomorphism is a bijection, continuous, and  $f^{-1}$  continuous.

e.g.  $\forall$  embedding  $f: M \to N$ ,  $f(M) \subset N$  naturally has the structure of an embedding submanifold  $\equiv$  regular submanifold. Useful, intrinsic definition of **embedded submanifold**  $\equiv$  regular submanifold.

Let manifold M, dimM = n, let  $k \in \mathbb{Z}^+$ , s.t. 0 < k < n.

A k-dim. embedded submanifold  $\equiv$  regular submanifold S is subset  $S \subset M$  s.t.  $\forall p \in S, \exists$  chart  $(U \subset M, \varphi : U \to \mathbb{R}^n \ni 0)$ , s.t.  $\varphi(S \cap U)$  is the intersection of a k-dim. plane with  $\varphi(U)$ .

(pairs  $(S \cap U, \varphi|_{S \cap U})$  form an atlas for differential structure on S)

Proof 1:

*Proof.* H subgroup of G, so

multiplication map  $H \times H \to H$ 

inversion map  $H \to H$ 

are restrictions of multiplication and inversion maps on G.

 $\square$  Since H regular submanifold, maps are smooth.

Recall H regular submanifold iff H immersive submanifold (i.e.  $H \leftarrow G$  is an immersion) and H topological subspace of G, i.e. submanifold topology on H is same as subspace topology.

Claim: H closed.

Let  $x_0 \in \overline{H}$ 

Let (U, x) be a chart adapted to H, whose domain contains e.

Let

$$\delta:G\times G\to G$$

$$\delta(g_1, g_2) = g_1^{-1} g_2$$

Choose open set V s.t.  $e \in V \subset \overline{V} \subset U$ .

By continuity map  $\delta$ , find open neighborhood O of identity e.s.t.  $O \times O \subset \delta^{-1}(V)$ 

If  $\{h_i\}$  sequence in H converging to  $x_0 \in \overline{H}$ , then  $x_0^{-1}h_i \to e$  and  $x_0^{-1}h_i \in O$  for all sufficiently large i.

Since  $h_i^{-1}h_i = (x_0^{-1}h_j)^{-1}x_0^{-1}h_i, h_i^{-1}h_i \in V$  for sufficiently large i, j.

For any sufficiently large fixed j,

$$\lim_{i \to 0} h_j^{-1} h_i = h_j^{-1} x_0 \in \overline{V} \subset U$$

Since U is domain of a single-slice chart,  $U \subset H$  closed in U.

Thus, since  $\forall h_j^{-1}h_i \in U \cap H$ ,  $h_j^{-1}x_0 \in U \cap H \subset H$ ,  $\forall$  sufficiently large j.  $\Longrightarrow x_0 \in H$ , and since  $x_0$  arbitrary, done.

Proof 2:

cf. 9.2 The Closed Subgroup Theorem I of 427 Notes<sup>4</sup>

*Proof.* Claim: Since H is an embedded submanifold  $\equiv$  regular submanifold,  $\exists$  neighborhood U of  $1, 1 \in G$ , s.t.  $U \cap H$  closed in U.

Let  $x_0 \in \overline{H}$ ,  $\overline{H} \equiv$  closure of  $x_0$ .

Then  $x_0U^{-1}\subseteq G$  is a neighborhood of  $x_0$  in G (since  $1\in U^{-1}$ ,  $x_01=x_0\in x_0U^{-1}$ )

$$\Longrightarrow x_0 U^{-1} \cap H \neq \emptyset$$

 $\forall x \in x_0 U^{-1} \cap H, \ x = x_0 U^{-1} \text{ for some } u \in U. \text{ Thus, } x^{-1} x_0 = u \in U.$ 

Now

 $L_{x^{-1}}: G \to G$  is a homeomorphism, so  $L_{x^{-1}}(H) = H$ . By continuity,  $L_{x^{-1}}(\overline{H}) = \overline{H}$ . Thus  $x^{-1}x_0 \in \overline{H}$ .

Claim:  $x^{-1}x_0 \in H \cap U$ .

Since  $x^{-1}x_0 \in \overline{H} \cap U$ ,  $\exists$  sequence  $\{h_i\} \subset H \cap U$  s.t.  $h_0 \to x^{-1}x_0$ .

But recall  $H \cap U$  closed in U, so  $x^{-1}x_0 \in H \cap U$ .

$$\implies x_0 \in xH = H, \quad \overline{H} \subseteq H$$

Thus H closed.

Claim: If H abstract subgroup of Lie group G, that's also an embedded submanifold  $\equiv$  regular submanifold, then H is a Lie subgroup.

Recall that by definition, Lie group has group multiplication and inverse map to be  $C^{\infty}$ . Then, just show group multiplication is  $C^{\infty}$ , first.

Since G is a Lie group, then

$$\mu: G \times G \to G$$

$$\mu(x,y) = xy$$

is  $C^{\infty}$  (by definition).

Then  $\mu: G \times G \to G$  cont.

Consider subgroup  $H \subseteq G$  and  $\mu: H \times H \to H$ .

Since  $H \times H \subseteq G \times G$ ,  $\forall (x, y) \in H \times H$  (fix  $(x, y) \in H \times H$ ),  $\forall$  neighborhood V of  $\mu(x, y) = xy$ ,  $V \subset G$ ,  $\exists$  neighborhood U of (x, y) s.t.  $\mu(U) \subseteq V$  (by  $\mu: G \times G \to G$  cont.).

Since H embedded submanifold  $\equiv$  regular submanifold of G,

 $\exists$  neighborhood  $V' \subseteq V$  of  $xy \in G$ , coordinate map  $\varphi : V' \to \mathbb{R}^n$   $(n = \dim G)$  s.t.

$$\varphi(H \cap V') = \varphi(V') \cap (\mathbb{R}^k \times \{0\})$$

where  $k = \dim H$ 

(since H is a k-dim. embedded submanifold  $\equiv$  regular submanifold,  $H \subseteq G$ , s.t.  $\forall p \in H$ ,  $\exists$  chart  $(V \subset G, \varphi : U \to \mathbb{R}^n \ni 0)$ , s.t.  $\varphi(U \cap V) = \varphi(V) \cap (\mathbb{R}^k \times \{0\})$ ).

Now

$$\varphi \circ \mu : \mu^{-1}(V') \cap U \to \mathbb{R}^n \text{ is } C^{\infty}, \text{ and } \varphi \circ \mu(\mu^{-1}(V') \cap U) \subseteq \mathbb{R}^k \times \{0\}$$

Let projection  $\pi: \mathbb{R}^n \to \mathbb{R}^k$  be the standard projection,

$$\pi \circ \varphi \circ \mu : \mu^{-1}(V') \cap U \to \mathbb{R}^k$$
 is  $C^{\infty}$ 

 $\Longrightarrow \mu \text{ is } C^{\infty}$ 

From Chapter 4 "Lie Groups and Lie Algebras" of Kosmann-Schwarzbach (2010) [12] While Proposition 2.6 of Kosmann-Schwarzbach (2010) [12] states that

$$\det(\exp(X)) = \exp(\operatorname{tr} X)$$

here are some other resources online that gave further discussion on the characteristic polynomial,  $det(A - \lambda 1)$  and the different terms of it, called Newton identities:

- http://scipp.ucsc.edu/~haber/ph116A/charpoly\_11.pdf
- http://math.stackexchange.com/questions/1126114/how-to-find-this-lie-algebra-proof-that-mathfraksl-is-trace

• http://mathoverflow.net/questions/131746/derivative-of-a-determinant-of-a-matrix-field

**Theorem 15** (5.1 [12]). Consider  $\mathfrak{g} = \{X = \gamma'(0) | \gamma : 1 \to G \text{ of class } C^1, \gamma(0) = 1\}$ Let Lie group G

- (i)  $\mathfrak{g}$  vector subspace of  $\mathfrak{gl}(n,\mathbb{R})$
- (ii)  $X \in \mathfrak{g}$  iff  $\forall t \in \mathbb{R}$ ,  $\exp(tX) \in G$
- (iii) if  $X \in \mathfrak{g}$ , if  $g \in G$ , then  $gXg^{-1} \in \mathfrak{g}$
- (iv)  $\mathfrak{g}$  closed under matrix commutator, i.e. if  $X,Y \in \mathfrak{g}$ ,  $[X,Y] \in \mathfrak{g}$

Proof. (i

(ii) If  $\exp(tX) \in G$ , then  $X \frac{d}{dt} \exp(tX)\big|_{t=0} \in \mathfrak{g}$  (by def.) If  $X \in \mathfrak{g}$ , then by def.,  $X = \frac{d}{dt}\gamma(t)\big|_{t=0}$  with  $\gamma(t) \in G$ . Now Taylor expand;  $\forall k \in \mathbb{Z}^+$ 

$$\gamma\left(\frac{t}{k}\right) = 1 + \frac{t}{k}X + O\left(\frac{1}{k^2}\right) = \exp\left(\frac{t}{k}X + O\left(\frac{1}{k^2}\right)\right)$$
$$\Longrightarrow \left(\gamma\left(\frac{t}{k}\right)\right)^k = \exp\left(tX\right)$$
$$\gamma\left(\frac{t}{k}\right) \in G \quad \forall k \in \mathbb{Z}^+$$

G closed subgroup, so  $\lim_{k\to\infty} (\gamma\left(\frac{t}{k}\right))^k = \exp(tX) \in G$ 

- (iii)
- (iv)

<sup>4</sup>https://faculty.math.illinois.edu/~lerman/519/s12/427notes.pdf

**Definition 27** (Lie algebra). Lie algebra  $\mathfrak{g}$ , tangent space to G at 1, i.e.  $\mathfrak{g} := T_1G$  is called Lie algebra of Lie group G.

$$\mathfrak{g} := \{X = \gamma'(0) | \gamma : 1 \to G \text{ of class } C^1, \gamma(0) = 1\} = T_1 G$$

This is based on Proposition 5.3 of Kosmann-Schwarzbach (2010) [12].

For Lie group

$$U(n) = \{ U \in GL(n, \mathbb{C}) | UU^{\dagger} = 1 \}$$

If  $X \in \mathfrak{u}(n)$ , then  $\exp(tX) \in U(n)$ . Then

$$\exp(tX)\exp(tX)^{\dagger} = (1+tX+O(t^2))(1+tX^{\dagger}+O(t^2)) = 1+t(X+X^{\dagger})+O(t^2) = 1 \forall t \in \mathbb{R} \Longrightarrow X+X^{\dagger} = 0$$

i.e.  $X \in \mathfrak{u}(n)$  is an anti-Hermitian complex  $n \times n$  matrix.

$$\mathfrak{u}(n) = \{ X \in \mathfrak{gl}(n, \mathbb{C}) | X + X^{\dagger} = 0 \}$$

Physicists: X = iA and so  $A - A^{\dagger}$ .  $A \in \mathfrak{u}(n)$  is a Hermitian complex  $n \times n$  matrix.

$$\mathfrak{u}(n) = \{ A \in \mathfrak{gl}(n,\mathbb{C}) | A - A^{\dagger} = 0 \}$$

Regardless,  $\dim_{\mathbb{R}}\mathfrak{u}(n)=n^2=2n^2-n^2$ 

For Lie group

$$SU(n) = \{ U \in GL(n, \mathbb{C}) | UU^{\dagger} = 1, \det U = 1 \}$$

Then

$$\mathfrak{su}(n) = \{ X \in \mathfrak{gl}(n, \mathbb{C}) | X + X^{\dagger} = 1, \operatorname{tr} X = 0 \}$$

is the Lie algebra of traceless anti-Hermitian complex  $n \times n$  matrices, and that

$$\dim_{\mathbb{R}}\mathfrak{su}(n) = n^2 - 1$$

In summary,

From Chapter 5 "Lie Groups SU(2) and SO(3)" of Kosmann-Schwarzbach (2010) [12],

8.0.1. Bases of su(2), Subsection 1.1 of Chapter 5of Kosmann-Schwarzbach (2010) [12]. Recall that

$$\mathfrak{su}(n) = \{ X \in \mathfrak{gl}(n,\mathbb{C}) | X + X^{\dagger} = 0, \text{tr} X = 0 \}$$

$$\exp(tX) \downarrow$$

$$SU(n) = \{ U \in GL(n,\mathbb{C}) | UU^{\dagger} = 1, \det U = 1 \}$$

$$\dim_{\mathbb{R}}\mathfrak{su}(n) = n^2 - 1$$

and so

$$\begin{split} \mathfrak{su}(2) &= \{X \in \mathfrak{gl}(2,\mathbb{C}) | X + X^\dagger = 0, \mathrm{tr} X = 0\} \\ &\qquad \qquad \exp{(tX)} \\ \downarrow \\ SU(2) &= \{U \in GL(n,\mathbb{C}) | UU^\dagger = 1, \mathrm{det} U = 1\} \end{split}$$

$$\dim_{\mathbb{R}}\mathfrak{su}(2)=3$$

Also, recall that  $\mathfrak{g} \subseteq \mathfrak{gl}(n,\mathbb{C})$  is a vector subspace (15) and that  $X \in \mathfrak{g}$  iff  $\forall t \in \mathbb{R}$ ,  $\exp(tX) \in G$ . if  $X \in \mathfrak{g}$ , if  $g \in G$ , then  $gXg^{-1} \in \mathfrak{g}$ 

 $\mathfrak{g}$  closed under  $\mathfrak{g} \times \mathfrak{g} \to \mathfrak{g}$ 

$$(X,Y) \mapsto [X,Y]$$

and so with  $\mathfrak g$  as a vector space, we can have a choice of bases.

$$\xi_1 = \frac{i}{2} \begin{pmatrix} 1 \\ 1 \end{pmatrix}$$
(a) 
$$\xi_2 = \frac{1}{2} \begin{pmatrix} -1 \\ 1 \end{pmatrix}$$

$$\xi_3 = \frac{i}{2} \begin{pmatrix} 1 \\ -1 \end{pmatrix}$$
satisfying

$$[\xi_k, \xi_l] = \epsilon_{klm} \xi_m$$

(b) Physics
$$\sigma_1 = -2i\xi_1 = \begin{pmatrix} 1 \\ 1 \end{pmatrix}$$

$$\sigma_2 = 2i\xi_2 = \begin{pmatrix} -i \\ i \end{pmatrix}$$

$$\sigma_3 = -2i\xi_3 = \begin{pmatrix} 1 \\ -1 \end{pmatrix}$$

satisfying

$$[\sigma_k, \sigma_l] = 2i\epsilon_{klm}\sigma_m$$

EY: 20151001 Sage Math 6.8 doesn't run on Mac OSX El Capitan: I suspect that it's because in Mac OSX El Capitan, /usr cannot be modified anymore, even in an Administrator account. The TUG group for MacTeX had a clear, thorough, and useful (i.e. copy UNIX commands, paste, and run examples) explanation of what was going on:

http://tug.org/mactex/elcapitan.html

So keep in mind that my code for Sage Math is for Sage Math 6.8 that doesn't run on Mac OSX El Capitan. I'll also use sympy in Python as an alternative and in parallel.

One can check in sympy the traceless anti-Hermitian (or Hermitian) property of the bases and Pauli matrices, and the commutation relations (see groups.py):

import itertools
from itertools import product, permutations

```
import sympy
from sympy import I, LeviCivita
from sympy import Rational as Rat
from sympy.physics.matrices import msigma # <class 'sympy.matrices.dense.MutableDenseMatrix'>
def commute(A,B):
commute = commute(A,B)
commute takes the commutator of A and B
return (A*B - B*A)
def xi(i):
xi = xi(i)
xi is a function that returns the independent basis for
Lie algebra su(2) \setminus equiv su(2, \mathbb{C}) of Lie group SU(2) of
traceless anti-Hermitian matrices, based on msigma of sympy
cf. http://docs.sympy.org/dev/_modules/sympy/physics/matrices.html#msigma
if i not in [1,2,3]:
raise IndexError("Invalid_Pauli_index")
elif i==1:
return I/Rat(2)*msigma(1)
elif i==2:
return -I/Rat(2)*msigma(2)
elif i == 3:
return I/Rat(2)*msigma(3)
## check anti-Hermitian property and commutation relations with xi
# xi is indeed anti-Hermitian
xi(1) = -xi(1).adjoint() # True
xi(2) = -xi(2).adjoint() # True
xi(3) = -xi(3).adjoint() # True
# xi obeys the commutation relations
for i, j in product ([1,2,3], repeat = 2): print i, j
for i, j in product([1,2,3], repeat=2): print i, j, "\t_Commutator:_", commute(xi(i),xi(j))
## check traceless Hermitian property and commutation relations with Pauli matrices
# Pauli matrices i.e. msigam is indeed traceless Hermitian
msigma(1) == msigma(1).adjoint() # True
msigma(2) == msigma(2).adjoint() # True
msigma(3) == msigma(3).adjoint() # True
msigma(1).trace() = 0 # True
msigma(2).trace() = 0 # True
msigma(3).trace() = 0 # True
# Pauli matrices obey commutation relation
print "For_Pauli_matrices, _the_commutation_relations_are_:\n"
for i, j in product([1,2,3],repeat=2): print i, j, "\t_Commutator:_", commute(msigma(i),msigma(j))
for i,j,k in permutations([1,2,3],3): print "Commute:_", i,j,k, msigma(i), msigma(j),
":_and_is_2*i_of_", msigma(k), commute(msigma(i), msigma(j)) == 2*I*msigma(k)*LeviCivita(i,j,k)
```

And finally the traceless property of the Pauli matrices:

8.1. **Spin.** Let's follow the development by Baez and Muniain (1994) on pp. 175 of the Section II.1 "Lie Groups", the second (II) chapter on "Symmetry" [9].

Let  $V = \mathbb{C}^2$ , G = SU(2). Then consider the graded algebra of polynomials on  $V = \mathbb{C}^2 \ni (x,y)$ 

$$P(V) = \bigoplus_{k=0}^{\infty} P^{(k)}(V) = \bigoplus_{\substack{j=0\\2j \in \mathbb{Z}}}^{\infty} P^{(2j)}(V) = \bigoplus_{\substack{j=0\\j \in \mathbb{Z}}}^{\infty} P^{(2j)}(V) \oplus \bigoplus_{\substack{j=1/2\\2j \text{ odd}}}^{\infty} P^{(2j)}(V)$$

 $P^{(2j)}(V) \equiv \text{vector space of complex polynomials of degree } 2j$ 

and recall this representation on  $P^{(2j)}(V)$ 

$$\begin{split} &\rho^{(2j)}:G\to \operatorname{End}(P^{(2j)}(V))\\ &\rho^{(2j)}:P^{(2j)}(V)\to P^{(2j)}(V)\\ &\rho^{(2j)}(g)(f)=f\circ\rho(g^{-1}) \text{ where }\rho\text{ is the fundamental representation of }G=SU(2)\\ &\rho^{(2j)}(g)(f)(v)=f\circ\rho(g^{-1})(v)\quad\forall\, f\in P^{(2j)}(V),\,\forall\, v\in V=\mathbb{C}^2 \end{split}$$

Note,  $\dim P^{(2j)} = {2j+2-1 \choose 2-1} = 2j+1$ 

**Exercise 21.** [9]  $spin-\theta$  Consider the trivial representation  $\tau$ :

Clearly,  $P^{(0)}(V) = \mathbb{C}$ , since  $P^{(0)}(V)$  consists of polynomials of constants in  $\mathbb{C}$ .

Consider  $c_0 \in \mathbb{C}$ ,  $f = k_0 \in P^{(0)}(V)$ 

Consider 
$$c_0 \in \mathbb{C}, \ f = k_0 \in F$$
  
 $\rho^{(0)}(g)(f) = f \circ \rho(g^{-1}) = k_0$ 

$$\implies \rho^0(g)T(c_0) = T \circ \tau(g)c_0 = T(c_0).$$
 Let  $T = 1_{\mathbb{C}} = 1_{P^0(V)}$ 

So  $\rho^{(0)}(g) = \tau(g) = 1$ . To representations  $\rho^{(0)}$  and trivial representation  $\tau$  on G are equivalent.

**Exercise 22.** [9]  $spin-\frac{1}{2}$  For spin- $\frac{1}{2}$ ,  $j=\frac{1}{2}$ , 2j=1.

$$\forall f \in P^{(1)}(V), V = \mathbb{C}^2$$
. So in general form,  $f(x,y) = ax + by \in P^{(1)}(V), \begin{pmatrix} x \\ y \end{pmatrix} \in V = \mathbb{C}^2$ 

Recall the fundamental representation  $\rho:G\to GL(2,\mathbb{C})\equiv GL(\mathbb{C}^2)$   $\rho(g):\mathbb{C}^2\to\mathbb{C}^2$   $\rho(g)=g$ 

So consider T such that

$$\begin{array}{c|c}
\mathbb{C}^2 & \xrightarrow{T} & P^{(1)}(V) \\
\rho(g) \downarrow & & \downarrow \rho^{(1)}(g) \\
\mathbb{C}^2 & \xrightarrow{T} & P^{(1)}(V)
\end{array}$$

Consider  $\forall v \in \mathbb{C}^2$ ,  $v = \begin{pmatrix} x \\ y \end{pmatrix}$ , then

$$\rho(g)v = gv = \begin{bmatrix} ax + by \\ cx + dy \end{bmatrix}$$

For notation, let  $U \in G = SU(2)$  s.t.  $UU^{\dagger} = 1$ .

Consider  $(\rho^{(2j)}(U)(f))(x) = f(U^{-1}x), \forall x \in \mathbb{C}^2$ .

Choose 
$$f(x,y)=x$$
. So for  $f(x,y)=Ax+By$ ,  $A=1,B=0$ . Choose  $U=\begin{pmatrix} a & b \\ -\overline{b} & \overline{a} \end{pmatrix}$  so  $U^{-1}=\begin{pmatrix} \overline{a} & -b \\ \overline{b} & a \end{pmatrix}$ . Then

$$U^{-1}x = \begin{pmatrix} \overline{a}x - by \\ \overline{b}x + ay \end{pmatrix}$$

$$(\rho^{(1)}(U)(f))(x) = f(U^{-1}x) = \overline{a}x - by$$
  
$$(\rho^{(1)}(U)(f))(x) = f(U^{-1}x) = \overline{b}x + ay \text{ for } f(x,y) = y$$

Let f(x,y) = Ax + By

$$(\rho^{(1)}(U)(f))(x) = f(U^{-1}x) = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (\overline{a}x - by)A + (\overline{b}x + ay)B = (A\overline{a} + B\overline{b})x + (Ba - Ab)y = (A\overline{a} + B\overline{b})x + (A\overline{a}$$

which was calculated with the assistance of Sage Math:

$$\begin{array}{l} sage: \ U\_try1 = \ Matrix(\ [[a.conjugate(),-b],[b.conjugate(),a\ ]\ ]) \\ sage: \ f1(\ U\_try1*X).coefficient(x) \\ A*conjugate(a) + B*conjugate(b) \\ sage: \ f1(\ U\_try1*X).coefficient(y) \\ B*a - A*b \\ \end{array}$$

Treating  $P^{(1)}(\mathbb{C}^2)$  as a vector space, in its matrix formulation, then  $f(x,y) = Ax + By \in P^{(1)}(\mathbb{C}^2)$  is treated as  $\begin{bmatrix} A \\ B \end{bmatrix}$ , then Note  $C_{gg'} = C_gC_{g'} \equiv C(gg') = C(g) \circ C(g')$  and so  $(\rho^{(1)}(U)f)$  is

$$\implies \begin{bmatrix} \overline{a} & \overline{b} \\ -b & a \end{bmatrix} \begin{bmatrix} A \\ B \end{bmatrix} = \begin{bmatrix} A\overline{a} + B\overline{b} \\ -Ab + Ba \end{bmatrix}$$

so conclude in general that  $\rho^{(1)}(U) = (U^{\dagger})^T$ .

Now, as Kosmann-Schwarzbach (2010) [12] says, on pp. 13, Chapter 2 Representations of Finite Groups, "Two representations  $(E_1, \rho_1)$  and  $(E_2, \rho_2)$  are equivalent if and only if there is a basis  $B_1$  of  $E_1$  and a basis  $B_2$  of  $E_2$  such that for every  $g \in G$ , the matrix of  $\rho_1(q)$  in the basis  $B_1$  is equal to the matrix of  $\rho_2(q)$  in the basis  $B_2$ . In particular, if the representations  $(E_1, \rho_1)$ and  $(E_2, \rho_2)$  are equivalent, then  $E_1$  is isomorphic to  $E_2$ ." So we need a change of basis between  $\rho(U) = U$  and  $\rho^{(1)}(U)$ . What's the linear transformation T s.t.

$$T^{-1}\rho^{(1)}(U)T = U$$
?

By intuition,

$$T = \sigma_x \sigma_z \equiv \sigma_1 \sigma_3$$

where  $\sigma_i$ 's are Pauli matrices.

Indeed.

sage: Paulimat[3] \* Paulimat[1]\*U\_try\*Paulimat[1] \* Paulimat[3] [conjugate(a) conjugate(b)]

Then  $\rho^{(1)}(U) \circ T = TU$ , so this  $T = \sigma_1 \sigma_3$  is an "intertwining operator" between  $\rho^{(1)}(U)$  and fundamental representation  $\rho(U) = U$ , with  $T = \begin{bmatrix} & -1 \\ 1 & \end{bmatrix}$ , and  $T^{-1} = \begin{bmatrix} & 1 \\ -1 & \end{bmatrix}$ 

T is an isomorphism between  $\mathbb{C}^2$  and  $P^{(1)}(\mathbb{C}^2)$ . So fundamental representation  $\rho$  of G = SU(2) is equivalent to  $\rho^{(1)}(U)$  on  $P^{(1)}(\mathbb{C}^2).$ 

Exercise 23. [9] (Also from Exercise 2.6 of Kosmann-Schwarzbach (201) [12])

Let  $(E,\pi)$  representation of group G.

 $\forall g \in G, \xi \in E^*, x \in E, \text{ set } \langle \pi^*(g)(\xi), x \rangle = \langle \xi, \pi(g^{-1})(x) \rangle$ 

dual (or contragredient) of  $\pi$ ,  $\pi^*: G \to \operatorname{End}(E^*)$ ,  $\pi^*$  is a representation, since

$$\langle \pi^*(gh)(\xi), x \rangle = \langle \xi, \pi((gh)^{-1})(x) \rangle = \langle \xi, \pi(h^{-1}g^{-1})(x) \rangle = \langle \xi, \pi(h^{-1})\pi(g^{-1})(x) \rangle = \langle \xi, \pi(h^{-1})(\pi(g^{-1})(x)) \rangle = \langle \pi^*(h)(\xi), \pi(g^{-1})(x) \rangle = \langle \pi^*(g)\pi^*(h)(\xi), x \rangle$$

since this is true,  $\forall x \in E, \forall \xi \in E^*, \pi^*(gh) = \pi^*(g)\pi^*(h)$ .

dual  $\pi^*$  of  $\pi$  is a representation.

8.2. Adjoint Representation. I will first follow Sec. 7.3 The Adjoint Representation of Ch. 4 Lie Groups and Lie Algebras of Kosmann-Schwarzbach (201) [12]).

The conjugation action  $C_a: G \to G$  is defined as

$$C_g: G \to G$$
  
 $C_g: h \mapsto ghg^{-1}$ 

So

$$C: G \to \operatorname{Aut}(G)$$
$$Cg = C_g$$

Now define the adjoint action of g as the differential or push forward of  $C_q$ :

$$\operatorname{Ad}_g := D_1 \mathcal{C}_g \equiv (\mathcal{C}_g)_{*1} \equiv (\mathcal{C}_g)_*|_{g=1}$$
 (adjoint action of  $g$ )

$$Ad(g) \equiv Ad_g$$

$$\mathcal{C}_{g'} = \mathcal{C}_g \mathcal{C}_{g'} \equiv \mathcal{C}(gg') = \mathcal{C}(g) \circ \mathcal{C}(g') \text{ and so}$$

$$\xrightarrow{D_1} \mathrm{Ad}_{gg'} = \mathrm{Ad}_g \circ \mathrm{Ad}_{g'}$$

Kosmann-Schwarzbach (201) [12]) claims, because  $Ad_q = 1_{\mathfrak{q}}$  when g = 1,

 $Ad: G \to GL(\mathfrak{g})$  is a representation of G on  $\mathfrak{g}$ . (EY: 20160505???)

 $Ad: q \mapsto Ad_q$ 

**Definition 28.** representation Ad of G on  $V = \mathfrak{g}$  is called adjoint representation of Lie group G.

Denote adjoint representation of Lie algebra g, ad.

By definition,  $Ad_{\exp(tX)} = \exp(tad_X)$ 

cf. Prop. 7.8 of Kosmann-Schwarzbach (201) [12])

**Proposition 12.** (1) Let A invertible matrix,  $A \in Lie \ group \ G$ .

Let X matrix s.t.  $X \in \mathfrak{g}$ . Then

$$Ad_A(X) = AXA^{-1}$$

(2) Let  $X, Y \in \mathfrak{g}$ . Then

$$ad_X(Y) = [X, Y]$$

(3) Let  $X, Y \in \mathfrak{g}$ . Then

$$ad_{[X,Y]} = [ad_X, ad_Y]$$

*Proof.* (1) By def.,  $\forall B \in G$ ,  $C_A(B) = ABA^{-1}$ , and thus

$$\operatorname{Ad}_{A}(X) = \left. \frac{d}{dt} A \exp(tX) A^{-1} \right|_{t=0} = AXA^{-1}$$

(2)

$$\operatorname{ad}_{X}(Y) = \frac{d}{dt}\operatorname{Ad}_{\exp(tX)}(Y)\bigg|_{t=0} = \frac{d}{dt}\exp(tX)Y\exp(tX)\bigg|_{t=0} =$$
$$= XY - YX = [X, Y]$$

(3) Use Jacobi identity:

$$[A, [B, C]] + [B, [C, A]] + [C, [A, B]] = 0$$
 or  $[[A, B], C] = [A, [B, C]] - [B, [A, C]]$ 

$$ad_{[X,Y]}C = [[X,Y],C] = [X,[Y,C]] - [Y,[X,C]] = [X,ad_YC] - [Y,ad_XC]$$
 and that  $ad_Xad_YC = [X,[Y,C]] \Longrightarrow ad_{[X,Y]}C = [ad_X,ad_Y]C$ 

9. Lie Groups and their Lie Algebras

Lie groups and their Lie algebras - Lec 13 - Frederic Schuller, Schuller (2015) [28]

- 9.1. Chapter 4. Lie Theory, Schuller (2015) [28].
- 9.1.1. 4.1 Lie groups, Schuller (2015) [28].

**Definition 29** (Lie group, Schuller (2015) [28]). A lie group  $(G, \cdot)$  is

### Part 5. Cohomology; Stoke's Theorem

#### 10. Stoke's Theorem

**Theorem 16** (Stoke's Theorem). Let M be oriented, smooth n-manifold with boundary, let  $\omega$  be a compactly supported smooth (n-1)-form on M, or if  $\omega \in A_c^{n-1}(M)$ , Then

$$\int_{M} d\omega = \int_{\partial M} \omega$$

If  $\partial M = \emptyset$ , then  $\int_{\partial M} \omega = 0$  $\int_{\partial M} \omega$  interpreted as  $\int_{\partial M} i^*_{\partial M} \omega = \int_{\partial M} i^* \omega$  so

(24) 
$$\int_{M} d\omega = \int_{\partial M} i^{*}(\omega)$$

where inclusion  $i: \partial M \hookrightarrow M$ 

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*Proof.* Begin with very special case: Suppose  $M=\mathbb{H}^n$  (upper half space),  $\partial M=\mathbb{R}^{n-1}$ 

 $\omega$  has compact support, so  $\exists R > 0$  s.t. supp $\omega \subseteq \text{rectangle } A = [-R, R] \times \cdots \times [-R, R] \times [0, R]$ .  $\forall \omega \in A_c^{n-1}(\mathbb{H}^n)$ 

(25) 
$$\omega = \sum_{i=1}^{n} (-1)^{j-1} f_j dx^1 \wedge \dots \wedge \widehat{dx}^j \wedge \dots \wedge dx^n \equiv \sum_{i=1}^{n} \omega_i dx^1 \wedge \dots \wedge \widehat{dx}^i \wedge \dots \wedge dx^n$$

with Conlon (2008) [17] and John Lee (2012) [3]'s notation, respectively, and where  $f_i$  has compact support.

$$i^*\omega = (f_1 \circ i)dx^2 \wedge \dots \wedge dx^n \in A_c^{n-1}(\partial \mathbb{H}^n)$$

$$d\omega = \sum_{i=1}^{n} d\omega_{i} \wedge dx^{1} \wedge \dots \wedge \widehat{dx}^{i} \wedge \dots \wedge dx^{n} = \sum_{i,j=1}^{n} \frac{\partial \omega_{i}}{\partial x^{j}} dx^{j} \wedge dx^{1} \wedge \dots \wedge \widehat{dx}^{i} \wedge \dots \wedge dx^{n} =$$

$$= \sum_{i=1}^{n} (-1)^{i-1} \frac{\partial \omega_{i}}{\partial x^{i}} dx^{1} \wedge \dots \wedge dx^{n}$$

i.e. (for another notation)

$$d\omega = \left(\sum_{j=1}^{n} \frac{\partial f_j}{\partial x^j}\right) dx^1 \wedge \dots \wedge dx^n \in A_c^n(\mathbb{H}^n)$$

$$d\omega = \left(\sum_{j=1}^{n} \frac{\partial f_j}{\partial x^j}\right) dx^1 \wedge \dots \wedge dx^n \in A_c^n(\mathbb{H}^n)$$

$$\int_{\mathbb{H}^n} d\omega = \sum_{i=1}^n (-1)^{i-1} \int_A \frac{\partial \omega_i}{\partial x^i} dx^1 \wedge \dots \wedge dx^n = \sum_{i=1}^n (-1)^{i-1} \int_0^R \int_{-R}^R \dots \int_{-R}^R dx^1 \dots dx^n \frac{\partial \omega_i}{\partial x^i} (x)$$

We can change order of integration in each term so to do  $x^i$  integration first.

By fundamental thm. of calculus, terms for which  $i \neq n$  reduce to

$$\sum_{i=1}^{n-1} (-1)^{i-1} \int_0^R \int_{-R}^R \cdots \int_{-R}^R \frac{\partial \omega_i}{\partial x^i}(x) dx^1 \dots dx^n = \sum_{i=1}^{n-1} (-1)^{i-1} \int_0^R \int_{-R}^R \cdots \int_{-R}^R \frac{\partial \omega_i}{\partial x^i}(x) dx^i dx^1 \dots \widehat{dx}^i \dots dx^n = \sum_{i=1}^{n-1} (-1)^{i-1} \int_0^R \int_{-R}^R \cdots \int_{-R}^R [\omega_i(x)]_{x^i=-R}^{x^i=R} dx^1 \dots \widehat{dx}^i \dots dx^n = 0$$

because we've chosen R large enough that  $\omega = 0$  when  $x^i = \pm R$ .

#### Part 6. Prástaro

Prástaro (1996) [13]

10.0.1. Affine Spaces. cf. Sec. 1.2 - Affine Spaces of Prástaro (1996) [13]

**Definition 30** (affine space).

$$\textit{affine space} \qquad \quad (M,\mathbf{M},\alpha)$$

with

$$M \equiv set (set of pts.)$$

 $\mathbf{M} \equiv vector\ space\ (space\ of\ free\ vectors)$  $\alpha \equiv \mathbf{M} \times M \to M \equiv translation\ operator$ 

$$\alpha: (v, p) \mapsto p' \equiv p + v$$

Note:  $\alpha$  is a transitive action and without fixed pts. (free).

i.e.  $\forall p \in M$ ,

$$\forall \text{ pt. } O \in M, \ \alpha : (v, O) \mapsto O' \equiv O + v, \ \alpha(\cdot, O) \equiv \alpha_O \equiv \alpha(O). \ \alpha_O(v) = O' = O + \mathbf{v} \qquad \forall O' \in M, \ \exists \mathbf{v} \in \mathbf{M} \text{ s.t. } O' = O + \mathbf{v} \\ \Longrightarrow M \equiv \mathbf{M}.$$

$$\forall (O, \{e_i\})_{1 \le i \le n}$$
, where  $\{e_i\}$  basis of  $\mathbf{M}, M \equiv \mathbf{M} = \mathbb{R}^n$  so isomorphism  $M \simeq \mathbb{R}^n$ 

i.e.  $\alpha$  is without fixed pts., meaning,

Given pointed space (M, O), where base pt.  $O \in M$ , we can associate  $\forall p \in M$ , vector  $\mathbf{x} \in \mathbf{M}$ , by 1-to-1 mapping  $M \to \mathbf{M}$ . So for

$$\alpha : \mathbf{M} \times M \to M$$
  
 $\alpha(\mathbf{x}, p) = p' = p + \mathbf{x}$ 

Consider

$$\alpha(\mathbf{x}, O) = p = \alpha_O(\mathbf{x}) = p \Longrightarrow \exists \alpha_O^{-1}(p) = \mathbf{x} \in \mathbf{M}$$

- (1) tangent space of M in  $p \in M$  is vector space  $T_pM \equiv (\mathbf{M}, p) \cong M$
- (2) If M Euclidean space, affine space  $(M, \mathbf{M}, \alpha)$  is Euclidean
- (3) Call dim. of affine space  $(M, \mathbf{M}, \alpha)$ , dim. of  $\mathbf{M} \equiv \dim \mathbf{M}$

 $\{\mathbf{e}_i\}$  basis of  $\mathbf{M}$ 

**Definition 31.**  $(O, \{e_i\}) \equiv affine frame.$ 

 $\forall$  affine frame  $(O, \{e_i\})$ ,  $\exists$  coordinate system  $x^{\alpha} : M \to \mathbb{R}$ , where  $x^{\alpha}(p)$  is  $\alpha$ th component, in basis  $\{e_i\}$ , of vector p - O

**Proposition 13** (1.6, Prástaro (1996) [13]).  $\forall O \in M$ , we have canonical identification  $M \equiv \mathbf{M}$ , since

$$\alpha_O^{-1}: M \to \mathbf{M}$$
  $\alpha_O: \mathbf{M} \to M$   $\alpha_O: \mathbf{x} = \alpha(\mathbf{x}, O) = p$ 

Furthermore.

 $\forall$  affine frame  $(O, \{e_i\})_{1 \le i \le d}$ , where  $\{e_i\}$  basis of M,

 $\exists isomorphism M \cong \mathbb{R}^d$ ,

Then,  $\forall (O, \{\mathbf{e}_i\})_{1 \leq i \leq d}$ ,

 $\exists coordinate system x^{\alpha}: M \to \mathbb{R},$ 

where  $x^{\alpha}(p) = \alpha th$  component, in basis  $\{e_i\}$ , of vector p - O.

**Theorem 17** (1.4 Prástaro (1996) [13]). Let  $(x^{\alpha})$ ,  $(\overline{a}^{\alpha})$  2 coordinate systems correspond to affine frames  $(O, \{e_i\})$ ,  $(\overline{O}, \{\overline{e}_i\})$ , respectively.

$$\overline{x}^{\alpha} = A^{\alpha}_{\beta} x^{\beta} + y^{\alpha}$$

where

$$y^{\alpha} \in \mathbb{R}^n, \qquad A^{\alpha}_{\beta} \in GL(n; \mathbb{R})$$

**Definition 32** (1.10 Prástaro (1996) [13]).

(28) 
$$A(n) \equiv Gl(n, \mathbb{R}) \times \mathbb{R}^n$$

affine group of dim. n

**Theorem 18** (1.5). symmetry group of n-dim. affine space, called affine group A(M) of M.  $\exists$  isomorphism,

(29) 
$$A(M) \simeq A(n), \qquad f \mapsto (f^{\alpha}_{\beta}, y^{\alpha}); \qquad f^{\alpha} \equiv x^{\alpha} \circ f = f^{\alpha}_{\beta} x^{\beta} + y^{\alpha}$$

cf. Eq. 1.4 Prástaro (1996) [13]

**Definition 33** (metric). Let smooth manifold M, dimM = n,  $\forall p \in M$ ,  $\exists vector space T_nM$ , and so for

(30) 
$$g_p(T_pM)^2 \to \mathbb{R}$$
$$g_p: (X_p, Y_p) \mapsto g_p(X_p, Y_p) \in \mathbb{R}$$

with  $g_p$  being bilinear, symmetric (in  $X_p, Y_p$ ), nondegenerate (i.e. if  $g_p(X_p, Y_p) = 0$ , then  $X_p$  or  $Y_p = 0$ ) Note that

$$g \in \Gamma((TM \otimes TM)^*)$$

and that for 
$$X = X^{i} \frac{\partial}{\partial x^{i}}$$
 so 
$$Y = Y^{i} \frac{\partial}{\partial x^{i}}$$

$$g(X,Y) = g_{ij}X^iY^j$$

Now for

$$F: M \to N$$
 
$$DF \equiv F_*: T_p M \to T_{F(p)} N$$
 
$$F: x \mapsto y = y(x)$$
 
$$DF: X_p \mapsto (DF)(X^j \frac{\partial}{\partial x^j}) = X^j \frac{\partial y^i}{\partial x^j} \frac{\partial}{\partial y^i}$$

$$(F^*g')(X,Y) = (F^*g')(X^i \frac{\partial}{\partial x^i}, Y^j \frac{\partial}{\partial x^j}) = (F^*g')_{ij}X^iY^j = g'(F_*X, F_*Y) =$$

### Part 7. Connections

11. Connections of Vector Bundles

[31]

**Definition 34** (Connection in a vector bundle). connection in a vector bundle  $\pi: E \to M$  over  $C^{\infty}$  manifold M, is a bilinear map

$$\nabla: \mathfrak{X}(M) \times \Gamma(E) \to \Gamma(E)$$

satisfying

(i) 
$$\nabla_{fX} s = f \nabla_X s$$

(ii) 
$$\nabla_X(fs) = f\nabla_X s + (Xf)s$$
 where  $f \in C^{\infty}(M)$ ,  $X \in \mathfrak{X}(M)$ ,  $s \in \Gamma(E)$ 

 $\nabla_X s$  is covariant derivative of s relative to X (for Morita (2001)[31])

Claim: Any vector bundle admits a connection.

e.g. product bundle  $M \times \mathbb{R}^n$ . Let  $x_1, \ldots, x_n$  be canonical coordinates in  $\mathbb{R}^n$ . Take frame field  $(s_1, \ldots, s_n)$ , where  $s_i(p) = \frac{\partial}{\partial x^i}$  Set  $\nabla_X s_i = 0$   $(i = 1, \ldots, m)$   $\forall$  vector space X,

 $\forall s = \sum_i a_i s_i, \forall X \in \mathfrak{X}(M), \text{ set}$ 

$$\nabla_X s = \sum_{i=1}^n (Xa_i)s_i$$

For this connection  $\nabla_X s$  is the partial derivative in direction of X if s is considered  $\mathbb{R}^n$ -valued function on M. Call it **trivial** connection in product bundle.

Indeed,

$$\nabla_X s = \nabla_{X^i \frac{\partial}{\partial x^i}} \left( s^m \frac{\partial}{\partial x^m} \right) = X^i \nabla_{\frac{\partial}{\partial x^i}} \left( s^m \frac{\partial}{\partial x^m} \right) = X^i \left( \nabla_{\frac{\partial}{\partial x^i}} s^m \right) \frac{\partial}{\partial x^m} + X^i s^m \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^m} =$$

$$= X^i \left( \frac{\partial}{\partial x^i} s^m \right) \frac{\partial}{\partial x^m} + X^i s^m \Gamma^q_{mi} \frac{\partial}{\partial x^q} = X^i \frac{\partial s^m}{\partial x^i} \frac{\partial}{\partial x^m} + 0$$

if  $\Gamma^q_{mi} = 0$  at a chosen point p.

For arbitrary vector bundle  $\pi: E \to M$ , take locally finite open covering  $\{U_{\alpha}\}_{{\alpha}\in A}$  s.t.  $\pi^{-1}(U_{\alpha})$  trivial. Denote  $\nabla^{\alpha}$  trivial connection  $\forall \pi^{-1}(U_{\alpha})$ . Let  $\{f_{\alpha}\}$  be a partition of unity for covering  $U_{\alpha}$ , define

$$\nabla_X s := \sum_{\alpha} f_{\alpha} \nabla_X^{\alpha} s$$

Verify this defines connection in E:

$$\nabla_X(gs) = \sum_{\alpha} f_{\alpha} \nabla_X^{\alpha}(gs) = \sum_{\alpha} f_{\alpha} \left[ X^i \left( \frac{\partial g}{\partial x^i} \right) s + X^i g \frac{\partial s^m}{\partial x^i} \frac{\partial}{\partial x^m} \right] =$$

$$= g \sum_{\alpha} f_{\alpha} \nabla_X^{\alpha} s + \sum_{\alpha} f_{\alpha}(Xg) s$$

**Proposition 14** (5.18 Morita (2001)[31]). Let  $\nabla_i$  ( $1 \le i \le k$ ) be k connections in a given vector bundle. Then  $\forall$  linear combination  $\sum_{i=1}^k t_i \nabla_i$ , where  $t_1 + \cdots + t_k = 1$  is a connection.

Proof. TODO: Ex. 5.5

### Part 8. Holonomy

**Definition 35** (Conlon, 10.1.2). If  $X, Y \in \mathfrak{X}(M)$ ,  $M \subset \mathbb{R}^m$ , Levi-Civita connection on  $M \subset \mathbb{R}^m$ 

(31) 
$$\nabla : \mathfrak{X}(M) : \mathfrak{X}(M) \to \mathfrak{X}(M)$$
$$\nabla_X Y := p(D_X Y)$$

with

$$D_X Y := \sum_{j=1}^m X(Y^j) \frac{\partial}{\partial x^j} = \sum_{i,j=1}^m X^i \frac{\partial Y^j}{\partial x^i} \frac{\partial}{\partial x^j} \qquad \forall X = \sum_{i=1}^m X^i \frac{\partial}{\partial x^i},$$
$$\forall Y = \sum_{i=1}^m Y^i \frac{\partial}{\partial x^i},$$

$$\nabla_{fX}Y = f(D_{fX}Y) = p(fD_XY) = fpD_XY = f\nabla_XY$$

$$\nabla_X f Y = p(D_X f Y) = p\left(\sum_{i,j=1}^m \left(X^i f \frac{\partial Y^j}{\partial x^i} + X^i Y^j \frac{\partial f}{\partial x^i}\right) \frac{\partial}{\partial x^j}\right) = f \nabla_X Y + p \sum_{j=1}^m X(f) Y^j \frac{\partial}{\partial x^j} = f \nabla_X Y + X(f) p(Y)$$

**Definition 36** (Conlon, 10.1.4; Christoffel symbols).

**Definition 37** (torsion).

$$T:\mathfrak{X}(M)\in\mathfrak{X}(M) o\mathfrak{X}(M) \ T(X,Y)=
abla_XY-
abla_YX-[X,Y]$$

If T = 0,  $\nabla$  torsion-free or symmetric.

$$T(fX,Y) = f\nabla_X Y - (f\nabla_Y X + Y(f)X) - \{(fXY - (Y(f)X + fYX)\} = fT(X,Y)$$
  
$$T(X,fY) = f\nabla_X Y + X(f)Y - f\nabla_Y X - \{((X(f)Y + fXY) - fYX\} = fT(X,Y)$$

Thus, T(X,Y)  $C^{\infty}(M)$ -bilinear.

 $T \in \tau_1^2(M)$ .

 $T(v, w) \in T_x M$  defined,  $\forall v, w \in T_x M, \forall x \in M$ .

Thus, torsion is a **tensor**.

Exercise 10.1.7 Conlon (2008)[17]...

If T(X,Y)=0,

$$T(e_i, e_j) = \Gamma_{ii}^k e_k - \Gamma_{ij}^k e_k - 0 = 0 \Longrightarrow \Gamma_{ii}^k = \Gamma_{ij}^k$$

If  $\Gamma_{ij}^k = \Gamma_{ii}^k$ ,  $T(e_i, e_j) = 0$ .

Exercise 10.1.8, Conlon (2008)[17].

If  $M \subset \mathbb{R}^m$  smoothly embedded submanifold,  $\forall \frac{\partial}{\partial x^j}, \frac{\partial}{\partial x^i} \in T_x M$ , spanning  $T_x M$ , consider  $\frac{\partial}{\partial x^j} = X_j^k \frac{\partial}{\partial \widetilde{x}^k}, \frac{\partial}{\partial x^i} = X_i^k (\widetilde{x}) \frac{\partial}{\partial \widetilde{x}^k}$ 

$$\begin{split} \nabla_{\frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^i} &= p D_{X^k_j \frac{\partial}{\partial \widetilde{x}^k}} X^l_i \frac{\partial}{\partial \widetilde{x}^l} = p \left( X^k_j \frac{\partial X^l_i}{\partial \widetilde{x}^k} \frac{\partial}{\partial \widetilde{x}^l} \right) = X^k_j p \left( \frac{\partial X^l_i}{\partial \widetilde{x}^k} \frac{\partial}{\partial \widetilde{x}^l} \right) \\ \nabla_{\frac{\partial}{\partial x^i}} \frac{\partial}{\partial x^j} &= X^k_i p \left( \frac{\partial X^l_j}{\partial \widetilde{x}^k} \frac{\partial}{\partial \widetilde{x}^l} \right) \end{split}$$

If  $X \in \mathfrak{X}(M)$ , smooth  $s : [a, b] \to M$ , then  $\forall s(t)$ ,

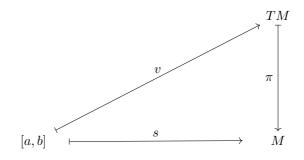
$$X'_{s(t)} = \nabla_{\dot{s}(t)} X \in T_{s(t)} M$$

In fact, it's often natural to consider fields  $X_{s(t)}$  along s, parametrized by parameter t, allowing

$$X_{s(t_1)} \neq X_{s(t_2)}$$

each of  $s(t_1) = s(t_2)$ .

**Definition 38** (10.1.9). Let smooth  $s : [a, b] \to M$ . Vector field along s is smooth  $v : [a, b] \to TM$  s.t.



commutes.

Note that  $v \in \mathfrak{X}(s) \subset \mathfrak{X}(M)$ 

e.g.  $(Y|s)(t) = Y_{s(t)}$ , restriction of  $Y \in \mathfrak{X}(M)$  to s. e.g.  $\dot{s}(t) \in \mathfrak{X}(M)$ .

 $\forall v, w \in \mathfrak{X}(s), v + w \in \mathfrak{X}(s),$ 

$$(fv + gv)(t) := (f(s(t)) + g(s(t)))v(t) = f(s(t))v(t) + g(s(t))v(t) = (f + g)v(t)$$

Likewise.

$$f(v+w) = fv + fw$$

 $\mathfrak{X}(s)$  is a real vector space and  $C^{\infty}[a,b]$ -module.

**Definition 39** (10.1.10). Let conection  $\nabla$  on M. Associated covariant derivative is operator

$$\frac{\nabla}{dt}\mathfrak{X}(s) \to \mathfrak{X}(s)$$

 $\forall$  smooth s on M, s.t.

- (1)  $\frac{\nabla}{dt} \mathbb{R}$ -linear
- (2)  $\binom{\nabla}{dt}(fv) = \frac{df}{dt}v + f\frac{\nabla}{dt}v, \ \forall f \in C^{\infty}[a,b], \ \forall v \in \mathfrak{X}(s)$ (3) If  $Y \in \mathfrak{X}(M)$ , then

$$\frac{\nabla}{dt}(Y|s)(t) = \nabla_{\dot{s}(t)}Y \in T_{s(t)}M, \quad a \le t \le b$$

**Theorem 19** (Conlon Thm. 10.1.11[17]).  $\forall$  connection  $\nabla$  on M,  $\exists$ ! associated covariant derivative  $\frac{\nabla}{dt}$ 

*Proof.* Consider arbitrary coordinate chart  $(U, x^1 \dots x^n)$ .

Consider smooth curve  $s:[a,b]\to U$ .

Let  $v \in \mathfrak{X}(s)$ ,  $v(t) = v^{i}(t) \frac{\partial}{\partial x^{i}}$ ;  $\dot{s}(t) = s^{j} \frac{\partial}{\partial x^{j}}$ 

$$\frac{\nabla v}{dt} = \frac{dv^{i}(t)}{dt} \frac{\partial}{\partial x^{i}} + v^{i}(t) \frac{\nabla}{dt} \frac{\partial}{\partial x^{i}} = \frac{dv^{i}}{dt} \frac{\partial}{\partial x^{i}} + v^{i} \nabla_{\dot{s}(t)} \frac{\partial}{\partial x^{i}} = \dot{v}^{i} \frac{\partial}{\partial x^{i}} + v^{i} \dot{s}^{j} \Gamma_{ij}^{k} \frac{\partial}{\partial x^{k}} = \left(\dot{v}^{k} + v^{i} \dot{s}^{j} \Gamma_{ij}^{k}\right) \frac{\partial}{\partial x^{k}}$$

This is an explicit, local formula in terms of connection, proving uniqueness

Existence:  $\forall$  coordinate chart  $(U, x^1 \dots x^n)$ ,  $(\dot{v}^k + v^i \dot{s}^j \Gamma^k_{ij}) \frac{\partial}{\partial x^k} =: \frac{\nabla v}{dt}$ 

$$\frac{\nabla}{dt}(fv) = \dot{f}v^k + f\dot{v}^k + fv^i\dot{s}^j = \dot{f}v + f\frac{\nabla v}{dt}$$

If f constant, then  $\frac{\nabla}{dt}$  is  $\mathbb{R}$ -linear.

**Definition 40** (10.1.12 Conlon (2008)[17]). Let  $(M, \nabla)$ . Let  $v \in \mathfrak{X}(s)$  for smooth  $s : [a, b] \to M$ . If  $\frac{\nabla v}{dt} \equiv 0$  on s, then v is **parallel** along s.

**Theorem 20** (10.1.13). Let  $(M, \nabla)$ , smooth  $s : [a, b] \to M$ ,  $c \in [a, b]$ ,  $v_0 \in T_{s(c)}M$ . Then  $\exists !$  parallel field  $v \in \mathfrak{X}(s)$  s.t.  $v(c) = v_0$ .

v parallel transport along s.

Proof.

$$\dot{s}(t) = \dot{s}^{j}(t)e_{j}$$

$$v(t) = v^{i}(t)e_{i}$$

$$v_{0} = a^{i}e_{i}$$

$$0 = \left(\frac{dv^{k}}{dt}(t) + v^{i}(t)\dot{s}^{j}(t)\Gamma_{ij}^{k}(s(t))\right)e_{k}$$

or equivalently

(34) 
$$\frac{dv^k}{dt} = -v^i \dot{s}^j \Gamma^k_{ij}, \qquad 1 \le k \le n \qquad (10.1)$$

with initial conditions  $v^k(c) = a^k$ ,  $1 \le k \le n$ .

By existence and uniqueness of solutions of O.D.E.

 $\exists \epsilon > 0 \text{ s.t. } \exists ! \text{ solutions } v^k(t). \text{ For } c - \epsilon < t < c + \epsilon.$ 

In fact, these ODEs being linear in  $v^k$ , by ODE theory (Appendix C, Thm. C.4.1).

 $\nexists$  restriction on  $\epsilon$ , so  $\exists ! v^k(t) \quad \forall t \in [a, b], 1 \leq k \leq n$ 

11.1. Principal bundle, vector bundle case for parallel transport. Recall the 2 different forms or viewpoints for Liealgebra valued 1-forms, or vector-valued 1-forms, or sections of 1-form-valued endomorphisms:

$$\omega^k_{\ i\mu}dx^\mu \equiv \omega^k_{\ i} \in \Omega^1(M,\mathfrak{gl}(n,\mathbb{F})) = \Gamma(\mathfrak{gl}(n,\mathbb{R} \otimes T^*M|_U))$$

for  $i, k = 1 \dots n = \dim E$ 

 $\mu = 1 \dots d = \dim E$ 

Now

$$D_X \mu = X^{\mu} D_{\frac{\partial}{\partial x^{\mu}}} \mu = X^{\mu} \left[ \left( \frac{\partial}{\partial x^{\mu}} \mu^k \right) e_k + \mu^i \omega_{i\mu}^k e_k \right] = \left( X(\mu^k) + \mu^i \omega_i^k(X) \right) e_k = \left( d\mu^k(X) + \mu^i \omega_i^k(X) \right) e_k$$

So then define

(35) 
$$D: \Gamma(E) \to \Gamma(E) \otimes \Gamma(T^*M)$$

$$D\mu = D(\mu^i e_i) = e_k (d\mu^k + \mu^i \omega_i^k) \equiv (d+A)\mu$$

Also, D can be defined for this case:

$$D: \Gamma(\operatorname{End}(E)) \to \Gamma(\operatorname{End}E) \otimes \Gamma(T^*M)$$

Let  $\sigma = \sigma^i_{i} e_i \otimes e^j \in \Gamma(\text{End}(E))$ 

(36) 
$$D\sigma = D(\sigma^{i}_{j}e_{i}) \otimes e^{j} + \sigma^{i}_{j}e_{i} \otimes D^{*}e^{j} = \left(d\sigma^{k}_{j} + \sigma^{i}A^{k}_{i}\right)e_{k} \otimes e^{j} + \sigma^{i}_{j}e_{i} \otimes \left(A^{*}\right)^{j}_{k}e^{k} = \left(d\sigma^{k}_{j} + \sigma^{i}_{j}A^{k}_{i}\right)e_{k} \otimes e^{j} + \sigma^{k}_{i}e_{j} \otimes \left(-A^{i}_{j}\right)e^{j} = \left(d\sigma^{k}_{j} + [A,\sigma]^{k}_{j}\right)e_{k} \otimes e^{j}$$

cf. Def. 4.1.4 of Jost (2011), pp. 138.

For  $\mu \in \Gamma(E)$ , smooth  $s: [a, b] \to M$ ,  $X(t) = \dot{s}(t)$ ,

$$(37) D_{\dot{s}(t)}\mu = \dot{s}^{\mu}D_{\frac{\partial}{\partial x^{\mu}}}\mu = \dot{s}^{\mu}\left[\frac{\partial\mu^{k}}{\partial x^{\mu}}e_{k} + \mu^{i}\omega_{i\mu}^{k}e_{k}\right] = \left[\dot{s}^{\mu}\frac{\partial\mu^{k}}{\partial x^{\mu}} + \dot{s}^{\mu}\mu^{i}\omega_{i\mu}^{k}\right]e_{k} = \frac{d}{dt}\mu(s(t)) + \mu^{i}\dot{s}^{\mu}\omega_{i\mu}^{k}e_{k}$$

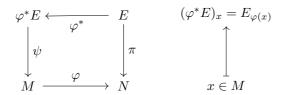
Let  $D_{\dot{s}(t)}\mu = 0$ . Then,

(38) 
$$\frac{d}{dt}\mu(s(t)) = -\mu^i \dot{s}^\mu \omega^k_{i\mu} e_k$$

Recall, given vector bundle  $E \xrightarrow{\pi} N$ , given  $\varphi: M \to N$ , then pullback

$$\varphi^* E \to M$$

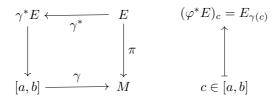
i.e.



i.e. if  $s \in \Gamma(E)$ ,

$$\varphi^* s = s \circ \varphi \in \Gamma(\varphi^* E)$$

Thus,



For

$$\dot{v}^k = -v^i \dot{s}^j \Gamma^k_{ij}$$

$$v^k(c) = v_0^k \qquad 1 \le k \le m$$

$$\dot{v} = -v^i \dot{s}^j \Gamma_{ij}$$

$$(v + w) = -(v^i + w^i) \dot{s}^j \Gamma_{ij} (v + w)(c) = v(c) + w(c) = v_0 + w_0$$

so  $v + w \in \mathfrak{X}(s)$  is parallel transport of  $v_0 + w_0$ .

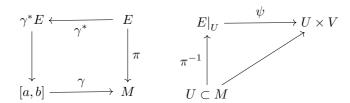
Likewise,  $\forall a \in \mathbb{F}, av \in \mathfrak{X}(s)$  is the parallel transport of  $av_0$ .

$$\dot{\mu}^k = -\mu^i \dot{s}^\mu \omega^k_{i\mu} = -\mu^i \omega^k_{i} (\dot{s}^\mu)$$

Suppose  $\gamma^* E$  trivialized over [a, b].

Closed interval is contractible, so this is always possible.

For chart  $(U, \varphi)$ ,



Consider

$$\varphi: [a, b] \times V \to \gamma^* E$$
  
$$\varphi(t, \cdot) = \gamma^* \circ \psi^{-1}(\gamma(t), \cdot)$$

 $\forall \mu \in \Gamma(E|_{x \in M}),$   $\mu = \mu^{i} e_{i}.$   $\varphi(t, e_{i}) = \epsilon_{i} \text{ is a basis for } \gamma^{*} E.$  $\forall \sigma \in \Gamma(\gamma^{*} E),$ 

$$\sigma = \sigma^i \epsilon_i, \quad \sigma^i : [a, b] \to \mathbb{F}$$

$$\nabla_{\frac{\partial}{\partial x^{\mu}}} \sigma = \frac{\partial \sigma^{k}}{\partial x^{\mu}} \epsilon_{k} + \omega^{k}_{j\mu} \sigma^{j} \epsilon_{k} = \left(\frac{\partial \sigma^{k}}{\partial x^{\mu}} + \omega^{k}_{,j\mu} \sigma^{j}\right) \epsilon_{k}$$

$$\nabla \sigma = \epsilon_{k} \otimes (d\sigma^{k} + \omega^{k}_{j\mu} dx^{\mu} \sigma^{j}) = \epsilon_{k} \otimes (d\sigma^{k} + \omega^{k}_{j} \sigma^{j})$$

$$\nabla_{\frac{d}{dt}} \sigma = \epsilon_{k} \otimes \left(\frac{d\sigma^{k}}{dt} + \omega^{k}_{j\mu} \dot{x}^{\mu} \sigma^{j}\right)$$

Now

$$\frac{d}{dt} = \dot{x}^{\nu} \frac{\partial}{\partial x^{\nu}}$$

Then  $\sigma$  parallel along  $\gamma$  if

$$\frac{d\sigma^k}{dt} + \omega^k_{\ j\mu} \dot{x}^\mu \sigma^j = 0$$

**Definition 41** (3.1.4 [18]). Parallel transport along  $\gamma$  is

(40) 
$$P_{\gamma}: E_{\gamma(a)} \to E_{\gamma(b)}$$
$$P_{\gamma}(v) \mapsto \sigma(b)$$

where  $\sigma \in \Gamma(\gamma^* E)$ ,  $\sigma$  unique and s.t.  $\sigma(a) = v$ .

**Lemma 1** (10.1.16[17]). holonomy

$$h_s: T_xM \to T_{x_0}M$$

if  $\nabla$  around piecewise smooth loop s is a linear transformation.

**Lemma 2** (10.1.18 Conlon (2008)[17]). Let piecewise smooth loop  $s:[a,b]\to M$  at  $x_0$ .

Let weak reparametrization  $\tilde{s} = s \circ r : [c, d] \to M$ .

If reparametrization is orientation-preserving, then  $h_{\tilde{s}} = h_s$ , If reparametrization is orientation-reversing, then  $h_{\tilde{s}} = h_s^{-1}$ ,

*Proof.* Without loss of generality, assume smooth s, r

$$\widetilde{s}(\tau) = s(r(\tau))$$

$$\widetilde{v}(\tau) = v(r(\tau))$$

$$\widetilde{u}^{j}(\tau) = \frac{dt}{d\tau}(\tau)u^{j}(r(\tau))$$

$$\frac{d\widetilde{v}^{k}}{d\tau}(\tau) = \frac{dr}{d\tau}(\tau)\frac{dv^{k}}{dt}(r(\tau))$$

$$\frac{d\widetilde{v}^{k}}{d\tau} = -\widetilde{v}^{i}\widetilde{u}^{j}\Gamma_{ij}^{k}$$

since

$$\begin{split} \frac{dv^k}{dt} &= -v^i u^j \Gamma^k_{ij}; \qquad 1 \leq k \leq n \\ v^k(c) &= a^k; \qquad 1 \leq k \leq a \\ \frac{dr}{d\tau} \frac{dv^k}{dt} &= -v^i \frac{dr}{d\tau} u^j \Gamma^k_{ij} = \frac{d\widetilde{v}^k}{d\tau} = -\widetilde{v}^i \widetilde{u}^j \Gamma^k_{ij} \end{split}$$

Thus, if r(c) = a, r(d) = b

$$h_{\widetilde{s}}(v_0) = \widetilde{v}(d) = v(b) = h_s(v_0)$$

If r(c) = a, r(d) = b, then

$$\widetilde{v}(c) = v(b) = h_s(v_0)$$

and

$$h_{\widetilde{s}}(h_s(v_0)) = h_{\widetilde{s}}(v(b)) = \widetilde{v}(d) = v(a) = v_0$$

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At this point, I will switch to my notation because it clarified to me, at least, what was going on, in that a holonomy  $h_s$  is invariant under orientation-preserving reparametrization, and its inverse is well-defined.

For  $\widetilde{s} = s \circ t : [c, d] \to M$ ,

piecewise smooth t is reparametrized, i.e.

$$(41) t: [c, d] \to [a, b]$$

Now,

$$\frac{d}{d\tau}\widetilde{s}(\tau) = \frac{d}{d\tau}\widetilde{s}(t(\tau)) = \dot{s}(t)\frac{dt}{d\tau}(\tau) \equiv \dot{s}\frac{dt}{d\tau}$$

$$v^k(t) = v^k(t(\tau)) = v^k(\tau)$$

$$\frac{dv^k}{d\tau}(t(\tau)) = \frac{dv^k}{dt}\frac{dt}{d\tau} = \frac{dt}{d\tau}(-v^i(\tau)\dot{s}^j(t)\Gamma^k_{ij}) = -v^i(\tau)\frac{d\widetilde{s}^j}{d\tau}\Gamma^k_{ij}$$

Consider

$$h_s(v_0) = v(b)$$

If 
$$t(c) = a$$
,  $t(d) = b$ 

$$h_{\widetilde{s}}(v_0) = \widetilde{v}(d) = v(t(d)) = v(b) = h_s(v_0)$$

If 
$$t(c) = b$$
,  $t(d) = a$ 

$$h_{\widetilde{s}}(h_s(v_0)) = h_{\widetilde{s}}(v(b)) = h_{\widetilde{s}}(v(t(c))) = h_{\widetilde{s}}(\widetilde{v}(c)) =$$
$$= \widetilde{v}(d) = v(t(d)) = v(a) = v_0$$

Thus,

$$h_{\widetilde{s}} = h_s^{-1}$$

I am working through Conlon (2008) [17], Clarke and Santoro (2012) [18], and Schreiber and Waldorf (2007)[19], concurrently, for holonomy.

### 12. Path Groupoid of a smooth manifold: generalization of paths

cf. Schreiber and Waldorf (2007)[19].

**Definition 42** (path). *path* is a smooth map  $\gamma:[0,1] \to M$ , between 2 pts.  $x,y \in M$ , which has a sitting instant; i.e. number  $0 < \epsilon < \frac{1}{2}$  s.t.

(42) 
$$\gamma(t) = \begin{cases} x & \text{for } 0 \le t < \epsilon \\ y & \text{for } 1 - \epsilon < t \le 1 \end{cases}$$

Denote the set of such paths by PM,

$$(43) \qquad PM \equiv \{ \gamma \in \Gamma(M) | \ smooth \ \gamma : [0,1] \rightarrow M \ s.t. \ \exists \ 0 < \epsilon < \frac{1}{2} \ s.t. \ \begin{cases} x & for \ 0 \leq t < \epsilon \\ y & for \ 1 - \epsilon < t \leq 1 \end{cases} \}$$

cf. Def. 2.1. of Schreiber and Waldorf (2007)[19] Define *composition*:

Given paths  $\gamma_1, \gamma_2$ ;  $\gamma_1(0) = x$ ,  $\gamma_2(0) = y$ ,  $\gamma_1(1) = y$   $\gamma_2(1) = z$  define composition to be path

(44) 
$$(\gamma_2 \circ \gamma_1)(t) := \begin{cases} \gamma_1(2t) & \text{for } 0 \le t \le \frac{1}{2} \\ \gamma_2(2t-1) & \text{for } \frac{1}{2} \le t \le 1 \end{cases}$$

 $\gamma_2 \circ \gamma_1$  smooth since  $\gamma_1, \gamma_2$  both constant near gluing pt., due to sitting instants  $\epsilon_1, \epsilon_2$ , respectively. Define *inverse*:

(45) 
$$\gamma^{-1} : [0,1] \to M$$

$$\gamma^{-1}(t) := \gamma(1-t)$$

(so that 
$$\gamma^1(t) = \begin{cases} y & \text{for } 1 - \epsilon < 1 - t \le 1 \text{ or } 0 \le t < \epsilon \\ x & \text{for } 0 \le 1 - t < \epsilon \text{ or } 1 - \epsilon < t \le 1 \end{cases}$$

**Definition 43** (thin homotopy equivalent). 2 paths  $\gamma_1$ ,  $\gamma_2$  s.t.  $\gamma_1(0) = \gamma_2(0) = x$ ,  $\gamma_1, \gamma_2$  are thin homotopy equivalent,  $\gamma_1(1) = \gamma_2(1) = y$ 

if  $\exists$  smooth  $h: [0,1] \times [0,1] \rightarrow M$  s.t.

(1) 
$$\exists 0 < \epsilon < \frac{1}{2} \text{ with}$$

(a) 
$$h(s,t) = x$$
 for  $0 \le t < \epsilon$   
 $h(s,t) = y$  for  $1 - \epsilon < t \le 1$   
(b)

(c) 
$$h(s,t) = \gamma_1(t)$$
 for  $0 \le s < \epsilon$   
 $h(s,t) = \gamma_2(t)$  for  $1 - \epsilon < s \le 1$ 

(2) differential of h has at most rank 1 everywhere, i.e.

(46) 
$$rank(dh|_{(s,t)}) \le 1 \qquad \forall (s,t) \in [0,1] \times [0,1]$$

cf. Def. 2.2. of Schreiber and Waldorf (2007)[19]

 $h(s,t) = \gamma_1(t)$  for  $0 \le s < \epsilon$  is the homotopy from  $\gamma_1$  to  $\gamma_2$ , i.e.  $h(0,t) = \gamma_1(t)$ 

$$h(s,t) = \gamma_2(t) \text{ for } 1 - \epsilon < s \le 1$$

$$h(1,t) = \gamma_2(t)$$

and define an equivalence relation on PM.

Note that for  $h:[0,1]\times[0,1]\to M$ ,

$$(Dh)|_{(s,t)} = \left[\frac{\partial h^i}{\partial s}, \frac{\partial h^i}{\partial t}\right]$$

 $P^1M \equiv \text{ set of thin homotopy classes of paths, i.e.}$   $P^1M = \{ [\gamma] | \gamma_1 \in PM, \text{ if } \exists \text{ smooth } h : [0,1] \times [0,1] \to M \text{ s.t. } h \text{ thin homotopy of } \gamma_1 \text{ and } \gamma_2, \gamma_1 \sim \gamma_2 \}$   $\text{pr} : PM \to P^1M \text{ is projection to classes.}$ 

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Denote thin homotopy class of path  $\gamma$ ,  $\gamma(0) = x$ , by  $\overline{\gamma}$ , or  $[\gamma]$ .  $\gamma(1) = y$ 

12.1. Reparametrization of thin homotopies. Let  $\beta : [0,1] \to [0,1], \ \beta(0) = 0.$   $\beta(1) = 1$ 

Then  $\forall$  path  $\gamma$ ,  $\gamma(0) = x$ ,  $\gamma \circ \beta$  is also a path  $\gamma \circ \beta(0) = x$  and

$$\gamma(1) = y$$

$$\gamma \circ \beta(1) = y$$

$$h(s,t) := \gamma(t\beta(1-s) + \beta(t)\beta(s))$$

defines a homotopy from  $\gamma$  to  $\gamma \circ \beta$ .

$$\gamma_1 \circ \gamma_2 \in PM \stackrel{\operatorname{pr}}{\longmapsto} [\gamma_1 \circ \gamma_2] = [\gamma_1][\gamma_2] \in P^1M$$

Composition of thin homotopy classes of paths obeys following rules:

**Lemma 3.**  $\forall$  path  $\gamma$ ,  $\gamma(0) = x$ 

$$\gamma(1) = y$$

(1) 
$$\overline{\gamma} \circ \overline{id_x} = \overline{\gamma} = \overline{id_y} \circ \overline{\gamma} \equiv [\gamma] 1_x = [\gamma] = 1_y[\gamma]$$

(2) for paths  $\gamma'$ ;  $\gamma'(0) = y$ ,  $\gamma''(0) = z$ 

$$\gamma'(1) = z \qquad \gamma''(1) = w$$

$$(\overline{\gamma}'' \circ \overline{\gamma}') \circ \overline{\gamma} = \overline{\gamma}'' \circ (\overline{\gamma}' \circ \overline{\gamma}) \equiv ([\gamma''][\gamma'])[\gamma] = [\gamma'']([\gamma'][\gamma])$$

$$(3) \ \overline{\gamma} \circ \overline{\gamma}^{-1} = \overline{id_y} \ and \ \overline{\gamma^{-1}} \circ \overline{\gamma} = \overline{id_x} \equiv [\gamma][\gamma^{-1}] = 1_y \ and \ [\gamma^{-1}][\gamma] = 1_x$$

cf. Lemma 2.3. of Schreiber and Waldorf (2007)[19]

**Definition 44** (path groupoid).  $\forall$  smooth manifold M, consider category whose set of objects is M,

whose set of morphisms is  $P^1M$ , where class  $[\gamma]$ ,  $[\gamma](0) = x$  is a morphism from x to y and

$$[\gamma](1) = y$$

composition  $[\gamma_1][\gamma_2] = [\gamma_1 \circ \gamma_2] \in P^1M$  Lemma 3 are axioms of a category, 3rd. property says  $\forall$  morphism is invertible. Hence, we've defined a groupoid, called **path groupoid** of M,  $\mathcal{P}_1(M)$ .

So

$$\mathrm{Obj}(\mathcal{P}_1(M)) = M$$

$$\operatorname{Mor}(\mathcal{P}_1(M)) = P^1 M$$

 $\forall$  smooth  $f: M \to N$ , denote functor  $f_*$ 

$$(50) f_*: \mathcal{P}_1(M) \to \mathcal{P}_1(N)$$

with

$$f_*(x) = f(x)$$
$$(f_*)([\gamma]) := [f \circ \gamma]$$

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If 
$$\gamma \sim \gamma'$$
, for  $f \circ \gamma$ ,  $f \circ \gamma'$ ,

 $f \circ h(s,t)$  with  $f \circ h(0,t) = f \circ \gamma(t)$ .

$$f \circ h(1,t) = f \circ \gamma'(t)$$

so  $f \circ h$  is a thin homotopy between  $f \circ \gamma$ ,  $f \circ \gamma'$  and so  $[f \circ \gamma]$  well-defined.

## Part 9. Complex Manifolds

EY: 20170123 I don't see many good books on Complex Manifolds for physicists other than Nakahara's. I will supplement this section on Complex Manifolds with external links to the notes of other courses that I found useful to myself.

Complex Manifolds - Lecture Notes Koppensteiner (2010) [14]

Lectures on Riemannian Geometry, Part II: Complex Manifolds by Stefan Vandoren

Vandoren (2008) [15]

### Part 10. Jets, Jet bundles, h-principle, h-Prinzipien

cf. Eliashberg and Misahchev (2002) [20]

cf. Ch. 1 Jets and Holonomy, Sec. 1.1 Maps and sections of Eliashberg and Misahchev (2002) [20].

Visualize  $f: \mathbb{R}^n \to \mathbb{R}^q$  as graph  $\Gamma_f \subset \mathbb{R}^n \times \mathbb{R}^q$ .

Consider this graph as image of  $\mathbb{R}^n \to \mathbb{R}^n \times \mathbb{R}^q$ , i.e.

$$x \mapsto (x, f(x))$$

 $\mathbb{R}^n \to \mathbb{R}^n \times \mathbb{R}^q$  is called section (by mathematicians),

$$x \mapsto (x, f(x))$$

is called *field* or  $\mathbb{R}^q$ -valued field (by physicists).

cf. Ch. 1 Jets and Holonomy, Sec. 1.2 Coordinate definition of jets of Eliashberg and Misahchev (2002) [20].

**Definition 45** (r-jet). Given (smooth)  $f: \mathbb{R}^n \to \mathbb{R}^q$ , given  $x \in \mathbb{R}^n$ .

r-jet of f at x - sequence of derivatives of f, up to order r,  $\equiv$ 

(51) 
$$J_f^r(x) = (f(x), f'(x) \dots f^{(r)}(x))$$

 $f^{(q)}$  consists of all partial derivatives  $D^{\alpha}f$ ,  $\alpha = (\alpha_1 \dots \alpha_n)$ ,  $|\alpha| = \alpha_1 + \dots + \alpha_n = s$ , ordered lexicographically. e.g. q = 1,

 $f: \mathbb{R}^n \to \mathbb{R}.$ 

1-jet of f at  $x = J_f^1(x) = (f(x), f^{(1)}(x))$ .

$$f^{(1)}(x) = \{D^{\alpha}f | \alpha = (\alpha_1 \dots \alpha_n), |\alpha| = \alpha_1 + \dots + \alpha_n = 1\} = \left(\frac{\partial f}{\partial x^1}, \frac{\partial f}{\partial x^2}, \dots \frac{\partial f}{\partial x^n}\right)$$

Let  $d_r = d(n,r) = \text{number of all partial derivatives } D^{\alpha} \text{ of order } r \text{ of function } \mathbb{R}^n \to \mathbb{R}.$ 

Consider r-jet  $J_f^r(x)$  of map  $f: \mathbb{R}^n \to \mathbb{R}^q$  as pt. of space  $\mathbb{R}^q \times \mathbb{R}^{qd_1} \times \mathbb{R}^{qd_2} \times \cdots \times \mathbb{R}^{qd_r} = \mathbb{R}^{qN_r}$ , where  $N_r = N(n,r) = 1 + d_1 + d_2 + \cdots + d_r$ , i.e.

$$J_f^r(x) = (f(x), f^{(1)}(x), \dots f^{(r)}(x)) \in \mathbb{R}^q \times \mathbb{R}^{qd_1} \times \dots \times \mathbb{R}^{qd_r} = \mathbb{R}^{qN_r}$$

#### Exercise 1.

Given order r, consider n-tuple of (positive) integers  $(r_1, r_2 \dots r_n)$  s.t.  $r_1 + r_2 + \dots + r_n = r$ , and  $r_k \ge 0$ . Imagine  $r_k =$  occupancy number, num ber of balls in kth cell.  $(r_1 \dots r_n)$  describes a positive ocnfiguration of occupancy numbers,

with indistinguishable balls; 2 distributions are distinguishable only if corresponding n-tuples  $(r_1 \dots r_n)$  not identical.

Represent balls by stars, and indicate n cells by n spaces between n+1 bars.

With n+1 bars, r stars, 2 bars are fixed. n-1 bars and r stars to arrange linearly, so a total of n-1+r objects to arrange. where f(p)=0 and p is at origin of these coordinates. r stars indistinguishable amongst themselves, so choose r out of n-1+r to be stars.

$$\implies d_r = d(n,r) = \binom{n-1+r}{r}$$

Use induction (cf. Ch. 4 Binomial Coefficients).

$$N_0 = N(n,0) = \binom{n-1+0}{0} = 1$$

$$N_1 = N(n,1) = 1 + \binom{n-1+1}{1} = 1 + n = \frac{(n+1)!}{n!1!}$$

Induction step:

$$N_{r-1} = N(n, r-1) = \sum_{k=1}^{r-1} d_k + 1 = \binom{n+r-1}{r-1}$$

and so

$$N_{r} = N(n,r) = \sum_{k=1}^{r} d_{k} + 1 = \sum_{k=1}^{r} {n-1+k \choose k} + 1 = \sum_{k=1}^{r-1} {n-1+k \choose k} + {n-1+r \choose r} + 1 =$$

$$= {n+r-1 \choose r-1} + {n-1+r \choose r} = \frac{(n+r-1)!}{(r-1)!n!} + \frac{(n-1+r)!}{r!(n-1)!} = \frac{(n+r)!}{n!r!} = {n+r \choose r}$$

$$\mathbb{R}^{qN_{r}} \qquad J_{f}^{r}(x)$$

$$J_{f}^{r} \qquad J_{f}^{r}(x)$$

$$\mathbb{R}^{qN_{r}} \qquad J_{f}^{r}(x)$$

**Definition 46** (space of r-jets). space of r-jets of maps  $\mathbb{R}^n \to \mathbb{R}^q$  or space of r-jets of sections  $\mathbb{R}^n \to \mathbb{R}^n \times \mathbb{R}^q \equiv$ 

(53) 
$$J^{r}(\mathbb{R}^{n}, \mathbb{R}^{q}) = \mathbb{R}^{n} \times \mathbb{R}^{qN_{r}} \equiv \mathbb{R}^{n} \times \mathbb{R}^{q} \times \mathbb{R}^{qd_{1}} \times \mathbb{R}^{qd_{2}} \times \cdots \times \mathbb{R}^{qd_{r}}$$

e.q.  $J^1(\mathbb{R}^n, \mathbb{R}^q) = \mathbb{R}^n \times \mathbb{R}^q \times M_{q \times n}$ , where  $M_{q \times n} = \mathbb{R}^{qn}$  is the space of  $(q \times n)$ -matrices.

### Part 11. Morse Theory

#### 13. Morse Theory introduction from a physicist

I needed some physical motivation to understand Morse theory, and so I looked at Hori, et. al. [16]. cf. pp. 43, Sec. 3.4 Morse Theory, from Ch. 3. Differential and Algebraic Topology of Hori, et. al. [16]. Consider smooth  $f: M \to \mathbb{R}$ , with non-degenerate critical points.

If no critical values of f between a and b (a < b), then subspace on which f takes values less than a is deformation retract of subspace where f less than b, i.e.

$$\{x \in M | f(x) < b\} \times [0,1] \xrightarrow{F} \{x \in M | f(x) < b\}$$

 $\forall x \in M \text{ s.t. } f(x) < b,$ 

$$F(x,0) = x$$
  
 $F(x,1) \in \{x \in M | f(x) < a\}$  and  $F(a',1) = a'$   $\forall a' \in M \text{ s.t. } f(a') < a$ 

To show this, consider  $-\nabla f/|\nabla f|^2$ 

Morse lemma:  $\forall$  critical pt. p s.t.  $\exists$  choice of coordinates s.t.

$$f = -(x_1^2 + x_2^2 + \dots + x_n^2) + x_{n+1}^2 + \dots + x_n^2$$

• difference between

$$f^{-1}(\{x \le -\epsilon\}), f^{-1}(\{x \le +\epsilon\})$$

can be determined by local analysis and only depends on  $\mu$ ,  $\mu \equiv$  "Morse index" = number of negative eigenvalues of

Answer:

 $f^{-1}(\{x < +\epsilon\})$  can be obtained from  $f^{-1}(\{x < -\epsilon\})$  by "attaching  $\mu$ -cell" along boundary  $f^{-1}(0)$ 

• "attaching  $\mu$ -cell to X mean, take  $\mu$ -ball  $B_{\mu} = \{|x| \leq 1\}$  in  $\mu$ -dim. space, identity pts. on boundary  $S^{\mu-1}$  with pts. in the space X, through cont.  $f: S^{\mu-1} \to X$ , i.e. take

$$X \coprod B_{\mu}$$

with  $x \sim f(x) \quad \forall x \in \partial B_{\mu} = S^{\mu-1}$ .

• find homology of M,

f defines chain complex  $C_k^*$ , kth graded piece  $C^{\alpha_k}$ ,  $\alpha_k$  is number of critical pts. with index k.

$$\partial: C_p^k \to C_p^{k-1}$$
$$\partial x_a = \sum_b \Delta_{a,b} x_b$$

where  $\Delta_{a,b} :=$  signed number of lines of gradient flow from  $x_a$  to  $x_b$ , b labels pts. of index k-1.

Gradient flow line is path x(t) s.t.  $\dot{x} = \nabla(f)$ , with  $x(-\infty) = x_a$ 

$$x(+\infty) = x_b$$

- To define this number  $(\Delta_{a,b}?)$ , construct moduli space of such lines of flow (???)by intersecting outward and inward flowing path spaces from each critical point, and then show this moduli space is oriented, 0-dim. manifold (pts. with signs)
- $\partial^2 = 0$  proof

 $\partial$ , boundary of space of paths connecting critical points, whose index differs by 2 = union over compositions of pathsbetween critical pts. whose index differs by 1.

 $\implies$  coefficients of  $\partial^2$  are sums of signs of pts. in 0-dim. space, which is boundary of 1-dim. space.

These signs must therefore add to 0, so  $\partial^2 = 0$ .

Hori, et. al. [16] is good for physics, but there isn't much thorough, step-by-step explanations of the math. I will look at Hirsch (1997) [6] and Shastri (2011) [5] at the same time.

13.1. Introduction, definitions of Morse Functions, for Morse Theory. cf. Ch. 6, Morse Theory of Hirsch (1997) [6], Section 1. Morse Functions, pp. 143-

Recall for TM,  $T_xM \xrightarrow{\varphi} \mathbb{R}^n$ .

Cotangent bundle  $T^*M$  defined likewise:

$$T_x^*M \xrightarrow{\varphi} \text{dual vector space } (\mathbb{R}^n)^* = L(\mathbb{R}^n, \mathbb{R})$$

i.e.

(55)

$$T^*M = \bigcup_{x \in M} (M_x^*) \qquad M_x^* = L(M_x, \mathbb{R})$$

If chart  $(\varphi, U)$  on M, natural chart on  $T^*M$  is

$$T^*U \to \varphi(U) \times (\mathbb{R}^n)^*$$
  
 $\lambda \in M_x^* \mapsto (\varphi(x), \lambda \varphi_x^{-1})$ 

Projection map

$$p: T^* \to M$$
$$M_x^* \mapsto x$$

Let  $C^{r+1}$  map,  $1 \le r \le \omega$ ,  $f: M \to \mathbb{R}$ ,  $\forall x \in M$ , linear map  $T_x f: M_x \to \mathbb{R}$  belongs to  $M_x^*$ 

$$T_x f = Df_x \in M_x^*$$

Then

$$Df: M \to T^*M$$
  
 $x \mapsto Df_x = Df(x)$ 

is  $C^r$  section of  $T^*M$ .

Definition 47. critical point x of f is zero of Df, i.e.

$$Df(x) = 0$$

of vector space  $M_x^*$ .

Thus, set of critical pts. of f is counter-image of submanifold  $Z^* \subset T^*M$  of zeros. Note  $Z^* \approx M$ , codim. of  $Z^*$  is  $n = \dim M$ .

**Definition 48.** *Morse function* f *if*  $\forall$  *critical pts. of* f *are nondegenerate.* 

Note set of critical pts. closed discrete subset of M.

Let open  $U \subset \mathbb{R}^n$ , let  $C^2$  map  $g: U \to \mathbb{R}$ , critical pt.  $p \in U$  nondegenerate iff

- linear  $D(Dg)(p): \mathbb{R}^n \to (\mathbb{R}^n)^*$  bijective
- identify  $L(\mathbb{R}^n, (\mathbb{R}^n)^*)$  with space of bilinear maps  $\mathbb{R}^n \times \mathbb{R}^n \to \mathbb{R}$ ,  $\Longrightarrow$  equivalent to condition that symmetric bilinear  $D^2q(p): \mathbb{R}^n \times \mathbb{R}^n \to \mathbb{R}$  non-degenerate
- $n \times n$  Hessian matrix

$$\left[\frac{\partial^2 g}{\partial x^i \partial x^j}(p)\right]$$

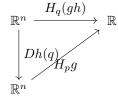
has rank n

Hessian of g at critical pt. p is quadratic form  $H_p f$  associated to bilinear form  $D^2 g(p)$ 

$$\implies H_p f(y) = D^2 g(p)(y, y) = \sum_{i,j} \frac{\partial^2 g}{\partial x^i \partial x^j}(p) y^i y^j$$

Let open  $V \subset \mathbb{R}^n$ , suppose  $C^2$  diffeomorphism  $h: V \to U$ .

Let  $q = h^{-1}(p)$ , so q is critical pt. of  $gh: V \to \mathbb{R}$ .



(quadratic) form  $(H_p f)$  invariant under diffeomorphisms.

Let 
$$C^2 f: M \to \mathbb{R}$$
.

 $\forall$  critical pt. x of f, define

Hessian quadratic form

$$H_x f: M_x \to \mathbb{R}$$

$$H_x f: M_x \xrightarrow{D\varphi_x} \mathbb{R}^n \xrightarrow{H_{\varphi(x)}(f\varphi^{-1})} \mathbb{R}$$

where  $\varphi$  is any chart at x.

Thus, critical pt. of a  $C^2$  real-valued function nondegenerate iff associated Hessian quadratic form is nondegenerate.

Let Q nondegenerate quadratic form on vector space E.

Q negative definite on subspace  $F \subset E$  if Q(x) < 0 whenever  $x \in F$  nonzero.

Index of  $Q \equiv \text{Ind}Q$ , is largest possible dim. of subspace on which Q is negative definite.

cf. 1.1. Morse's Lemma of Ch. 6, pp. 145, Morse Theory of Hirsch (1997) [6]

**Lemma 4** (Morse's Lemma). Let  $p \in M$  be nondegenerate critical pt. of index k of  $C^{r+2}$  map  $f: M \to \mathbb{R}$ ,  $1 \le r \le \omega$ . Then  $\exists C^r$  chart  $(\varphi, U)$  at p s.t.

(57) 
$$f\varphi^{-1}(u_1 \dots u_n) = f(p) - \sum_{i=1}^k u_i^2 + \sum_{i=k+1}^n u_i^2$$

Let  ${}^TQ \equiv Q^T$  denote transpose of matrix Q.

**Lemma 5.** Let  $A = diag\{a_1, \ldots, a_n\}$  diagonal  $n \times n$  matrix, with diagonal entries  $\pm 1$ . Then  $\exists$  neighborhood N of A in vector space of symmetric  $n \times n$  matrices,  $C^{\infty}$  map

$$P: N \to GL(n, \mathbb{R})$$

s.t. 
$$P(A) = I$$
, and if  $P(B) = Q$ , then  $Q^TBQ = A$ 

*Proof.* Let  $B = [b_{ij}]$  be symmetri matrix near A s.t.  $b11 \neq 0$  and  $b_{11}$  has same sign as  $a_1$ . Consider x = Ty where

$$x_1 = \left[ y_1 - \frac{b_{12}}{b_{11}} y_2 - \dots - \frac{b_{1n}}{b_{11}} y_n \right] / \sqrt{|b_n|}$$
  
$$x_k = y_k \text{ for } k = 2, \dots n$$

#### 14. Lagrange multipliers

From wikipedia:Lagrange multiplier, https://en.wikipedia.org/wiki/Lagrange\_multiplier, find local minima (maxima), pt.  $a \in N$ , s.t.  $\exists$  neighborhood U s.t.  $f(x) \geq f(a)$  ( $f(x) \leq f(a)$ )  $\forall x \in U$ .

For  $f: U \to \mathbb{R}$ , open  $U \subset \mathbb{R}^n$ , find  $x \in U$  s.t.  $D_x f \equiv Df(x) = 0$ , check if Hessian  $H_x f < 0$ .

Maxima may not exit since U open.

References:

Relative Extrema and Lagrange Multipliers

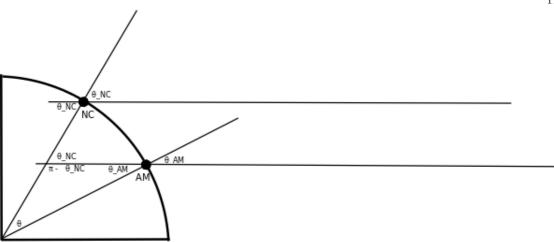
Other interesting links:

The Lagrange Multiplier Rule on Manifolds and Optimal Control of nonlinear systems

### Part 12. Classical Mechanics

Problem 1.4 The altitude angle of the Sun, Japan Physics Olympiad (2014). cf. Japan Physics Olympiad, (2014) [34]. Solution 1.4 The altitude angle of the Sun, Japan Physics Olympiad (2014).

$$NC \equiv \text{Niigata City},$$
  
 $AM \equiv \text{Amagi-san in Izu}.$ 



From all 3 internal angles of a triangle sum up to 180 degrees, or i.e.  $\pi$  radians,

$$\pi = \pi - \theta_{NC} + \theta_{AM} + \theta \Longrightarrow \theta_{NC} - \theta_{AM}$$

So we've shown indeed that  $\theta$  is the difference between altitude angles of the Sun at NC and AM, respectively. From

$$C = 2\pi R$$

$$l_M = \theta R$$

$$\frac{l_M}{l} = \frac{\theta R}{\frac{\pi}{2}R} \text{ or } \frac{l_M}{l} = \frac{2\theta}{\pi}$$

$$\implies \theta = \frac{\pi}{2} \frac{l_M}{l} = \frac{\pi}{2} \cdot \frac{334 \times 10^3 \, m}{10^7 m} \cdot \frac{180^{\circ}}{\pi} = 3^{\circ}$$

Advice, remember to draw a *cross-section* to make diagram clear, rather than worry about "the other transverse dimension". Also, draw the lines literally parallel to the Sun because of how far away the Sun is and how large the Sun is.

#### 15. STRUCTURE OF GALILEAN SPACE-TIME

cf. Sec. 3.1 - Structure of Galilean Space-Time of Prástaro (1996) [13].

Mechanics assumes a particular simple formulation if formulated with respect to some spacetime manifold.

In Galilean spacetime, it's possible to naturally recognize absolute objects, and others that depend on frames.

cf. Def. 3.1 of Prástaro (1996) [13]

**Definition 49** (Galilean spacetime structure). (1) Galilean spacetime structure :=  $(\mathcal{G}, g)$  where  $\mathcal{G}$  is (fiber bundle space-time)

(59) 
$$\mathcal{G} \equiv \{ \tau : M \to T \}$$

where

M = 4-dim. affine manifold (**space-time**); corresponding structure is  $(M, \mathbf{M}, \alpha)$ ,

- (2) T = 1-dim. affine space (time), corresponding affine structure is  $(T, T, \beta)$
- (3)  $\tau = surjective$  affine mapping, of constant rank 1, s.t.  $\forall p \in M$  associates its time  $\tau(p) \in T$

$$Put \mathbf{S} = ker(\underline{\tau}) \equiv ker(D\tau) \in M,$$

where  $\underline{\tau} \equiv D\tau$ , D is symbol of derivative.

Define

$$g: M \to vS_2^0(M) \equiv M \times S_2^0(\mathbf{S})$$
$$g(p) = (p, q) \equiv (p, Dq), \forall p \in M$$

where  $g \equiv Dg$  is a Euclidean structure on **S**. g is called vertical metric field.

Thus, given  $(M, \mathbf{M}, \alpha)$ ,  $\forall (O, \{\mathbf{e}_i\}_{1 \le i \le d}, \{\mathbf{e}_i\}_{i=1...d})$ , is basis of  $\mathbf{M}$ ,

$$M \cong \mathbb{R}^4$$
, and  $\exists \{x^\alpha : M = \mathbb{R}^4 \to \mathbb{R}\}_{\alpha=1...4}$ 

cf. 1.1.3 Principle of Relativity, pp. 9, 1 Basic Principles of Classical Mechanics, Arnold, Kozlov, Neishtadt (2006) [26]. direct product  $E^3 \times \mathbb{R}\{t\}$  (spacetime) has natural structure of affine space

Galilean group is by def. group of all affine transformations of  $E^3 \times \mathbb{R}$  that preserve time intervals and are isometries of space  $E^3 \forall$  fixed  $t \in \mathbb{R}$ .

Thus if

**Definition 50** (Galilean group). *Galilean group* is group of all affine transformations of  $E^3 \times \mathbb{R}$  that preserve time intervals and are isometries of space  $E^3 \forall$  fixed  $t \in \mathbb{R}$ .

Thus, if Galilean transformation g,

(60) 
$$g: E^3 \times \mathbb{R}\{t\} \to E^3 \times \mathbb{R}\{t\}$$
$$g: (s,t) \mapsto (s',t')$$

 $(1) t_{\alpha} - t_{\beta} = t'_{\alpha} - t'_{\beta}$ 

(2) if 
$$t_{\alpha} = t_{\beta}$$
, then  $|s_{\alpha} - s_{\beta}| = |s'_{\alpha} - s'_{\beta}|$ 

Consider  $(s_a, t_a), (s_b, t_b), t_a - t_b, |s_a - s_b|$ .

$$g(s_a, t_a) = (s'_a, t'_a)$$

$$g(s_b, t_b) = (s'_b, t'_b)$$

$$t'_a - t'_b = \pi_t g(s_a, t_a) - \pi_t g(s_b, t_b) = t_a - t_b$$

$$|s'_a - s'_b| = |\pi_s g(s_a, t_a) - \pi_s g(s_b, t_b)| = |s_a - s_b|$$

Galilean group g acts on  $\mathbb{R}^3\{\setminus\}\times\mathbb{R}\{t\}$ .

15.1. 3 examples of Galilean transformations. uniform motion with constant velocity v:

(61) 
$$g_1(\mathbf{r},t) = (\mathbf{r} + \mathbf{v}t, t)$$

translation of origin in spacetime

(62) 
$$g_2(\mathbf{r},t) = (\mathbf{r} + \mathbf{x}, t + \alpha)$$

rotation of coordinate axes

$$g_3(\mathbf{r},t) = (G\mathbf{r},t)$$

where  $G: \mathbb{R}^3 \to \mathbb{R}^3$  is orthonormal transformation.

**Proposition 15** (Galilean transformation decomposition).  $\forall$  galilean transformation  $g: \mathbb{R}^3 \times \mathbb{R} \to \mathbb{R}^3 \times \mathbb{R}$  can be uniquely represented as composition  $g_1g_2g_3$ 

cf. Prop. 1.1 of Arnold, Kozlov, Neishtadt (2006) [26]

*Proof.* Consider affine transformation of  $E^4 = E^3 \times \mathbb{R}$  of general form

$$\mathbf{x} = A\mathbf{x}' + \mathbf{v}t' + \mathbf{a}$$
$$t = \langle \mathbf{l}, \mathbf{x}' \rangle + kt' + s$$

 $A \in \operatorname{Mat}_{\mathbb{R}}(3,3), \mathbf{v}, \mathbf{a}, \mathbf{l} \in \mathbb{R}^3, k, s \in \mathbb{R}$ 

By galilean transformation (Def. 50),  $t_2 - t_1 = t_2' - t_1'$ , so

$$t_2 - t_1 = \langle \mathbf{l}, \mathbf{x}_2' \rangle + kt_2' + s - \langle \mathbf{l}, \mathbf{x}_1' \rangle - kt_1' - s = \langle \mathbf{l}, \mathbf{x}_2' - \mathbf{x}_1' \rangle + k(t_2' - t_1') = t_2' - t_1'$$

Since each of the 3 components of  $\mathbf{x}_2 - \mathbf{x}_1$  are arbitrary,  $\mathbf{l} = 0$ , k = 1 clearly. Or, let  $t = 2 = t_1$ . Then clearly  $\mathbf{l} = 0$ .

By galilean transformation (Def. 50),  $|\mathbf{x}_2 - \mathbf{x}_1| = |\mathbf{x}_2' - \mathbf{x}_1'|$ 

Let  $t_2 = t_1$ .

$$|\mathbf{x}_2 - \mathbf{x}_1| = |A\mathbf{x}_2' + \mathbf{v}t_2' + \mathbf{a} - A\mathbf{x}_1' - \mathbf{v}t_1' - \mathbf{a}| = |A(\mathbf{x}_2' - \mathbf{x}_1') + \mathbf{v}(t_2' - t_1')| = |A(\mathbf{x}_2' - \mathbf{x}_1')| = |\mathbf{x}_2' - \mathbf{x}_1'|$$

Now let  $\mathbf{y} = \mathbf{x}_2' - \mathbf{x}_1'$ .

$$|A\mathbf{v}|^2 = y^T A^T A y = y^T y$$

only if  $A^T A = 1$ . So A orthonormal.

Thus we've found the general form of Galilean transformations:

(64) 
$$\mathbf{x} = A\mathbf{x}' + \mathbf{v}t' + \mathbf{a}, \quad A \in O(3), \mathbf{v}, \mathbf{a} \in \mathbb{R}^3, s \in \mathbb{R}$$
$$t = t' + s$$

Since orthogonal matrices form a 3 parameter family,  $\dim A + \dim \mathbf{v} + \dim \mathbf{a} + \dim \mathbf{s} = 3 + 3 + 3 + 1 = 10$  independent parameters. Recall that  $q_1(\mathbf{r}, t) = (\mathbf{r} + \mathbf{v}t, t)$ ,  $q_2(\mathbf{r}, t) = (\mathbf{r} + \mathbf{a}, t + s)$ ,  $q_3(\mathbf{r}, t) = (A\mathbf{r}, t)$ , then  $q = q_1q_2q_3$ .

Introduce  $E^3$  a "fixed" reference frame by fixing pt.  $o \in E^3$  and choosing 3 mutually perpendicular axes. inertial frame - frame moving uniformly and rectilinearly with respect to fixed pt.  $o \in E^3$ ;  $\forall$  element of Galilean group transforms fixed frame into inertial frame.

action of galilean group on  $E^3 \times \mathbb{R}$  can be extended to action on  $E^3 \times \cdots \times E^3 \times \mathbb{R}$  by the rule: if  $q:(s,t) \to (s',t')$ , then

$$g(s_1, \dots s_n, t) = g(s'_1, \dots s'_n, t')$$

Galileo-Newton principle of relativity asserts that Newton's equations invariant under Galilean transformation group in inertial reference frame.

M=4-dim. affine manifold (spacetime). Recall definition of an affine manifold from wikipedia.

Affine manifold M is real manifold with charts  $\psi_i: U_i \to \mathbb{R}^n$ , s.t.  $\psi_i \circ \psi_j^{-1} \in \text{Aff}(\mathbb{R}^n) \ \forall i,j$  where  $\text{Aff}(\mathbb{R}^n)$  denotes Lie group of affine transformations;

Recall the definition of a Lie group of affine transformations,  $Aff(\mathbb{R}^n)$ , the "affine structure",  $\frac{5}{3}$ ,

**Definition 51** (Lie group of affine transformations, affine structure).

(65) 
$$Aff(\mathbb{R}^n) = \left\{ \begin{pmatrix} A & b \\ 0 & 1 \end{pmatrix} | A \in GL(n, \mathbb{R}), b \in \mathbb{R}^n \right\}$$

s.t.  $\begin{pmatrix} A & b \\ 0 & 1 \end{pmatrix}$  acts on real affine space  $\widetilde{\mathbb{R}^n}$  by

$$\begin{pmatrix} A & b \\ 0 & 1 \end{pmatrix} \begin{pmatrix} v \\ 1 \end{pmatrix} = \begin{pmatrix} Av + b \\ 1 \end{pmatrix}$$

where  $A \in GL(n, \mathbb{R}), b \in \mathbb{R}^n$ , where  $(v, 1)^T \in \widehat{\mathbb{R}}^n$ 

If changing notation, and keeping in mind that  $n=4, \psi_i \equiv y(v_{(u)}) \to \mathbb{R}^n$ 

$$\psi_i \equiv x : U_{(x)} \to \mathbb{R}^n$$
  
 $x(p) \in \mathbb{R}^n, x_i(p) \in \mathbb{R}$ 

Rewrite the Eq. 64 in matrix form:

$$\mathbf{x} = A\mathbf{x}' + \mathbf{v}t + \mathbf{b}$$

$$t = t' + s, \ A \in O(3), \ \mathbf{v}, \mathbf{b} \in \mathbb{R}^3, \ s \in \mathbb{R}$$

$$\begin{pmatrix} A & \mathbf{v} & \mathbf{b} \\ 0 & 1 & s \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \mathbf{x}' \\ t' \\ 1 \end{pmatrix} = \begin{pmatrix} \mathbf{x} \\ t \\ 1 \end{pmatrix} \text{ or } \begin{pmatrix} A & \mathbf{v} \\ 0 & 1 \end{pmatrix} \begin{pmatrix} \mathbf{x}' \\ t' \end{pmatrix} + \begin{pmatrix} \mathbf{b} \\ s \end{pmatrix} = \begin{pmatrix} \mathbf{x} \\ t \end{pmatrix}$$

Consider y being standard Euclidean coordinates in fixed frame o. x then is coordinates in any other inertial frame.

T=1-dim. affine space since affine structure  $(T, \mathbf{T}, \beta)$  is s.t.  $T \in \mathbb{R}, \ \mathbf{T} \in \mathbb{R}, \ \beta : \mathbf{T} \times T \to T$   $\beta(s, t') = t' + s = t$ 

Holm (2011) [21]

### 16. Fundamental Theorems of (Classical) Dynamics

cf. Sec. 3.4 - Fundamental Theorems of Dynamics of Prástaro (1996) [13].

cf. Thm. 3.20 of Prástaro (1996) [13]

**Theorem 21** (Momentum Theorem). Variation of the free part of momentum of the observed motion of 1 body, in time interval  $\Delta t \equiv [0, t]$  is equal to the corresponding impulse:

$$I[0,t] \equiv \int_0^t F dt \equiv \left(\int_0^t F^j dt\right) \mathbf{e}_j$$

where  $\{\mathbf{e}_j\}_{1\leq j\leq 3}$  is a fixed basis of **S** 

$$\overline{p}_{\psi}(t) - \overline{p}_{\psi}(0) = I[0, t]$$

Proof.

$$\overline{p}_{\psi} = \mu \ddot{m}_{\psi} = \overline{f}_{\psi} \Longrightarrow \dot{\overline{p}_{\psi}} = \dot{p}^{j} \mathbf{e}_{j} = F^{j} \mathbf{e}_{j} \Longrightarrow \int_{[0,t]} \dot{p}^{j} dt = \int_{[0,t]} F^{j} dt$$

<sup>5&</sup>quot; Affine structures on Lie algebras" by Michel Goze and Elisabeth Remm

16.1. Transformation of reference frames. cf. Sec. 4.2.1 "Transformation of reference frames", Kambe (2009) [22]. Let us take pt.  $\mathbf{y}_b$  fixed to body,

located at  $\mathbf{x} = X^1 \mathbf{e}_1 + X^2 \mathbf{e}_2 + X^3 \mathbf{e}_3 = X^i \mathbf{e}_i$  at t = 0,

where  $\mathbf{e}_i$  (i = 1, 2, 3) are orthonormal basis fixed to space  $F \equiv \text{nonrotating fixed space}$  (inertial space)

By transformation matrix  $A = (A^i_j) \in SO(3)$ , initial pt.  $\mathbf{x} = (X^i)$  mapped to current pt.  $\mathbf{y}_b$  at time t.

(69) 
$$\mathbf{y}_b(t) = A(t)\mathbf{x} = y^i(t)\mathbf{e}_i, \quad y^i(t) = A^i_{\ i}(t)X^j$$

cf. Eq. (4.14) of Kambe (2009) [22].

$$\mathbf{y}_b(t) = A(t)\mathbf{x} = A^i_b e_i \otimes e^k x^l e_l = A^i_b X^k e_i = y^i(t)e_i$$

EY: Observe that A(t),  $t \in [0, t]$  is a curve in SO(3).

On the other hand, relative to body frame  $F_B$ , which is the frame instantaneously fixed to the moving body. For the same pt.,  $\mathbf{y}_b(t)$ , fixed to body, is

$$\mathbf{Y}(=\mathbf{y}_b) = Y^1 \mathbf{b}_1 + Y^2 \mathbf{b}_2 + Y^3 \mathbf{b}_3 = Y^i \mathbf{b}_i$$

where  $\mathbf{b}_i$  (i = 1, 2, 3) are orthonormal basis fixed to body which coincided with  $\mathbf{e}_i$  (i = 1, 2, 3) at t = 0. From property of a rigid body, it's required that  $Y^i = X^i$  which don't change with t. According to Sec. 1.5.1(b) (Kambe (2009) [22]), basis transformation written as (1.40) Kambe (2009) [22], where

(70) 
$$\mathbf{b}_k(t) = B_k^i(t)\mathbf{e}_i \text{ hence } Y = Y^k\mathbf{b}_k = Y^kB_k^i\mathbf{e}_i$$

(4.10) Kambe (2009) [22], where  $B = (B^i_k) \in SO(3)$ , i.e.  $BB^T = 1$ . For the basis transformation derivation, recall the push forward transformation (cf. Kambe (2009) [22]).

$$\phi: M \to N$$

$$\phi_*: T_x M \to T_y N$$

$$y = \phi(x) = F(x)$$

$$Y = \phi_* X = F_* X = \frac{d}{dt} (F(p(t))) \Big|_{t=0}$$

$$Y^k = (\phi_* X)^k = \left(\frac{\partial F^k}{\partial x^j}\right) X^j$$

$$Y = \phi_* X = \phi_* \left[ X^j \frac{\partial}{\partial x^j} \right] = X^j \phi_* \left[ \frac{\partial}{\partial x^j} \right] = X^j \frac{\partial y^k}{\partial x^j} \frac{\partial}{\partial y^k} = Y^k \frac{\partial}{\partial y^k}$$

$$Y^k = \frac{\partial y^k}{\partial x^k} X^j$$

17. Fluid Mechanics, Fluid Flow

17.0.1. Courses to consider. MIT 16.01, 16.02

17.0.2. Videos to watch. Engineering MAE 130A. Intro to Fluid Mechanics. Lecture 01. UCI. Follow the rest of the videos.

17.1. **Buoyancy.** From Problem 1.3 of the Japan Physics Olympiad (2014) [34], Ch. 1, General Physics, Elementary Problems. **Problem 1.3.** The part of the iceberg above the sea.

$$\rho_{sw} = 1024 \,\mathrm{kg/}m^3$$
$$\rho_{ice} = 917 \,\mathrm{kg/}m^3$$

with  $sw \equiv \text{sea water.}$ 

Find the ratio of the volume of the part of the iceberg above the sea to whole volume of the iceberg.

#### Solution 1.3.

The buoyancy force exerted on the iceberg is equal to the weight of the seawater displaced by the iceberg.

$$\rho_{sw}gV_{is} = \rho_{ice}gV_{ice} \Longrightarrow \frac{\rho_{sw}}{\rho_{ice}} = \frac{V_{ice}}{V_{is}}$$

with iceberg in sea  $\equiv$  is.

iceberg above sea  $\equiv$  as

$$V_{as}/V_{ice} = \frac{V_{ice} - V_{is}}{V_{ice}} = 1 - \frac{V_{is}}{V_{ice}} = 1 - \frac{\rho_{ice}}{\rho_{sw}} = 1 - \frac{917 \,\text{kg/}m^3}{1024 \,\text{kg/}m^3} = \frac{1024 - 917}{1024} = \frac{107}{1024}$$

The buoyancy on a body equals the resultant force due to the pressure exerted by the surrounding fluid, i.e. The buoyancy force exerted on a body (iceberg) is equal to the resultant force due to pressure exerted by the surrounding fluid (seawater).

Pressure on body of volume V due to its surrounding fluid (whose density is  $\rho$ ) acts perpendicularly to the boundary surface between the body and the fluid.

Since fluid pressure at a deep location is greater than that at shallow location, resultant force due to pressure on boundary surface points upward. This resultant force is the buoyancy, F, acting on body.

Consider region of fluid with same volume V as body.

buoyancy  $F \equiv$  force exerted vertically on body by its surrounding fluid

 $\underbrace{=}_{\text{balance gravitational force acting on }V} \rho g V$ 

For a body floating in a fluid, the magnitude of the buoyancy force acting on the body is equal to the magnitude of the gravitational force on the fluid displaced by the part of the body submerged in the fluid.

17.2. Fluids as a State of Matter. cf. Pert (2013) [23], Ch. 1 Introduction, 1.1 Fluids as a State of Matter.

fluid - isotropic, (no directional preference; identical values of a property in all directions),

locally homogeneous, macroscopic material,

whose particles are free to move within constraints established by the dynamical laws of continuum physics.

requirement that fluid be a continuum - if volume of fluid successively subdivided into smaller elements, each element will remain structurally similar to its parent, and this subdivision can be carried out down to infinitesimal volumes.

**fluid particle** - fictitious particle fixed within fluid continuum and moving with flow velocity; represents average over large number of microscopic particles.

**fluid point** - fixed in fluid moving with flow velocity;

- fluid particle always situated at same fluid point.

**infinitesimal volume** - within continuum of fluid, large compared with microscopic scales, but small compared with macroscopic ones.

fluid is made up of discrete microscopic particles, namely molecules, distributed randomly, with distribution of velocities characteristic of fluid in termal equilibrium, typically Maxwell-Boltzmann distribution in a gas.

At typical experimental densities,

intermolecular separation extremely small and very much less than lab scale, i.e.  $l_s \ll L_0$ 

possible to average over small volumes which contain very large number of particles, yet very small on lab scale, N-V in  $V_0$ ,  $V_0^{1/3} \ll L_0$ 

⇒ allow us to recover continuum approximation

obtain terms which characterize fluid as bulk material

Typical average quantities:

Density,  $\rho$ : number of mass particles per unit volume.

temperature,  $\tau$ : average energy of random motion per particle in thermal equilibrium

pressure p average momentum flow associated with random motion per unit area.

flow velocity u - mean velocity of molecules averaging out random motion.

role of collisions amongst particles important in defining irreversibility through loss of correlation between particles.

mean free path  $l_0$  - particles collide on average after distance equal to  $l_0$  and time after collision interval

since fluid mechanics assumes fluid particles in thermal equilibrium and randomly distributed, this condition requires that spatial and temporal averages be taken to include large number of collisions; i.e. lab scale length  $L_0$  large compared with  $l_0$ , and time to collision interval

effects of collisions on fluid transport (momentum and energy) are averaged over thermal distribution to yield bulk properties of material, namely viscosity and thermal conduction, respectively.

theory: flow of basic fluid may be calculated using Newtonian mechanics, classical thermodynamics, viscosity and thermal conductivity values.

Conditions under which this theory may be applied:

- lab-scale lengths  $L_0$  must be large compared with intermolecule separation and mean free path  $l_0$
- characteristic lab times  $T_0$  must be large compared with collision interval
- fluid locally in thermal equilibrium

cf. 1.2 The Fundamental Equations for Flows of a Dissipationless Fluid. Pert (2013) [23]

inviscid limit - viscosity and thermal conduction are neglected

adiabatic flow - specific entropy of fluid particle constant in time

isentropic flow - specific entropy of each fluid particle has same initial value

many flows are both isentropic and adiabatic, e.g. ideal steady fluid flow, whose specific entropy an entry is everywhere constant.

Eulerian frame - lab frame, coordinates fixed in space and time, derivatives are usual partial derivatives

$$\left. \frac{\partial}{\partial t} \right|_{\mathbf{r}}, |\nabla|$$

Consider coordinate system  $(t, x^1, x^2, x^3)$  and so each vector field  $\mathbf{e}_j$  forming basis of tangent space at each pt. expressed as linear combination of coordinate vector fields

$$\frac{\partial}{\partial t}\Big|_{\mathbf{r}}, \frac{\partial}{\partial x^1}\Big|_{t}, \frac{\partial}{\partial x^2}\Big|_{t}, \frac{\partial}{\partial x^3}\Big|_{t},$$

in Lagrangian frame, spatial variation seen by particle due to its motion is absorbed into time derivative:

$$\frac{d}{dt} = \frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla$$

Rather, consider total time derivative of integral of n-form:

$$\frac{d}{dt} \int_{V_0} \omega = \int_{V_0} \mathcal{L}_{\frac{\partial}{\partial t} + \mathbf{v}} \omega$$

Volume form  $\Omega^n(M) = \bigwedge^n(T^*M)$ 

$$\operatorname{vol}^n \in \Gamma(\Omega^n(M)) = \Gamma(\bigwedge^n T^*M)$$

i.e. vol<sup>n</sup> is n-form, a section of the line bundle,  $\Omega^n(M) = \bigwedge^n T^*M$ 

Now

$$vol^{n} = \frac{\sqrt{g}}{n!} \epsilon_{i_{1}, \dots, i_{n}} dx^{i_{1}} \wedge \dots \wedge dx^{i_{n}}$$

Consider

$$\operatorname{vol}^{n+1} = \frac{\sqrt{g}}{(n+1)!} \epsilon_{i_0 i_1, \dots i_n} dx^{i_0} \wedge dx^{i_1} \wedge \dots \wedge dx^{i_n}$$

17.3. Lagrangian frame. cf. Pert (2013) [23] 1.3 Lagrangian Frame.

**Lagrangian frame** - frame of the moving particle;

Lagrangian reference frame considers fluid from point of view of an observer on a fluid particle.

fluid particle may be conveniently identified by coordinate set, which is fixed on the particle,

 $\Lambda = (\lambda, \mu, \nu)$  i.e. triad of numbers.

e.g.  $\Lambda$  maybe initial position of particle  $\mathbf{r}_0 = (x_0, y_0, z_0)$ , i.e. fluid particle labeled by some (time-independent) vector field  $\mathbf{x}_0 \equiv \Lambda$ .

⇒ position, velocity, thermodynamic state of particles are therefore functions of time alone.

 $\Longrightarrow$  leads to simple set of kinematic and dynamic relations governing motion of particle

$$\mathbf{r} = \mathbf{r}(t) = (x(t), y(t), z(t)) = (x^{1}(t), x^{2}(t), x^{3}(t))$$
$$\dot{\mathbf{r}} \equiv \dot{\mathbf{r}}(t) \equiv \frac{d\mathbf{r}}{dt} \equiv \frac{d\mathbf{r}}{dt}(t) = \mathbf{v} \equiv \mathbf{v}(t)$$
$$d\mathbf{r}$$

or

(71) 
$$\frac{d\mathbf{r}}{dt} = \mathbf{v}$$

$$\frac{d\mathbf{v}}{dt} = \frac{\mathbf{F}}{m}$$

cf. Eq. (1.2) of Pert (2013) [23], i.e.

flow described by function  $\mathbf{X}(\mathbf{x}_0,t)$  giving position of particle labeled at  $\mathbf{x}_0$  at time t.

17.4. Momentum conservation in Lagrangian Frame-Euler's Equation. Recall that a connection  $\nabla$  on smooth manifold  $(M, \mathcal{O}, \mathcal{A})$  is a map that takes a pair consisting of a vector (field) X and (p, q)-tensor field T and sends them to a (p, q)-tensor (field)  $\nabla_X T$ , i.e.

$$\nabla: \mathfrak{X}(M) \times (\otimes^p TM \otimes^q T^*M) \to \otimes^p TM \otimes^q T^*M$$

.t.

$$\begin{split} \nabla_X f &= Xf \quad \forall \, f \in C^\infty(M) \\ \nabla_X (T+S) &= \nabla_X T + \nabla_X S \\ \nabla_X (T(\omega,Y)) &= (\nabla_X T)(\omega,T) + T(\nabla_X \omega,Y) + T(\omega,\nabla_X Y) \\ \nabla_{fX+Z} T &= f \nabla_X T + \nabla_Z T, \quad f \in C^\infty(M) \quad \text{``Leibnitz''} \text{ rule} \end{split}$$

cf. Schuller, International Winter School on Gravity and Light held central lectures.

Remember that in the Lagrangian frame, we are an observer following an individual fluid parcel or i.e. "fluid particle" as it moves through space and time. This fluid particle will obey Newton's 2nd. law:  $\nabla_u u = \frac{F}{m} \iff m \cdot \mathfrak{a} = F$ .

$$\nabla_{u}u = \nabla_{u^{\mu}}\frac{\partial}{\partial x^{\mu}}(u^{\nu}\frac{\partial}{\partial x^{\nu}}) = u^{\mu}\nabla_{\frac{\partial}{\partial x^{\mu}}}(u^{\nu}\frac{\partial}{\partial x^{\nu}}) = u^{\mu}\left(\nabla_{\frac{\partial}{\partial x^{\mu}}}u^{\nu}\right)\frac{\partial}{\partial x^{\nu}} + u^{\mu}u^{\nu}\nabla_{\frac{\partial}{\partial x^{\mu}}}\frac{\partial}{\partial x^{\nu}} = u^{\mu}\frac{\partial u^{\nu}}{\partial x^{\mu}}\frac{\partial}{\partial x^{\nu}} + u^{\mu}u^{\nu}\Gamma_{\mu\nu}^{\gamma}\frac{\partial}{\partial x^{\gamma}}$$

Let  $u^0 = 1$ 

$$\nabla_u u = \frac{\partial u^a}{\partial t} \frac{\partial}{\partial x^a} + u^b \frac{\partial u^a}{\partial x^b} \frac{\partial}{\partial x^a} + u^\mu u^\nu \Gamma^a_{\mu\nu} \frac{\partial}{\partial x^a}$$

where  $\Gamma_{ba}^{0} = 0$  since  $\nabla dt = 0$  and we choose a stratified atlas  $\mathcal{A}_{\text{stratified}}$ 

Then let

$$\int \rho \operatorname{vol}^n \nabla_u u = F_{\text{tot}} = \int f^i \operatorname{vol}^n \otimes e_i$$

From Sec. 2.4 Equations of Motion Neglecting Viscosity, pp. 38, Sabersky, et. al. (1998) [24], the principal forces are those due to pressures acting on surfaces and to forces acting directly on a mass of the particle:

$$\int_{\partial V} p dS^{i} \otimes e_{i} = \int_{V} \mathbf{d}(p dS^{i} \otimes e_{i}) = \int_{V} \frac{\partial p}{\partial x^{j}} dx^{j} \wedge \left(\frac{\sqrt{g}}{(n-1)!} \epsilon^{i}_{i_{2}, \dots i_{n}} dx^{i_{2}} \wedge \dots \wedge dx^{i_{n}}\right) \otimes e_{i} =$$

$$\int_{V} \frac{\partial p}{\partial x^{j}} g^{ij} \operatorname{vol}^{n} \otimes e_{i} = \int_{V} (\operatorname{grad} p)^{i} \operatorname{vol}^{n} \otimes e_{i}$$

$$\Longrightarrow \int \rho \operatorname{vol}^{n} \nabla_{u} u = -\int_{V} (\operatorname{grad} p)^{i} \operatorname{vol}^{n} \otimes e_{i} + \int_{V} f^{i} \operatorname{vol}^{n} \otimes e_{i}$$

$$\Longrightarrow \rho(\nabla_{u} u)^{i} = -(\operatorname{grad} p)^{i} + f^{i}$$

$$(\nabla_{u} u)^{i} = -\frac{1}{\rho} (\operatorname{grad} p)^{i} + f^{i}$$

describes the motion of a perfect fluid.

Assuming a non-rotating, inertial frame (so that we can set the Christoffel symbols to 0),

(73) 
$$\frac{\partial}{\partial t}u^{i} + u^{b}\frac{\partial u^{i}}{\partial x^{b}} = -\frac{1}{\rho}(\operatorname{grad} p)^{i} + f^{i}$$

which describes the motion of the perfect fluid in an inertial ref. frame. Compare this to Eqns. (2.10a- 2.10c) of Sabersky (1998) [24]. Sabersky says that no assumptions were made about density so these results apply to compressible fluid as well as incompressible ones. Also Eqns. 72, 73 are called either Euler's Equation by Pert (2013) [23] or Eulerian equation by Sabersky, et. al. (1998) [24]

17.5. Mass Conservation for Fluid Flow, Continuum media. The mass of fluid in some volume  $V_0 \subset N$  is  $\int_{V_0} \rho \text{vol}^n$ , where  $\rho$  is fluid density,  $\rho \in C^{\infty}(N)$ .

The total mass of fluid flowing out of volume  $V_0$  is

$$\frac{d}{dt} \int_{V_0} \rho \text{vol}^n = \int_{V_0} \mathcal{L}_{\frac{\partial}{\partial t} + \mathbf{u}}(\rho \text{vol}^n) = \int_{V_0} \frac{\partial}{\partial t} \rho \text{vol}^n + \int_{V_0} \mathcal{L}_u \rho \text{vol}^n$$

$$\int_{V_0} \mathcal{L}_u \rho \text{vol}^n = \int_{V_0} di_{\mathbf{u}} \rho \text{vol}^n + i_{\mathbf{u}} d\rho \text{vol}^n = \int_{V_0} di_{\mathbf{u}} \rho \text{vol}^n + 0 = \int_{V_0} di_{\mathbf{u}} \rho \text{vol}^n = \int_{\partial V_0} i_{\mathbf{u}} \rho \text{vol}^n$$

Now

(72)

$$i_u \operatorname{vol}^n = i_u \frac{\sqrt{g}}{n!} \epsilon_{i_1 \dots i_n} dx^{i_1} \wedge \dots \wedge dx^{i_n}$$

 $i_n dx^{i_1} \wedge \cdots \wedge dx^{i_n} = u^{i_1} dx^{i_2} \wedge \cdots \wedge dx^{i_n} - dx^{i_1} \wedge u^{i_2} dx^{i_3} \wedge \cdots \wedge dx^{i_n} + \cdots + (-1)^{n+1} dx^{i_1} \wedge \cdots \wedge dx^{i_{n-1}} u^{i_n} = \epsilon^{i_1 \dots i_n}_{i_1 \dots i_n} u^{j_1} dx^{j_2} \wedge \cdots$  In the reason for f to be a 1-form is that we integrate differential forms, we don't integrate vectors

We can also rewrite Eq. 74 to be a "surface differential":

(75) 
$$\int_{V_0} \mathbf{d}i_{\mathbf{u}}\rho \operatorname{vol}^n = \int_{\partial V_0} i_{\mathbf{u}}\rho \operatorname{vol}^n = \int_{\partial V_0} \rho \frac{\sqrt{g}}{(n-1)!} u^{j_1} \epsilon_{j_1 j_2 \dots j_n} dx^{j_2} \wedge \dots \wedge dx^{j_n} \equiv \int_{\partial V_0} \rho \mathbf{u} \cdot d\mathbf{S} \equiv \int_{\partial V_0} \rho \langle \mathbf{u}, d\mathbf{S} \rangle$$

If  $\sqrt{g} = 1$ , n = 2,

$$i_u \text{vol}^2 = (u^1 dx^2 - u^2 dx^1) = u \cdot n_1 dx^2 + u \cdot n_2 dx^1 = u \cdot n dS$$

with  $n_1 = e_1$  and  $n_2 = -e_2$ .

$$di_{u}\rho \operatorname{vol}^{n} =$$

$$= \frac{\partial(\sqrt{g}\rho u^{j_{1}})}{\partial x^{k}} \frac{\epsilon_{j_{1}...j_{n}}}{(n-1)!} dx^{k} \wedge dx^{j_{2}} \wedge \cdots \wedge dx^{j_{n}} = \frac{\partial(\sqrt{g}\rho u^{k})}{\partial x^{k}} \frac{\epsilon_{j_{1}...j_{n}}}{n!} dx^{j_{1}} \wedge \cdots \wedge dx^{j_{n}} = \frac{1}{\sqrt{g}} \frac{\partial(\sqrt{g}\rho u^{k})}{\partial x^{k}} \operatorname{vol}^{n} =$$

$$= \frac{\partial(\rho u^{k})}{\partial x^{k}} \operatorname{vol}^{n} + \rho u^{k} \frac{\partial \ln \sqrt{g}}{\partial x^{k}} \operatorname{vol}^{n} = \operatorname{div}(\rho u) \operatorname{vol}^{n} + \rho u^{k} \frac{\partial \ln \sqrt{g}}{\partial x^{k}} \operatorname{vol}^{n}$$

Now if  $\sqrt{g} = 1$ , then

$$\frac{d}{dt} \int_{V_0} \rho \operatorname{vol}^n = \int_{V_0} \frac{\partial \rho}{\partial t} \operatorname{vol}^n + \int_{V_0} di_u \rho \operatorname{vol}^n = \int_{V_0} \frac{\partial \rho}{\partial t} \operatorname{vol}^n + \int_{V_0} \operatorname{div}(\rho u) \operatorname{vol}^n \Longrightarrow \frac{\partial \rho}{\partial t} + \operatorname{div}(\rho u) = 0$$

which is the so-called mass continuity equation.  $j = \rho u$  is the mass flux density.

Thus,

(mass conservation)

(76) 
$$\dot{m} = m(t) := \int_{V_0} \rho \operatorname{vol}^n, V_0 \subset N$$

$$\dot{m} \equiv \frac{d}{dt} m(t) = \int_{V_0} \left( \frac{\partial \rho}{\partial t} \operatorname{vol}^n + \mathbf{d} i_{\mathbf{u}} \rho \operatorname{vol}^n \right) = \left[ \int_{V_0} \frac{\partial \rho}{\partial t} \operatorname{vol}^n + \int_{\partial V_0} \rho \mathbf{u} \cdot d\mathbf{S} \right]$$

$$\text{if } \sqrt{g} = 1, \text{ and } \dot{m} = 0, \text{ then}$$

$$\frac{\partial \rho}{\partial t} + \operatorname{div}(\rho \mathbf{u}) = 0$$

Sec. 1.4 Eulerian Frame, 1.4.1 Conservation of Mass-Equation of Continuity, pp. 8 of Pert (2013) [23] says that Eq. 76 is for

From wikipedia, recall that the Eulerian specification focuses on specific locations in space through which fluid flow as time passes. Visualize by sitting on bank of river and watch water pass fixed location. In the Eulerian specification of field, field represented as a function of position x, time t, e.g.  $\mathbf{u}(\mathbf{x},t) \equiv$  flow velocity.

TODO: 20190804 Frankel (2012) [30] in pp. 138 and onwards, for Sec. 4.3. Differentiation of Integrals posed the rightful question, "How does one compute the rate of change of an integral when the domain of integration is also changing?" Revisit the derivation from a Lie derivative and 1-parameter flow point of view.

Force should not be represented by a vector but rather by a 1-form. Then

$$f \in \Omega^{1}(N), f = f(t, \mathbf{x}) = f_{j} dx^{j} \ j = 1, \dots, \mathbf{x} = (x^{1}, \dots, x^{n}), \dim N = n$$

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(77) 
$$W = \int_C f \qquad \text{(line integral)}$$

If f conservative,  $\exists$  scalar U = U(x);  $x \in N$  s.t. f = -dU.

Let  $\mathbf{u} = u^j(t, \mathbf{x}) \frac{\partial}{\partial x^j} \in \mathfrak{X}(N)$  generate flow  $\phi_t$ .

The Lie derivative along vector field  $\mathbf{u}$ ,  $\mathcal{L}_{\mathbf{u}}$ , can be calculated in at least 2 ways: from the definition:

$$\mathcal{L}_{\mathbf{u}}\omega = \frac{d}{dt} \mid_{t=0} \phi_t^* \omega$$

or Cartan's magic formula:

(78) 
$$\mathcal{L}_{\mathbf{u}} \operatorname{vol}^{n} = (i_{\mathbf{u}} \mathbf{d} + \mathbf{d} i_{\mathbf{u}}) \operatorname{vol}^{n} = 0 + \mathbf{d} i_{\mathbf{u}} \operatorname{vol}^{n} = \mathbf{d} i_{\mathbf{u}} \left( \frac{\sqrt{g}}{n!} \epsilon_{i_{1} \dots i_{n}} dx^{i_{1}} \wedge \dots \wedge dx^{i_{n}} \right) = \\ = \mathbf{d} \left( \frac{\sqrt{g}}{(n-1)!} \epsilon_{j_{1} \dots j_{n}} u^{j_{1}} dx^{j_{2}} \wedge \dots \wedge dx^{j_{n}} \right) = \frac{\partial (\sqrt{g} u^{j_{1}})}{\partial x^{k}} \frac{\epsilon_{j_{1} \dots j_{n}}}{(n-1)!} dx^{k} \wedge dx^{j_{2}} \wedge \dots \wedge dx^{j_{n}} = \\ = \frac{\partial (\sqrt{g} u^{j_{1}})}{\partial x^{k}} \frac{\epsilon_{j_{1} \dots j_{n}}}{(n-1)!} \epsilon_{k_{1}, \dots, k_{n}}^{k_{j_{2} \dots j_{n}}} dx^{k_{1}} \wedge \dots \wedge dx^{k_{n}} \underbrace{\left(\frac{\sqrt{g}}{\sqrt{g}}\right)}_{=1} = \frac{1}{\sqrt{g}} \frac{\partial (\sqrt{g} u^{k})}{\partial x^{k}} \operatorname{vol}^{n} = \\ = (\operatorname{div} \mathbf{u}) \operatorname{vol}^{n}$$

cf. https://math.stackexchange.com/questions/2566381/lie-derivative-of-volume-form?rq=1

Note that

$$\operatorname{div}: \mathfrak{X}(N) \to C^{\infty}(N)$$
  
$$\mathcal{L}_{\mathbf{n}}: \Omega^{n}(N) \to \Omega^{n}(N)$$

Frankel (2012) [30] on pp. lvii of the "Elasticity and Stresses" section offered this analogy: "While work in particle mechanics pairs a force covector  $(f_i)$  with a contravariant tangent vector  $(dx^i/dt)$  to a curve, work done by traction in elasticity pairs the contravariant stress force 2-form S with the covector valued deformation 1-form E, to yield a scalar valued 3-form.

Instead of pushing along a curve (line) particle to do work, what does it mean to do work on a moving fluid element? One could possibly try to consider the deformation of a volume element under fluid flow. Frankel (2012) [30] in Sec. 4.2. The Lie Derivative of a Form, on pp. 132, asks "If a flow deforms some attribute, say volume, how does one measure the deformation?  $\mathcal{L}_{\mathbf{u}} \text{vol}^n$  "is the *n*-form that reads off the rate of change of volume of a parallelopiped spanned by *n* vectors that are pushed forward by the flow  $\phi_t$ ." So, "in other words,  $\mathcal{L}_{\mathbf{u}} \text{vol}^n$  measures how volumes are changing under the flow  $\phi_t$  generated by  $\mathbf{X}$ " on pp. 133 on Frankel (2012) [30].

Then rate of work done on a fluid element could be the following:

$$W = \int_{V} \mathcal{L}_{\mathbf{u}}(p \text{vol}^{n}) = \int_{V} \mathbf{d}i_{\mathbf{u}} p \text{vol}^{n} = \int_{\partial V} i_{\mathbf{u}} p \text{vol}^{n} = \int_{\partial V} p \mathbf{u} \cdot d\mathbf{S}$$

TODO: work on a fluid element?

### 17.6. Bernoulli's equation.

17.6.1. Simple cases. cf. Ch. 3 The Bernoulli Equation, Sec. 3.2 "A Simple Form of the Bernoulli Equation", pp. 90, Sabersky, et. al. (1998) [24]

Start from Eq. 72. So

$$(\nabla_u u)^i = \frac{-1}{\rho} (\operatorname{grad} p)^i + f^i$$

Suppose  $f^i=0$  (no external forces),  $\frac{\partial u^i}{\partial t}=0$  (steady flow). Suppose  $\Gamma^a_{\mu\nu}=0$  (Cartesian coordinates? TODO: find out the conditions when this is true).

Suppose the flow is irrotational (TODO: find out the condition for irrotational flow; is it when having streamlines or ?). For the curl, recall that an external derivative **d** takes a 1-form into a 2-form. So consider

$$\mathbf{u} \in \mathfrak{X}(V), \mathbf{u} \xrightarrow{\flat} u_i dx^i = g_{ij} u^j dx^i$$
$$d\mathbf{u}^{\flat} \in \Omega^2(V), \ \mathbf{d}\mathbf{u}^{\flat} = \frac{\partial u_j}{\partial x^i} dx^i \wedge dx^j = \frac{\partial (g_{ij} u^j)}{\partial x^a} dx^a \wedge dx^i$$

If we have the condition that the flow is irrotational, i.e.

$$\mathbf{du}^{\flat} = 0$$

then

$$\mathbf{d}\mathbf{u}^{\flat}(\mathbf{u}) = \frac{\partial(u_i)}{\partial x^a} u^a dx^i - \frac{\partial(u_i)}{\partial x^a} dx^a u^i = 0 \Longrightarrow u^a \frac{\partial(u_i)}{\partial x^a} dx^i = u^i \frac{\partial u_i}{\partial x^a} dx^a$$

Take the exterior derivative d of the inner product of a vector with itself in the most general case on a metric manifold:

$$d(u_i u^i) = d(g_{ij} u^j u^i) = \frac{\partial g_{ij}}{\partial x^k} u^j u^i dx^k + g_{ij} \frac{\partial u^j}{\partial x^k} u^i dx^k + g_{ij} u^j \frac{\partial u^i}{\partial x^k} dx^k =$$

$$= 2u_j \frac{\partial u^j}{\partial x^k} dx^k + u^i u^j \frac{\partial g_{ij}}{\partial x^k} dx^k$$

Since the Christoffel symbols are the connection coefficients of a metric-compatible affine connection (i.e. they can be used to construct a connection  $\nabla$  s.t.  $\nabla g = 0$ ), then <sup>6</sup>

(80) 
$$\nabla_i g_{jk} = \frac{\partial g_{jk}}{\partial r^i} - \Gamma^s_{ji} g_{sk} - \Gamma^s_{ki} g_{js} = 0$$

and so

(81) 
$$u^i u^j \frac{\partial g_{ij}}{\partial x^k} = u^i u^j \left[ \Gamma^s_{ik} g_{sj} + \Gamma^s_{jk} g_{is} \right] = u^i u_s \Gamma^s_{ik} + u_s u^j \Gamma^s_{jk} = 2u_s u^i \Gamma^s_{ik}$$

Then

$$\frac{1}{2}\mathbf{d}(u^2) = u_j \frac{\partial u^j}{\partial x^k} dx^k + u_s u^i \Gamma^s_{ik} dx^k$$

Eq. 82 only depended upon the fact that the connection  $\nabla$  is *metric-compatible*. So if we suppose that the Christoffel symbols  $\Gamma_{ik}^s$  are 0, and have irrotational flow, as well,

$$\mathbf{d}(u^2/2) = u_j \frac{\partial u^j}{\partial x^k} dx^k + 0 = u^a \frac{\partial u_i}{\partial x^a} dx^i$$

Now consider the RHS of Eq. 72, with  $f_i = 0$ ,

$$\frac{-1}{\rho}(\operatorname{grad} p)^i \xrightarrow{\flat} \frac{-1}{\rho}(\operatorname{grad} p)^i g_{ij} dx^j = \frac{-1}{\rho} \frac{\partial p}{\partial x^k} g^{ki} g_{ij} dx^j = \frac{-1}{\rho} \frac{\partial p}{\partial x^j} dx^j = \frac{-1}{\rho} \mathbf{d} p$$

So if the fluid is incompressible, then  $\rho$  is constant and so we can write

(83) 
$$\mathbf{d}\left(\frac{u^2}{2} + \frac{p}{\rho}\right) = 0$$

for any irrotational, incompressible, fluid flow with no (external) body forces.

<sup>6&</sup>quot;Prove Christoffel Symbol Identity", Physics StackExchange

17.6.2. Venturi tube example. cf. Example 3.1 in Sabersky, et. al. (1998) [24]

Consider flow of liquid through the tube shown. Determine pressure difference between points 1 and 2 as a function of flow  $A_t$ . rate Q, density  $\rho$ . Velocity assumed constant over cross sections, and pressure will also be assumed uniform over each of these 2 areas  $A_1$ ,  $A_2$ . Also assume there's no viscosity.

Solution Example 3.1 in Sabersky, et. al. (1998) [24].

Since density is constant, continuity equation requires that for Q the volume flow rate,

flow rate
$$Q = u_1 A_1 = u_2 A_2$$

$$\frac{u_1^2}{2} + \frac{p_1}{\rho_1} = \frac{u_2^2}{2} + \frac{p_2}{\rho_2} \Longrightarrow \frac{Q^2}{2A_1^2} + -\frac{Q^2}{2A_2^2} = \frac{-(p_1 - p_2)}{\rho} \text{ or } p_1 - p_2 = \frac{+\rho Q^2}{2A_2^2} \left(1 - \frac{A_2^2}{A_1^2}\right)$$

$$\frac{Q^2}{2} \left(\frac{1}{A_1^2} - \frac{1}{A_2^2}\right) = \frac{-(p_1 - p_2)}{\rho}$$

17.6.3. Conservative body forces with Bernoulli equation. cf. Sec. 3.3 in Sabersky, et. al. (1998) [24] Consider U = U(x), force potential.

$$-\mathbf{d}U = \frac{-\partial U}{\partial x^i} dx^i = f_i dx^i$$
$$\mathbf{d} \left(\frac{u^2}{2} + \frac{p}{\rho} + U\right) = 0$$

17.6.4. Barotropic Fluids ( $\rho = \rho(p)$ ) with Bernoulli equation. cf. Sec. 3.4 in Sabersky, et. al. (1998) [24]

Barotropic fluid: density is single-valued function of p. In the absence of heat transfer and absence of entropy change, by thermodynamics this can be shown to be barotropic.

$$\mathbf{d}\left(\frac{u^2}{2}\right) = -\frac{1}{\rho}\mathbf{d}p - \mathbf{d}U \text{ or } \mathbf{d}\left(\frac{u^2}{2} + U\right) + \frac{1}{\rho}\mathbf{d}p$$
Suppose  $i = i(p)$  s.t.  $di = \frac{\partial i}{\partial p}dp = \frac{1}{\rho(p)} = \frac{di}{dp}dp$ . So
$$\frac{di}{dp} = \frac{1}{\rho(p)} \text{ or } i = \int \frac{1}{\rho}dp + \text{ const}$$

$$\implies \mathbf{d}\left(\frac{u^2}{2} + U + i\right) = 0$$

In special case of isentropic flow of perfect gas,  $\frac{P}{\rho^{\nu}} = \text{const.}$  and  $i = \int \frac{dp}{p^{1/\nu}} = \frac{kp^{1-1/\nu}}{1-1/\nu} = \text{enthalpy.}$ 

17.6.5. Nonsteady flow  $\frac{\partial u}{\partial t}$  with Bernoulli equation. First, use the musical isomorphism to transform the (1,0) tensor:

$$\frac{\partial u^i}{\partial t} \frac{\partial}{\partial x^i} \xrightarrow{\flat} g_{ij} \frac{\partial u^i}{\partial t} dx^j$$

Along a curve 
$$\gamma = \gamma^i(t) = x^i(t)$$
 s.t. Then 
$$\dot{\gamma} = \dot{\gamma}^i(t) \frac{\partial}{\partial x^i} = u^i(t, \gamma(t)) \frac{\partial}{\partial x^i}$$
 
$$\int_{\gamma} g_{ij} \frac{\partial u^i}{\partial t} dx^j = \int g_{ij} \frac{\partial u^i}{\partial t} \dot{\gamma}^j(t) dt = \int g_{ij} \frac{\partial u^i}{\partial t} u^j dt = \int \frac{\partial ||u||}{\partial t} dt$$

The last step can be made only if the metric g is time-independent.

cf. Example 3.2 in Sabersky, et. al. (1998) [24]

Consider the efflux of liquid of density  $\rho$  through an orifice of area  $A_0$  at the bottom of a large tank of cross-sectional area  $A_1$ .

The depth of liquid at the given instant is y.

pressure exerted on the liquid in the tank is maintained at  $p_t$ ,

pressure outside the orifice is the atmospheric pressure  $p_a$ .

The action of gravity is to be considered, but frictional effects may be neglected. Velocity of efflux  $u_j$  is to be determined for these conditions.

efflux - something given off in or as if in a stream (Webster dictionary)

Solution Example 3.2 in Sabersky, et. al. (1998) [24].

In Sec. 3.8(b) flows through orifices discussed in great detail. For this example,

assume the orifice is carefully designed so streamlines at orifice exit are parallel and thus orifice area  $A_0$  and cross-sectional area of jet  $A_i$  are equal there.

cf. Example 3.3 in Sabersky, et. al. (1998) [24]. Example of a problem in which nonsteady terms cannot be neglected.

### 18. Thermodynamics

Let  $\Sigma$  be a (topological) manifold. Suppose U is a global coordinate on  $\Sigma$ :

(First Law of Thermodynamics (energy conservation))

(84) 
$$dU = Q - W \text{ or } Q = dU + W$$

where  $dU, Q, W \in \Omega^1(\Sigma)$  (i.e. dU, Q, W are 1-forms over manifold  $\Sigma$ ).

Consider a path in  $\Sigma$ ,  $\gamma$ ,  $\gamma : \mathbb{R} \to \Sigma$ , and using a chart  $(U, S^1, \dots, S^n)$  (e.g.  $n = 1, S^1 = v$  for volume)

$$\gamma(t) = (U(t), S^{1}(t), \dots, S^{n}(t))$$
$$\dot{\gamma} \in \mathfrak{X}(\Sigma), \, \dot{\gamma} = \dot{U}\frac{\partial}{\partial U} + \dot{S}^{i}\frac{\partial}{\partial S^{i}}$$

Now

$$dU(\dot{\gamma}) = \dot{\gamma}(U) = \dot{U}\frac{\partial}{\partial U}U + 0 = \dot{U}$$
$$Q(\dot{\gamma}) = Q(t)dt\left(\dot{\gamma}\frac{\partial}{\partial t}\right) = Q(t)\dot{\gamma}$$
$$\Longrightarrow \dot{U} = Q(t)\dot{\gamma} - W(t)\dot{\gamma}$$

Recall that for enthalpy  $H, H := U + pV, H = H(\sigma, p)$ 

TODO 20190804 Derive and check convection form of enthalpy against both Kittle and Kroemer plus thermodynamics and Sonntag, et. al.

Incomplete:

$$dH = Q + Vdp$$
 
$$W = pdV = pdV + Vdp - Vdp = d(pV) - Vdp =$$
 
$$dE = Q - W + \mu dH$$
 
$$dE = Q - W + \hat{h}dN$$

 $U \to E$  notation is to promote the internal energy U to include kinetic and potential energies, so that possibly, E = U + K.E. + P.E., or, i.e., E includes internal energy and mechanical energy.

$$\begin{split} G &= \mu N \\ G &= U - \tau \sigma + pV \xrightarrow{U \to E} G = E - \tau \sigma + pV \Longrightarrow \tau \check{h} \\ dE &= Q + W + \mu dN = Q + W + (\check{h} - \tau \check{\sigma}) dN \\ Q &= \tau d\sigma = \tau d(N\check{\sigma}) = \tau N d\check{\sigma} + \tau \check{\sigma} dN \end{split}$$

 $\tau N d\check{\sigma}$  is the entropy change due to change in entropy per particle; i.e. **conduction term**  $\tau \check{\sigma} dN$  is entropy change due to change in number of particles, i.e. **convection term** 

$$dE = Q + W + \check{h}dN - \tau\check{\sigma}dN = \tau Nd\check{\sigma} + \tau\check{\sigma}dN + W + \check{h}dN - \tau\check{\sigma}dN = \tau Nd\check{\sigma} + W + \check{h}dN$$

m(t) = MN

Assume only 1 chemical species:

$$\begin{split} \check{Q} := \tau N d\check{\sigma} & \qquad \widehat{h} := \frac{H}{m} = \frac{H}{MN} \\ \Longrightarrow dE = \check{Q} + W + \check{h} dN \frac{M}{M} = \check{Q} + W + \widehat{h} dm \end{split}$$

The Gibbs free energy equilibrium is given by,

$$dG = \mu dN - \sigma d\tau + V dp$$

For a throttling valve (don't all valves throttle?) the pressure drop is accounted for by the Gibbs free energy, *not* by an isenthalpy condition.

### Part 13. Classical Mechanics applications

cf. Arnold, Kozlov, Neishtadt (2006) [26]. If known forces  $\mathbf{F}_1 \dots \mathbf{F}_n$  acts on points, then

(85) 
$$\sum_{i=1}^{n} \langle m_i \ddot{\mathbf{r}}_i - \mathbf{F}_i, \xi_i \rangle = 0$$

cf. Eq. (1.26) of Arnold, Kozlov, Neishtadt (2006) [26], where  $\xi_1, \ldots \xi_n$  are arbitrary tangent vectors to  $M, \xi_i, \ldots \xi_n \in TM$ .  $\sum_{i=1}^{n} \langle m_i \ddot{\mathbf{r}}_i - \mathbf{F}_i, \xi_i \rangle$  called "general equation of dynamics" or d'Alembert-Lagrange principle.

#### 19. Orbital Mechanics

References: Arnold, Kozlov, Neishtadt (2006) [26] Choquet-Bruhat (2009) [29], pp. 39, 3. General relativity Bate, Mueller, White (1971) [33], Sec. 1.2, pp. 5, https://en.wikipedia.org/wiki/N-body\_problem Abraham and Marsden (2008) [27], Ch. 9 Prástaro (1996) [13], Remark 3.20 many-bodies dynamics, pp. 346

19.1. **Conic sections.** The following is an amalgamation or summary of the development in Ch. 13, Sections 13.18-22, from pp. 497, Apostol (1991) [32]

**Definition 52** (General definition of conic sections (vector-formulated definition)). Given line L, called the directrix, pt. F, the focus, such that  $F \notin L$ .

eccentric e > 0.

 $d = distance \ of \ L \ from \ F, \ then$ 

the conic section  $C = \{X\}$  (set of all points) such that

(86) 
$$||X - F|| = ed(X, L)$$

Now,

e < 1 ellipse,

e=1 parabola.

e > 1 hyperbola.

Further, if N unit normal to L, then  $\forall P \in L$ ,

$$d(X,L) = |(X-P) \cdot N|$$

If  $(X - P) \cdot N \ge 0$ ,  $X \in positive half plane (negative half-plane)$ .

**Theorem 22** (Further conic section development). Building upon Def. 52,

If  $F \in positive\ half\ plane\ determined\ by\ N$ , then  $F - dN \in L$ .

If  $F \in negative \ half \ plane \ determined \ by \ N$ , then  $F + dN \in L$ .

Choose this convenient, specific point:  $F + dN = P \in L$   $\Longrightarrow ||X - F|| = e|(X - (F + dN)) \cdot N|$ 

If 
$$F = 0$$
,  $||X|| = e|X \cdot N - d| = r = e(d - r\cos\theta)$ . Then

(88) 
$$\frac{ed}{r} = 1 + e\cos\theta \text{ or } r = \frac{ed}{1 + e\cos\theta} \text{ for } F = 0$$

Consider symmetry about the origin (i.e. of  $X, \exists -X$ ).

As a caveat, let  $A = e(d + F \cdot N)$ . Then

$$||X - F|| = e|(X - (F + dN)) \cdot N| = ed(X, L) = e|X \cdot N - \frac{A}{e}|$$

Then

(89) 
$$||X - F||^2 = ||X||^2 - 2X \cdot F + ||F||^2 = e^2((X \cdot N)^2 - 2\frac{A}{e}X \cdot N + \left(\frac{A}{e}\right)^2)$$

Using the summetry about the origin,  $X \to -X$ .

$$\implies X \cdot F = AeX \cdot N \text{ or } X \cdot (F - AeN) = 0$$

Then F = AeN. Note that  $A = e(F \cdot N \pm d)$  (for F in negative (positive) half-plane). Then

$$Ae = e^{2}(Ae + d) \text{ or } \frac{A}{e} = Ae + d \text{ or } A\left(\frac{1}{e} - e\right) = d \text{ or } A = \frac{ed}{1 - e^{2}}$$
$$\implies ||X - AeN|| = e|(X - (AeN + dN)) \cdot N| = e|X \cdot N - (Ae + d)|$$

Let X = aN. Plug this into Eq. 89.

$$a^{2} - 2aAe + A^{2}e^{2} = e^{2}\left(a^{2} - \frac{2A}{e}a + \left(\frac{A}{e}\right)^{2}\right) = e^{2}a^{2} - 2Aae + A^{2} \Longrightarrow a = A$$
$$\Longrightarrow a = \frac{ed}{1 - e^{2}}$$

Thus, we've shown that A is equal to the semi-major axis a and that

$$(90) F = aeN$$

Thus, the semi-major axis is obtained when  $X = \pm aN$  (vertices), so that  $||X||^2 + (ae)^2 = e^2(X \cdot N)^2 + a^2$  is satisfied. Likewise, the semi-minor axis is obtained if  $X = \pm bN_{\perp}$ , so that

(91) 
$$b^{2} + (ae)^{2} = 0 + a^{2} \Longrightarrow b = a\sqrt{1 - e^{2}}$$

Compare this against (26.6) of Greiner (2004) [25]:

(92) 
$$r = \frac{ed}{1 + e\cos\theta}, \quad e = \frac{C\widehat{L}_0^2}{GM}, \quad d = 1/C$$

cf. Eq. (26.32) of Greiner (2004) [25].

Investigate which physical quantities (e.g. E, L) eccentricity e depends on.

$$u = \frac{GM}{\widehat{L_0}^2} + C\cos(\theta - \theta_0) \Longrightarrow \frac{du}{d\theta} = -C\sin(\theta - \theta_0)$$

$$\frac{1}{2}mL_0^2 \left[ C^2\sin^2(\theta - \theta_0) + \left(\frac{GM}{L_0^2}\right)^2 + 2C\frac{GM}{L_0^2}\cos(\theta - \theta_0) + C^2\cos^2(\theta - \theta_0) \right] =$$

$$\Longrightarrow \qquad = \frac{1}{2}m\widehat{L_0}^2 \left[ C^2 + \left(\frac{GM}{\widehat{L_0}^2}\right)^2 + 2C\frac{GM}{\widehat{L_0}^2}\cos(\theta - \theta_0) \right] =$$

$$= E - V\left(\frac{1}{u}\right) = E + GMmu = E + \frac{(GM)^2}{\widehat{L_0}^2}m + CGMm\cos(\theta - \theta_0)$$

$$\Longrightarrow \frac{1}{2}m\widehat{L_0}^2C^2 = E + \frac{1}{2}\frac{(GM)^2}{\widehat{L_0}^2}m$$

Solve for the dummy constant C:

(93) 
$$C = \sqrt{\frac{2E}{m\hat{L}_0^2} + \frac{(GM)^2}{\hat{L}_0^4}}$$

cf. Eq. (26.36) of Greiner (2004) [25].

From Eq. (26.32) of Greiner (2004) [25] or Eq. 92,

$$e = \frac{C\widehat{L}_0^2}{GM} = \sqrt{\frac{2E}{m\widehat{L}_0^2} + \frac{(GM)^2}{\widehat{L}_0^4}} \left(\frac{\widehat{L}_0^2}{GM}\right)$$

So

(94) 
$$e = \sqrt{\frac{2E\hat{L}_0^2}{m(GM)^2} + 1}$$

cf. Eq. (26.37) of Greiner (2004) [25].

Hence, shape of the path depends on total energy E, angular momentum  $L_0 = m\widehat{L}_0$  of moving body (recall that  $\widehat{L} := \mathbf{r} \times \mathbf{v}$ ).

It holds that for a

parabola 
$$e=1,$$
 
$$E=0$$
 ellipse  $0< e<1,$  
$$E<0,$$
 
$$\frac{(GM)^2m}{-2\widehat{L}_0^2}< E<0$$
 circle  $e=0$  
$$E=\frac{-(GM)^2m}{2\widehat{L}_0^2}$$
 hyperbola  $e=1,$  
$$E>0$$

The general form of initial energy  $E_0$  is, recall, give by

$$E_0 = \frac{1}{2}mv^2 + V(r) = \frac{1}{2}mv^2 - \frac{GMm}{r_0} = m\left(\frac{v^2}{2} - \frac{GM}{r_0}\right)$$

Notice that for the case of the parabola that  $E_0 = 0$  when

$$(95) v = \sqrt{\frac{2GM}{r_0}}$$

Eq. 95 is the expression for the escape velocity.

Knowing what we know now, let's rewrite some of the expressions from before:

$$r(\theta) = \frac{\hat{L}_{0}^{2}/GM}{1 + \frac{C\hat{L}_{0}^{2}}{GM}\cos(\theta - \theta_{0})}$$
$$\hat{L}_{0}^{2} := r_{0}^{2}\dot{\theta}_{0}$$
$$C = \sqrt{\frac{2E}{m\hat{L}_{0}^{2}} + \frac{(GM)^{2}}{\hat{L}_{0}^{4}}}$$

19.2. Path parameters (i.e. orbital parameters). cf. pp. 262, Greiner (2004) [25].

$$a = \frac{1}{2}(r(\theta = 0) + r(\theta = \pi)) = \frac{1}{2} \left[ \frac{\widehat{L}_0^2/GM}{1 + e} + \frac{\widehat{L}_0^2/GM}{1 - e} \right] = \frac{\widehat{L}_0^2/GM}{1 - e^2}$$

The addition above for a is needed because we center origin at a focus, for this expression.

(96) 
$$a = \frac{\hat{L}_0^2/GM}{\frac{-2E\hat{L}_0^2}{m(GM)^2}} = \frac{GMm}{-2E}$$

cf. Eq. (26.39) of Greiner (2004) [25].

(97) 
$$b = \sqrt{a^2 - f^2} = \sqrt{a^2 - e^2 a^2} = a\sqrt{1 - e^2} = a\sqrt{\frac{-2E\hat{L}_0^2}{m(GM)^2}} = \hat{L}_0\sqrt{\frac{-m}{2E}}$$

cf. Eq. (26.40) of Greiner (2004) [25]

We can also calculate the perigee (point of closest approach to mass M):

(98) 
$$r(\theta = 0) = \frac{\widehat{L}_0^2 / GM}{1 + e} = \frac{\widehat{L}_0^2 / GM}{1 + \sqrt{\frac{2E\widehat{L}_0^2}{m(GM)^2} + 1}}$$

Given Eq. 96 for a, plug in for E:

(99) 
$$a = \frac{-GMm}{2E} = -\frac{GMm}{2\left(\frac{1}{2}mv_0^2 - \frac{GMm}{r}\right)} = \frac{-GM}{v_0^2 - \frac{2GM}{r}} = \frac{-GM}{v_0^2\left(1 - \frac{2GM}{rv^2}\right)}$$

Notice when the denominator is 0. It is when  $v_0$  is the the escape velocity (cf. Eq. 95).

19.2.1. Eccentricity from  $r_0$ ,  $\mathbf{v}_0$ . Consider being given  $\mathbf{r}$ ,  $\mathbf{v}$ , with  $\mathbf{r}$  in a coordinate frame with its origin at the center of mass (or, if M large enough, from the focus where M is centered at).

The first step is to use a coordinate *frame* (set of orthonormal unit vectors) with  $\mathbf{r}$ ,  $\mathbf{v}$  forming a *plane*. Note this important fact from geometry: a plane is spanned by 2 non-collinear vectors. So  $\mathbf{r}$ ,  $\mathbf{v}$  better not be parallel to each other.

Let

$$\mathbf{v} = v^x \mathbf{e}_x + v^y \mathbf{e}_y + v^z \mathbf{e}_z = v^r \mathbf{e}_r + v^{\varphi} \mathbf{e}_{\varphi}$$
$$\mathbf{r} = x \mathbf{e}_x + y \mathbf{e}_y + z \mathbf{e}_z = r \mathbf{e}_r$$

Now

$$(\mathbf{r} \times \mathbf{v})_i = \epsilon_{iki} r_i v_k$$

To get the vector *orthogonal* to the plane that is formed by  $\mathbf{r}, \mathbf{v}$ ,

mogorial to the plane that is formed by 
$$\mathbf{r}, \mathbf{v}$$
,
$$((\mathbf{r} \times \mathbf{v}) \times \mathbf{r})_m = \epsilon_{ijm} \epsilon_{jki} r_j v_k r_l = \epsilon_{ilm} \epsilon_{ijk} r_j v_k r_l = \\ = r^2 v_m - v_k r_k r_m = r^2 v_m - v_k r_k r_m \qquad \text{or } ((\mathbf{r} \times \mathbf{v}) \times \mathbf{r}) = r^2 \mathbf{v} - (\mathbf{v} \cdot \mathbf{r}) \mathbf{r}$$

Then

$$\widehat{L}_0 = r^2 \dot{\theta} = r \left( \mathbf{v} \cdot \frac{((\mathbf{r} \times \mathbf{v}) \times \mathbf{r})}{|(\mathbf{r} \times \mathbf{v}) \times \mathbf{r}|} \right) = r \left( \frac{r^2 v^2 - (\mathbf{v} \cdot \mathbf{r})^2}{|r^2 \mathbf{v} - (\mathbf{v} \cdot \mathbf{r})\mathbf{r}|} \right)$$

Thus, we can find the following:

$$e^{2} = 1 + \frac{2E\hat{L}_{0}^{2}}{m(GM)^{2}} = 1 + \frac{2\left(\frac{1}{2}mv_{0}^{2} - \frac{GMm}{r}\right)\hat{L}_{0}^{2}}{m(GM)^{2}} = 1 + \frac{\left(v_{0}^{2} - \frac{2GMm}{r}\right)\hat{L}_{0}^{2}}{(GM)^{2}}$$
$$a = \frac{-GMm}{2E} = -\frac{GMm}{2\left(\frac{1}{2}mv_{0}^{2} - \frac{GMm}{r}\right)} = \frac{-GM}{v_{0}^{2} - \frac{2GM}{r}} = \frac{-GM}{v_{0}^{2}\left(1 - \frac{2GM}{rv_{0}^{2}}\right)}$$

TODO: Calculate period of revolution.

Here are some calculations (forms) that can help with further calculations:

(100) 
$$\widehat{L}_0^2 = b^2 \left(\frac{2E}{-m}\right) \text{ or } \frac{\widehat{L}_0^2}{GM} = b^2 \left(\frac{2E}{-GMm}\right) = \frac{b^2}{a} = a(1 - e^2)$$

using Eq. 91, for  $e^2 = 1 - (\frac{b}{a})^2$ .

19.3. **Path in coordinates.** For **r** centered at focus F and **r**, **v** expressed in coordinates in the plane of the conic section path, i.e. in the plane formed by **r**, **v**, with the axes expressed as  $\frac{\partial}{\partial \mathbf{r}}$ ,  $\frac{\partial}{\partial \mathbf{r}}$ ,

(101) 
$$\mathbf{r} = r\cos\theta \frac{\partial}{\partial x} + r\sin\theta \frac{\partial}{\partial y}$$

Compare this to Eq. (2.5-1) of Sec. 2.5.1 "Expressing r and v in the Perifocal System" of Bate, Mueller, White (1971) [33]. Recall the equation for the conic section in polar coordinates, with the coordinates centered at focus F:

(102) 
$$r = \frac{a(1 - e^2)}{1 + e \cos \theta}$$

Compare this to Eq. (1.5-4) of Bate, Mueller, White (1971) [33].

(103) 
$$\mathbf{v} = \dot{\mathbf{r}} = (\dot{r}\cos\theta - r\dot{\theta}\sin\theta)\frac{\partial}{\partial r} + (\dot{r}\sin\theta + r\dot{\theta}\cos\theta)\frac{\partial}{\partial u}$$

Compare this to the equation on pp. 72, of Sec. 2.5.1 "Expressing r and v in the Perifocal System" of Bate, Mueller, White (1971) [33].

Now

Now

$$\dot{r} = \frac{a(1 - e^2)}{(1 + e\cos\theta)^2} e\sin\theta \\ \dot{\theta} = \frac{r^2}{a(1 - e^2)} e\sin\theta \\ \dot{\theta} = \frac{\widehat{L}_0 e\sin\theta}{r^2} = \frac{\widehat{L}_0 e\sin\theta}{a(1 - e^2)} = \frac{\widehat{L}_0 e\sin\theta}{\frac{-GMm}{2E} \left(\frac{L_0^2}{\left(\frac{(GM)^2m}{-2E}\right)}\right)} = \frac{1}{2E} \left(\frac{L_0^2}{\left(\frac{(GM)^2m}{-2E}\right)}\right) = \frac{1}{2E}$$

$$\Longrightarrow \dot{r} = \frac{GM}{\hat{L}_0} e \sin \theta$$

Compare this to Eq. (2.5-2) of Sec. 2.5.1 "Expressing r and v in the Perifocal System" of Bate, Mueller, White (1971) [33]. Then

$$\Rightarrow \frac{a(1-e^2)}{(1+e\cos\theta)^2}\dot{\theta} = \frac{GM}{\widehat{L}_0} \text{ or}$$

$$r\dot{\theta} = \frac{GM}{(1+e\cos\theta)}$$

(105) 
$$r\dot{\theta} = \frac{GM}{\hat{L}_0} (1 + e\cos\theta)$$

Compare this to Eq. (2.5-3) on pp. 73 of Sec. 2.5.1 "Expressing r and v in the Perifocal System" of Bate, Mueller, White (1971) [33].

Then

$$\mathbf{v} = \frac{GM}{\widehat{L}_0} \left[ (e \sin \theta \cos \theta - (1 + e \cos \theta) \sin \theta) \frac{\partial}{\partial x} + (e \sin^2 \theta + (1 + e \cos \theta) \cos \theta) \frac{\partial}{\partial y} \right] =$$

(106) 
$$\mathbf{v} = \frac{GM}{\widehat{L}_0} \left[ \sin \theta \frac{\partial}{\partial x} + (e + \cos \theta) \frac{\partial}{\partial y} \right] = \frac{GM}{\widehat{L}_0} \left[ \frac{y}{r} \frac{\partial}{\partial x} + (e + \frac{x}{r}) \frac{\partial}{\partial y} \right]$$

Compare this to Eq. (2.5-4) on pp. 73 of Sec. 2.5.1 "Expressing r and v in the Perifocal System" of Bate, Mueller, White (1971) [33].

Consider this ansatz from subsection 4.4.3 "The f and g expressions from Ch. 4 section on "Position and velocity - A function of time" in Bate, Mueller, White (1971) [33].

Because r, v lie in the plane (i.e. coplanar) containing the entire conic section trajectory, then make the following ansatz:

(107) 
$$\mathbf{r} = f\mathbf{r}_0 + g\mathbf{v}_0$$
$$\dot{\mathbf{r}} =: \mathbf{v} = \dot{f}\mathbf{r}_0 + \dot{g}\mathbf{v}_0$$

Solve for f and g by taking the appropriate cross products:

(108) 
$$\mathbf{r} \times \mathbf{v}_0 = f\mathbf{r}_0 \times \mathbf{v}_0$$
$$\mathbf{r} \times \mathbf{r}_0 = q\mathbf{v}_0 \times \mathbf{r}$$

19.4. Eccentricity anomaly; Eccentricity anomaly coordinates. Let the point of interest be P. Choose the origin for the coordinates (x, y) to have its origin at the center of the ellipse.

Let point C be the point at the center of the ellipse (origin), let point A be the projection of point P onto the x-axis, point F be the point at the focus.

Recall that the length from the center C to the focus,  $\overline{\text{CF}}$ , is ae, where a is the semi-major axis, and eccentricity is e. The eccentricity anomaly E (it's just an angle) is given by

$$cos E = \frac{x}{a}$$

Remember E is not the angle formed by triangle CAP, but rather, first take the line through P and A. This line is perpendicular to the x-axis. Then intersect that line with a circle, centered at C, of radius a. Then, the line from C to that intersection forms an angle E with the x-axis.

By using the ellipse equation and trigonometry,

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1 \Longrightarrow \cos^2 E + \frac{y^2}{b^2} = 1 \text{ or } \frac{y^2}{b^2} = \sin^2 E \text{ or }$$

$$\sin E = \frac{y}{b}$$

Now

$$\overline{AF} = \overline{CF} - \overline{CA} = ae - a\cos E$$

Use the right triangle PAF and the hypotenuse being r (distance from the focus F to point P) to find

$$\overline{AF} = r \cos(\pi - \theta) = -r \cos \theta$$

So

$$\cos \theta = \frac{a(\cos E - e)}{r}$$

Then, plug into the conic section equation, Eq. 88,

$$r = \frac{ed}{1 + e\cos\theta} = \frac{a(1 - e^2)}{1 + e\cos\theta} = \frac{b^2/a}{1 + e\cos\theta} = \frac{b^2/a}{1 + e\left(\frac{a(\cos E - e)}{r}\right)}$$

$$\implies 1 = \frac{b^2/a}{r + e(a(\cos E - e))}$$

$$\implies r = \frac{b^2}{a} - e(a(\cos E - e)) = \frac{b^2}{a} + ae^2 - ea\cos E = a(1 - e\cos E)$$

$$(111) r = a(1 - e\cos E)$$

Compare this to Eq. (4.6-9) on pp. 215 of Sec. 4.6 in "The Kepler Problem" of Bate, Mueller, White (1971) [33]. For the coordinates with the origin at the focus F, it is given by

(112) 
$$x = a(\cos E - e)$$
$$y = b\sin E = a\sqrt{1 - e^2}\sin E$$

Compare this to Eqns. (4.6-7), (4.6-8) on pp. 215 of Sec. 4.6 in "The Kepler Problem" of Bate, Mueller, White (1971) [33]. Indeed, for E=0, x=a-ae (the distance from the focus F to the vertex is a-ae),  $E=\pi/2, x=-ae$  (the distance from the focus F and back to the center C) and for  $E=\pi, x=-a-ae$  (the distance to the other vertex).

19.4.1. Path in terms of eccentric anomaly. Take the time derivative of Eq. 112:

(113) 
$$\dot{x} = -a\sin E\dot{E}$$

$$\dot{y} = a\sqrt{1 - e^2}\cos E\dot{E}$$

To determine  $\dot{E}$ , consider both expressions for the radial distance r for Eq. 111 and Eq. 88:

$$r = \frac{a(1 - e^2)}{1 + e\cos\theta} = \frac{b}{1 + e\cos\theta}$$

$$\Rightarrow \dot{r} = \frac{-b}{(1 + e\cos\theta)^2} (-e\sin\theta)\dot{\theta} = \frac{be\sin\theta\dot{\theta}}{(1 + e\cos\theta)^2} =$$

$$\dot{r} = \frac{r^2e}{b}\sin\theta\dot{\theta}$$

$$r = a(1 - e\cos E)$$

$$\Rightarrow \dot{r} = +ae\sin E\dot{E}$$

$$\Rightarrow ae\sin E\dot{E} = \frac{r^2e}{b}\sin\theta\dot{\theta} \text{ or } a\sin E\dot{E} = \frac{r^2}{b}\sin\theta\dot{\theta} = \frac{\hat{L}_0\sin\theta}{b} = \frac{\hat{L}_0y}{b}$$

$$\Rightarrow \dot{E} = \frac{\hat{L}_0}{ab}\frac{y}{r\sin E} = \frac{\hat{L}_0}{ar} = \frac{b}{a^{3/2}r}$$

Compare the expression for  $\dot{E}$  with Eq. (4.6-11) on pp. 216 of Bate, Mueller, White (1971) [33].

Armed with an expression for  $\dot{E}$ , plug this back into the expression for  $\dot{r}$ :

(114) 
$$\dot{r} = ae \sin E \left(\frac{\hat{L}_0}{ar}\right) = \frac{e \sin E \hat{L}_0}{r}$$

Compare this to the equation on pp. 216, "Position and Velocity - a function of time" in Ch. 4 of Bate, Mueller, White (1971) [33].

Use the expressions above in Eqns. 108. For  $\mathbf{r}, \mathbf{r}_0$  centered at focus F:

$$xv_{0,y} - yv_{0,x} = a(\cos E - e)b\cos E_0 \dot{E}_0 - b\sin E(-a\sin E_0 \dot{E}_0) =$$

$$= ab\dot{E}_0 \left[ (\cos E - e)\cos E_0 + \sin E\sin E_0 \right] = \frac{b^2}{\sqrt{ar}} \left[ \cos (E - E_0) - e\cos E_0 \right] = \frac{\hat{L}^2 \left( \frac{-m}{2E} \right)}{\sqrt{GM} \sqrt{\frac{m}{-2E}}r} \left[ \cos (E - E_0) - e\cos E_0 \right] =$$

$$= \frac{\hat{L}_0 b}{r\sqrt{GM}} \left[ \cos (E - E_0) - e\cos E_0 \right]$$

Comparing this against the "right-hand side (RHS)" with  $f(\mathbf{r}_0 \times \mathbf{v}_0)$ , the RHS has a magnitude of  $f\widehat{L}_0$  in the same direction as  $(\mathbf{r} \times \mathbf{v}_0)$ . Thus

(115) 
$$f = \frac{b}{r\sqrt{GM}} \left[\cos\left(E - E_0\right) - e\cos E_0\right]$$

For the expression involving q in Eqns. 108,

$$\mathbf{r} \times \mathbf{r}_0 = xy_0 - yx_0 = a(\cos E - e)b\sin E_0 - b\sin E_0(\cos E_0 - e) = -ab[(\cos E_0)\sin E - (\cos E - e)\sin E_0]$$

with

$$ab = a\widehat{L}_0\sqrt{\frac{-m}{2E}} = a\widehat{L}_0\left(\frac{a}{GM}\right)^{1/2} = \left(\frac{a^3}{GM}\right)^{1/2}\widehat{L}_0$$

For the RHS with  $g\mathbf{v}_0 \times \mathbf{r}_0$ ,  $\mathbf{v}_0 \times \mathbf{r}_0$ , by the right-hand rule or the right-hand orientation, it is in the *opposite* direction as  $\mathbf{r} \times \mathbf{r}_0$ . Thus, apply a negative side on the RHS.

(116) 
$$\Longrightarrow g = \left(\frac{a^3}{GM}\right)^{1/2} \left[ (\cos E_0 - e) \sin E - (\cos E - e) \sin E_0 \right]$$

### Part 14. General Relativity

#### Part 15. WE Heraeus International Winter School on Gravity and Light

The International Winter School on Gravity and Light held *central lectures* given by Dr. Frederic P. Schuller. These lectures on General Relativity and Gravity are unequivocally and undeniably, the best and most lucid and well-constructed lecture series on General Relativity and Gravity. The mathematical foundation from topology and differential geometry from which General Relativity arises from is solid, well-selected in rigor. The lectures themselves are well-thought out and clearly explained.

Even more so, the International Winter School provided accompanying Tutorial Sessions for each of the lectures. I had given up hopes in seeing this component of the learning process ever be put online so that anyone and everyone in the world could learn through the Tutorial process as well. I was afraid that nobody would understand how the Tutorial or "Office Hours" session was important for students to digest and comprehend and work out-doing exercises-the material presented in the lectures. This International Winter School gets it and shows how online education has to be done, to do it in an excellent manner, moving forward.

For anyone who is serious about learning General Relativity and Gravity, I would simply point to these video lectures and tutorials.

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What I want to do is to build upon the material presented in this International Winter School. Why it's important to me, and to the students and practicing researchers out there, is that the material presented takes the student from an introduction and recall  $\mathcal{O}_S | := \{U \cap S | U \in \mathcal{O}_M\}$ . to the research frontier. That is the stated goal of the International Winter School. I want to dig into and help contribute to the cutting edge in research and this entire program with lectures and tutorials appears to be the most direct and sensible route  $\implies f|_S$  cont. directly to being able to do research in General Relativity and Gravity. -EY 20150323

20200321 update: Someone collected all the lecture videos and tutorials:

http://mathswithphysics.blogspot.com/2016/07/the-we-heraeus-international-winter.html

#### 20. Lecture 1: Topology

### 20.1. Lecture 1: Topological Spaces.

**Definition 53.** Let M be a set.

A topology  $\mathcal{O}$  is a subset  $\mathcal{O} \subseteq \mathcal{P}(M)$ ,  $\mathcal{P}(M)$  power set of M: set of all subsets of M. satisfying

- (i)  $\emptyset \in \mathcal{O}$ .  $M \in \mathcal{O}$
- (ii)  $U \in \mathcal{O}$ ,  $V \in \mathcal{O} \Longrightarrow U \cap V \in \mathcal{O}$
- (iii)  $U_{\alpha} \in \mathcal{O}, \quad \alpha \in \mathcal{A} \implies (\bigcup_{\alpha \in \mathcal{A}} U_{\alpha}) \in \mathcal{O}$
- $\mathcal{O}$ } utterly useless

Definition 54.  $\mathcal{O}_{standard} \subseteq \mathcal{P}(\mathbb{R}^d)$ 

EY: 20150524

I'll fill in the proof that  $\mathcal{O}_{\text{standard}}$  is a topology.

Proof.  $\emptyset \in \mathcal{O}_{\text{standard}}$ 

since  $\forall p \in \emptyset, \exists r \in \mathbb{R}^+$ :  $\mathcal{B}_r(p) \subseteq \emptyset$  (i.e. satisfied "vacuously")

Suppose  $U, V \in \mathcal{O}_{\text{standard}}$ .

Let  $p \in U \cap V$ . Then  $\exists r_1, r_2 \in \mathbb{R}^+$  s.t.  $\mathcal{B}_{r_1}(p) \subseteq U$ 

$$\mathcal{B}_{r_2}(p) \subseteq V$$

Let  $r = \min\{r_1, r_2\}.$ 

Clearly  $\mathcal{B}_r(p) \subset U$  and  $\mathcal{B}_r(p) \subset V$ . Then  $\mathcal{B}_r(p) \subset U \cap V$ . So  $U \cap V \in \mathcal{O}_{\text{standard}}$ .

Suppose,  $U_{\alpha} \in \mathcal{O}_{\text{standard}}, \forall \alpha \in \mathcal{A}.$ 

Let  $p \in \bigcup_{\alpha \in A} U_{\alpha}$ . Then  $p \in U_{\alpha}$  for at least  $1 \alpha \in A$ .

 $\exists r_{\alpha} \in \mathbb{R}^+ \text{ s.t. } \mathcal{B}_{r_{\alpha}}(p) \subseteq U_{\alpha} \subseteq \bigcup_{\alpha \in A} U_{\alpha}. \text{ So } \bigcup_{\alpha \in A} U_{\alpha} \in \mathcal{O}_{\text{standard}}$ 

### 20.2. 2. Continuous maps.

### 20.3. 3. Composition of continuous maps.

### 20.4. **4. Inheriting a topology.** EY: 20150524

I'll fill in the proof that given f continuous (cont.), then the restriction of f onto a subspace S is cont. If you want a reference, check out Klaus Jänich [?, pp. 13, Ch. 1 Fundamental Concepts, Sec. Continuous Maps]

If cont.  $f: M \to N$ ,  $S \subseteq M$ , then  $f|_S$  cont.

*Proof.* Let open  $V \subseteq N$ , i.e.  $V \in \mathcal{O}_N$  i.e. V in the topology  $\mathcal{O}_N$  of N.

$$f|_{S}^{-1}(V) = \{ m \in M | f|_{S}(m) \in V \}$$

Now 
$$f^{-1}(V) = \{ m \in M | f(m) \in V \}.$$
  
So  $f^{-1}(V) \cap S = f|_S^{-1}(V)$ 

```
Now f cont. So f^{-1}(V) \in \mathcal{O}_N.
so f^{-1}(V) \cap S = f|_{S}^{-1}(V) \in \mathcal{O}_{S} i.e. f|_{S}^{-1}(V) open
```

TOPOLOGY TUTORIAL SHEET

filename: main.pdf

The WE-Heraeus International Winter School on Gravity and Light: Topology

EY: 20150524

What I won't do here is retype up the solutions presented in the Tutorial (cf. https://youtu.be/\_XkhZQ-hNLs): the presenter did a very good job. If someone wants to type up the solutions and copy and paste it onto this LaTeX file, in the spirit of open-source collaboration, I would encourage this effort.

Instead, what I want to encourage is the use of as much CAS (Computer Algebra System) and symbolic and numerical computation because, first, we're in the 21st century, second, to set the stage for further applications in research. I use Python and Sage Math alot, mostly because they are open-source software (OSS) and fun to use. Also note that the structure of Sage Math modules matches closely to Category Theory.

In checking whether a set is a topology, I found it strange that there wasn't already a function in Sage Math to check each of the axioms. So I wrote my own; see my code snippet, which you can copy, paste, edit freely in the spirit of OSS here, titled

gist github ernestyalumni topology.sage

Download topology.sage

Loading topology.sage, after changing into (with the usual Linux terminal commands, cd, ls) by

```
sage: load(''topology.sage'')
```

Exercise 2: Topologies on a simple set.

Question Does  $\mathcal{O}_1 := \dots$  constitute a topology  $\dots$ ?.

Solution: Yes, since we check by typing in the following commands in Sage Math:

```
Axiom2check(O_1) # True
Axiom3check(0_1) # True
```

Question What about  $\mathcal{O}_2 \dots$ ?.

Solution: No since the 3rd. axiom fails, as can be checked by typing in the following commands in Sage Math:

```
emptyset in 0_2
Axiom2check(O_2) # True
Axiom3check(0_2) # False
```

### 21. Lecture 2: Topological Manifolds

**Lecture 2: Manifolds.** Topological spaces:  $\exists$  so man that mathematicians cannot even classify them.

For spacetime physics, we may focus on topological spaces  $(M, \mathcal{O})$  that can be charted, analogously to how the surface of the earth is charted in an atlas.

21.1. Topological manifolds.

**Definition 55.** A topological space  $(M, \mathcal{O})$  is called a d-dimensional topological method if  $\forall p \in M : \exists U \in \mathcal{O}, U \ni p : \exists x : U \subseteq M \to x(U) \subseteq \mathbb{R}^d$   $(M, \mathcal{O}), (\mathbb{R}^d, \mathcal{O}_{std})$ 

(i) x invertible:

$$x^{-1}: x(U) \to U$$

- (ii) x continuous
- (iii)  $x^{-1}$  continuous
- 21.2. Terminology.
- 21.3. 3. Chart transition maps. Imagine 2 charts (U, x) and (V, y) with overlapping regions.
- 21.4. **4. Manifold philosophy.** Often it is desirable (or indeed the way) to define properties ("continuity") of real-world object (" $\mathbb{R} \xrightarrow{\gamma} M$ ") by judging suitable coordinates not on the "real-world" object itself, but on a chart-representation of that real world object.

EY's add-ons. This lecture gives me a good excuse to review Topology and Topological Manifolds from a mathematician's point of view. I find John M. Lee's Introduction to Topological Manifolds book good because it's elementary and thorough and it's fairly recent (2010) so it's up to date [?]. See my notes and solutions for the book; it's a file titled LeeJM\_IntroTopManifolds\_sol.pdf of which I'll try to keep the pdf and LaTeX file available for download on my ernestyalumni Google Drive (so try to search for it on Google).

TUTORIAL TOPOLOGICAL MANIFOLDS

filename: Sheet\_1.2.pdf

Exercise 4: Before the invention of the wheel.

Another one-dimensional topological manifold. Another one? Consider set  $F^1 := \{(m, n) \in \mathbb{R}^2 | m^4 + n^4 = 1\}$ , equipped with subset topology  $\mathcal{O}_{\text{std}}|_{\mathcal{D}^1}$ 

Question  $x: F^1 \to \mathbb{R}$  is what?.

Solution. EY: 20150525 The tutorial video https://youtu.be/ghfEQ3u\_B6g is really good and this solution is how I'd write it, but it's really the same (I needed the practice).

$$x: F^1 \to \mathbb{R}$$
$$(m,n) \mapsto m$$

If m = 0,  $n^4 = 1$  so  $n = \pm 1$  so it's not injective.

Let the closed n-dim. upper half-space  $\mathbb{H}^n \subset \mathbb{R}^1$ . Then

$$\mathbb{H}^n = \{(x_1 \dots x_n) \in \mathbb{R}^n | x_n \ge 0\}$$
$$\operatorname{int}\mathbb{H}^n = \{(x_1 \dots x_n) \in \mathbb{R}^n | x_n > 0\}$$
$$-\mathbb{H}^n = \{(x_1 \dots x_n) \in \mathbb{R}^n | x_n \le 0\}$$
$$-\operatorname{int}\mathbb{H}^n = \{(x_1 \dots x_n) \in \mathbb{R}^n | x_n < 0\}$$

Question This map x may be made injective by restricting its domain to either of 2 maximal open subsets of  $F^1$ . Which ones?.

Solution.

$$U_{+} = F^{1} \cap \operatorname{int} \mathbb{H}^{2}$$

$$U_{-} = F^{1} \cap -\operatorname{int} \mathbb{H}^{2}$$

Look at

$$x^4 = 1 - n^4$$
$$\Longrightarrow x = \pm (1 - n^4)^{1/4}$$

Then for

$$x_{+}^{-1}: (-1,1) \subseteq \mathbb{R} \to U_{+}$$

$$m \mapsto (m, (1-m^{4})^{1/4})$$

$$x_{-}^{-1}: (-1,1) \subseteq \mathbb{R} \to U_{-}$$

$$m \mapsto (m, -(1-m^{4})^{1/4})$$

 $x_{+},x_{-}$  injective (since left inverse exists).

Question Construct injective y.

Solution .

Let

$$V_{+} = F^{1} \cap \operatorname{int} \mathbb{H}^{1}$$
$$V_{-} = F^{1} \cap -\operatorname{int} \mathbb{H}^{1}$$

Then

$$y_{+}: V_{+} \to (-1,1) \subseteq \mathbb{R}$$
$$(m,n) \mapsto n$$
$$y_{-}: V_{-} \to (-1,1) \subseteq \mathbb{R}$$
$$(m,n) \mapsto n$$

Question Construct inverse  $y^{-1}$ . Solution .

For

$$y_{+}^{-1}: (-1,1) \subseteq \mathbb{R} \to V_{+}$$

$$n \mapsto ((1-n^{4})^{1/4}, n)$$

$$y_{-}^{-1}: (-1,1) \subseteq \mathbb{R} \to V_{-}$$

$$n \mapsto (-(1-n^{4})^{1/4}, n)$$

 $y_+,y_-$  injective (since left inverse exists).

Note 
$$(-1,0) \notin U_+, U_-$$
  
 $(1,0) \notin U_+, U_-$   
and

$$(0,1) \notin V_+, V_-$$
  
 $(0,-1) \notin V_+, V_-$ 

Question construct transition map  $x \circ y^{-1}$ .

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Solution.

$$x_{+}y_{+}^{-1}: (0,1) \subseteq \mathbb{R} \to (0,1) \subseteq \mathbb{R}$$

$$n \mapsto (1 - n^{4})^{1/4}$$

$$x_{-}y_{+}^{-1}: (-1,0) \subseteq \mathbb{R} \to (0,1) \subseteq \mathbb{R}$$

$$n \xrightarrow{y_{+}^{-1}} ((1 - n^{4})^{1/4}, n) \xrightarrow{x_{-}} (1 - n^{4})^{1/4}$$

$$x_{+}y_{-}^{-1}: (0,1) \subseteq \mathbb{R} \to (-1,0) \subseteq \mathbb{R}$$

$$n \mapsto -(1 - n^{4})^{1/4}$$

$$x_{-}y_{-}^{-1}: (-1,0) \subseteq \mathbb{R} \to (-1,0) \subseteq \mathbb{R}$$

$$n \mapsto -(1 - n^{4})^{1/4}$$

Question ... Does the collection of these domains and maps form an atlas of  $F^1$ ?.

Yes, with atlas

$$\mathcal{A} = \{ \begin{matrix} (U_+, x_+) & (V_+, y_+) \\ (U_-, x_-) & (V_-, y_-) \end{matrix} \}$$

Clearly

$$U_{+} \cup U_{-} \cup V_{+} \cup V_{-} = (F^{1} \cap \operatorname{int}\mathbb{H}^{2}) \cup (F^{1} \cap -\operatorname{int}\mathbb{H}^{2}) \cup (F^{1} \cap \operatorname{int}\mathbb{H}^{1}) \cup (F^{1} \cap -\operatorname{int}\mathbb{H}^{1}) =$$

$$= F^{1} \cap \mathbb{R}^{2} \setminus \{(0,0)\} = F^{1}$$

and (the point is that)  $x_+, y_+$  are homeomorphisms of open sets of  $F^1$  onto open sets of 1 dim.  $\mathbb{R}^1$  (namely  $(-1,1) \subseteq \mathbb{R}^1$ ), and so A is an atlas of  $F^1$ .

### 22. Lecture 3: Multilinear Algebra

### Lecture 3: Multilinear Algebra (International Winter School on Gravity and Light 2015)

We will **not** equip space(time) with a vector space structure. Do you know where

Moreover, the tangent spaces  $T_nM$  (lecture 5) smooth manifolds (Lecture 4)

- Beneficial to first study vector spaces abstractly for two reason
- (ii) tensor technique are most easily understood in an abstract setting.

(i) for construction of  $T_pM$  one needs an intermediate vector space  $C^{\infty}(M)$ 

### 22.1. Vector spaces.

**Definition 56.** A vector space (V, +, -) is

- (i) a set V
- (ii)  $+: V \times V \to V$ "addition"
- (iii)  $\cdot : \mathbb{R} \times V \to V$ "s-multiplication" EY: 20160317 s for "scalar"

satisfying:

CANIADDU

$$C^{+}: \qquad v+w=w+v$$
 
$$A^{+}: \qquad (u+v)+w=u+(v+w)$$
 
$$N^{+}: \qquad \exists \, 0 \in V: \forall \, v \in V: v+0=v$$
 
$$I^{+}: \qquad \forall \, v \in V: \exists \, (-v) \in V: v+(-v)=0$$
 
$$A: \qquad \lambda \cdot (\mu+v)=(\lambda \cdot \mu) \cdot v \qquad (\forall \, \lambda, \mu \in \mathbb{R})$$
 
$$D: \qquad (\lambda+\mu) \cdot v=\lambda \cdot v+\mu \cdot v$$
 
$$D: \qquad \lambda \cdot v+\lambda \cdot w=\lambda \cdot (v+w)$$
 
$$U: \qquad 1 \cdot v=v$$

Terminology. An element of a vector space is often referred to, informally as a vector.

Example. def. set of polynomials (fixed) degree  $\mathcal{P} := \{p : (-1, +1) \to \mathbb{R} | p(x) = \sum_{n=0}^{N} p_n \cdot x^n \}$ 

Thought bubble: is 
$$\square$$
 a vector? 
$$\square(x) = x^2$$
No  $\square \in \mathcal{P}$ .  $+ : \mathcal{P} \times \mathcal{P} \to \mathcal{P}$ 

$$(p,q) \mapsto p + q$$
where  $(p+q)(x) = p(x) +_{\mathbb{R}} q(x)$ 

$$\cdot : \mathbb{R} \times \mathcal{P} \to \mathcal{P}$$

$$(\lambda,p) \mapsto \lambda \cdot p$$
where  $(\lambda \cdot p)(x) := \lambda \cdot_{\mathbb{R}} p(x)$ 
Thought bubble:  $\square$  a vector? Yes, but who cares?  $(\mathcal{P},+,\cdot)$  is a vector space.

22.2. Linear maps. These are the structure-respecting maps between vector spaces.

EY: 20160316 out of tradition, they're called "linear" maps

**Definition 57.**  $(V, +_V, \cdot_V)$  and  $(W, +_W, \cdot_W)$  vector spaces Then a map

$$\varphi: V \to W$$

is called **linear** if

(i) 
$$\varphi(v +_V \widetilde{v}) = \varphi(v) +_W \varphi(\widetilde{v})$$

(ii) 
$$\varphi(\lambda \cdot_V v) = \lambda \cdot_W \varphi(v)$$

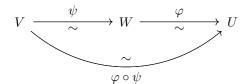
linear:

 $\square \in \mathcal{P}$ 

(i) 
$$\delta(p+q) = (p+pq)' = p'+pq' = \delta(p)+p\delta(q)$$

(ii) 
$$\delta(\lambda p) = (\lambda p)' = \lambda \cdot p' = \lambda \cdot \delta(p)$$

Notation:  $\varphi: V \to W$  linear  $\iff: \varphi: V \xrightarrow{\sim} W$ 



22.2.1. Example\*.  $\delta \circ \delta : \mathcal{P} \xrightarrow{\sim} \mathcal{P}$ 

22.3. Vector space of Homomorphisms. fun fact:  $(V, +, \cdot)$   $(W, +, \cdot)$  vector spaces  $def. \operatorname{Hom}(V, W) := \{ \varphi : V \xrightarrow{\sim} W \}$  set.

We can make this into a vector spaces.

$$\oplus : \operatorname{Hom}(V, W) \times \operatorname{Hom}(V, W) \to \operatorname{Hom}(V, W)$$
$$(\varphi, \psi) \mapsto \varphi \oplus \psi$$

where 
$$(\varphi \otimes \psi)(v) := \varphi(v) +_W \psi(v)$$
  
  $\otimes : \dots$  similarly.  
  $(\operatorname{Hom}(V, W), \oplus, \otimes)$  is a vector space.

22.3.1. Example\*. 
$$\operatorname{Hom}(\mathcal{P}, \mathcal{P})$$
 is a vector space.  $\delta \in \operatorname{Hom}(\mathcal{P}, \mathcal{P})$ 

 $\delta \circ \delta \in \operatorname{Hom}(\mathcal{P}, \mathcal{P})$ 

$$\underbrace{\delta \circ \cdots \circ \delta}_{M} \in \operatorname{Hom}(\mathcal{P}, \mathcal{P})$$

$$\Longrightarrow 5 \circ \delta \oplus_{\operatorname{Hom}(\mathcal{P},\mathcal{P})} \delta \circ \delta \in \operatorname{Hom}(\mathcal{P},\mathcal{P})$$

# 22.4. Dual vector space. heavily used special case:

 $(V, +, \cdot)$  vector space:

Definition 58.

$$V^* := \{ \varphi : V \xrightarrow{\sim} \mathbb{R} \} = Hom(V, \mathbb{R})$$

$$\underbrace{(V^*, \oplus, \otimes)}_{\textit{dual vector space (to V)}} \textit{is a vector space}$$

Terminology:  $\varphi \in V^*$  is called, informally, a covector.

Example. 
$$I: \mathcal{P} \xrightarrow{\sim} \mathbb{R}$$
  
i.e.  $I \in \mathcal{P}^*$   

$$\underline{\det}. \ I(p) := \int_0^1 dx p(x)$$

$$I(p+q) = \int_0^1 dx \underbrace{(p+q)(x)}_{p(x)+q(x)}$$

$$= \cdots = I(q) + I(p)$$
i.e.  $I = \int_0^1 dx$ 

THE DIFFERENTIAL GEOMETRY DIFFERENTIAL TOPOLOGY DUMP

22.5. Tensors.

**Definition 59.** Let  $(V, +, \cdot)$  be a vector space.

An (r,s)-tensor T over V  $r,s \in \mathbb{N}_0$  is a multi-linear map

$$T: \underbrace{V^* \times \cdots \times V^*}_r \times \underbrace{V \times \cdots \times V}_s \overset{\overset{\sim}{\underset{\sim}{:}} \}r+s}{\overset{\sim}{\underset{\sim}{:}}} \mathbb{R}$$

22.5.1. Example. : T (1,1)-tensor

$$T(\varphi + \psi, v) = T(\varphi, v) + T(\psi, v) \qquad T(\varphi, v + w) = T(\varphi, v) + T(\varphi, w)$$

$$T(\lambda \varphi, v) = \lambda \cdot T(\varphi, v) \qquad T(\varphi, \lambda \cdot v) = \lambda T(\varphi, v)$$

$$T(\varphi + \psi, v + w) =$$

$$= t(\varphi, v) + T(\varphi, w) + T(\psi, v) + t(\psi, w)$$

Excursion: Given 
$$T: V^* \times V \xrightarrow{\sim} \mathbb{R}$$

$$\phi_T: V \xrightarrow{\sim} (V^*)^* = V$$

Define 
$$v \mapsto \underbrace{T(\cdot, v)}_{V^* \xrightarrow{\sim} \mathbb{R}}$$
  
Given  $\phi : V \xrightarrow{\sim} V$ 

Construct 
$$T_{\phi}: V^* \times V \xrightarrow{\sim} \mathbb{R}$$
  
 $(\varphi, v) \mapsto \varphi(\phi(v))$   
 $\Longrightarrow \text{ given } T: T = T_{\varphi_T}$   
 $\text{ given } \phi: \phi = \phi_{T_+}$ 

Example. 
$$g: P \times P \xrightarrow{\sim} \mathbb{R}$$
  $(p,q) \mapsto \int_{-1}^{1} dx p(x) q(x)$  is a  $(0,2)$ -tensor over  $P$ . Info: If  $T \in \operatorname{Hom}(V,W)$ 

22.6. Vectors and covectors as tensors.

Theorem 23. (including proof)

"covector"  $\varphi \in V^* \iff \varphi : V \xrightarrow{\sim} \mathbb{R} \iff \varphi(0,1)$ -tensor.

Theorem 24.  $v \in V = (V^*)^* \iff v : V^* \xrightarrow{\sim} \mathbb{R} \iff v \text{ is } (1,0)\text{-tensor.}$ 

22.7. Bases.

**Definition 60.**  $(V, +, \cdot)$  vector space.

A subset  $B \subset V$  is called

a basis if

Thought bubble: Hamel (L.A.) EY: 20160316 Hamel basis, Linear Algebra

$$\forall v \in V \quad \exists \quad \underline{finite} \quad \underbrace{F}_{\{f_1, \dots, f_n\}} \subset B : \exists ! \underbrace{v^1, v^2, \dots, v^n}_{\in \mathbb{R}}, \qquad v = v^1 f_1 + \dots + v^n f_n$$

23. Lecture 4: Differentiable Manifolds

**Definition 61.** If  $\exists$  basis  $\mathcal{B}$  with finitely many elements, say d many, then we call d =: dimV This is well-defined.

Remark:  $(V, +, \cdot)$  be a finite-dim. vector space.

Having chosen a basis  $e_1, \ldots, e_n$  of  $(V, +, \cdot)$  we may uniquely associate

(Thought bubble: this requires a chosen basis)

 $v \mapsto (v^1, \dots, v^n)$  called the components of v w.r.t. chosen basis

where:  $v^{1}e_{1} + \cdots + v^{n}e + n = v$ 

22.8. Basis for the dual space. choose Basis  $e_1, \ldots, e_n$  for V

can choose Basis  $\epsilon^1, \ldots, \epsilon^n$  for  $V^*$ 

However, more economical to require

once  $e_1, \ldots, e_n$  on V has been chosen, that

$$\epsilon^a(e_b) = \delta^a_b$$

This uniquely determines choice of  $\epsilon^1, \ldots, \epsilon^n$  from choice of  $e_1, \ldots, e_n$ 

**Definition 62.** If a basis  $\epsilon^1, \ldots, \epsilon^n$  of  $V^*$  satisfies this, it is called <u>the</u> **dual basis** (of the dual space)

Example: 
$$P(N=3)$$

$$e_0(x) = 1$$

$$e_1(x) = x$$

$$e_0, e_1, e_2, e_3 \text{ basis if } e_2(x) = x^2 \{e_a(x) := x^a \}$$

$$e_3(x) = x^3$$

$$\epsilon^0, \epsilon^1, \epsilon^2, \epsilon^3 \text{ dual basis } \epsilon^a := \frac{1}{24} \frac{\partial^a}{\partial x^a} \Big|_{x=0}$$

Proof.  $\epsilon^a(e_b) = \delta^a_b$ 

22.9. Components of tensors. Let T be an (r, s)-tensor on a finite-dim. vs. V. Then define the  $(r + s)^{\dim V}$  many real numbers.

$$\underbrace{T^{i_1\dots i_r}_{j_1\dots j_s}}_{\in\mathbb{P}} := T(\epsilon^{i_1},\dots,\epsilon^{i_r},e_{j_1},e_{j_2},\dots,e_{j_s})$$

 $i_1 \dots i_r, j_1 \dots j_s \in \{1, \dots, \dim V\}$ 

Thought bubble:  $\underbrace{T^{i_1...i_r}_{j_1...j_s}}_{\in \mathbb{R}}$  are the <u>components</u> of the tensor w.r.t. chosen basis

Useful: Knowing components (and basis) one can reconstruct the entire tensor.

Example. T(1,1)- tensor

$$T^{i}_{i} := T(\epsilon^{i}, e_{i})$$

reconstruc

Teconstruct 
$$T(\varphi, v) = T(\sum_{i=1}^{\dim V} \varphi_i \epsilon^i, \sum_{j=1}^{\dim V} v^j e_j) \qquad \varphi_i \in \mathbb{R}, v^j \in \mathbb{I}$$

$$= \sum_{i=1}^{\dim V} \sum_{j=1}^{\dim V} \varphi_i v^j \underbrace{T(\epsilon^i, e_j)}_{T^i_j}$$

$$=: \varphi_i v^j T^i_j$$

so far: top. mfd. 
$$(M, \mathcal{O})$$
  
 
$$\dim M = d$$

we wish to define a notion of differentiable

curves  $\mathbb{R} \to M$ function  $M \to \mathbb{R}$ maps  $M \to N$ 

23.1. 1. Strategy. choose a chart (U, x)

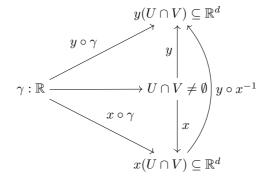
 $\gamma: \mathbb{R} \to M$  portion of curve in chart domain

$$x: \mathbb{R} \xrightarrow{U} U$$
 $x \circ \gamma \downarrow x$ 
 $x(U) \in \mathbb{I}$ 

idea. try to "lift" the undergraduate notion of differentiability of a curve on  $\mathbb{R}^d$  to a notion of

differentiability of a curve on M

<u>Problem</u> Can this be well-defined under change of chart?



 $x \circ \gamma$  undergraduate differentiable ("as a map  $\mathbb{R} \to \mathbb{R}^{d}$ ")

$$\underbrace{y \circ \gamma}_{\text{maybe only continuous, but not undergraduate differentiable}} = \underbrace{\underbrace{(y \circ x^{-1})}_{\mathbb{R}^d \to \mathbb{R}^d}}_{\text{continuous}} \circ \underbrace{\underbrace{(x \circ \gamma)}_{\mathbb{R}^d \to \mathbb{R}^d}}_{\text{undergrad differentiable}} = y \circ (x^{-1} \circ x) \circ \underbrace{(x \circ \gamma)}_{\mathbb{R}^d \to \mathbb{R}^d}$$

At first sight, strategy does not work out.

23.2. **2. Compatible charts.** In section 1, we used any imaginable charts on the top. mfd.  $(M, \mathcal{O})$ . To emphasize this, we may say that we took U and V from the maximal atlas  $\mathcal{A}$  of  $(M, \mathcal{O})$ .

**Definition 63.** Two charts (U, x) and (V, y) of a top. mfd. are called  $\Re$ -compatible if either

- (a)  $U \cap V = \emptyset$  or
- (b)  $U \cap V \neq \emptyset$

chart transition maps have undergraduate & property.

 $EY: 20151109 \text{ e.g. since } \mathbb{R}^d \to \mathbb{R}^d$ , can use undergradate  $\mathfrak{B}$  property such as continuity or differentiability.

$$y \circ x^{-1} : x(U \cap V) \subseteq \mathbb{R}^d \to y(U \cap V) \subseteq \mathbb{R}^d$$
$$x \circ y^{-1} : y(U \cap V) \subseteq \mathbb{R}^d \to x(U \cap V) \subseteq \mathbb{R}^d$$

Philosophy:

**Definition 64.** An atlas  $A_{\circledast}$  is a  $\circledast$ -compatible atlas if any two charts in  $A_{\circledast}$  are  $\circledast$ -compatible.

**Definition 65.** A 
$$\circledast$$
-manifold is a triple  $(\underbrace{M,\mathcal{O}}_{top.\ mfd}, \mathcal{A}_{\circledast})$   $\mathcal{A}_{\circledast} \subseteq \mathcal{A}_{maximal}$ 

€	undergraduate 🏶	
$C^0$	$C^0(\mathbb{R}^d \to \mathbb{R}^d) =$	continuous maps w.r.t. O
$C^1$	$C^1(\mathbb{R}^d \to \mathbb{R}^d) =$	differentiable (once) and is continuous
$C^k$	, ,	k-times continuously differentiable
$D^k$		k-times differentiable
:		
$C^{\infty}$	$C^\infty(\mathbb{R}^d o\mathbb{R}^d)$	
IU	, ,	
$C^{\omega}$	$\exists$ multi-dim. Taylor exp.	
$\mathbb{C}^{\infty}$	satisfy Cauchy-Riemann equations, pair-wise	

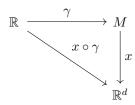
EY: 20151109 Schuller says:  $C^k$  is easy to work with because you can judge k-times cont. differentiability from existence of all partial derivatives and their continuity. There are examples of maps that partial derivatives exist but are not  $D^k$ , k-times differentiable.

**Theorem 25** (Whitney). Any  $C^{k\geq 1}$ -atlas,  $A_{C^{k\geq 1}}$  of a topological manifold contains a  $C^{\infty}$ -atlas.

Thus we may w.l.o.g. always consider  $C^{\infty}$ -manifolds, "smooth manifolds", unless we wish to define Taylor expandibility/complex differentiability . . .

EY: 20151109 Hassler Whitney <sup>7</sup>

**Definition 66.** A smooth manifold  $(\underbrace{M, \mathcal{O}}_{top.\ mfd.}, \underbrace{\mathcal{A}}_{C^{\infty}-atlas})$ 



EY: 20151109 Schuller was explaining that the trajectory is real in M; the coordinate maps to obtain The wild world of 4-manifolds

coordinates is  $x \circ \gamma$ 

# 23.3. 4. Diffeomorphisms. $M \xrightarrow{\phi} N$

If M, N are naked sets, the structure preserving maps are the bijections (invertible maps) e.g.  $\{1, 2, 3\} \rightarrow \{a, b\}$ 

**Definition 67.**  $M \cong_{set} N$  (set-theoretically) isomorphic if  $\exists$  bijection  $\phi: M \to N$ 

Examples.  $\mathbb{N} \cong_{\text{set}} \mathbb{Z}$ 

 $\mathbb{N} \cong_{\text{set}} \mathbb{Q}$  (EY: 20151109 Schuller says from diagonal counting)

 $\mathbb{N} \cong_{\operatorname{set}} \mathbb{R}$ 

Now  $(M, \mathcal{O}_M) \cong_{\text{top}} (N, \mathcal{O}_N)$  (topl.) isomorphic = "homeomorphic"  $\exists$  bijection  $\phi : M \to N$   $\phi, \phi^{-1}$  are continuous.

 $(V,+,\cdot)\cong_{\mathrm{vec}}(W,+_w,\cdot_w)$  (EY: 20151109 vector space isomorphism) if

 $\exists$  bijection  $\phi: V \to W$  linearly

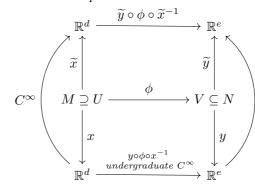
finally

**Definition 68.** Two  $C^{\infty}$ -manifolds

 $(M, \mathcal{O}_M, \mathcal{A}_M)$  and  $(N, \mathcal{O}_N, \mathcal{A}_N)$  are said to be **diffeomorphic** if  $\exists$  bijection  $\phi : M \to N$  s.t.

$$\phi: M \to N$$
$$\phi^{-1}: N \to M$$

are both  $C^{\infty}$ -maps



**Theorem 26.** #= number of  $C^{\infty}$ -manifolds one can make out of a given  $C^{0}$ -manifolds (if any) - up to diffeomorphisms.

dimM	#	
1	1	Morse-Radon theorems
2	1	$Morse ext{-}Radon\ theorems$
3	1	$Morse ext{-}Radon\ theorems$
4	uncountably infinitely many	
5	finite	$surgery\ theory$
6	finite	$surgery\ theory$
:	finite	surgery theory

EY: 20151109 cf. http://math.stackexchange.com/questions/833766/closed-4-manifolds-with-uncountably-many-differe

### TUTORIAL 4 DIFFERENTIABLE MANIFOLDS

EY: 20151109 The gravity-and-light.org website, where you can download the tutorial sheets and the full length videos for the tutorials and lectures, are no longer there. = (

Hopefully, the YouTube video will remain: https://youtu.be/FXPdKxOq1KA?list=PLFeEvEPtX\_ORQ1ys-7VIsK1BWz7RX-FaL

Exercise 1: True or false?. These basic questions are designed to spark discussion and as a self-test.

Tick the correct statements, but not the incorrect ones!

- (a) The function  $f: \mathbb{R} \to \mathbb{R}, \ldots$ 
  - •
  - ..., defined by  $f(x) = |x^3|$ , lies in  $C^3(\mathbb{R} \to \mathbb{R})$ .

<sup>&</sup>lt;sup>7</sup>http://mathoverflow.net/questions/8789/can-every-manifold-be-given-an-analytic-structure

EY: 20151109 Solution 1a3. For 
$$f: \mathbb{R} \to \mathbb{R}$$
,  $f(x) = |x^3| = \begin{cases} x^3 & \text{if } x \ge 0 \\ -x^3 & \text{if } x < 0 \end{cases}$ 

$$f'(x) = \begin{cases} 3x^2 & \text{if } x \ge 0 \\ -3x^2 & \text{if } x < 0 \end{cases}$$
$$f''(x) = \begin{cases} 6x & \text{if } x \ge 0 \\ -6x & \text{if } x < 0 \end{cases}$$

Thus,

$$f(x) = |x^3| \in C^1(\mathbb{R}) \text{ but } f(x) \notin C^2(\mathbb{R}) \subseteq C^3(\mathbb{R})$$

(b)

(c)

### Short Exercise 4: Undergraduate multi-dimensional analysis.

A good notation and basic results for partial differentiation.

For a map  $f: \mathbb{R}^d \to \mathbb{R}$  we denote by the map  $\partial_i f: \mathbb{R}^d \to \mathbb{R}$  the partial derivative with respect to the *i*-th entry.

Question: Given a function

$$f: \mathbb{R}^3 \to \mathbb{R}; (\alpha, \beta, \delta) \mapsto f(\alpha, \beta, \delta) := \alpha^3 \beta^2 + \beta^2 \delta + \delta$$

calculate the values of the following derivatives:

Solution :.

- $(\partial_2 f)(x,y,z) =$
- $(\partial_1 f)(\Box, \circ, *) =$
- $(\partial_1 \partial_2 f)(a,b,c) =$
- $(\partial_3^2 f)(299, 1222, 0) =$

EY: 20151110

For 
$$f(\alpha, \beta, \delta) := \alpha^3 \beta^2 + \beta^2 \delta + \delta$$
, or  $f(x, y, z) = x^3 y^2 + y^2 z + z$ ,

$$(\partial_2 f) = 2(x^3 y + yz)$$
$$(\partial_1 f) = 3x^2 y^2$$
$$(\partial_1 \partial_2 f) = 6x^2 y$$

 $(\partial_3^2 f) = 0$ 

and so

- $\bullet (\partial_2 f)(x, y, z) = 2(x^3y + yz)$
- $(\partial_1 f)(\Box, \circ, *) = 3\Box^2 \circ^2$
- $(\partial_1 \partial_2 f)(a, b, c) = 6a^2b$
- $(\partial_3^2 f)(299, 1222, 0) = 0$

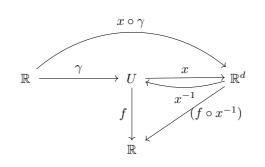
### Exercise 5: Differentiability on a manifold.

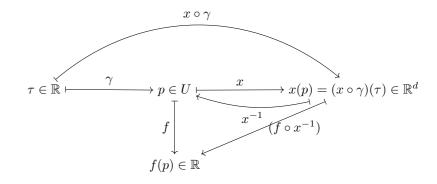
How to deal with functions and curves in a chart

Let  $(M, \mathcal{O}, \mathcal{A})$  be a smooth d-dimensional manifold. Consider a chart (U, x) of the atlas  $\mathcal{A}$  together with a smooth curve  $\gamma : \mathbb{R} \to U$  and a smooth function  $f : U \to \mathbb{R}$  on the domain U of the chart.

Question: Draw a commutative diagram containing the chart domain, chart map, function, curveand the respective represen-

tatives of the function and the curve in the chart. **Solution**:





**Question:** Consider, for d=2,

$$(x \circ \gamma)(\lambda) := (\cos(\lambda), \sin(\lambda))$$
 and  $(f \circ x^{-1})((x, y)) := x^2 + y^2$ 

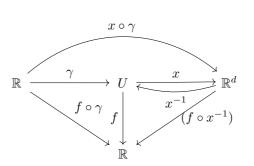
Using the chain rule, calculate

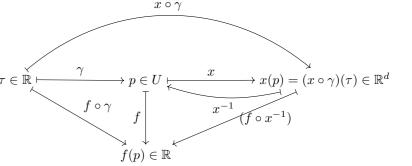
$$(f \circ \gamma)'(\lambda)$$

explicitly.

Solution:.

EY : 20151109 Indeed, the domains and codomains of this  $f\gamma$  mapping makes sense, from  $\mathbb{R} \to \mathbb{R}$  for





$$(f \circ \gamma)'(\lambda) = (Df) \cdot \dot{\gamma}(\lambda) = \frac{\partial f}{\partial x^j} \dot{\gamma}^j(\lambda) = 2x(-\sin \lambda) + 2y\cos \lambda = 2(-\cos \lambda \sin \lambda + \sin \lambda \cos \lambda) = 0$$

24. Lecture 5: Tangent Spaces

lead question: "what is the velocity of a curve  $\gamma$  point p?

### 24.1. Velocities.

**Definition 69.**  $(M, \mathcal{O}, \mathcal{A})$  smooth mfd. curve  $\gamma : \mathbb{R} \to M$  at least  $C^1$ .

Suppose  $\gamma(\lambda_0) = p$ 

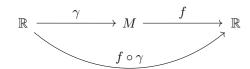
The **velocity** of  $\gamma$  p is the linear map

$$v_{\gamma,p}: C^{\infty}(M) \xrightarrow{\sim} \mathbb{R}$$

 $C^{\infty}(M) := \{ f : M \to \mathbb{R} | f \text{ smooth function } \} \text{ equipped with } (f \oplus g)(p) := f(p) + g(p)$  $(\lambda \otimes g)(p) := \lambda \cdot g(p)$ 

 $\sim$  denotes linear map on top of  $\rightarrow$ .

$$f \mapsto v_{\gamma,p}(f) := (f \circ \gamma)'(\lambda_0)$$



intuition

Schuller says: children run around the world. Temperature function as temperature contour lines. You feel the temperature. You observe the rate of change of temperature as you run around. f is temperature.

$$\underline{\text{past}} : \text{``} \underbrace{v^i(\partial_i f) = (\underbrace{v^i \partial_i}) f}_{\text{vector}}$$

### 24.2. Tangent vector space.

**Definition 70.** For each point  $p \in M$  def the set "tangent space  $\neq_0 M$  p"

$$T_pM := \{v_{\gamma,p}|\gamma \text{ smooth curves }\}$$

picture

rather M than (embedded) p  $T_pM$  EY: 20151109 see https://youtu.be/pepU\_7NJSGM?t=12m38s for the picture Observation:  $T_pM$  can be made into a vector space.

$$\bigoplus : T_p M \times T_p M \to \\
(v_{\gamma,p} \oplus v_{\delta,p})(\underbrace{f}_{\in C^{\infty}(M)}) := v_{\gamma,p}(f) +_{\mathbb{R}} v_{\delta,p}(f) \\
\odot : \mathbb{R} \times T_p M \to \operatorname{Hom}(C^{\infty}(\mathbb{R}), \mathbb{R}) \\
(\alpha \odot v_{\gamma,p})(f) := \alpha \cdot_{\mathbb{R}} v_{\gamma,p}(f)$$

Remains to be shown that

(i)  $\exists \sigma \text{ curve} : v_{\gamma,p} \oplus v_{\delta,p} = v_{\sigma,p}$ (ii)  $\exists \tau \text{ curve} : \alpha \odot v_{\gamma,p} = v_{\tau,p}$ 

Claim: 
$$\tau : \mathbb{R} \to M$$
 where  $\mu_{\alpha} : \mathbb{R} \to \mathbb{R}$  , does the trick.  $\mapsto \tau(\lambda) := \gamma(\alpha\lambda + \lambda_0) = (\gamma \circ \mu_{\alpha})(\lambda)$   $r \mapsto \alpha \cdot r + \lambda_0$   $\tau(0) = \gamma(\lambda_0) = p$ 

$$v_{\tau,p} := (f \circ \tau)'(0) = (f \circ \gamma \circ \mu_{\alpha})'(0)$$
$$= (f \circ \gamma)'(\lambda_0) \cdot \alpha =$$
$$= \alpha \cdot v_{\gamma,p}$$

Now for the sum:

 $v_{\gamma,p} \oplus v_{\delta,p} \stackrel{?}{=} v_{\sigma,p}$ 

make a choice of chart (U, x) In cloud: ill definition alarm bells.

and define:

Claim:

$$\sigma: \mathbb{R} \to M$$

$$\sigma(\lambda) := x^{-1} \underbrace{((x \circ \gamma)(\lambda_0 + \lambda))}_{\mathbb{R} \to \mathbb{R}^d} + (x \circ \delta)(\lambda_1 + \lambda) - (x \circ \gamma)(\lambda_0))$$

does the trick.

Proof. Since:

$$\sigma_x(0) = x^{-1}((x \circ \gamma)(\lambda_0) + (x \circ \delta)(\lambda_1) - (x \circ \gamma)(\lambda_0))$$
  
=  $\delta(\lambda_1) = p$ 

Now:

$$v_{\sigma_{x},p}(f) := (f \circ \sigma_{x})'(0) =$$

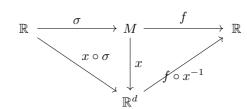
$$= \underbrace{((f \circ x^{-1}) \circ (x \circ \sigma_{x}))'(\gamma)}_{\mathbb{R}^{d} \to \mathbb{R}} = \underbrace{(x \circ \sigma_{x})'(0)}_{(x \circ \gamma)'(\lambda_{0}) + (x \circ \delta)'(\lambda_{1})} \cdot (\partial_{i}(f \circ x^{-1})) (x(\underline{\sigma}(0))) =$$

$$= (x \circ \gamma)'(\lambda_{0})(\partial_{i}(f \circ x^{-1}))(x(p)) + (x \circ \delta)(\lambda_{1})(\partial_{i}(f \circ x^{-1}))(x(p))$$

$$= (f \circ \gamma)'(\lambda_{0}) + (f \circ \delta)'(\lambda_{1}) =$$

$$= v_{\gamma,p}(f) + v_{\delta,p}(f) \quad \forall f \in C^{\infty}(M)$$

$$v_{\gamma,p} \oplus v_{\delta,p} = v_{\sigma,p}$$



picture: (cf. https://youtu.be/pepU\_7NJSGM?t=39m5s)

$$\gamma: \mathbb{R} \to M$$
$$\delta: \mathbb{R} \to M$$

$$(\gamma \oplus)(\lambda) := \gamma(\lambda) + \delta(\lambda)$$

EY: 20151109 Schuller says adding trajectories is chart dependent, bad. Adding velocities is good.

### 24.3. Components of a vector wrt a chart.

**Definition 71.** Let 
$$(U, x) \in \mathcal{A}_{smooth}$$
.

$$Let \begin{array}{c} \gamma : \mathbb{R} \to U \\ \gamma(0) = p \end{array}$$

$$v_{\gamma,p}(f) := (f \circ \gamma)'(0) = (\underbrace{(f \circ x^{-1})}_{\mathbb{R}^d \to \mathbb{R}} \circ \underbrace{(x \circ \gamma)}_{\mathbb{R} \to \mathbb{R}^d})'(0)$$
$$= \underbrace{(x \circ \gamma)^{i'}(0)}_{\dot{\gamma}_x^i(0)} \cdot \underbrace{(\partial_i (f \circ x^{-1}))(x(p))}_{=:(\frac{\partial f}{\partial x^i})_p}$$

think cloud  $f: M \to \mathbb{R}$ 

$$= \boxed{\dot{\gamma}_x^i(0) \cdot \left(\frac{\partial}{\partial x^i}\right)_p} f \quad \forall f \in C^{\infty}(M)$$

 $\therefore$  as a map.

$$v_{\gamma,p} = \underbrace{v_{\chi,p}}_{use\ of\ chart\ "components\ of\ the\ velocity\ v_{\gamma,p}"} \underbrace{\left(\frac{\partial}{\partial x^i}\right)}_{basis\ elements\ of\ the\ T_pM\ wrt\ which\ the\ components\ need\ to\ be\ understood.}$$

Picture: https://youtu.be/pepU\_7NJSGM?t=1h16s

### 24.4. 4. Chart-induced basis.

**Definition 72.** 
$$(U, x) \in \mathcal{A}_{smooth}$$
  
 $the \left(\frac{\partial}{\partial x^1}\right)_p, \dots, \left(\frac{\partial}{\partial x^d}\right)_p \in T_p U \subseteq T_p M$   
 $constitute \ a \ basis \ of \ T_p U$ 

*Proof.* remains: linearly independent

$$\lambda^{i} \left( \frac{\partial}{\partial x^{i}} \right)_{p} \stackrel{!}{=} 0$$

$$\Longrightarrow \lambda^{i} \left( \frac{\partial}{\partial x^{i}} \right)_{p} (x^{j}) = \lambda^{i} \partial_{i} (\underbrace{x^{j} \circ x^{-1}})(x(p)) = \qquad \begin{aligned} x^{j} \circ x^{-1} &: \mathbb{R}^{d} \to \mathbb{R} \\ (\alpha^{1}, \dots, \alpha^{d}) \mapsto \alpha^{j} \end{aligned}$$

$$= \lambda^{i} \delta_{i}^{j} = \lambda^{j} \qquad j = 1, \dots, d$$

in cloud:  $x^j:U\to\mathbb{R}$  differentiable

Corollary 1.  $dimT_nM=d=dimM$ 

Terminology: 
$$X \in T_pM \to \exists \gamma : \mathbb{R} \to M : X = v_{\gamma,p}$$
 and  $\exists \underbrace{X_1^1, \dots, X_d^d}_{\in \mathbb{R}} : X = X^i \left(\frac{\partial}{\partial x^i}\right)_p$ 

24.5. **5.** Change of vector components under a change of chart. **X** vector does not change under change of chart. Let (U, x) and (V, y) be overlapping charts and  $p \in U \cap V$ . Let  $X \in T_n M$ 

$$X_{(y)}^{i} \cdot \left(\frac{\partial}{\partial y^{i}}\right)_{p} \underbrace{=}_{V(y)} X \underbrace{=}_{(U,x)} X_{x}^{i} \left(\frac{\partial}{\partial x^{i}}\right)_{p}$$

to study change of components formula:

$$\left(\frac{\partial}{\partial x^{i}}\right)_{p} f = \partial_{i}(f \circ x^{-1})(x(p)) = 
= \partial_{i}(\underbrace{(f \circ y^{-1})}_{\mathbb{R}^{d} \to \mathbb{R}} \circ \underbrace{(y \circ x^{-1})}_{\mathbb{R}^{d} \to \mathbb{R}^{d}})(x(p)) 
= (\partial_{i}(y^{i} \circ x^{-1}))(x(p)) \cdot (\partial_{j}(f \circ y^{-1}))(y(p)) = 
= \left[\left(\frac{\partial y^{p}}{\partial x^{i}}\right)_{p} \cdot \left(\frac{\partial f}{\partial y^{j}}\right)_{p} f \right] 
\implies X_{(x)}^{i} \left(\frac{\partial y^{j}}{\partial x^{i}}\right)_{p} \left(\frac{\partial}{\partial y^{j}}\right)_{p} = X_{(y)}^{j} \left(\frac{\partial}{\partial y^{j}}\right)_{p} 
\implies X_{(y)}^{j} = \left(\frac{\partial y^{j}}{\partial x^{i}}\right)_{p} X_{(x)}^{i}$$

24.6. 6. Cotangent spaces.  $T_nM = V$ trivial  $(T_n M)^* := \{ \varphi : T_n M \xrightarrow{\sim} \mathbb{R} \}$ Example:  $f \in C^{\infty}(M)$ 

$$(df)_p: T_pM \xrightarrow{\sim} \mathbb{R}$$
  
 $X \mapsto (df)_p(X)$ 

i.e.  $(df)_p \in T_p M^*$  $(\overline{df})_p$  called the gradient of f  $p \in M$ . Calculate components of gradient w.r.t. chart-induced basis (U, x)

$$((df)_p)_j := (df)_p \left( \left( \frac{\partial}{\partial x^j} \right)_p \right)$$
$$= \left( \frac{\partial f}{\partial x^j} \right)_p = \partial_j (f \circ x^{-1})(x(p))$$

**Theorem 27.** Consider chart  $(U, x) \Longrightarrow x^i : U \to \mathbb{R}$ <u>Claim</u>:  $(dx^1)_p, (dx^2)_p, \dots, (dx^d)_p$  basis of  $T_p^*M$  $\implies$  In fact: dual basis:

$$(dx^a)_p\left(\left(\frac{\partial}{\partial x^b}\right)_p\right) = \left(\frac{\partial x^a}{\partial x^b}\right)_p = \dots = \delta_b^a$$

24.7. 7. Change of *components* of a covector under a change of chart:

$$\underbrace{T_p^* M}_{\ni \omega} \text{ with } \omega_{(y)} (dy^j)_p = \omega = \omega_{(x)i} (dx^i)_p$$

$$\Longrightarrow \boxed{\omega_{(y)i} = \frac{\partial x^j}{\partial y^i} \omega_{(x)j}}$$

# cf. Lecture 6: Fields (International Winter School on Gravity and Light 2015)

So far:

 $T_pM$   $\vdots$   $T^*M$ 

now

in Thought Cloud: theory of bundles

### 24.8. Bundles.

### **Definition 73.** A bundle is a triple

 $E \xrightarrow{\pi} M$ 

E smooth manifold "total space"  $\pi$  smooth map (surjective) "projection map" M smooth manifold "base space"

Example E = cylinder M = circle

### Definition 74.

$$E \xrightarrow{T} M$$

bundle.

 $p \in M$ define **fibre over** p:=  $preim_{\pi}(\{p\})$ 

**Definition 75.** A section  $\sigma$  of a bundle

$$\begin{array}{c}
E \\
\pi \downarrow \int \sigma \\
M
\end{array}$$

require  $\pi \circ \sigma = id_M$ 

Schuller says: in quantum mechanics, Aside:  $\psi: M \to \mathbb{C}$ 

### 24.9. Tangent bundle of smooth manifold. $(M, \mathcal{O}, \mathcal{A})$ smooth manifold

(a) as a set 
$$TM := \dot{\bigcup}_{p \in M} T_p M$$

(b) surjective  $\pi: TM \to M$  the unique point  $p \in M, X \in T_pM$ 

THE DIFFERENTIAL GEOMETRY DIFFERENTIAL TOPOLOGY DUMP

(c) Construct topology on TM that is the coarsest topology such that  $\pi$  (just) continuous. ("initial topology with respect

$$\mathcal{O}_{TM} := \{ \operatorname{preim}_{\pi}(U) | \mathcal{U} \in \mathcal{O} \}$$

Show: Tutorial  $\mathcal{O}_{TM}$  Schuller says this is shown in the tutorial  $(TM, \mathcal{O}_{TM})$ 

Construction of a  $C^{\infty}$ -atlas on TM from the  $C^{\infty}$ -atlas  $\mathcal{A}$  on M.

$$\mathcal{A}_{TM} := \{ (T\mathcal{U}, \xi_x) | (U, x) \in \mathcal{A} \}$$

where

$$\xi_x : T\mathcal{U} \to \mathbb{R}^{2 \cdot \dim M} \\
X \mapsto \underbrace{((x^1 \circ \pi)(X), \dots, (x^d \circ \pi)(X), (dx^1)_{\pi(X)}(X), \dots, (dx^d)_{\pi(X)}(X))}_{(U,x) - \text{ coords of } \pi(X) \ (d \text{ many})} (dx^1)_{\pi(X)}(X), \dots, (dx^d)_{\pi(X)}(X))$$

where  $X \in T_{\pi(X)}M$  $X = X_{(x)}^{i} \left(\frac{\partial}{\partial x^{i}}\right)_{\pi(X)}$ 

$$(dx^j)_{\pi(X)}(X) = (dx^j)_{\pi(X)} \left( X^i_{(x)} \left( \frac{\partial}{\partial x^i} \right)_{\pi(X)} \right) =$$

$$= X^i_{(x)} \delta^j_i = X^j_{(x)}$$

 $= X_{(x)}^{i} \delta_{i}^{j} = X_{(x)}^{j}$ <u>Write</u>  $\xi_{x}^{-1} : \xi_{x}(TU) \subseteq \mathbb{R}^{2\dim M} \to TU$ 

$$(\alpha^1, \dots, \alpha^d, \beta^1, \dots, \beta^d) := \beta^i \left(\frac{\partial}{\partial x^i}\right)_{\underbrace{x^{-1}(\alpha^1, \dots, \alpha^d)}_{\pi(X)}}$$

Check:

$$(\xi_y \circ \xi_x^{-1})(\alpha^1, \dots, \alpha^d, \beta^1, \dots, \beta^d) =$$

$$= \xi_y \left( \beta^i \left( \frac{\partial}{\partial x^i} \right)_{x^{-1}(\alpha^1, \dots, \alpha^d)} \right)$$

$$= \left( \dots, (y^i \circ \pi)(\beta^m \cdot \left( \frac{\partial}{\partial x^m} \right)_{x^{-1}(\alpha^1, \dots, \alpha^d)}), \dots, \dots (dy^i)_{x^{-1}(\alpha^1, \dots, \alpha^d)} \left( \beta^m \left( \frac{\partial}{\partial x^m} \right)_{x^{-1}(\alpha^1, \dots, \alpha^d)} \right), \dots \right) =$$

$$= (\dots, (y^i \circ x^{-1})(\alpha^1, \dots, \alpha^d), \dots, \dots, \underline{\beta}^m (dy^i)_{x^{-1}(\alpha^1, \dots, \alpha^d)} \left( \left( \frac{\partial}{\partial x^m} \right)_{x^{-1}(\alpha^1, \dots, \alpha^d)} \right) \right)$$

$$\beta^m \left( \frac{\partial y}{\partial x^m} \right)_{x^{-1}(\alpha^1, \dots, \alpha^d)}$$

Check transition map:  $(U, x), (V, y), U \cap V \neq 0$  $\left(\frac{\partial y}{\partial x^m}\right)_{x^{-1}(\alpha^1...\alpha^d)} = \partial_m(y^i \circ x^{-1})(x \circ (x^{-1}(\alpha^1...\alpha^d))) = \partial_m(y^i \circ x^{-1})(\alpha^1...\alpha^d) \text{ smooth.}$ upshot

$$TM$$
  $\xrightarrow{\pi}$   $M$ 

bundle, called the tangent bundle.

### 3. Vector fields.

**Definition 76.** A smooth vector field  $\chi$  is a smooth map, (where)

$$TM$$
 $\pi \downarrow \chi$ 
 $M$ 

Example:

$$TS^1$$

$$\downarrow^{\pi}$$
 $S^1$ 

4. The  $C^{\infty}(M)$ -module  $\Gamma(TM)$ .

$$C^{\infty}(M)$$
-module  $\leftarrow (C^{\infty}(M), +, \cdot)$  (satisfies)  $C^+, A^+, N^+, I^+, C^+, A^+, N^+, D^+$ . Not a field. A ring.

$$\mathbf{set} \ \Gamma(TM) = \{ \chi \quad M \to TM | \text{ smooth section } \}$$

$$(\chi \oplus \widetilde{\chi})(f) := (\chi f) \underbrace{+}_{C^{\infty}(M)} \widetilde{\chi}(f)$$

$$(\underbrace{g}_{C^{\infty}(M)} \odot \xi)(f) := \underbrace{g}_{C^{\infty}(M)} \cdot \chi(f)$$

$$\chi: M \to TM$$
$$p \mapsto \chi(p)$$
$$\chi f: M \to \mathbb{R}$$
$$p \mapsto \chi(p)f$$

$$(\Gamma(TM), \oplus, \odot) C^{\infty}(M)$$
 - module

upshot: set of all smooth vector fields can be made into a  $C^{\infty}(M)$ -module.

#### Fact

- (1)  $ZFC \Longrightarrow$  every vector space has a basis. (You have to have C axiom of choice in set theory)
- (2) no such result exists for modules.

This is a shame, because otherwise, we could have chosen (for any manifolds) vector fields,

$$\chi_{(1)}, \ldots, \chi_{(d)} \in \Gamma(TM)$$

and would be able to write every vector field  $\Xi$ 

$$\chi = \underbrace{f^i}_{\text{component functions}} \cdot \chi_{(i)}$$

### Simple counterexample

Schuller says: Take a sphere, Morse Theorem, every smooth vector field must vanish at 2 pts. "mustn't choose a global basis"

$$\underbrace{\text{However}\colon \frac{\partial}{\partial x^i} \colon U \xrightarrow{\text{smooth}} TU}_{p \mapsto \left(\frac{\partial}{\partial x^i}\right)_p}$$

24.10. **Tensor fields.** so far

 $\Gamma(M)$  = "set of vector fields"  $C^{\infty}(M)$ -module  $\Gamma(T^*M)$  = "covector fields"  $C^{\infty}(M)$ -module

**Definition 77.** An (r, s)-tensor field T is a multi-linear map

$$T: \underbrace{\Gamma(T^*M) \times \cdots \times \Gamma(T^*M)}_{r} \times \Gamma(TM) \times \cdots \times \Gamma(TM) \xrightarrow{\sim} C^{\infty}(M)$$

Example:  $f \in C^{\infty}(M)$ 

$$df: \Gamma(TM) \xrightarrow{\sim} C^{\infty}(M)$$
$$\chi \mapsto df(\chi) := \chi[f]$$

df (0,1)-T.F. (tensor field) where  $(\chi f)(\underbrace{p}) := \underbrace{\chi(p)}_{\in T_p M} f$ can check: df is  $C^{\infty}$ -linear

### Lecture 7: Connections

cf. Lecture 7: Connections (International Winter School on Gravity and Light 2015)

So far: saw that a vector field X can be used to provide a directional derivative

$$\nabla_X f := X f$$

of a function  $f \in C^{\infty}(M)$ .

Remark: from now on: consider mostly vector fields.

Notational overkill?

$$\nabla_X f = X f = (df)(X)$$

In Thought Bubble:  $\nabla_X(f \cdot g) = X(fg) = (Xf) \cdot g + fX(g)$  Product rule, because it's a derivative not quite:

$$X: C^{\infty}(M) \to C^{\infty}(M)$$
$$df: \Gamma(TM) \to C^{\infty}(M)$$
$$\nabla_X: C^{\infty}(M) \to C^{\infty}(M)$$

$$\nabla_X : C^{\infty}(M) \longrightarrow C^{\infty}(M)$$

$$\vdots \downarrow \qquad \qquad \vdots \downarrow$$

$$\nabla_X : \frac{TM^p \otimes T^*M^q \text{ i.e.}}{\binom{p}{q} \text{ tensor field}} \xrightarrow{\qquad \qquad} \frac{TM^p \otimes T^*M^q \text{ i.e.}}{\binom{p}{q} \text{ tensor field}}$$

1. Directional derivatives of tensor fields. We formulate a wish list of properties which the  $\nabla_X$  acting on a tensor field should have.

In Thought Bubble: Any remaining freedom in choosing  $\nabla$ , will need to be provided as additional structure beyond  $(M, \mathcal{O}, \mathcal{A})$ 

Definition 78 (connection). In Thought Bubble: linear connection, covariant derivatve, affine connection

A connection  $\nabla$  on a smooth manifold  $(M, \mathcal{O}, \mathcal{A})$  is a map that takes a pair consisting of a vector (field) X and a (p,q)-tensor field T and sends them to a (p,q)-tensor (field)  $\nabla_X T$  satisfying

- (i)  $\nabla_X f = Xf \quad \forall f \in C^{\infty}(M)$
- (ii)  $\nabla_X(T+S) = \nabla_X T + \nabla_X S$
- (iii)  $\nabla_X (\underbrace{T(\omega, Y)}_{\in C^{\infty}(M)}) = (\nabla_X T)(\omega, Y) + T(\nabla_X \omega, Y) + T(\omega, \nabla_X Y)$

In Thought Bubble: for (1,1)-TFT, but analogously for any (p,q)-TF

"Leibnitz" rule.

(iv) 
$$\nabla_{fX+Z}T = f\nabla_XT + \nabla_ZT$$
  
 $f \in C^{\infty}(M)$ 

A manifold with connection is quadruple  $(M, \mathcal{O}, \mathcal{A}, \nabla)$ 

Remark:  $\nabla_X$  is the extension of X.

$$\nabla$$
 — " — of  $d$ 

2. New structure on  $(M, \mathcal{O}, \mathcal{A})$  required to fix  $\nabla$ . Q: How much freedom do we have in choosing such a structure. Consider X, Y vector fields

$$(\underbrace{T}_{(p,q)} \otimes \underbrace{S}_{(r,s)})(\omega, \dots, X, \dots) := T(\omega, \dots, X, \dots) \underbrace{C}_{\infty(M)} S(\dots, \dots)$$

$$\nabla_X (T \otimes S) = (\nabla_X T) \otimes S + T \otimes (\nabla_X S)$$

**Definition 79** (Connection coefficient functions).  $(M, \mathcal{O}, \mathcal{A}, \nabla)$ ,  $(\mathcal{U}, x) \in \mathcal{A}$ .

Then the connection coefficient functions ("T"s) with respect to (wrt) (U,x) on the  $(\dim(M))^3$  many functions

$$\Gamma^{i}_{jk} : \mathcal{U} \to \mathbb{R}$$

$$p \mapsto \left( dx^{i} \left( \nabla_{\frac{\partial}{\partial x^{k}}} \frac{\partial}{\partial x^{j}} \right) \right) (p)$$

Thus:

$$(\nabla_X Y)^i = X^m \left( \frac{\partial}{\partial x^m} Y^i \right) + \Gamma^i_{nm} \underbrace{\cdot}_{C^{\infty}(M)} Y^n X^m$$

Remark: On a chart domain U, choice of the  $(\dim M)^3$  functions  $\Gamma^i_{jk}$  suffices to fix the action of  $\nabla$  on a vector field.

Fortunately, the same  $(\dim M)^3$  functions fix the action of  $\nabla$  on any tensor field. key observation:

$$\nabla_{\frac{\partial}{\partial x^m}}(dx^i) = \sum_{jm}^i dx^j$$

out now

$$\begin{split} & \nabla_{\frac{\partial}{\partial x^m}}(\underbrace{dx^i\left(\frac{\partial}{\partial x^j}\right)}) = \frac{\partial}{\partial x^m}(\delta^i_{\ j}) = 0 \\ & \underbrace{\qquad \qquad \qquad } \\ & = \left(\nabla_{\frac{\partial}{\partial x^m}}dx^i\right)\left(\frac{\partial}{\partial x^j}\right) + dx^i(\underbrace{\nabla_{\frac{\partial}{\partial x^m}}\frac{\partial}{\partial x^j}}_{\Gamma^q_{jm}\frac{\partial}{\partial x^q}}) = 0 \\ & \Longrightarrow \left(\nabla_{\frac{\partial}{\partial x^m}}dx^i\right)\left(\frac{\partial}{\partial x^j}\right) = -\Gamma^i_{jm} \end{split}$$

Summary so far:

$$(\nabla_X Y)^i = X(Y^i) \underbrace{+}_{\text{act on vector field}} \Gamma^i_{jm} Y^j X^m$$
$$(\nabla_X \omega)_i = X(\omega_i) + -\Gamma^j_{im} \omega_j X^m$$

Note that for the immediately above expression for  $(\nabla_X Y)^i$ , in the second term on the right hand side,  $\Gamma^i_{jm}$  has the last entry at the bottom, m going in the direction of X, so that it matches up with  $X^m$ . This is a good mnemonic to memorize the index positions of  $\Gamma$ .

similarly, by further application of Leibnitz

T a (1,2)-TF (tensor field)

$$(\nabla_X T)^i_{jk} = X(T^i_{jk}) + \Gamma^i_{sm} T^s_{jk} X^m - \Gamma^s_{jm} T^i_{sk} X^m - \Gamma^s_{km} T^i_{js} X^m$$

Question: If in a Euclidean space, the  $\Gamma$ s all vanish in a (then existing) global chart.

Answer: Yes, but: What is a Euclidean space:

 $(M = \mathbb{R}^n, \mathcal{O}_{st}, \mathcal{A})$  smooth manifold.

Assume  $(\mathbb{R}^n, \mathrm{id}_{\mathbb{R}^n}) \in \mathcal{A}$  and

$$(\Gamma^{i}_{(x)})_{jk} = dx^{i} \left( (\nabla_{\underline{\mathbf{E}}})_{\frac{\partial}{\partial x^{k}}} \frac{\partial}{\partial x^{j}} \right) \stackrel{!}{=} 0$$

Intuition:

 $\mathbb{R}^2$ :  $\nabla_{\text{Euclidean}}$ 

 $\mathbb{R}^2$ :  $\nabla_{\text{Hyperbolic}}$ 

**Definition 80.** X vector field on  $(M, \mathcal{O}, \mathcal{A}, \nabla)$ .

 $div(X) := \left(\nabla_{\frac{\partial}{\partial x^i}} X\right)^i$ 

Then divergence of 
$$X$$
 is the function:

Claim: chart-independent.

### 3. Change of $\Gamma$ 's under change of chart.

 $(U,x), (V,y) \in \mathcal{A} \text{ and } U \cap V \neq \emptyset$ 

$$\Gamma^{i}_{jk}(y) := dy^{i} \left( \nabla_{\frac{\partial}{\partial y^{k}}} \frac{\partial}{\partial y^{j}} \right) = \frac{\partial y^{i}}{\partial x^{q}} dx^{q} \left( \nabla_{\frac{\partial x^{p}}{\partial y^{k}}} \frac{\partial}{\partial x^{p}} \frac{\partial x^{s}}{\partial y^{j}} \frac{\partial}{\partial x^{s}} \right)$$

Note  $\nabla_{fX}$  is  $C^{\infty}$ -linear for fX covector  $dy^i$  is  $C^{\infty}$ -linear in its argument

$$\implies \Gamma^{i}_{jk}(y) = \frac{\partial y^{i}}{\partial x^{q}} dx^{q} \left( \frac{\partial x^{p}}{\partial y^{k}} \left[ \left( \nabla_{\frac{\partial}{\partial x^{p}}} \frac{\partial x^{s}}{\partial y^{j}} \right) \frac{\partial}{\partial x^{s}} + \frac{\partial x^{s}}{\partial y^{j}} \left( \nabla_{\frac{\partial}{\partial x^{p}}} \frac{\partial}{\partial x^{s}} \right) \right] \right) =$$

$$= \frac{\partial y^{i}}{\partial x^{q}} \underbrace{\frac{\partial x^{p}}{\partial y^{k}} \frac{\partial}{\partial x^{p}}}_{\frac{\partial}{\partial x^{p}}} \underbrace{\frac{\partial x^{s}}{\partial y^{j}} \delta^{q}_{s} + \frac{\partial y^{i}}{\partial x^{q}} \frac{\partial x^{p}}{\partial y^{k}} \frac{\partial x^{s}}{\partial y^{j}} \Gamma^{q}_{sp}(x)$$

in summary:

(117) 
$$\Gamma_{jk}^{i}(y) = \frac{\partial y^{i}}{\partial x^{q}} \frac{\partial^{2} x^{q}}{\partial y^{j} \partial y^{k}} + \frac{\partial y^{i}}{\partial x^{q}} \frac{\partial x^{s}}{\partial y^{i}} \frac{\partial x^{p}}{\partial y^{k}} \Gamma_{sp}^{q}(x)$$

Eq. (117) is the change of connection coefficient function under the change of chart  $(U \cap V, x) \to (U \cap V, y)$ 

### **4. Normal Coordinates.** Let $p \in M$ of $(M, \mathcal{O}, \mathcal{A}, \nabla)$

Then one can construct a chart (U, x) with  $p \in U$  such that

$$\Gamma(x)^{i}_{(ik)}(p) = 0$$

at the point p. Not necessarily in any neighborhood.

*Proof.* Let (V, y) be any chart (with)  $p \in V$ .

Thus, in general:  $\Gamma(y)^{i}_{ik} \neq 0$ 

Then consider a new chart (U, x) to which one transits by virtue of

$$(x \circ y^{-1})^i(\alpha^1, \dots, \alpha^d) := \alpha^i - \frac{1}{2}\Gamma(y)^i_{(jk)}(p)\alpha^j\alpha^k$$

 $p = x^{-1}(\alpha^1, \dots, \alpha^d)$ 

$$\left( \frac{\partial x^i}{\partial y^j} \right)_p = \partial_j (x^i \circ y^{-1}) = \delta_j^i - \Gamma(y)^i{}_{mj}(p) \alpha^m |_{\alpha=0} = \delta_j^i$$

$$\frac{\partial x^i}{\partial y^k \partial y^j}(p) = -\Gamma(y)^i{}_{kj}(p)$$

$$\Longrightarrow \Gamma(x)_{jk}^{i}(p) = \Gamma(y)_{jk}^{i}(p) - \Gamma(y)_{kj}^{i}(p) = 0$$
$$= \Gamma(y)_{[jk]}^{i}(p) = T(y)_{jk}^{i}$$

Terminology: (U, x) is called a **normal coordinate chart** of  $\nabla$  at  $\mathbf{p} \in \mathbf{M}$ .

Tutorial 7 Connections. Exercise 1.: True or false?

- (a)  $\bullet \nabla_{fX}Y = f\nabla_XY$  by definition so  $\nabla_{fX} = f\nabla_X$  i.e.  $\nabla_X$  is  $C^{\infty}(M)$ -linear in X
  - $f \in C^{\infty}(M)$  is a (0,0)-tensor field.  $\nabla_X f = Xf \equiv X(f)$  by definition.
  - If the manifold is flat, I'm assuming that means that the manifold is globally a Euclidean space, and by definition,  $\Gamma = 0$ .

$$\nabla_X Y = X^j \frac{\partial}{\partial x^j} (Y^i) \frac{\partial}{\partial x^i} + \Gamma^i_{jk} Y^k X^k \frac{\partial}{\partial x^i} = X^j \frac{\partial Y^i}{\partial x^j} \frac{\partial}{\partial x^i} + 0$$

and similarly for any (p,q)-tensor field, i.e.

$$\nabla_X T = X^j \frac{\partial T^{i_1 \dots i_p}_{j_1 \dots j_q}}{\partial x^j}$$

 $\nabla_X f = X^j \frac{\partial f}{\partial x^j} = X \cdot \operatorname{grad}(f)$ 

•  $\forall (U, x) \in \mathcal{A}$ , locally (after working out the first few cases, and doing induction, one can look up the expression for the local form; I found it in Nakahara's **Geometry**, **Topology and Physics**, Eq. 7.26, and it needs to be modified for the convention of order of bottom indices for  $\Gamma$ :

$$\nabla_{\nu} t_{\mu_{1}...\mu_{q}}^{\lambda_{1}...\lambda_{p}} = \partial_{\nu} t_{\mu_{1}...\mu_{q}}^{\lambda_{1}...\lambda_{p}} + \Gamma_{\kappa\nu}^{\lambda_{1}} t_{\mu_{1}...\mu_{q}}^{\kappa\lambda_{2}...\lambda_{p}} + \dots + \Gamma_{\kappa\nu}^{\lambda_{p}} t_{\mu_{1}...\mu_{q}}^{\lambda_{1}...\lambda_{p-1}\kappa} - \Gamma_{\mu_{1}\nu}^{\kappa} t_{\kappa\mu_{2}...\mu_{q}}^{\lambda_{1}...\lambda_{p}} - \dots - \Gamma_{\mu_{q}\nu}^{\kappa} t_{\mu_{1}...\mu_{q-1}\kappa}^{\lambda_{1}...\lambda_{p}}$$

Clearly,  $\nabla_X$  is uniquely fixed  $\forall p \in M$  by choosing each of the  $(\dim M)^3$  many connection coefficient functions  $\Gamma$ .

(b)  $\bullet \nabla : \mathfrak{X}(M) \to \mathfrak{X}(M)$ 

 $\nabla: (p,q)$ -tensor field  $\mapsto (p,q)$ -tensor field

- By definition,  $\nabla$  satisfies the Leibniz rule.
- •
- •

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Exercise 2. : Practical rules for how  $\nabla$  acts Torsion-free covariant derivative boils down to a connection coefficient function  $\Gamma$  that is symmetric in the bottom indices.

$$\nabla_X f = X(f) = X^i \frac{\partial f}{\partial x^i}$$

 $(\nabla_X Y)^a = X^i \frac{\partial Y^a}{\partial x^i} + \Gamma^a_{jk} Y^j X^k$ 

$$(\nabla_X \omega)_a = X^i \frac{\partial \omega_a}{\partial x^j} - \Gamma^i_{ak} \omega_i X^k$$

 $(\nabla_m T)^a_{\ bc} = \frac{\partial}{\partial x^m} (T^a_{\ bc}) + \Gamma^a_{\ im} T^i_{bc} - \Gamma^i_{bm} T^a_{ic} - \Gamma^j_{cm} T^a_{bj}$ 

$$(\nabla_{[m} A)_{n]} = (\nabla_{m} A)_{n} - (\nabla_{n} A)_{m} = \frac{\partial A_{n}}{\partial x^{m}} - \Gamma_{nm}^{i} A_{i} - \left(\frac{\partial A_{m}}{\partial x^{n}} - \Gamma_{mn}^{i} A_{i}\right) = \frac{\partial A_{m}}{\partial x^{m}} - \frac{\partial A_{m}}{\partial x^{n}}$$

$$(\nabla_m \omega)_{nr} = rac{\partial \omega_{nr}}{\partial x^m} - \Gamma^i_{nm} \omega_{ir} - \Gamma^i_{rm} \omega_{ni}$$

#### Exercise 3. : Connection coefficients

### Question .

The connection coefficient functions  $\Gamma$  in chart  $(U \cap V, y)$  is given, in terms of chart  $(U \cap V, x)$  as follows: Recall Eq. (117)

$$\Gamma^{i}_{jk}(y) = \frac{\partial y^{i}}{\partial x^{q}} \frac{\partial^{2} x^{q}}{\partial y^{j} \partial y^{k}} + \frac{\partial y^{i}}{\partial x^{q}} \frac{\partial x^{s}}{\partial y^{j}} \frac{\partial x^{p}}{\partial y^{k}} \Gamma^{q}_{sp}(x)$$

LECTURE 8: PARALLEL TRANSPORT & CURVATURE (INTERNATIONAL WINTER SCHOOL ON GRAVITY AND LIGHT 2015)

# **Experiment.** $\nabla_{v_n} Y \stackrel{!}{=} 0$

(please move in this fashion)

round sphere 
$$(S^2, \mathcal{O}, \mathcal{A}, \nabla)$$

ellipsoid

$$(S^2, \mathcal{O}, \mathcal{A}, \widetilde{\nabla})$$

ellisoid embedded in  $\mathbb{R}^3$  with  $\mathbb{R}^3$  Euclidean connection  $(\mathbb{R}^3, \nabla_E)$ 

1. Parallelity of vector fields.  $(M, \mathcal{O}, \mathcal{A}, \nabla)$  manifold with connection.

(1) A vector field X on M is said to be parallely transported along a **Definition 81** (parallely transported, parallel). smooth curve  $\gamma: \mathbb{R} \to M$ 

$$\nabla_{v_{\gamma}} X = 0$$

i.e.

$$\left(\nabla_{v_{\gamma,\gamma(\lambda)}} X\right)_{\gamma(\lambda)} = 0$$

(2) A slightly weaker condition than "parallely transported" is "parallel":

$$(\nabla_{v_{\gamma,\gamma(\lambda)}} X)_{\gamma(\lambda)} = \mu(\lambda) X_{\gamma(\lambda)}$$

for  $\mu: \mathbb{R} \to \mathbb{R}$ 

Example: Euclidean plane  $(\mathbb{R}^2, \mathcal{O}, \mathcal{A}, \nabla_E)$ ,

parallely transported X along  $\gamma$ , parallel X along  $\gamma$ , not parallel, not parallely transported.

2. Autoparallely transported curves. Thought bubble: self- is "Auto"

Child: "How do I get to the toy store." Adult: "Just follow your nose."

(Move in direction of nose means move along curve whose tangent vector is your nose. Means move in the straightest line. We are not talking about shortest line because we do not have a notion of distance.)

**Definition 82.** A (smooth) curve  $\gamma : \mathbb{R} \to M$  is called autoparallely transported if

$$\nabla_{v_{\gamma}} v_{\gamma} = 0$$

$$\iff \left(\nabla_{v_{\gamma,\gamma(\lambda)}} v_{\gamma}\right)_{\gamma(\lambda)} = 0$$

Remark: Define weaker notion of autoparallel curve

Thought bubble: !

$$\nabla_{v_{\gamma}} v_{\gamma} = \mu v_{\gamma}$$

Example. Euclidean plane  $(\mathbb{R}^2, \mathcal{O}, \mathcal{A}, \nabla_E)$ 

autoparallely transported (uniform straight motion, Newton's First law), autoparallel (straight motion) but not autoparallely transported,

### 3. Autoparallel equation.

Autoparallely transported curve  $\gamma$ 

Consider that portion of the curve that lies in U, where  $(U, x) \in \mathcal{A}$ .

$$\nabla_{v_{\gamma}} v_{\gamma} = 0$$

Express  $\nabla_{v_{\gamma}} v_{\gamma} = 0$  in terms of chart representation.

$$(\nabla_{v_{\gamma}} v_{\gamma}) = \left(\nabla_{\dot{\gamma}^m(x)\left(\frac{\partial}{\partial x^m}\right)_{\gamma(\lambda)}} \dot{\gamma}^n(x) \frac{\partial}{\partial x^n}\right)$$

where

$$v_{\gamma,\gamma(\lambda)} = \dot{\gamma}^m(x) \cdot \left(\frac{\partial}{\partial x^m}\right)$$
  
 $\gamma^m(x) := x^m \circ \gamma$ 

$$\Rightarrow (\nabla_{v_{\gamma}} v_{\gamma}) =$$

$$= \underbrace{\dot{\gamma}^{m} \frac{\partial \dot{\gamma}^{q}}{\partial x^{m}}}_{\ddot{\gamma}(x)^{m}} \cdot \frac{\partial}{\partial x^{q}} + \dot{\gamma}^{m} \dot{\gamma}^{n} \Gamma^{q}_{nm} \frac{\partial}{\partial x^{q}}$$

20200306 EY: For further clarification.

$$\dot{\gamma}(t) = v(\gamma(t)) \in T_{\gamma(t)}U$$

$$\dot{\gamma}(t) = v^{i}(\gamma^{1}(t), \dots, \gamma^{n}(t))$$

$$\frac{d}{dt}\dot{\gamma}^{i}(t) = \frac{d}{dt}v^{i}(\gamma^{1}(t), \dots, \gamma^{n}(t)) = \frac{d}{dt}v^{i}(x^{1}, \dots, x^{n})$$

$$= \frac{\partial v^{i}}{\partial x^{j}}(x)\dot{x}^{j} = \dot{\gamma}^{j}\frac{\partial \dot{\gamma}^{i}}{\partial x^{j}}$$

in summary:

(119) 
$$\ddot{\gamma}_{(x)}^{m}(\lambda) + (\Gamma_{(x)}^{m})_{ab}(\gamma(\lambda))\dot{\gamma}_{(x)}^{a}(\lambda)\dot{\gamma}_{(x)}^{b}(\lambda) = 0$$

chart expression of the condition that  $\gamma$  be autoparallely transported.

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Examples.

(a) Euclidean plane,  $U = \mathbb{R}^2$ ,  $x = \mathrm{id}_{\mathbb{R}}$ ,  $\Gamma(x)^i_{jk} = 0$ 

$$\implies \ddot{\gamma}(x)^m = 0 \implies \gamma(x)^m(\lambda) = a^m \lambda + b^m a, b \in \mathbb{R}^d$$

(b) round sphere  $(S^2, \mathcal{O}, \mathcal{A}, \nabla_{\text{round}})$ 

Consider a chart:

$$x(p) = (\theta, \varphi)$$
$$\theta \in (0, \pi)$$
$$\varphi \in (0, 2\pi)$$

all other  $(\Gamma^1_{11} = 0)$  vanish.

$$\Gamma(x)_{22}^{1}(x^{-1}(\theta, p)) := -\sin\theta\cos\theta \quad \Gamma(x)_{21}^{2} = \Gamma(x)_{12}^{2} := \cot\theta$$

Sloppy notation (classical mechanics)

$$x^{1}(p) = \theta(p)$$
$$x^{2}(p) = \varphi(p)$$

autoparallel equation

$$\begin{split} \ddot{\theta} + \Gamma^1_{22} \dot{\varphi} \dot{\varphi} &= 0 \\ \ddot{\varphi} + 2 \Gamma^2_{12} \dot{\theta} \dot{\varphi} &= 0 \\ \\ \Longleftrightarrow & \begin{vmatrix} \ddot{\theta} - \sin{(\theta)} \cos{(\theta)} \dot{\varphi} \dot{\varphi} &= 0 \\ \ddot{\varphi} + 2 \cot{(\theta)} \dot{\theta} \dot{\varphi} &= 0 \end{vmatrix} \end{split}$$

Solve: e.g.  $\theta(\lambda) = \frac{\pi}{2}$   $\varphi(\lambda) = \underline{\omega} \cdot \lambda + \varphi_0$ Similarly,

 $\nabla_{\mathrm{potato}}, (S^2, \mathcal{O}, \mathcal{A})$ 

**3. Torsion.** Q: Can one use  $\nabla$  to define tensors on  $(M, \mathcal{O}, \mathcal{A}, \nabla)$ ?

**Definition 83.** The torsion of a connection  $\nabla$  is the (1,2)-tensor field

$$(120) \hspace{3.1em} T(\omega,X,Y) := \omega(\nabla_X Y - \nabla_Y X - [X,Y])$$

(Inside a cloud)

[X,Y] vector field defined by

$$[X,Y]f := X(Yf) - Y(Xf)$$

*Proof.* check T is  $C^{\infty}$ -linear in each entry

$$T(f \cdot \omega, X, Y) = f \cdot \omega(\dots) = fT(\omega, X, Y)$$

$$T(\omega + \psi, X, Y) = \dots = T(\omega, X, Y) + T(\psi, X, Y)$$

$$T(\omega, fX, Y) = \omega(\nabla_{fX}Y - \nabla_{Y}(fX) - [fX, Y]) =$$

$$= \omega(f\nabla_{X}Y - (Yf)X - f\nabla_{Y}X - f[X, Y] + (Yf)X) = fT(\omega, X, Y)$$

since

$$[fX,Y]g = fX(Yg) - Y(fXg) =$$

$$= fX(Yg) - (Yf)(Xg) - fY(Xg) =$$

$$= f \cdot [X,Y]g - (Yf)Xg =$$

$$= (f \cdot [X,Y] - (Yf)X)g$$

$$T(\omega, X, Y) = -T(\omega, Y, X) \quad \checkmark$$

**Definition 84.** A  $(M, \mathcal{O}, \mathcal{A}, \nabla)$  is called torsion-free if T = 0

In a chart

$$T^{i}_{ab} := T\left(dx^{i}, \frac{\partial}{\partial x^{a}}, \frac{\partial}{\partial x^{b}}\right) = dx^{i}(\dots)$$
  
=  $\Gamma^{i}_{ab} - \Gamma^{i}_{ba} = 2\Gamma^{i}_{[ab]}$ 

From now on, in these lectures, we only use torsion-free connections.

### 4. Curvature.

**Definition 85.** The Riemann curvature of a connection  $\nabla$  is the (1,3)-tensor field

(121) 
$$Riem(\omega, Z, X, Y) := \omega(\nabla_X \nabla_Y Z - \nabla_Y \nabla_X Z - \nabla_{[X,Y]} Z)$$

*Proof.* do it:  $C^{\infty}$ -linear in each slot.

<u>Tutorials</u> Riem $^{i}_{jab} = \dots$ 

Algebraic relevance of Riem.

$$(\nabla_X \nabla_Y Z - \nabla_Y \nabla_X Z) = R(\cdot, Z, X, Y) + \nabla_{[X,Y]} Z$$

In one chart (U, x)

$$(\underbrace{\nabla_a)}_{=\nabla_{\frac{\partial}{\partial x^a}}} \nabla_b Z - \nabla_b \nabla_a Z)^m = \operatorname{Riem}_{nab}^m Z^n + \nabla_{\underbrace{\frac{\partial}{\partial x^a}, \frac{\partial}{\partial x^b}}_{=0}} Z^n$$

$$\left| (\nabla_a \nabla_b Z)^m - (\nabla_b \nabla_a Z)^m = \operatorname{Riem}_{nab}^m Z^n \right|$$

Geometric significance.

$$\frac{\text{Idea}:}{[X,Y]=0}$$

If 
$$[X, Y] \neq 0$$

$$(\delta Z)^m = \dots = R^m_{nab} X^a Y^b Z^n \delta s \delta t + \mathcal{O}(\delta s^2 \delta t, \delta t^2 \delta s)$$

Holonomy

### Tutorial 8 Parallel transport & Curvature

#### Exercise 1.

### Exercise 2. : Where connection coefficients appear

It was suggested in the tutorial sheets and hinted in the lecture that the following should be committed to memory.

Question : Recall the autoparallel equation for a curve  $\gamma$ .

(a) 
$$\nabla_{v_{\gamma}} v_{\gamma} = 0$$

(b) 
$$\nabla_{v_{\gamma}} v_{\gamma} = \nabla_{\dot{\gamma} \frac{\partial}{\partial x^{\mu}}} v_{\gamma} = \dot{\gamma}^{\nu} \nabla_{\partial_{\nu}} v_{\gamma} = \dot{\gamma}^{\nu} \left[ \frac{\partial v_{\gamma}^{\mu}}{\partial x^{\nu}} + \Gamma_{\mu\nu}^{\rho} v_{\gamma}^{\mu} \right] \frac{\partial}{\partial x^{\rho}} = \dot{\gamma}^{\nu} \left[ \frac{\partial \dot{\gamma}^{\rho}}{\partial x^{\nu}} + \Gamma_{\mu\nu}^{\rho} \dot{\gamma}^{\mu} \right] \frac{\partial}{\partial x^{\rho}} = 0$$

$$\Longrightarrow \left[ \ddot{\gamma}^{\rho} + \Gamma_{\mu\nu}^{\rho} \dot{\gamma}^{\mu} \dot{\gamma}^{\nu} \right]$$
as, for example, for  $F(x(t))$ ,
$$\frac{dF(x(t))}{dt} = \dot{x} \frac{\partial F}{\partial x} = \frac{d}{dt} F$$
so that
$$\dot{\gamma}^{\nu} \frac{\partial v_{\gamma}^{\mu}}{\partial x^{\nu}} = \frac{d}{d\lambda} v_{\gamma}^{\mu} = \frac{d^{2}}{d\lambda^{2}} \gamma^{\mu}$$

Question: Determine the coefficients of the Riemann tensor with respect to a chart (U,x).

Recall this manifestly covariant definition

$$Riem(\omega, Z, X, Y) = \omega(\nabla_X \nabla_Y Z - \nabla_Y \nabla_X Z - \nabla_{[X,Y]} Z)$$

We want  $R^{i}_{jab}$ .

$$\nabla_{X}\nabla_{Y}Z = \nabla_{X}((Y^{\mu}\frac{\partial}{\partial x^{\mu}}Z^{\rho} + \Gamma^{\rho}_{\mu\nu}Z^{\mu}Y^{\nu})\frac{\partial}{\partial x^{\rho}}) = (X^{\alpha}\frac{\partial}{\partial x^{\alpha}}(Y^{\mu}\frac{\partial}{\partial x^{\mu}}Z^{\rho} + \Gamma^{\rho}_{\mu\nu}Z^{\mu}Y^{\nu}) + \Gamma^{\rho}_{\alpha\beta}(Y^{\mu}\frac{\partial}{\partial x^{\mu}}Z^{\alpha} + \Gamma^{\alpha}_{\mu\nu}Z^{\mu}Y^{\nu})X^{\beta})\frac{\partial}{\partial x^{\rho}}(Y^{\mu}\frac{\partial}{\partial x^{\mu}}Z^{\alpha} + \Gamma^{\alpha}_{\mu\nu}Z^{\mu}Y^{\nu})X^{\beta})\frac{\partial}{\partial x^{\mu}}(Y^{\mu}\frac{\partial}{\partial x^{\mu}}Z^{\alpha} + \Gamma^{\alpha}_{\mu\nu}Z^{\mu}Y^{\nu})X^{\beta})\frac{\partial}{\partial x^{\mu}}(Y^{\mu}\frac{\partial}{\partial x^{\mu}}Z^{\alpha} + \Gamma^{\alpha}_{\mu\nu}Z^{\mu}Y^{\nu})X^{\beta})\frac{\partial}{\partial x^{\mu}}(Y^{\mu}\frac{\partial}{\partial x^{\mu}}Z^{\alpha} + \Gamma^{\alpha}_{\mu\nu}Z^{\mu}Y^{\nu})X^{\beta})\frac{\partial}{\partial x^{\mu}}(Y^{\mu}\frac{\partial}{\partial x^{\mu}}Z^{\mu}Y^{\nu})X^{\beta}$$

For  $X = \partial_a$ ,  $Y = \partial_b$ ,  $Z = \partial_j$ , then the partial derivatives of the coefficients of the input vectors become zero.

$$\Longrightarrow \nabla_{\partial_a} \nabla_{\partial_b} \partial_j = \frac{\partial}{\partial x^a} (\Gamma^i_{jb}) + \Gamma^i_{\alpha a} \Gamma^{\alpha}_{jb}$$

Now

$$[X,Y]^{i} = X^{j} \frac{\partial}{\partial x^{j}} Y^{i} - Y^{j} \frac{\partial X^{i}}{\partial x^{j}}$$

For coordinate vectors,  $[\partial_i, \partial_j] = 0 \ \forall i, j = 0, 1 \dots d$ 

Thus

$$\boxed{R^i_{\ jab} = \frac{\partial}{\partial x^a} \Gamma^i_{jb} - \frac{\partial}{\partial x^b} \Gamma^i_{ja} + \Gamma^i_{\alpha a} \Gamma^\alpha_{jb} - \Gamma^i_{\alpha b} \Gamma^\alpha_{ja}}$$

Question :Ric(X, Y) := Riem<sup>m</sup><sub>amb</sub> $X^aY^b$  define (0, 2)-tensor?.

Yes, transforms as such:

EY developments. I roughly follow the spirit in Theodore Frankel's The Geometry of Physics: An Introduction Second Ed. 2003, Chapter 9 Covariant Differentiation and Curvature, Section 9.3b. The Covariant Differential of a Vector Field. P.S. EY: 20150320 I would like a copy of the Third Edition but I don't have the funds right now to purchase the third edition: go to my tilt crowdfunding campaign, http://ernestyalumni.tilt.com, and help with your financial support if you can or send me a message on my various channels and ernestyalumni gmail email address if you could help me get a hold of a digital or hard copy as a pro bono gift from the publisher or author.

The spirit of the development is the following:

"How can we express connections and curvatures in terms of forms?" -Theodore Frankel.

From Lecture 7, connection  $\nabla$  on vector field Y, in the "direction" X,

$$\nabla_{\frac{\partial}{\partial x^k}} Y = \left(\frac{\partial Y^i}{\partial x^k} + \Gamma^i_{jk} Y^j\right) \frac{\partial}{\partial x^i}$$

Make the ansatz (approche, impostazione) that the connection  $\nabla$  acts on Y, the vector field, first:

$$\nabla Y(X) = \left(X^k \frac{\partial Y^i}{\partial x^k} + \Gamma^i_{jk} Y^j X^k\right) \frac{\partial}{\partial x^i} = X^k \left(\nabla_{\frac{\partial}{\partial x^k}} Y\right)^i \frac{\partial}{\partial x^i} = (\nabla_X Y)^i \frac{\partial}{\partial x^i} = \nabla_X Y$$

Now from Lecture 7, Definition for  $\Gamma$ ,

$$dx^{i}\left(\nabla_{\frac{\partial}{\partial x^{k}}}\frac{\partial}{\partial x^{j}}\right) = \Gamma_{jk}^{i}$$

Make this ansatz (approche, impostazine)

$$\nabla \frac{\partial}{\partial x^{j}} = \left(\Gamma^{i}_{jk} dx^{k}\right) \otimes \frac{\partial}{\partial x^{i}} \in \Omega^{1}(M, TM) = T^{*}M \otimes TM$$

where  $\Omega^1(M,TM) = T^*M \otimes TM$  is the set of all TM or vector-valued 1-forms on M, with the 1-form being the following:

$$\Gamma^{i}_{jk}dx^{k} = \Gamma^{i}_{j} \in \Omega^{1}(M)$$
  $i = 1 \dots \dim(M)$   
 $j = 1 \dots \dim(M)$ 

So  $\Gamma^i_i$  is a dim $M \times \text{dim}M$  matrix of 1-forms (EY !!!).

Thus

$$\nabla Y = (d(Y^i) + \Gamma^i_j Y^j) \otimes \frac{\partial}{\partial x^i}$$

So the connection is a (smooth) map from TM to the set of all vector-valued 1-forms on M,  $\Omega^1(M,TM)$ , and then, after "eating" a vector Y, yields the "covariant derivative":

$$\nabla: TM \to \Omega^{1}(M, TM) = T^{*}M \otimes TM$$

$$\nabla: Y \mapsto \nabla Y$$

$$\nabla Y: TM \to TM$$

$$\nabla Y(X) \mapsto \nabla Y(X) = \nabla_{X}(Y)$$

Now

$$\left[\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}\right] f = \frac{\partial}{\partial x^i} \left(\frac{\partial}{\partial x^j}\right) - \frac{\partial}{\partial x^j} \left(\frac{\partial}{\partial x^i}\right) = 0$$

(this is okay as on  $p \in (U, x)$ ; x-coordinates on same chart (U, x))

EY: 20150320 My question is when is this nontrivial or nonvanishing (i.e. not equal to 0).

$$[e_a, e_b] = ?$$

for a frame  $(e_c)$  and would this be the difference between a tangent bundle TM vs. a (general) vector bundle? Wikipedia helps here. cf. wikipedia, "Connection (vector bundle)"

$$\nabla : \Gamma(E) \to \Gamma(T^*M \otimes E) = \Omega^1(M, E)$$
$$\nabla e_a = \omega_{ab}^c f^b \otimes e_c$$

 $f^b \in T^*M$  (this is the dual basis for TM and, note, this is for the manifold, M

$$\nabla_{f_b} e_a = \omega_{ab}^c e_c \in E$$
$$\omega_a^c = \omega_{ab}^c f^b \in \Omega^1(M)$$

is the connection 1-form, with  $a, c = 1 \dots \dim V$ . EY: 20150320 This V is a vector space living on each of the fibers of E. I know that  $\Gamma(T^*M \otimes E)$  looks like it should take values in E, but it's meaning that it takes vector values of V. Correct me if I'm wrong: ernestyalumni at gmail and various social media.

Let  $\sigma \in \Gamma(E)$ ,  $\sigma = \sigma^a e_a$ 

$$\nabla \sigma = (d\sigma^c + \omega_{ab}^c \sigma^a f^b) \otimes e_c \text{ with}$$

$$d\sigma^c = \frac{\partial \sigma^c}{\partial x^b} f^b$$

$$\Longrightarrow \nabla_X \sigma = \left( X^b \frac{\partial \sigma^c}{\partial x^b} + \omega_{ab}^c \sigma^a X^b \right) e_c = X^b \left( \frac{\partial \sigma^c}{\partial x^b} + \omega_{ab}^c \sigma^a \right) e_c$$

Lecture 9: Newtonian spacetime is curved!

**Axiom 1** (Newton I:). A body on which no force acts moves uniformly along a straight line

**Axiom 2** (Newton II:). Deviation of a body's motion from such uniform straight motion is effected by a force, reduced by a factor of the body's reciprocal mass.

### Remark:

- (1) 1st axiom in order to be relevant must be read as a measurement prescription for the geometry of space ...
- (2) Since gravity universally acts on every particle, in a universe with at least two particles, gravity must not be considered a force if Newton I is supposed to remain applicable.

### 24.11. Laplace's questions. Laplace \* 1749

Q: "Can gravity be encoded in a curvature of space, such that its effects show if particles under the influence of (no other) force we postulated to move along straight lines in this curved space?"

Answer: No!

*Proof.* gravity is a force point of view

$$m\ddot{x}^{\alpha}(t) = F^{\alpha}(x(t))$$
  
 $m\ddot{x}^{\alpha}(t) = \underbrace{mf^{\alpha}(x(t))}_{\text{Pa}}$ 

 $-\partial_{\alpha} f^{\alpha} = 4\pi G \rho$  (Poisson)  $\rho$  mass density of matter

$$m\ddot{x}^{\alpha}(t) = \underbrace{mf^{\alpha}}_{F^{\alpha}}(x(t))$$

True?

(EY: 20150330) You know this,  $F = Gm_1m_2/r^2$ 

Yes

weak equivalence principle

$$\ddot{x}^{\alpha}(t) - f^{\alpha}(x(t)) = 0$$

Laplace asks: Is this  $(\ddot{x}(t))$  of the form

$$\ddot{x}^{\alpha}(t) + \Gamma^{\alpha}_{\beta\gamma}(x(t))\dot{x}^{\beta}(t)\dot{x}^{\gamma}(t) = 0$$

Conclusion: One cannot find  $\Gamma$  s such that Newton's equation takes the form of an autoparallel.

Question (from audience) We can evaluate the autoparallel equation pointwise?! But at each point, we can set the Gammas

Then, one should be able to write Newton's second law in the usual form?

Prof. Schuller: you (observer) fall with the mass (i.e. accelerated reference frame) and so you transform Γ's to be 0. The problem with this is if you do the same experiment in the North pole and fall with the body. If someone else at the South Pole does the same experiment at the same time, with that same transformation (of reference frames), the effect of gravity cannot be transformed out.

In a homogeneous gravitational field, you can possibly transform away gravity,  $\Gamma = 0$ . But in an inhomogeneous gravitational field, no.

2. The full wisdom of Newton I. use also the information from Newton's first law that particles (no force) move uniformly introduce the appropriate setting to talk about the difference easily

insight: in spacetime | uniform & straight motion in space | is simply straight motion

So let's try in spacetime:

let  $x: \mathbb{R} \to \mathbb{R}^3$ 

be a particle's trajectory in space  $\longleftrightarrow$  worldline (history) of the particle  $t\mapsto (t,x^1(t),x^2(t),x^3(t)):=$  $:= (X^{0}(t), X^{1}(t), X^{2}(t), X^{3}(t))$ 

That's all it takes:

Trivial rewritings:

$$\dot{X}^{0} = 1$$

$$\Rightarrow \begin{bmatrix} \ddot{X}^0 & = 0 \\ \ddot{X}^\alpha - f^\alpha(X(t)) \cdot \dot{X}^0 \cdot \dot{X}^0 & = 0 \end{bmatrix} \quad (\alpha = 1, 2, 3) \Rightarrow \begin{bmatrix} a = 0, 1, 2, 3 \\ \ddot{X}^a + \Gamma^a_{bc} \dot{X}^b \dot{X}^c = 0 \end{bmatrix}$$
 antoparallel eqn in spacetime

Yes, choosing  $\Gamma^0_{ab} = 0$ 

$$\Gamma^{\alpha}_{\beta\gamma} = 0 = \Gamma^{\alpha}_{0\beta} = \Gamma^{\alpha}_{\beta0}$$

$$\Gamma^{\alpha}_{\beta\gamma} = 0 = \Gamma^{\alpha}_{0\beta} = \Gamma^{\alpha}_{\beta0}$$
only:
$$\Gamma^{\alpha}_{00} \stackrel{!}{=} -f^{\alpha}$$

Question: Is this a coordinate-choice artifact?

No, since  $R^{\alpha}_{0\beta0} = -\frac{\partial}{\partial x^{\beta}} f^{\alpha}$  (only non-vanishing components) (tidal force tensor, – the Hessian of the force component)

Ricci tensor  $\Longrightarrow R_{00} = R_{0m0}^m = -\partial_{\alpha} f^{\alpha} = 4\pi G \rho$ 

Poisson:  $-\partial_{\alpha} f^{\alpha} = 4\pi G \cdot \rho$ 

writing:  $T_{00} = \frac{1}{2}\rho$ 

$$\Longrightarrow R_{00} = 8\pi G T_{00}$$

Einstein in 1912  $R_{ab} \gg 8\pi G T_{ab}$ 

<u>Conclusion</u>: Laplace's idea works in spacetime

Remark

$$\Gamma^{\alpha}_{00} = -f^{\alpha}$$

$$R^{\alpha}_{\beta\gamma\delta} = 0 \qquad \alpha, \beta, \gamma, \delta = 1, 2, 3$$

$$\boxed{R_{00} = 4\pi G\rho}$$

Q: What about transformation behavior of LHS of

$$\underbrace{\ddot{x}^a + \Gamma^a_{bc} \dot{X}^b \dot{X}^c}_{\text{=}a^a \text{"acceleration } \underline{\text{vector}}} = 0$$

### 3. The foundations of the geometric formulation of Newton's axiom. new start

Definition 86. A Newtonian spacetime is a quintuple

$$(M, \mathcal{O}, \mathcal{A}, \nabla, t)$$

where  $(M, \mathcal{O}, \mathcal{A})$  4-dim. smooth manifold

$$t: M \to \mathbb{R}$$
 smooth function

(i) "There is an absolute space"

$$(dt)_p \neq 0 \qquad \forall p \in M$$

(ii) "absolute time flows uniformly"

$$\nabla dt = 0$$
 everywhere space of (0, 2)-tensor fields

 $\nabla dt$  is a (0,2)-tensor field

(iii) add to axioms of Newtonian spacetime  $\nabla = 0$  torsion free

**Definition 87.** absolute space at time  $\tau$ 

$$S_{\tau} := \{ p \in M | t(p) = \tau \}$$

$$\xrightarrow{dt \neq 0} M = \prod S_{\tau}$$

**Definition 88.** A vector  $X \in T_pM$  is called

(a) future-directed if

(b) spatial if

$$dt(X) = 0$$

(c) past-directed if

picture

Newton I: The worldline of a particle under the influence of no force (gravity isn't one, anyway) is a future-directed autoparallel i.e.

$$\nabla_{v_X} v_X = 0$$
$$dt(v_X) > 0$$

and (iii)  $\nabla$  is torsion-free.

Newton II:

$$\nabla_{v_X} v_X = \frac{F}{m} \Longleftrightarrow m \cdot \mathfrak{a} = F$$

where F is a spatial vector field:

$$dt(F) = 0$$

Convention: restrict attention to atlases  $\mathcal{A}_{\text{stratefied}}$  whose charts  $(\mathcal{U}, x)$  have the property

$$x^{0}: \mathcal{U} \to \mathbb{R}$$

$$x^{1}: \mathcal{U} \to \mathbb{R}$$

$$\vdots \qquad \qquad x^{0} = t|_{\mathcal{U}} \qquad \Longrightarrow 0$$

$$0 \text{ "absolute time flows uniformly" } \nabla dt$$

$$0 = \nabla_{\frac{\partial}{\partial x^{a}}} dx^{0} = -\Gamma_{ba}^{0} \qquad a = 0, 1, 2, 3$$

$$x^{3}$$

Let's evaluate in a chart  $(\mathcal{U}, x)$  of a stratified atlas  $\mathcal{A}_{\text{sheet}}$ : Newton II:

$$\nabla_{v_X} v_X = \frac{F}{m}$$

in a chart.

$$(X^{0})'' + \underline{\Gamma^{0}_{cd}(X^{\alpha})'(X^{b})^{\text{fstratified atlas}}} = 0$$

$$(X^{\alpha})'' + \Gamma^{\alpha}_{\gamma\delta}X^{\gamma'}X^{\delta'} + \Gamma^{\alpha}_{00}X^{0'}X^{0'} + 2\Gamma^{\alpha}_{\gamma0}X^{\gamma'}X^{0'} = \frac{F^{\alpha}}{m} \qquad \alpha = 1, 2, 3$$

EY: 20160623: where the factor of 2 comes from torsion free  $\nabla$ 

$$\Longrightarrow (X^0)''(\lambda) = 0 \Longrightarrow X^0(\lambda) = a\lambda + b$$
 constants  $a, b$  with  $X^0(\lambda) = (x^0 \circ X)(\lambda) \stackrel{\text{stratified}}{=} (t \circ X)(\lambda)$ 

convention parametrize worldline by absolute time

$$\frac{d}{d\lambda} = a\frac{d}{dt}$$

$$a^2\ddot{X}^{\alpha} + a^2\Gamma^{\alpha}_{\ \gamma\delta}\dot{X}^{\gamma}\dot{X}^{\delta} + a^2\Gamma^{\alpha}_{\ 00}\dot{X}^{0}\dot{X}^{0} + 2a^2\Gamma^{\alpha}_{\ \gamma0}\dot{X}^{\gamma}\dot{X}^{0} = \frac{F^{\alpha}}{m}$$

$$\Longrightarrow \underbrace{\ddot{X}^{\alpha} + \Gamma^{\alpha}_{\ \gamma\delta}\dot{X}^{\gamma}\dot{X}^{\delta} + \Gamma^{\alpha}_{\ 00}\dot{X}^{0}\dot{X}^{0} + 2\Gamma^{\alpha}_{\ \gamma0}\dot{X}^{\gamma}\dot{X}^{0}}_{a^{\alpha}} = \frac{1}{a^2}\frac{F^{\alpha}}{m}$$

Student question: How can you get the Lagrange equations in this formalism?

Schuller says: In all systems they take the form like this:

$$L = \underbrace{T}_{\frac{m}{2}g(v_x, v_x)} + V$$

where  $v_x$  is velocity to the spatial curves x.

Schuller says:  $\frac{m}{2}\sqrt{g(v_x,v_x)}$ , parameter invariance, you get straight curves; if you take  $\sqrt{\text{away}}$ ,  $\frac{m}{2}g(v_x,v_x)$  you get straight, uniform curves.

Schuller says: Only in inertial systems, which is defined as:

$$\nabla_{\frac{\partial}{\partial x^a}}\frac{\partial}{\partial x^b}=0$$

inertial frame, non-accelerating, non-rotating.

Schuller says: Apart from g, a 2-tensor, you need to know  $\omega$ .

Schuller references 2nd. vear Mechanik in German: https://www.video.uni-erlangen.de/course/id/272

### 25. Lecture 10: Metric Manifolds

### cf. Lecture 10: Metric Manifolds (International Winter School on Gravity and Light 2015)

We establish a structure on a smooth manifold that allows one to assign vectors in each tangent space a length (and an angle between vectors in the same tangent space).

From this structure, one can then define a notion of length of a curve.

Then we can look at shortest curves.

Requiring then that the shortest curves coincide with the straightest curves (wrt  $\nabla$ ) will result in  $\nabla$  being determined by the 25.2. Signature. Linear algebra: metric structure.

$$g \overset{\text{straight=short}}{\overset{T=0}{\leadsto}} \nabla \leadsto \text{Riem}$$

### 25.1. Metrics.

**Definition 89.** A metric q on a smooth manifold  $(M, \mathcal{O}, \mathcal{A})$  is a (0,2)-tensor field satisfying

- (i) symmetry  $g(X,Y) = g(Y,X) \quad \forall X,Y \ vector \ fields$
- (ii) non-degeneracy: the musical map

"flat" 
$$\flat : \Gamma(TM) \to \Gamma(T^*M)$$
  
 $X \mapsto \flat(X)$ 

where 
$$b(X)(Y) := g(X,Y)$$
  
 $b(X) \in \Gamma(T^*M)$   
In thought bubble:  $b(X) = g(X,\cdot)$   
... is a  $C^{\infty}$ -isomorphism in other words, it is invertible.

Remark: 
$$(\flat(X))_a$$
 or  $X_a$ 

$$(\flat(X))_a := g_{am}X^m$$
Thought bubble:  $\flat^{-1} = \sharp$ 

$$\flat^{-1}(\omega)^a := g^{am}\omega_m$$

$$\flat^{-1}(\omega)^a := (g^{"-1"})^{am}\omega_m \Longrightarrow \text{not needed. (all of this is not needed)}$$

**Definition 90.** The (2,0)-tensor field  $g^{"-1"}$  with respect to a metric g is the symmetric

$$g^{"-1"}: \Gamma(T^*M) \times \Gamma(T^*M) \to C^{\infty}(M)$$
$$(\omega, \sigma) \mapsto \omega(\flat^{-1}(\sigma)) \qquad \flat^{-1}(\sigma) \in \Gamma(TM))$$

chart: 
$$g_{ab} = g_{ba}$$

$$(g^{-1})^{am}g_{mb} = \delta_b^a$$
Example:  $(S^2, \mathcal{O}, \mathcal{A})$ 
chart  $(\mathcal{U}, x)$ 

$$\varphi \in (0, 2\pi)$$
 $\theta \in (0, \pi)$ 
define the metric

$$g_{ij}(x^{-1}(\theta,\varphi)) = \begin{bmatrix} R^2 & 0\\ 0 & R^2 \sin^2 \theta \end{bmatrix}_{ij}$$

$$R \in \mathbb{R}^+$$

"the metric of the round sphere of radius R"

(1, 1) tensor has eigenvalues

(0,2) has signature (p,q) (well-defined)

$$(+ + +) 
(+ + -) 
(+ - -) 
(- - -) 
d + 1 if  $p + q = \dim V$$$

**Definition 91.** A metric is called

**Riemannian** if its signature is  $(++\cdots+)$ 

**Lorentzian** if  $(+-\cdots-)$ 

25.3. Length of a curve. Let  $\gamma$  be a smooth curve.

Then we know its veloctiy  $v_{\gamma,\gamma(\lambda)}$  at each  $\gamma(\lambda) \in M$ .

**Definition 92.** On a Riemannian metric manifold  $M, \mathcal{O}, \mathcal{A}, g$ , the **speed** of a curve at  $\gamma(\lambda)$  is the number

 $A^a_m v^m = \lambda v^a$ 

$$(\sqrt{g(v_{\gamma}, v_{\gamma})})_{\gamma(\lambda)} = s(\lambda)$$

F. Schuller: "I feel the need for speed." -Top Gun.

(I feel the need for speed, then I feel the need for a metric)

Aside: 
$$[v^a] = \frac{1}{T}$$
  
 $[g_{ab}] = L^2$   
 $[\sqrt{g_{ab}v^av^b}] = \sqrt{\frac{L^2}{T^2}} = \frac{L}{T}$ 

**Definition 93.** Let  $\gamma:(0,1)\to M$  a smooth curve.

Then the **length** of  $\gamma$  is the number

$$\mathbb{R}\ni L[\gamma]:=\int_0^1 d\lambda s(\lambda)=\int_0^1 d\lambda \sqrt{(g(v_\gamma,v_\gamma))_{\gamma(\lambda)}}$$

F. Schuller: "velocity is more fundamental than speed, speed is more fundamental than length'

Example: reconsider the round sphere of radius R

Consider its equator:

$$\theta(\lambda) := (x^1 \circ \gamma)(\lambda) = \frac{\pi}{2}$$

$$\varphi(\lambda) := (x^2 \circ \gamma)(\lambda) = 2\pi\lambda^3$$

$$\theta'(\lambda) = 0$$

$$\varphi'(\lambda) = 6\pi\lambda^2$$

on the same chart  $g_{ij} = \begin{bmatrix} R^2 & \\ & R^2 \sin^2 \theta \end{bmatrix}$ 

F.Schuller: do everything in this chart

$$L[\gamma] = \int_0^1 d\lambda \sqrt{g_{ij}(x^{-1}(\theta(\lambda), \varphi(\lambda)))(x^i \circ \gamma)'(\lambda)(x^j \circ \gamma)'(\lambda)} = \int_0^1 d\lambda \sqrt{R^2 \cdot 0 + R^2 \sin^2(\theta(\lambda)) 36\pi^2 \lambda^4} =$$

$$= 6\pi R \int_0^1 d\lambda \lambda^2 = 6\pi R [\frac{1}{3}\lambda^3]_0^1 = 2\pi R$$

**Theorem 28.**  $\gamma:(0,1)\to M$  and

 $\sigma:(0,1)\to(0,1)$  smooth bijective and increasing "reparametrization"

$$L[\gamma] = L[\gamma \circ \sigma]$$

 $Proof. \Longrightarrow Tutorials$ 

### 25.4. Geodesics.

**Definition 94.** A curve  $\gamma:(0,1)\to M$  is called a **geodesic** on a Riemannian manifold  $(M,\mathcal{O},\mathcal{A},g)$  if its a stationary curve with respect to a length functional L.

Thought bubble: in classical mechanics, deform the curve a little,  $\epsilon$  times this deformation, to first order, it agrees with  $L[\gamma]$ 

**Theorem 29.**  $\gamma$  geodesic iff it satisfies the Euler-Lagrange equations for the Lagrangian

$$\mathcal{L}:TM \to \mathbb{R}$$
$$X \mapsto \sqrt{g(X,X)}$$

In a chart, the Euler Lagrange equations take the form:

$$\left(\frac{\partial \mathcal{L}}{\partial \dot{x}^m}\right) \cdot - \frac{\partial \mathcal{L}}{\partial x^m} = 0$$

F.Schuller: this is a chart dependent formulation

here:

$$\mathcal{L}(\gamma^i, \dot{\gamma}^i) = \sqrt{g_{ij}(\gamma(\lambda))\dot{\gamma}^i(\lambda)\dot{\gamma}^j(\lambda)}$$

Euler-Lagrange equations:

$$\frac{\partial \mathcal{L}}{\partial \dot{\gamma}^{m}} = \frac{1}{\sqrt{\dots}} g_{mj}(\gamma(\lambda)) \dot{\gamma}^{j}(\lambda) 
\left(\frac{\partial \mathcal{L}}{\partial \dot{\gamma}^{m}}\right)^{\cdot} = \left(\frac{1}{\sqrt{\dots}}\right)^{\cdot} g_{mj}(\gamma(\lambda)) \cdot \dot{\gamma}^{j}(\lambda) + \frac{1}{\sqrt{\dots}} \left(g_{mj}(\gamma(\lambda)) \ddot{\gamma}^{j}(\lambda) + \dot{\gamma}^{s}(\partial_{s} g_{mj}) \dot{\gamma}^{j}(\lambda)\right)$$

Thought bubble: reparametrize  $g(\dot{\gamma}, \dot{\gamma}) = 1$  (it's a condition on my reparametrization) By a clever choice of reparametrization  $\left(\frac{1}{\sqrt{-}}\right)^{\cdot} = 0$ 

$$\frac{\partial \mathcal{L}}{\partial \gamma^m} = \frac{1}{2\sqrt{\dots}} \partial_m g_{ij}(\gamma(\lambda)) \dot{\gamma}^i(\lambda) \dot{\gamma}^j(\lambda)$$

putting this together as Euler-Lagrange equations:

$$g_{mj}\ddot{\gamma}^j + \partial_s g_{mj}\dot{\gamma}^s\dot{\gamma}^j - \frac{1}{2}\partial_m g_{ij}\dot{\gamma}^i\dot{\gamma}^j = 0$$

Multiply on both sides  $(g^{-1})^{qm}$ 

$$\ddot{\gamma}^{q} + (g^{-1})^{qm} (\partial_{i} g_{mj} - \frac{1}{2} \partial_{m} g_{ij}) \dot{\gamma}^{i} \dot{\gamma}^{j} = 0$$

$$\ddot{\gamma}^{q} + (g^{-1})^{qm} \frac{1}{2} (\partial_{i} g_{mj} + \partial_{j} g_{mi} - \partial_{m} g_{ij}) \dot{\gamma}^{i} \dot{\gamma}^{j} = 0$$

geodesic equation for  $\gamma$  in a chart.

$$g^{-1})^{qm} \frac{1}{2} (\partial_i g_{mj} + \partial_j g_{mi} - \partial_m g_{ij}) =: \Gamma^q_{ij} (\gamma(\lambda))$$

Thought bubble:  $\left(\frac{\partial \mathcal{L}}{\partial \xi_x^{a+\dim M}}\right)_{\sigma(x)}^{\cdot} - \left(\frac{\partial \mathcal{L}}{\partial x i_x^a}\right)_{\sigma(x)} = 0$ 

**Definition 95.** "Christoffel symbol" L.C. $\Gamma$  are the connection coefficient functions of the so-called Levi-Civita connection L.C. $\nabla$ 

We usually make this choice of  $\nabla$  if g is given.

$$(M, \mathcal{O}, \mathcal{A}, g) \to (M, \mathcal{O}, \mathcal{A}, g, \overset{\text{L.C.}}{\nabla})$$

$$\xrightarrow{\text{abstract way:}} \nabla g = 0 \text{ and } T = 0 \text{ (torsion)}$$

$$\xrightarrow{} \nabla = \overset{\text{L.C.}}{\nabla}$$

**Definition 96.** (a) The Riemann-Christoffel curvature is defined by

$$R_{abcd} := g_{am} R^m_{bcd}$$

(b) Ricci:  $R_{ab} = R^m_{amb}$ Thought bubble: with a metric, L.C. $\nabla$ 

(c) (Ricci) scalar curvature:

$$R = g^{ab}R_{ab}$$

Thought bubble:  $^{L.C.}\nabla$ 

**Definition 97.** <u>Einstein curvature</u>  $(M, \mathcal{O}, \mathcal{A}, g)$ 

$$G_{ab} := R_{ab} - \frac{1}{2}g_{ab}R$$

Convention:  $g^{ab} := (g^{-1''})^{ab}$ 

F. Schuller: these indices are not being pulled up, because what would you pull them up with (student) Question: Does the Einstein curvature yield new information?

Answer:

$$g^{ab}G_{ab} = R_{ab}g^{ab} - \frac{1}{2}g_{ab}g^{ab}R = R - \delta^a_a R = R - \frac{1}{2}\dim M R = (1 - \frac{d}{2})R$$

Tutorial 9: Metric manifolds. Exercise 3: Levi-Civita Connection. Suppose torsion-free T=0 and metric-compatible connection  $\nabla g=0$ 

Question Recall T = 0 on a chart.

$$\Gamma_{ba}^{c} = \frac{1}{2} (g^{-1})^{cm} \left( \frac{\partial g_{bm}}{\partial x^{a}} + \frac{\partial g_{ma}}{\partial x^{b}} - \frac{\partial g_{ab}}{\partial x^{m}} \right)$$

or

$$\Gamma_{bc}^{a} = \frac{1}{2} (g^{-1})^{am} \left( \frac{\partial g_{bm}}{\partial x^{c}} + \frac{\partial g_{mc}}{\partial x^{b}} - \frac{\partial g_{bc}}{\partial x^{m}} \right)$$

### 26. Symmetry

EY: 20150321 This lecture tremendously and lucidly clarified, for me at least, what a symmetry of the Lie algebra is, and in comparing structures  $(M, \mathcal{O}, \mathcal{A})$  vs.  $(M, \mathcal{O}, \mathcal{A}, \nabla)$ , clarified differences, and asking about differences is a good way to learn, the difference between  $\mathcal{L}$  and  $\nabla$ , respectively.

Feeling that the round sphere

$$(S^2, \mathcal{O}, \mathcal{A}, g^{\text{round}})$$

has rotational symmetry, while the potato

$$(S^2, \mathcal{O}, \mathcal{A}, q^{\text{potato}})$$

does not.

26.1.

26.2. Important

26.3. Flow of a complete vector field. Let  $(M, \mathcal{O}, \mathcal{A})$  smooth X vector field on M

**Definition 98.** A curve  $\gamma: I \subseteq \mathbb{R} \to M$  is called an <u>integral curve of X</u>

$$v_{\gamma,\gamma(\lambda)} = X_{\gamma(\lambda)}$$

**Definition 99.** A vector filed X is complete if all integral curves have  $I = \mathbb{R}$  EY: 20150321 (i.e. domain is all of  $\mathbb{R}$ )

Ex. minute 48:30 EY: reall good explanation by F.P.Schuller; take a pt. out for an incomplete vector field.

**Theorem 30.** compactly supported smooth vector field is complete.

**Definition 100.** The flow of a complete vector field X is a 1-parameter family

$$h^X = \mathbb{R} \times M \to M$$

where  $\gamma_p: \mathbb{R} \to M$  is the integral curve of X with

$$\gamma(0) = p$$

Then for fixed  $\lambda \in \mathbb{R}$ 

$$h_{\lambda}^{X}: M \to M \text{ smooth}$$

picture  $h_{\lambda}^{X}(S) \neq S(\text{ if } X \neq 0)$ 

26.4. Lie subalgebras of the Lie algebra  $(\Gamma(TM), [\cdot, \cdot])$  of vector fields.

(a)  $\Gamma(TM) = \{ \text{ set of all vector fields } \}$   $C^{\infty}(M)$ -module =  $\mathbb{R}$ -vector space

$$\Longrightarrow [X,Y] \in \Gamma(TM)$$
  $[X,Y]f := X(Yf) - Y(Xf)$ 

- (i) [X, Y] = -[Y, X]
- (ii)  $[\lambda X + Z, Y] = \lambda [X, Y] + [Z, Y]$
- (iii) [X, [Y, Z]] + [Z, [X, Y]] + [Y, [Z, X]] = 0 $(\Gamma(TM), [\cdot, \cdot])$  Lie algebra
- (b) Let  $X_1 \dots X_s$  for s (many) vector fields on M, such that

Tutorial 11 Symmetry. Exercise 1.: True or false?

- (a)  $\phi^*: T^*N \to T*M$  i.e.  $\phi^*\nu(X) = \nu(\phi_*X)$  for smooth  $\phi: M \to N$ , so the pullback of a covector  $\nu \in T^*N$  maps to a covector in T\*M.
  - •
- (b)
- (b) (c)

Exercise 2.: Pull-back and push-forward

Question. Let's check this locally

$$\phi^*(df)(X) = (df)(\phi_*X) = (df)(X^i \frac{\partial y^j}{\partial x^i} \frac{\partial}{\partial y^j}) = X^i \frac{\partial y^j}{\partial x^i} \frac{\partial f}{\partial y^j} \text{ where } \phi_*X = X^i \frac{\partial y^j}{\partial x^i} \frac{\partial}{\partial y^j}$$
$$d(\phi^*f)(X) = d(f(\phi))(X) = \frac{\partial f}{\partial y^j} \frac{\partial y^j}{\partial x^i} dx^i(X) = X^i \frac{\partial y^j}{\partial x^i} \frac{\partial f}{\partial y^j}$$

So

$$\boxed{\phi^*(df) = d(\phi^*f)} \qquad \forall \, p \in M, \, \forall \, X \in \mathfrak{X}(M)$$

The big idea is that this is a showing of the **naturality** of the pullback  $\phi^*$  with d, i.e. that this commutes:

$$\Omega^{1}(M) \xleftarrow{\phi^{*}} \Omega^{1}(N)$$

$$d \uparrow \qquad \qquad d \uparrow$$

$$C^{\infty}(M) \xleftarrow{\phi^{*}} C^{\infty}(N)$$

Question

$$(\phi_*)_b^a := (dy^a)(\phi_*(\frac{\partial}{\partial x^b}))$$
Let  $g \in C^{\infty}(N)$ 

$$\phi_*\left(\frac{\partial}{\partial x^b}\right)g = \frac{\partial x^b}{g}\phi(p) = \frac{\partial}{\partial x^b}g\phi x^{-1}x(p) = \frac{\partial}{\partial x^b}(gyy^{-1}\phi x^{-1})(x) =$$

$$= \frac{\partial}{\partial x^b}(gy^{-1}(y\phi x^{-1}(x(p)))) = \frac{\partial g}{\partial y}\Big|_y \frac{\partial y^a}{\partial x^b}\Big|_x = \frac{\partial y^a}{\partial x^b}\frac{\partial g}{\partial y^a}$$

$$\phi_*\left(\frac{\partial}{\partial x^b}\right) = \frac{\partial y^a}{\partial x^b}\frac{\partial}{\partial y^a}$$

$$(\phi_*)_b^a = \frac{\partial y^a}{\partial x^b}$$

Question .

Then

and so

### Exercise 3. :Lie derivative-the pedestrian way

Question. While it is true that  $\forall p \in S^2$ , for  $x(p) = (\theta, \varphi)$ , and  $(yix^{-1})(\theta, \varphi) = (y^1, y^2, y^3) \in \mathbb{R}^3$  and that, at this point

p,  $(y^1)^2/a^2 + (y^2)^2/b^2 + (y^3)^2/c^3 = 1$ , this doesn't imply (EY: 20150321 I think) that, globally, it's an ellipsoid (yet). In the S2.sage familiar charts given,

spherical chart  $(U, x) \in \mathcal{A}$  and

$$(\mathbb{R}^3, y = \mathrm{id}_{\mathbb{R}^3}) \in \mathcal{B}$$

it looks like an ellipsoid, but change to another choice of charts, and it could look something very different.

#### Question .

Equip  $(\mathbb{R}^3, \mathcal{O}_{st}, \mathcal{B})$  with the Euclidean metric q, and pullback q.

Note that the pullback of the inclusion from  $\mathbb{R}^3$  onto  $S^2$  for the Euclidean metric is the following:

$$i^*g\left(\frac{\partial}{\partial\theta^i},\frac{\partial}{\partial\theta^j}\right) = g\left(i_*\frac{\partial}{\partial\theta^i},i_*\frac{\partial}{\partial\theta^j}\right) = g\left(\frac{\partial x^a}{\partial\theta^i}\frac{\partial}{\partial x^a},\frac{\partial x^b}{\partial\theta^j}\frac{\partial}{\partial x^b}\right) = g_{ab}\frac{\partial x^a}{\partial\theta^i}\frac{\partial x^b}{\partial\theta^j}$$

With  $q_{ab} = \delta_{ab}$ , the usual Euclidean metric, this becomes the following:

$$g_{ij}^{\text{ellipsoid}} = \frac{\partial x^a}{\partial \theta^i} \frac{\partial x^a}{\partial \theta^j}$$

At this point, one should get smart (we are in the 21st century) and use some sort of CAS (Computer Algebra System). I like Sage Math (version 6.4 as of 20150322). I also like the Sage Manifolds package for Sage Math.

I like Sage Math for the following reasons:

- Open source, so it's open and freely available to anyone, which fits into my principle of making online education open and freely available to anyone, anytime
- Sage Math structures everything in terms of Category Theory and Categories and Morphisms naturally correspond to Classes and Class methods or functions in Object-Oriented Programming in Python and they've written it that way

and I like Sage Manifolds for roughly the same reasons, as manifolds are fit into a category theory framework that's written into the Python code. e.g.

```
sage: S2 = Manifold(2, 'S^2', r'\mathbb{S}^2', start_index=1) ; print S2
sage: print S2
2-dimensional manifold 'S^2'
sage: type(S2)
<class 'sage.geometry.manifolds.manifold.Manifold_with_category'>
```

With code (I've provided for convenience; you can make your own as I wrote it based upon to example of  $S^2$  on the sagemanifolds documentation website page), load it and do the following:

cf. https://github.com/ernestyalumni/diffgeo-by-sagemnfd/blob/master/S2.sage

```
http://sagemanifolds.obspm.fr/examples.html
 sage: load("S2.sage")
 sage: U_ep = S2.open_subset('U_{ep}')
 sage: eps.<the,phi> = U_ep.chart()
 sage: a = var(\a")
 sage: b = var(\b")
  sage: c = var("c")
 sage: inclus = S2.diff_mapping(R3, {(eps, cart): [ a*cos(phi)*sin(the), b*sin(phi)*sin(the), c*cos(the) ]} , name="inc",latex_name=r'\mathcal{i}')
 sage: inclus.pullback(h).display()
 inc_*(h) = (c^2*sin(the)^2 + (a^2*cos(phi)^2 + b^2*sin(phi)^2)*cos(the)^2) \text{ dthe*dthe } - (a^2 - b^2)*cos(phi)*cos(the)*sin(phi)*sin(the) \text{ dthe*dphi}
 - (a^2 - b^2)*\cos(phi)*\cos(the)*\sin(phi)*\sin(the) \ dphi*dthe + (b^2*\cos(phi)^2 + a^2*\sin(phi)^2)*\sin(the)^2 \ dphi*dphi + (b^2*\cos(phi)^2 + a^2*\sin(phi)^2) \ dphi*dphi + (b^2*\cos(phi)^2 + a^2*\cos(phi)^2) \ 
 sage: inclus.pullback(h)[2,2].expr()
 (b^2*cos(phi)^2 + a^2*sin(phi)^2)*sin(the)^2
```

A new open subset  $U_{\rm ep}$  was declared in  $S^2$ , a new chart  $(U_{\rm ep}, (\theta, \phi))$  was declared, the constants, a, b, c, were declared, and the inclusion map given in the problem

$$y \circ \mathfrak{i} \circ x^{-1} : (\theta, \phi) \mapsto (a \cos \phi \sin \theta, b \sin \phi \sin \theta, c \cos \theta)$$

Then the pullback of the inclusion map  $\rangle$  was done on the Euclidean metric h, defined earlier in the file

. Then one can access the components of this metric and do, for example,

```
simplify_full(),full_simplify(), reduce_trig()
```

on the expression.

In Python, I could easily do this, and give an answer quick in LaTeX:

```
sage: for i in range(1,3):
          for j in range(1,3):
               print inclus.pullback(h)[i,j].expr()
. . . . :
               latex(inclus.pullback(h)[i,j].expr() )
. . . . :
c^2*\sin(the)^2 + (a^2*\cos(phi)^2 + b^2*\sin(phi)^2)*\cos(the)^2
(EY: I'll suppress the LaTeX output but this sage math function gives you LaTeX code)
  and so
```

$$i^*g = c^2 \sin(the)^2 + \left(a^2 \cos(\phi)^2 + b^2 \sin(\phi)^2\right) \cos(the)^2 d\theta \otimes d\theta +$$

$$-2\left(a^2 - b^2\right) \cos(\phi) \cos(the) \sin(\phi) \sin(the) d\theta \otimes d\phi +$$

$$+\left(b^2 \cos(\phi)^2 + a^2 \sin(\phi)^2\right) \sin(the)^2 d\phi \otimes d\phi$$

### Question .

```
sage: polar_vees = eps.frame()
sage: X_1 = - sin(phi) * polar_vees[1] - cot( the ) * cos(phi) * polar_vees[2]
sage: X_2 = \cos(phi) * polar_vees[1] - \cot(the) * \sin(phi) * polar_vees[2]
sage: X_3 = polar_vees[2]
sage: X_2.lie_der(X_1).display()
(cos(the)^2 - 1)/sin(the)^2 d/dphi
sage: X_3.lie_der(X_1).display()
cos(phi) d/dthe - cos(the)*sin(phi)/sin(the) d/dphi
sage: X_3.lie_der(X_2).display()
sin(phi) d/dthe + cos(phi)*cos(the)/sin(the) d/dphi
  Indeed, one can check on a scalar field f_{\text{eps}} \in C^{\infty}(S^2):
sage: f_eps = S2.scalar_field({eps: function('f', the, phi ) }, name='f')
sage: (X_1(X_2(f_{eps})) - X_2(X_1(f_{eps}))).display()
U {ep} --> R
(the, phi) \mid -- \rangle -D[1](f) (the, phi)
sage: X_2.lie_der(X_1) == -X_3
True
sage: X_3.lie_der(X_1) == X_2
sage: X_3.lie_der(X_2) == -X_1
True
```

66

 $\Longrightarrow |[X_i, X_j] = -\epsilon_{ijk} X_k$ 

So  $\operatorname{span}_{\mathbb{R}}\{X_1, X_2, X_3\}$  equipped with [, ] constitute a Lie subalgebra on  $S^2$  (It's closed under [, ]

27. Integration

27.1.

27.2.

### 27.3. Volume forms.

**Definition 101.** On a smooth manifold  $(M, \mathcal{O}, \mathcal{A})$ a (0, dim M)-tensor field  $\Omega$  is called a volume form if

- (a)  $\Omega$  vanishes nowhere (i.e.  $\Omega \neq 0 \ \forall p \in M$ )
- (b) totally antisymmetric

$$\Omega(\ldots,\underbrace{X}_{ith},\ldots,\underbrace{Y}_{jth}\ldots) = -\Omega(\ldots,\underbrace{Y}_{ith},\ldots,\underbrace{X}_{jth}\ldots)$$

In a chart:

$$\Omega_{i_1...i_d} = \Omega_{[i_1...i_d]}$$

Example  $(M, \mathcal{O}, \mathcal{A}, g)$  metric manifold  $\overline{\text{construct}}$  volume form  $\Omega$  from gIn any chart: (U, x)

$$\Omega_{i_1...i_d} := \sqrt{\det(g_{ij}(x))} \epsilon_{i_1...i_d}$$

where Levi-Civita symbol  $\epsilon_{i_1...i_d}$  is <u>defined</u> as  $\epsilon_{123...d} = +1$ 

$$\epsilon_{1...d} = \epsilon_{[i_1...i_d]}$$

*Proof.* (well-defined) Check: What happens under a change of charts

$$\Omega(y)_{i_1...i_d} = \sqrt{\det(g(y)_{ij})} \epsilon_{i_1...i_d} = 
= \sqrt{\det(g_{mn}(x) \frac{\partial x^m}{\partial y^i} \frac{\partial x^n}{\partial y^j})} \frac{\partial y^{m_1}}{\partial x^{i_1}} \dots \frac{\partial y^{m_d}}{\partial x^{i_d}} \epsilon_{[m_1...m_d]} = 
= \sqrt{|\det g_{ij}(x)|} \left| \det\left(\frac{\partial x}{\partial y}\right) \right| \det\left(\frac{\partial y}{\partial x}\right) \epsilon_{i_1...i_d} = \sqrt{\det g_{ij}(x)} \epsilon_{i_1...i_d} \operatorname{sgn}\left(\det\left(\frac{\partial x}{\partial y}\right)\right)$$

EY: 20150323

Consider the following:

$$\begin{split} \Omega(y)(Y_{(1)}\dots Y_{(d)}) &= \Omega(y)_{i_1\dots i_d}Y_{(1)}^{i_1}\dots Y_{(d)}^{i_d} = \\ &= \sqrt{\det(g_{ij}(y))}\epsilon_{i_1\dots i_d}Y_{(1)}^{i_1}\dots Y_{(d)}^{i_d} = \\ &= \sqrt{\det(g_{mn}(x))\frac{\partial x^m}{\partial y^i}\frac{\partial x^n}{\partial y^j}}\epsilon_{i_1\dots i_d}\frac{\partial y^{i_1}}{\partial x^{m_1}}\dots\frac{\partial y^{i_d}}{\partial x^{m_d}}X^{m_1}\dots X^{m_d} = \\ &= \sqrt{\det(g_{mn}(x))\frac{\partial x^m}{\partial y^i}\frac{\partial x^n}{\partial y^j}}\det\left(\frac{\partial y}{\partial x}\right)\epsilon_{m_1\dots m_d}X^{m_1}\dots X^{m_d} = \\ &= \sqrt{\det(g_{mn}(x))}\left|\det\left(\frac{\partial x}{\partial y}\right)\right|\det\left(\frac{\partial y}{\partial x}\right)\epsilon_{m_1\dots m_d}X^{m_1}\dots X^{m_d} = \\ &= \sqrt{\det(g_{mn}(x))}\epsilon_{m_1\dots m_d}\mathrm{sgn}\left(\det\left(\frac{\partial x}{\partial y}\right)\right)X^{m_1}\dots X^{m_d} = \mathrm{sgn}(\det\left(\frac{\partial x}{\partial y}\right))\Omega_{m_1\dots m_d}(x)X^{m_1}\dots X^{m_d} \end{split}$$

If  $\det\left(\frac{\partial y}{\partial x}\right) > 0$ ,

$$\Omega(y)(Y_{(1)}...Y_{(d)}) = \Omega(x)(X_{(1)}...X_{(d)})$$

This works also if Levi-Civita symbol  $\epsilon_{i_1...i_d}$  doesn't change at all under a change of charts. (around 42:43 https: //youtu.be/2XpnbvPy-Zg)

Alright, let's require. restrict the smooth atlas Ato a subatlas ( $\mathcal{A}^{\uparrow}$  still an atlas)

$$\mathcal{A}^{\uparrow} \subseteq \mathcal{A}$$

s.t.  $\forall (U, x), (V, y)$  have chart transition maps  $y \circ x^{-1}$ 

$$x \circ y^{-1}$$

s.t.  $\det\left(\frac{\partial y}{\partial x}\right) > 0$ such  $\mathcal{A}^{\uparrow}$  called an **oriented** atlas

$$(M, \mathcal{O}, \mathcal{A}, g) \Longrightarrow (M, \mathcal{O}, \mathcal{A}^{\uparrow}, g)$$

Note: associated bundles. Note also:  $\det\left(\frac{\partial y^b}{\partial x^a}\right) = \det(\partial_a(y^bx^{-1}))$   $\frac{\partial y^b}{\partial x^a}$  is an endomorphism on vector space V.

 $\varphi:V\to V$ 

 $\det \varphi$  independent of choice of basis

q is a (0,2) tensor field, not endomorphism (not independent of choice of basis)  $\sqrt{|\det(q_{ij}(y))|}$ 

**Definition 102.**  $\Omega$  be a volume form on  $(M, \mathcal{O}, \mathcal{A}^{\uparrow})$  and consider chart (U, x)

 $\square$  Definition 103.  $\omega_{(X)} := \Omega_{i_1...i_d} \epsilon^{i_1...i_d}$  same way  $\epsilon^{12...d} = +1$ 

one can show

$$\omega_{(y)} = \det\left(\frac{\partial x}{\partial y}\right)\omega_{(x)} \qquad scalar \ density$$

### 27.4. Integration on one chart domain U.

### Definition 104.

(122) 
$$\int_{U} f : \stackrel{(U,y)}{=} \int_{y(U)} d^{d}\beta \omega_{(y)}(y^{-1}(\beta)) f_{(y)}(\beta)$$

*Proof.*: Check that it's (well-defined), how it changes under change of charts

$$\int_{U} f : \stackrel{(U,y)}{=} \int_{y(U)} d^{d}\beta \omega_{(y)}(y^{-1}(\beta)) f_{(y)}(\beta) = \underbrace{=}_{(U,y)} \int_{x(U)} \int d^{d}\alpha \left| \det \left( \frac{\partial y}{\partial x} \right) \right| f_{(x)}(\alpha) \omega_{(x)}(x^{-1}(\alpha) \det \left( \frac{\partial x}{\partial y} \right) = \underbrace{=}_{x(U)} d^{d}\alpha \omega_{(x)}(x^{-1}(x)) f_{(x)}(\alpha)$$

On an oriented metric manifold  $(M, \mathcal{O}, \mathcal{A}^{\uparrow}, g)$ 

$$\int_{U} f := \int_{x(U)} d^{d}\alpha \underbrace{\sqrt{\det(g_{ij}(x))(x^{-1}(\alpha))}}_{\sqrt{g}} f_{(x)}(\alpha)$$

### 27.5. Integration on the entire manifold.

### 28. Lecture 13: Relativistic spacetime

Recall, from Lecture 9, the definition of Newtonian spacetime

$$(M,\mathcal{O},\mathcal{A},\nabla,t) \qquad \begin{array}{l} \nabla \text{ torsion free} \\ t \in C^{\infty}(M) \\ dt \neq 0 \\ \nabla dt = 0 \quad \text{ (uniform time)} \end{array}$$

and the definition of relativistic spacetime (before Lecture )

$$\nabla$$
 torsion-free

$$(M, \mathcal{O}, \mathcal{A}^{\uparrow}, \nabla, g, T)$$
  $g$ 

g Lorentzian metric(+---)

T time-orientation

#### 28.1. Time orientation.

**Definition 105.**  $(M, \mathcal{O}, \mathcal{A}^{\uparrow}, g)$  a Lorentzian manifold. Then a time-orientation is given by a vector field T that

- (i) does **not** vanish anywhere
- (ii) g(T,T) > 0

Newtonian vs. relativistic

Newtonian

X was called future-directed if

 $\forall p \in M$ , take half plane, half space of  $T_pM$  also stratified atlas so make planes of constant t straight relativistic

half cone  $\forall p, q \in M$ , half-cone  $\subseteq T_pM$ 

This definition of spacetime

Question

I see how the cone structure arises from the new metric. I don't understand however, how the T, the time orientation, comes in

Answer

$$(M, \mathcal{O}, \mathcal{A}, g) g \stackrel{(}{\leftarrow} +---)$$

requiring q(X,X) > 0, select cones

T chooses which cone

This definition of spacetime has been made to enable the following physical postulates:

- (P1) The worldline  $\gamma$  of a massive particle satisfies
  - (i)  $g_{\gamma(\lambda)}(v_{\gamma,\gamma(lambda)}, v_{\gamma,\gamma(\lambda)}) > 0$
  - (ii)  $g_{\gamma(\lambda)}(T, v_{\gamma,\gamma(\lambda)}) > 0$
- (P2) Worldlines of <u>massless</u> particles satisfy
  - (i)  $g_{\gamma(\lambda)}(v_{\gamma,\gamma(\lambda)},v_{\gamma,\gamma(\lambda)})=0$
  - (ii)  $g_{\gamma(\lambda)}(T, v_{\gamma,\gamma(\lambda)}) > 0$

picture: spacetime:

Answer (to a question) T is a smooth vector field, T determines future vs. past, "general relativity: we have such a time orientation; smoothness makes it less arbitrary than it seems" -FSchuller,

Claim: 9/10 of a metric are determined by the cone

spacetime determined by distribution, only one-tenth error

28.2. Observers.  $(M, \mathcal{O}, \mathcal{A}^{\uparrow}, \nabla, g, T)$ 

**Definition 106.** An observer is a worldline  $\gamma$  with

$$g(v_{\gamma}, v_{\gamma}) > 0$$
$$g(T, v_{\gamma}) > 0$$

together with a choice of basis

$$v_{\gamma,\gamma(\lambda)} \equiv e_0(\lambda), e_1(\lambda), e_2(\lambda), e_3(\lambda)$$

of each  $T_{\gamma(\lambda)}M$  where the observer worldline passes, if  $g(e_a(\lambda), e_b(\lambda)) = \eta_{ab} = \begin{bmatrix} 1 & & & \\ & -1 & & \\ & & -1 & \\ & & & -1 \end{bmatrix}_{ab}$ 

 $\underline{\mathit{precise}}$ :  $\mathit{observer} = \underline{\mathit{smooth}}$   $\mathit{curve}$  in the frame bundle LM  $\mathit{over}$  M

28.2.1. Two physical postulates.

(P3) A clock carried by a specific observer  $(\gamma, e)$  will measure a time

$$\tau := \int_{\lambda_0}^{\lambda_1} d\lambda \sqrt{g_{\gamma(\lambda)}(v_{\gamma,\gamma(\lambda)}, v_{\gamma,\gamma(\lambda)})}$$

between the two "events"

$$\gamma(\lambda_0)$$
 "start the clock"

and

$$\gamma(\lambda_1)$$
 "stop the clock"

Compare with Newtonian spacetime:

$$t(p) = 7$$

Thought bubble: proper time/eigentime  $\tau$ 

$$\begin{split} M &= \mathbb{R}^4 \\ \mathcal{O} &= \mathcal{O}_{\mathrm{st}} \\ &\underbrace{Application/\mathrm{Example.}}_{\mathcal{A} \ni (\mathbb{R}^4, \mathrm{id}_{\mathbb{R}^4})} \\ g &: g_{(x)ij} = \eta_{ij} \quad ; \qquad T^i_{(x)} = (1, 0, 0, 0)^i \\ &\Longrightarrow \Gamma^i_{(x) \ jk} = 0 \text{ everywhere} \end{split}$$

$$\Longrightarrow (M, \mathcal{O}, \mathcal{A}^{\uparrow}, g, T, \nabla)$$
 Riemm = 0

 $\implies$  spacetime is flat

This situation is called special relativity.

Consider two observers:

$$\gamma: (0,1) \to M$$

$$\gamma^{i}_{(x)} = (\lambda, 0, 0, 0)^{i}$$

$$\delta: (0,1) \to M$$

$$\alpha \in (0,1) : \delta^{i}_{(x)} = \begin{cases} (\lambda, \alpha\lambda, 0, 0)^{i} & \lambda \leq \frac{1}{2} \\ (\lambda, (1-\lambda)\alpha, 0, 0)^{i} & \lambda > \frac{1}{2} \end{cases}$$

let's calculate:

$$\tau_{\gamma} := \int_{0}^{1} \sqrt{g_{(x)ij} \dot{\gamma}_{(x)}^{i} \dot{\gamma}_{(x)}^{j}} = \int_{0}^{1} d\lambda 1 = 1$$

$$\tau_{\delta} := \int_{0}^{1/2} d\lambda \sqrt{1 - \alpha^{2}} + \int_{1/2}^{1} \sqrt{1^{2} - (-\alpha)^{2}} = \int_{0}^{1} \sqrt{1 - \alpha^{2}} = \sqrt{1 - \alpha^{2}}$$

Note: piecewise integration

idea.

2 little mirrors

#### (P4) Postulate

Let  $(\gamma, e)$  be an observer, and

 $\delta$  be a massive particle worldline that is parametrized s.t.  $g(v_{\gamma}, v_{\gamma}) = 1$  (for parametrization/normalization convenience) Suppose the observer and the particle *meet* somewhere (in spacetime)

$$\delta(\tau_2) = p = \gamma(\tau_1)$$

This observer measures the 3-velocity (spatial velocity) of this particle as

(123) 
$$v_{\delta} : \epsilon^{\alpha}(v_{\delta,\delta(\tau_2)})e_{\alpha} \qquad \alpha = 1, 2, 3$$

where  $\epsilon^0$ ,  $\boxed{\epsilon^1, \epsilon^2, \epsilon^3}$  is the unique dual basis of  $e_0$ ,  $\boxed{e_1, e_2, e_3}$ 

There might be a major correction to Eq. (123) from the Tutorial 14: Relativistic spacetime, matter, and Gravitation, see the second exercise, Exercise 2, third question:

(124) 
$$v := \frac{\epsilon^{\alpha}(v_{\delta})}{\epsilon^{0}(v_{\delta})} e_{\alpha}$$

Consequence: An observer  $(\gamma, e)$  will extract quantities measurable in his laboratory from objective spacetime quantities always like that.

Ex: F Faraday (0, 2)-tensor of electromagnetism:

$$F(e_a, e_b) = F_{ab} = \begin{bmatrix} 0 & E_1 & E_2 & E_3 \\ -E_1 & 0 & B_3 & -B_2 \\ -E_2 & -B_3 & 0 & B_1 \\ -E_3 & B_2 & -B_1 & 0 \end{bmatrix}$$

observer frame  $e_a, e_b$ 

$$E_{\alpha} := F(e_0, e_{\alpha})$$

 $B^{\gamma} := F(e_{\alpha}, e_{\alpha}) \epsilon^{\alpha\beta\gamma}$  where  $\epsilon^{123} = +1$  totally antisymmetric

28.3. Role of the Lorentz transformations. Lorentz transformations emerge as follows:

Let  $(\gamma, e)$  and  $(\widetilde{\gamma}, \widetilde{e})$  be observers with  $\gamma(\tau_1) = \widetilde{\gamma}(\tau_2)$ 

(for simplicity  $\gamma(0) = \tilde{\gamma}(0)$ 

Now

$$e_0, \dots, e_1$$
 at  $\tau = 0$   
and  $\widetilde{e}_0, \dots, \widetilde{e}_1$  at  $\tau = 0$ 

both bases for the same  $T_{\gamma(0)}M$ 

Thus:  $\tilde{e}_a = \Lambda^b_{\ a} e_b$   $\Lambda \in GL(4)$ Now:

$$\eta_{ab} = g(\widetilde{e}_a, \widetilde{e}_b) = g(\Lambda_a^m e_m, \Lambda_b^n e_n) =$$

$$= \Lambda_a^m \Lambda_b^n \underbrace{g(e_m, e_n)}_{n}$$

i.e.  $\Lambda \in O(1,3)$ 

Result: Lorentz transformations relate the frames of any two observers at the same point.

" $\tilde{x}^{\mu} - \Lambda^{\mu}_{,\nu} x^{\nu}$ " is utter nonsense

Taking the clock postulate (P3) seriously, one better come up with a realistic clock design that supports the postulate. Tutorial video for this lecture, but I saw that the Tutorial sheet number 14 had the relevant topics. Go there.

### 29. Lecture 14: Matter

two types of matter

point matter

field matter

point matter

massive point particle

more of a phenomenological importance

field matter

electromagnetic field

more fundamental from the GR point of view

both classical matter types

29.1. **Point matter.** Our postulates (P1) and (P2) already constrain the possible particle worldlines.

But what is their precise law of motion, possibly in the presence of "forces",

(a) without external forces

$$S_{\text{massive}}[\gamma] := m \int d\lambda \sqrt{g_{\gamma(\lambda)}(v_{\gamma,\gamma(\lambda)}, v_{\gamma,\gamma(\lambda)})}$$

with:

$$g_{\gamma(\lambda)}(T_{\gamma(\lambda)}, v_{\gamma,\gamma(\lambda)}) > 0$$

dynamical law Euler-Lagrange equation

similarly

$$\begin{split} S_{\text{massless}}[\gamma,\mu] &= \int d\lambda \mu g(v_{\gamma,\gamma(\lambda)},v_{\gamma,\gamma(\lambda)}) \\ \delta_{\mu} & g(v_{\gamma,\gamma(\lambda)},v_{\gamma,\gamma(\lambda)}) = 0 \\ \delta_{\gamma} & \text{e.o.m.} \end{split}$$

Reason for describing equations of motion by actions is that composite systems have an action that is the sum of the actions of the parts of that system, possibly including "interaction terms."

Example.

$$S[\gamma] + S[\delta] + S_{\rm int}[\gamma, \delta]$$

### (b) presence of external forces

or rather presence of <u>fields</u> to which a particle "<u>couples</u>"

Example

$$S[\gamma;A] = \int d\lambda m \sqrt{g_{\gamma(\lambda)}(v_{\gamma,\gamma(\lambda)},v_{\gamma,\gamma(\lambda)})} + qA(v_{\gamma,\gamma(\lambda)})$$

where A is a **covector field** on M. A fixed (e.g. the electromagnetic potential)

Consider Euler-Lagrange eqns.  $L_{\text{int}} = qA_{(x)}\dot{\gamma}_{(x)}^m$ 

$$m(\nabla_{v_{\gamma}}v_{\gamma})_{a} + \underbrace{\left(\frac{\partial L_{\mathrm{int}}}{\partial \overset{\cdot}{\gamma_{(x)}^{m}}}\right)}_{*} - \underbrace{\frac{\partial L_{\mathrm{int}}}{\partial \gamma_{(x)}^{m}}}_{*} = 0 \Longrightarrow \underbrace{m(\nabla_{v_{\gamma}}v_{\gamma})^{a} = \underbrace{\frac{-qF_{m}^{a}\dot{\gamma}^{m}}{\int \nabla_{x_{\gamma}}v_{\gamma}^{m}}}_{\text{Lorentz force on a charged particle in an electromagnetic field}}$$

$$\frac{\partial L}{\partial \dot{\gamma}^{a}} = qA_{(x)a}, \qquad \left(\frac{\partial \dot{L}}{\partial \overset{\cdot}{\gamma_{m}}}\right) = q \cdot \frac{\partial}{\partial x^{m}}(A_{(x)m}) \cdot \dot{\gamma}_{(x)}^{m}$$

$$\frac{\partial L}{\partial \gamma^{a}} = q \cdot \frac{\partial}{\partial x^{a}}(A_{(x)m})\dot{\gamma}^{m}$$

$$* = q\left(\frac{\partial A_{a}}{\partial x^{m}} - \frac{\partial A_{m}}{\partial x^{a}}\right)\dot{\gamma}_{(x)}^{m} = q \cdot F_{(x)am}\dot{\gamma}_{(x)}^{m}$$

 $F \leftarrow \text{Faraday}$ 

$$S[\gamma] = \int (m\sqrt{g(v_{\gamma}, v_{\gamma})} + qA(v_{\gamma}))d\lambda$$

#### 29.2. Field matter.

**Definition 107.** Classical (non-quantum) field matter is any tensor field on spacetime where equations of motion derive from an action.

Example:

$$S_{\text{Maxwell}}[A] = \frac{1}{4} \int_{M} d^{4}x \sqrt{-g} F_{ab} F_{cd} g^{ac} g^{bd}$$

A(0,1)-tensor field

= thought cloud: for simplicity one chart covers all of M

- for  $\sqrt{-g} (+---)$ 

$$F_{ab} := 2\partial_{[a}A_{b]} = 2(\nabla_{[a}A)_{b]}$$

Euler-Lagrange equations for fields

$$0 = \frac{\partial \mathcal{L}}{\partial A_m} - \frac{\partial}{\partial x^s} \left( \frac{\partial \mathcal{L}}{\partial \partial_s A_m} \right) + \frac{\partial}{\partial x^s} \frac{\partial}{\partial x^t} \frac{\partial^2 \mathcal{L}}{\partial \partial_t \partial_s A_m}$$

Example ...

$$(\nabla_{\frac{\partial}{\partial x^m}}F)^{ma} = j^a$$

inhomogeneous Maxwell

thought bubble  $j = qv_{\gamma}$ 

$$\partial_{[a}F_{b]}-()$$

homogeneous Maxwell

Other example well-liked by textbooks

$$S_{\text{Klein-Gordon}}[\phi] := \int_{M} d^{4}x \sqrt{-g} [g^{ab}(\partial_{a}\phi)(\partial_{b}\phi) - m^{2}\phi^{2}]$$

 $\phi$  (0,0)-tensor field

29.3. Energy-momentum tensor of matter fields. At some point, we want to write down an <u>action</u> for the metric tensor field itself.

But then, this action  $S_{\text{grav}}[g]$  will be added to any  $S_{\text{matter}}[A, \phi, \dots]$  in order to describe the total system.

$$S_{\text{total}}[g, A] = S_{\text{grav}}[g] + S_{\text{Maxwell}}[A, g]$$

 $\delta A : \Longrightarrow \text{Maxwell's equations}$ 

$$\delta g_{ab} : \boxed{\frac{1}{16\pi G} G^{ab}} + (-2T^{ab}) = 0$$

G Newton's constant

$$G^{ab} = 8\pi G_N T^{ab}$$

**Definition 108.**  $S_{matter}[\Phi, g]$  is a matter action, the **so-called energy-momentum tensor** is

$$T^{ab} := \frac{-2}{\sqrt{-g}} \left( \frac{\partial \mathcal{L}_{matter}}{\partial g_{ab}} - \partial_s \frac{\partial \mathcal{L}_{matter}}{\partial \partial_s g_{ab}} + \dots \right)$$

- of  $\frac{-2}{\sqrt{g}}$  is Schrödinger minus (EY : 20150408 F.Schuller's joke? but wise) choose all sign conventions s.t.

$$T(\epsilon^0, \epsilon^0) > 0$$

Example: For  $S_{\text{Maxwell}}$ :

$$T_{ab} = F_{am}F_{bn}g^{mn} - \frac{1}{4}F_{mn}F^{mn}g_{ab}$$

 $T_{ab} \equiv T_{\text{Maxwell}ab}$ 

$$T(e_0, e_0) = \underline{E}^2 + \underline{B}^2$$
$$T(e_0, e_\alpha) = (E \times B)_\alpha$$

<u>Fact</u>: One often does not specify the fundamental action for some matter, but one is rather satisfied to assume certain properties / forms of

 $T_{ab}$ 

Example Cosmology: (homogeneous & isotropic) perfect fluid

of pressure p and density  $\rho$  modelled by

$$T^{ab} = (\rho + p)u^a u^b - pg^{ab}$$

radiative fluid

$$T_{ ext{Maxwell}}^{ab} g_{ab} = 0$$

$$T_{ab}^{ab} g_{ab} \stackrel{!}{=} 0$$

observe: 
$$T_{p,f.}^{ab}g_{ab} \stackrel{!}{=} 0$$
  
=  $(\rho + p)u^a u^b g_{ab} - p g^{ab}g_{ab}$ 

$$p = \frac{1}{3}\rho$$

Reconvene at 3 pm? (EY: 20150409 I sent a Facebook (FB) message to the International Winter School on Gravity and Light: there was no missing video; it continues on Lecture 15 immediately)

### Tutorial 14: Relativistic Spacetime, Matter and Gravitation. Exercise 2: Lorentz force law.

### Question electromagnetic potential.

### 30. Lecture 15: Einstein gravity

Recall that in Newtonian spacetime, we were able to reformulate the Poisson law  $\Delta \phi = 4\pi G_N \rho$  in terms of the Newtonian spacetime curvature as

$$R_{00} = 4\pi G_N \rho$$

 $R_{00}$  with respect to  $\nabla_{\text{Newton}}$ 

 $G_N$  = Newtonian gravitational constant

This prompted Einstein to postulate < 1915 that the relativistic field equations for the Lorentzian metric q of (relativistic) spacetime

$$R_{ab} = 8\pi G_N T_{ab}$$

However, this equation suffers from a problem

LHS: 
$$(\nabla_a R)^{ab} \neq 0$$

generically

RHS:

$$(\nabla_a T)^{ab} = 0$$

thought bubble: = formulated from an action

Einstein tried to argue this problem away.

Nevertheless, the equations cannot be upheld.

### 30.1. Hilbert. Hilbert was a specialist for variational principles.

To find the appropriate left hand side of the gravitational field equations, Hibert suggested to start from an action

$$S_{\text{Hilbert}}[g] = \int_{M} \sqrt{-g} R_{ab} g^{ab}$$

thought bubble = "simplest action"

aim: varying this w.r.t. metric  $q_{ab}$  will result in some tensor

$$G^{ab} = 0$$

# 30.2. Variation of $S_{\text{Hilbert}}$ .

$$0 \stackrel{!}{=} \underbrace{\delta}_{g_i} S_{\text{Hilbert}}[g] = \int_M \underbrace{\left[\delta\sqrt{-g}g^{ab}R_{ab} + \sqrt{-g}\delta g^{ab}R_{ab} + \sqrt{-g}g^{ab}\delta R_{ab}\right]}_{2}$$
and 
$$1 : \delta\sqrt{-g} = \frac{-(\det g)g^{mn}\delta g_{mn}}{2\sqrt{-g}} = \frac{1}{2}\sqrt{-g}g^{mn}\delta g_{mn}$$

thought bubble

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$$\delta \det(g) = \det(g)g^{mn}\delta g_{mn}$$
  
e.g. from  
 $\det(g) = \exp \operatorname{trln} g$ 

ad 2:  $g^{ab}g_{bc} = \delta^a_c$ 

$$\Longrightarrow (\delta g^{ab})g_{bc} + g^{ab}(\delta g_{bc}) = 0$$
$$\Longrightarrow \delta g^{ab} = -g^{am}g^{bn}\delta g_{mn}$$

ad 3:

$$\Delta R_{ab} = \delta \partial_b \Gamma^m_{am} - \delta \partial_m \Gamma^m_{ab} + \Gamma \Gamma - \Gamma \Gamma =$$

$$= \partial_b \delta \Gamma^m_{am} - \partial_m \delta \Gamma^m_{ab} =$$

$$= \nabla_b (\delta \Gamma)^m_{am} - \nabla_m (\delta \Gamma)^m_{ab}$$

$$\Longrightarrow \sqrt{-q} q^{ab} \delta R_{ab} = \sqrt{-q}$$

"if you formulate the variation properly, you'll see the variation  $\delta$  commute with  $\partial_b$ " EY: 20150408 I think one uses the integration at the bounds, integration by parts trick

 $\Gamma^{i}_{(x)jk} - \Gamma^{i}_{(x)jk}$  are the components of a (1, 2)-tensor.

Notation:  $(\nabla_b A)^i_{\ a} =: A^i_{\ i \cdot b}$ 

$$\Longrightarrow \sqrt{-g}g^{ab}\delta R_{ab}$$

$$\underset{\nabla g=0}{=} \sqrt{-g}(g^{ab}\delta \Gamma^m_{am})_{;b} - \sqrt{-g}(g^{ab}\delta \Gamma^m_{ab})_{;m} = \sqrt{-g}A^b_{;b} - \sqrt{-g}B^m_{,m}$$

Question: Why is the difference of coefficients a tensor?

Answer:

$$\Gamma^{i}_{(y)\ jk} = \frac{\partial y^{i}}{\partial x^{m}} \frac{\partial x^{m}}{\partial y^{j}} \frac{\partial x^{q}}{\partial y^{k}} \Gamma^{m}_{(x)\ ,nq} + \frac{\partial y^{i}}{\partial x^{m}} \frac{\partial^{2} x^{m}}{\partial y^{j} \partial y^{k}}$$

Collecting terms, one obtains

$$0 \stackrel{!}{=} \delta S_{\text{Hilbert}} = \int_{M} \left[ \frac{1}{2} \sqrt{-g} g^{mn} \delta g_{mn} g^{ab} R_{ab} - \sqrt{-g} g^{am} g^{bn} \delta g_{mn} R_{ab} + \underbrace{\left(\sqrt{-g} A^{a}\right)_{,a}}_{\text{surface term}} - \underbrace{\left(\sqrt{-g} B^{b}\right)_{,b}}_{\text{surface term}} \right]$$

$$= \int_{M} \sqrt{-g} \delta \underbrace{g_{mn}}_{\text{substant production}} \left[ \frac{1}{2} g^{mn} R - R^{mn} \right] \Longrightarrow G^{mn} = R^{mn} - \frac{1}{2} g^{mn} R$$

Hence Hilbert, from this "mathematical" argument, concluded that one may take

$$R_{ab} - \frac{1}{2}g_{ab}R = 8\pi G_N T_{ab}$$
 Einstein equations

$$S_{E-H}[g] = \int_{M} \sqrt{-g}R$$

30.3. Solution of the  $\nabla_a T^{ab} = 0$  issue. One can show ( $\rightarrow$  Tutorials) that the Einstein curvature

$$G_{ab} = R_{ab} - \frac{1}{2}g_{ab}R$$

satisfy the so-called <u>contracted</u> differential Bianchi identity

$$(\nabla_a G)^{ab} = 0$$

### 30.4. Variants of the field equations.

(a) a simple rewriting:

$$R_{ab} - \frac{1}{2}g_{ab}R = 8\pi G_N T_{ab} = T_{ab}$$

$$G_N = \frac{1}{8\pi}$$

Contract on both sides  $g^{ab}$ 

$$R_{ab} - \frac{1}{2}g_{ab}R = T_{ab}||g^{ab}$$

$$R - 2R = T := T_{ab}g^{ab}$$

$$\Rightarrow R = -T$$

$$\Rightarrow R_{ab} + \frac{1}{2}g_{ab}T = T_{ab}$$

$$\Leftrightarrow R_{ab} = (T_{ab} - \frac{1}{2}Tg_{ab}) =: \widehat{T}_{ab}$$

$$R_{ab} = \widehat{T}_{ab}$$

(b)

$$S_{E-H}[g] := \int_{M} \sqrt{-g}(R+2\Lambda)$$

thought bubble:  $\Lambda$  cosmological constant

History:

1915:  $\Lambda < 0$  (Einstein) in order to get a non-expanding universe

>1915:  $\Lambda = 0$  Hubble

today  $\Lambda > 0$  to account for an accelerated expansion

 $\Lambda \neq 0$  can be interpreted as a contribution

 $-\frac{1}{2}\Lambda q$  to the energy-momentum

"dark energy"

Question: surface terms scalar?

Answer: for a careful treatment of the surface terms which we discarded, see, e.g. E. Poisson, "A relativist's toolkit"

C.U.P. "excellent book"

Question: What is a constant on a manifold?

Answer:  $\int \sqrt{-g} \Lambda = \Lambda \int \sqrt{-g} 1$ 

[back to dark energy]

[Weinberg, QCD, calculated]

idea: 1 could arise as the vacuum energy of the standard model fields

 $\Lambda_{\rm calculated} = 10^{120} \times \Lambda_{\rm obs}$ 

"worst prediction of physics"

Tutorials: check that

- Schwarzscheld metric (1916)
- FRW metric
- pp-wave metric
- Reisner-Nordstrom

```
⇒ are solutions to Einstein's equations
```

```
in high school m\ddot{x}+m\omega^2x^2=0 x(t)=\cos{(\omega t)} \underline{\mathrm{ET}}\text{: [elementary tutorials]} study motion of particles & observers in Schwarzscheld S.T. \underline{\text{Satellite}}\text{: Marcus C. Werner} Gravitational lensing odd number of pictures Morse theory (EY:20150408 Morse Theory !!!) \underline{\text{Domenico Giulini}} Hamiltonian form Canonical Formulations Key to Quantum Gravity
```

### Tutorial 13 Schwarzschild Spacetime

EY: 20150408 I'm not sure which tutorial follows which lecture at this point.

The tutorial video is excellent itself. Here, I want to encourage the use of CAS to do calculations. There are many out there. Again, I'm partial to the Sage Manifolds package for Sage Math which are both open-source and based on Python. I'll use that here.

### Exercise 1. Geodesics in a Schwarzschild spacetime

### Question Write down the Lagrangian.

```
Load "Schwarzschild.sage" in Sage Math, which will always be available freely here https://github.com/ernestyalumni/diffgeo-by-sagemnfd/blob/master/Schwarzschild.sage:
```

. To get out the coefficients of L of the components of the tangent vectors to the curve, i.e.  $t', r', \theta', \phi'$ , denoted

```
sage: load("Schwarzschild.sage")
4-dimensional manifold 'M'
open subset 'U_sph' of the 4-dimensional manifold 'M'
Levi-Civita connection 'nabla_g' associated with the Lorentzian metric 'g' on the 4-dimensional manifold 'M'
and so on.
Look at the code and I had defined the Lagrangian to be
```

tp,rp,thp,php

```
in my .sage file, do the following:
sage: L.expr().coefficients(tp)[1][0].factor().full_simplify()
(2*G_N*M_0 - r)/r
sage: L.expr().coefficients(rp)[1][0].factor().full_simplify()
-r/(2*G_N*M_0 - r)
sage: L.expr().coefficients(php)[1][0].factor().full_simplify()
r^2
sage: L.expr().coefficients(thp)[1][0].factor().full_simplify()
r^2*sin(th)^2
```

Question There are 4 Euler-Lagrange equations for this Lagrangian. Derive the one with respect to the function  $t(\lambda)!$ .

```
sage: L.expr().diff(t)
0
```

This confirms that  $\frac{\partial L}{\partial t} = 0$ 

For  $\frac{d}{d\lambda} \frac{\partial L}{\partial t'}$ , then one needs to consider this particular workaround for Sage Math (computer technicality). One takes derivatives with respect to declared variables (declared with var) and then substitute in functions that are dependent upon  $\lambda$ , and then take the derivative with respect to the parameter  $\lambda$ . This does that:

sage: L.expr().diff( thp ).factor().subs( r == gamma1 ).subs( thp == gamma3.diff( tau ) ).subs( th == gamma3 ).diff(tau)\ ....: .factor()

2\*(2\*cos(gamma3(tau))\*gamma1(tau)\*D[0](gamma3)(tau)^2 + 2\*sin(gamma3(tau))\*D[0](gamma1)(tau)\*D[0](gamma3)(tau)

+ gamma1(tau)\*sin(gamma3(tau))\*D[0, 0](gamma3)(tau))\*gamma1(tau)\*sin(gamma3(tau))

## Question Show that the Lie derivative of g with respect to the vector fields $K_t := \frac{\partial}{\partial t}$ .

The first line defines the vector field by accessing the frame defined on a chart with spherical coordinates and getting the time vector. The second line is the Lie derivative of g with respect to this vector field.

sage: K\_t = espher[0]
sage: g.lie\_der(K\_t).display() # 0, as desired
0

EY: 20150410 My question is this:  $\forall X \in \Gamma(TM)$  i.e. X is a vector field on M, or, specifically, a section of the tangent bundle, then does

$$\mathcal{L}_X q = 0$$

instantly mean that X is a symmetry for (M, g)?  $\mathcal{L}_X g$  is interpreted geometrically as how g changes along the flow generated by X, and if it equals 0, then g doesn't change.

31.

32.

### 33. Canonical Formulation of GR I

Dynamical and Hailtonian formulation of General Relativity.

Purpose

- (1) formulate and solve initial-value problems
- (2) integrate Einstein's Equations by numerical codes
- (3) characterize degrees of freedom
- (4) characterize isolated systems, associated symmetry groups and conserved quantities, like Energy/Mass, Momenta (linear and angular), Poincaré charges
- (5) starting point for "canonical quantization" program.

How. We will rewrite Einstein's Eq. in form of a constrained Hamiltonian system.

 $\underbrace{R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R}_{G_{\mu\nu}} + \underbrace{\Lambda}_{\text{kosm.}const.} g_{\mu\nu} = \underbrace{k}_{\frac{8\pi G}{c^4}} T_{\mu\nu}$ 

(-+++)

 $T^{\mu\nu} = \begin{pmatrix} W & \frac{1}{c}S^m \\ cg^m & \mathbf{t}^{mn} \end{pmatrix}$ 

W = Energy density (1 component)

 $g^m = \text{Momentum density}, (3 \text{ components})$ 

 $S^m$  = Energy current-density (3 components)

 $\mathbf{t}^{mn} = \text{Momentum current-density (6 components)}$ 

$$T^{\mu\nu} = T^{\nu\mu} \Longrightarrow S^m = c^2 q^m$$

10 independent komp. (components)

Phys. dim. 
$$[T^{\mu\nu}] = \frac{J}{m^3}$$
  
 $[G^{\mu\nu}] = \frac{1}{m^2}$ 

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$$k = \frac{\text{curvature}}{\text{Energy } \cdot \text{ density}}$$

$$[k] = \frac{1}{m^2} / \frac{J}{m^3}, \, ^2k \quad \frac{\text{Curvature}}{\text{mass density}} = \left(\frac{1}{1.5 \, \text{AU}}\right)^2 / \text{ Density of water}$$

$$=\left(\frac{1}{10\,km}\right)^2$$
 / Nuclear density in core of neutron star  $\simeq 5\cdot 10^{17}\,kg/m^3$ 

If "Ein" for Einstein Tensor,  $G_{\mu\nu} = \text{Ein}\left(\frac{\partial}{\partial x^{\mu}}, \frac{\grave{\partial}}{\partial x^{\nu}}\right)$ 

$$\operatorname{Ein}(v, w) = \frac{1}{4} [\operatorname{Ein}(v + w, v + w) - \operatorname{Ein}(v - w, v - w)$$

$$\operatorname{Ein}(w, w) = -g(w, w) \sum_{v} \operatorname{Sec}$$

where  $\perp w$  is take the sum over any triple of mutually perp. 2-planes in  $\perp w$ 

$$\operatorname{Sec}(\operatorname{Span}\{v,w\}) = \frac{\operatorname{Riem}(v,w,v,w)}{[g(v,w)]^2 - g(v,v)g(w,w)}$$

"sectional curvature

Identity:  $\nabla_{\mu}G^{\mu\nu}=0$  (follows from twice-contracted II. Bianchi Identity

$$\sum_{\lambda\mu\nu \text{ cycl}} \nabla_{\lambda} R_{\alpha\beta\mu\nu} = 0 )$$

 $\underbrace{\partial_0 G^{0\nu}}_{\text{contains at most 1st time der.}} + \underbrace{\partial_k G^{k\nu} + \Gamma G + \Gamma G \equiv 0}_{\text{contains at most 2nd. time derivative}}$ 

 $\implies$  4 out of 10 Einstein Eq. do not evolve the fields but rather constrain the initial data. The space-space components (6 Eqns.) are the evolution Eqns.

10 Einstein Eq. - 4 constraints (underdetermined elliptic type)

\ - 6 evolution equations (undetermined hyperbolic type)

34.

35.

36.

### 37. Lecture 22: Black Holes

Only depends on Lectures 1-15, so does lecture on "Wednesday'

Schwarzschild solution also vacuum solution (from tutorial EY: oh no, must do tutorial)

Study the Schwarzschild as a vacuum solution of the Einstein equation:

 $m = G_N M$  where M is the "mass"

$$g = \left(1 - \frac{2m}{r}\right)dt \otimes dt - \frac{1}{1 - \frac{2m}{r}}dr \otimes dr - r^2(d\theta \otimes d\theta + \sin^2\theta d\varphi \otimes d\varphi)$$

in the so-called Schwarzschild coordinates

$$t$$
  $r$   $heta$ 

$$(-\infty,\infty)$$
  $(0,\infty)$   $(0,\pi)$   $(0,2\pi)$ 

What staring at this metric for a while, two questions naturally pose themselves:

(i) What exactly happens r = 2m?

$$\begin{array}{cccc} t & r & \theta & \varphi \\ (-\infty, \infty) & (0, 2m) \cup (2m, \infty) & (0, \pi) & (0, 2\pi) \end{array}$$

(ii) Is there anything (in the real world) beyond  $t \to -\infty$ ?

$$t \to +\infty$$

idea: Map of Linz, blown up

Insight into these two issues is afforded by stopping to stare.

Look at geodesic of g, instead.

### 37.1. Radial null geodesics. null - $g(v_{\gamma}, v_{\gamma}) = 0$

Consider null geodesic in "Schd"

$$S[\gamma] = \int d\lambda \left[ \left( 1 - \frac{2m}{r} \right) \dot{t}^2 - \left( 1 - \frac{2m}{r} \right)^{-1} \dot{r}^2 - r^2 (\dot{\theta}^2 + \sin^2 \theta \dot{\varphi}^2) \right]$$

with  $[\ldots] = 0$ 

and one has, in particular, the t-eqn. of motion:

$$\left(\left(1 - \frac{2m}{r}\right)\dot{t}\right)^{\cdot} = 0$$

 $\Longrightarrow$ 

$$\boxed{\left(1 - \frac{2m}{r}\right)\dot{t} = k} = \text{const.}$$

Consider radial null geodesics

$$\theta \stackrel{!}{=} \text{const.}$$
  $\varphi = \text{const.}$ 

From  $\square$  and  $\square$ 

$$\implies \dot{r}^2 = k^2 \leftrightarrow \dot{r} = \pm k$$

$$\implies r(\lambda) = \pm k \cdot \lambda$$

Hence, we may consider

$$\widetilde{t}(r) := t(\pm k\lambda)$$

Case A:  $\oplus$ 

$$\frac{d\widetilde{t}}{dr} = \frac{\dot{\widetilde{t}}}{\dot{r}} = \frac{k}{\left(1 - \frac{2m}{r}\right)k} = \frac{r}{r - 2m}$$

$$\Longrightarrow \widetilde{t}_+(r) = r + 2m \ln|r - 2m|$$

(outgoing null geodesics)

<u>Case b.</u>  $\pm$  (Circle around -, consider -):

$$\widetilde{t}_{-}(r) = -r - 2m \ln|r - 2m|$$

(ingoing null geodesics)

Picture

37.2. Eddington-Finkelstein. Brilliantly simple idea:

change (on the domain of the Schwarzschild coordinates) to different coordinates, s.t.

in those new coordinates,

ingoing null geodesics appear as straight lines, of slope -1

This is achieved by

$$\bar{t}(t,r,\theta,\varphi) := t + 2m \ln|r - 2m|$$

Recall: ingoing null geodesic has

$$\widetilde{t}(r) = -(r + 2m \ln|r - 2m|)$$
 (Schdcoords)

$$\iff \bar{t} - 2m \ln|r - 2m| = -r - 2m \ln|r - 2m| + \text{ const.}$$
$$\dot{t} = -r + \text{ const.}$$

(Picture)

outgoing null geodesics

$$\bar{t} = r + 4m \ln |r - 2m| + \text{const.}$$

Consider the new chart (V, g) while (U, x) was the Schd chart.

$$\underbrace{U}_{\text{Sold}} \bigcup \{ \text{ horizon } \} = V$$

"chart image of the horizon"

Now calculate the  $Schd\ metric\ g\ w.r.t.$  Eddington-Finkelstein coords.

$$\bar{t}(t, r, \theta, \varphi) = t + 2m \ln |r - 2m|$$
$$\bar{r}(t, r, \theta, \varphi) = r$$
$$\bar{\theta}(t, r, \theta, \varphi) = \theta$$
$$\bar{\varphi}(t, r, \theta, \varphi) = \varphi$$

EY: 20150422 I would suggest that after seeing this, one would calculate the metric by your favorite CAS. I like the Sage Manifolds package for Sage Math.

Schwarzschild\_BH.sage on github

Schwarzschild\_BH.sage on Patreon

Schwarzschild\_BH.sage on Google Drive

```
sage: load(''Schwarzschild_BH.sage'')
4-dimensional manifold 'M'
expr = expr.simplify_radical()
Levi-Civita connection 'nabla_g' associated with the Lorentzian metric 'g' on the 4-dimensional manifold 'M'
Launched png viewer for Graphics object consisting of 4 graphics primitives
```

Then calculate the Schwarzschild metric g but in Eddington-Finkelstein coordinates. Keep in mind to calculate the set of coordinates that uses  $\bar{t}$ , not  $\tilde{t}$ :

```
sage: gI.display()
gI = (2*m - r)/r dt*dt - r/(2*m - r) dr*dr + r^2 dth*dth + r^2*sin(th)^2 dph*dph
sage: gI.display( X_EF_I_null.frame())
gI = (2*m - r)/r dtbar*dtbar + 2*m/r dtbar*dr + 2*m/r dr*dtbar + (2*m + r)/r dr*dr + r^2 dth*dth + r^2*sin(th)^2 dph*dph
```

#### References

- [1] William Feller. An Introduction to Probability Theory and Its Applications, Vol. 1, 3rd Edition. 1968. ISBN-13: 978-0471257080
- [2] Jeffrey M. Lee. Manifolds and Differential Geometry, Graduate Studies in Mathematics Volume: 107, American Mathematical Society, 2009. ISBN-13: 978-0-8218-4815-9
- [3] John Lee, Introduction to Smooth Manifolds (Graduate Texts in Mathematics, Vol. 218), 2nd edition, Springer, 2012, ISBN-13: 978-1441999818
- [4] Victor Guillemin, Alan Pollack. Differential Topology. American Mathematical Society. 2010. ISBN-13: 978-0821851937 https://www.google.com/url?sa=t&rct=j&q=&esrc=s&source=web&cd=2&cad=rja&uact=8&ved=0ahUKEwjG96q9z63JAhWMLYgKHempDoMQFggmMAE&url=http%3A%2F%2Fwww.mat.unimi.it%2Fusers%2Fdedo%2Ftop%2520diff%2FGuillemin-Pollack\_Differential%2520topology.pdf&usg=AFQjCNF5im0H5xeXRSK60qzM7zT97sdIsw
- [5] Anant R. Shastri. Elements of Differential Topology. CRC Press. 2011. ISBN-13: 978-0415339209
- [6] Morris W. Hirsch, Differential Topology (Graduate Texts in Mathematics), Graduate Texts in Mathematics (Book 33), Springer (September 16, 1997). ISBN-13: 978-0387901480
- [7] Mikio Nakahara. Geometry, Topology and Physics. (Graduate Student Series in Physics) 2nd Edition. 2003. Taylor and Francis Group.
- [8] P.-A. Absil, R. Mahony, and R. Sepulchre. Optimization Algorithms on Matrix Manifolds. Princeton University Press. 2008. ISBN 978-0-691-13298-3 https://press.princeton.edu/titles/8586.html
- [9] John Baez, Javier P Muniain. Gauge Fields, Knots and Gravity (Series on Knots and Everything), World Scientific Publishing Company (October 24, 1994), ISBN-13: 978-9810220341
- [10] Joseph J. Rotman, Advanced Modern Algebra (Graduate Studies in Mathematics), (Book 114), American Mathematical Society; 2 edition, (August 10, 2010), ISBN-13: 978-0821847411
- [11] Andrew Baker, "Representations of Finite Groups", 07112011 http://www.maths.gla.ac.uk/~ajb/dvi-ps/groupreps.pdf
- [12] Yvette Kosmann-Schwarzbach, Groups and Symmetries: From Finite Groups, Springer, 2010. e-ISBN 978-0-387-78866-1 http://www.caam.rice.edu/~yad1/miscellaneous/References/Math/Groups/Groups%20and%20Symmetries.pdf
- [13] Agostino Prástaro. Geometry of PDEs and Mechanics. World Scientific Publishing Co. 1996. QC125.2.P73 1996 530.1'55353-dc20.
- Agostino Prástaro. Dipartimento di Metodi e Modelli Matematici per le Scienze Applicate, Università degli Studi di Roma "La Sapienza" [14] Clemens Koppensteiner. Complex Manifolds: Lecture Notes. http://www.caramdir.at/uploads/math/piii-cm/complex-manifolds.pdf
- [15] Stefan Vandoren. Lectures on Riemannian Geometry, Part II: Complex Manifolds. http://www.staff.science.uu.nl/~vando101/MRIlectures.pdf
- [16] Kentaro Hori (Author, Editor), Sheldon Katz (Editor), Albrecht Klemm (Editor), Rahul Pandharipande (Editor), Richard Thomas (Editor), Ravi Vakil (Editor), Eric Zaslow (Editor), Mirror Symmetry (Clay Mathematics Monographs, V. 1). Clay Mathematics Monographs (Book 1). American Mathematical Society (August 19, 2003). ISBN-10: 0821829556 ISBN-13: 978-0821829554 https://web.archive.org/web/20060919020706/http://math.stanford.edu/~vakil/files/mirrorfinal.pdf
- [17] Lawrence Conlon. Differentiable Manifolds (Modern Birkhäuser Classics). 2nd Edition. Birkhäuser; 2nd edition (October 10, 2008). ISBN-13: 978-0817647667
- [18] Andrew Clarke and Bianca Santoro. Holonomy Groups in Riemannian Geometry. arXiv:1206.3170 [math.DG]
- [19] Urs Schreiber and Konrad Waldorf. Parallel Transport and Functors. arXiv:0705.0452 [math.DG]
- [20] Y. Eliashberg, N. Mishachev, Introduction to the h-principle, Grad. Studies in Math. Vol 48 (AMS 2002)
- [21] Darryl D Holm. Geometric Mechanics Part II: Rotating, Translating and Rolling. 2nd. Ed. Imperial College Publishing. (October 31, 2011)
- [22] Tsutomu Kambe, Geometrical Theory of Dynamical Systems and Fluid Flows, Advanced Series in Nonlinear Dynamics: Volume 23, 2009, http://www.worldscientific.com/worldscibooks/10.1142/7418 ISBN: 978-981-4282-24-6 (hardcover)
- [23] G. J. Pert. Introductory Fluid Mechanics for Physicists and Mathematicians. 1st Edition. Wiley; 1 edition (May 28, 2013). ISBN-10: 1119944856. ISBN-13: 978-1119944850.
- [24] Rolf H. Sabersky, Allen J. Acosta, Edward G. Hauptmann, E.M. Gates. Fluid Flow: A First Course in Fluid Mechanics (4th Edition). Prentice Hall. (August 22, 1998). ISBN-13: 978-0135763728
- [25] Walter Greiner. Classical Mechanics: Point Particles and Relativity (Classical Theoretical Physics). Springer New York; 2004 edition (December 4, 2003). ISBN-13: 978-0387955865
- [26] Vladimir I. Arnold, Valery V. Kozlov, Anatoly I. Neishtadt. Mathematical Aspects of Classical and Celestial Mechanics. Third Edition. Springer. 2006.
- [27] Ralph Abraham and Jerrold E. Marsden. Foundations of Mechanics (Ams Chelsea Publishing) Hardcover May 21, 2008. American Mathematical Society; 2 edition (May 21, 2008). ISBN-10: 0821844385 ISBN-13: 978-0821844380
- [28] Frederic Schuller, Lectures on the Geometric Anatomy of Theoretical Physics, Sept. 2015.
- [29] Yvonne Choquet-Bruhat. General Relativity and the Einstein Equations (Oxford Mathematical Monographs) 1st Edition. Oxford Mathematical Monographs (February 4, 2009) ISBN-13: 978-0199230723
- [30] T. Frankel, The Geometry of Physics, Cambridge University Press, Third Edition, 2012.
- [31] Shigeyuki Morita. Geometry of Differential Forms (Translations of Mathematical Monographs, Vol. 201). American Mathematical Society (August 28, 2001). ISBN-10: 0821810456
- [32] Tom M. Apostol. Calculus, Vol. 1: One-Variable Calculus, with an Introduction to Linear Algebra. Wiley; 2nd edition (January 16, 1991).ISBN-13: 978-0471000051
- [33] Roger R. Bate, Donald D. Mueller, Jerry E. White. Fundamentals of Astrodynamics (Dover Books on Aeronautical Engineering). Publisher: Dover Publications; Revised ed. edition (June 1, 1971). ISBN-13: 978-0486600611
- [34] The Committee Of Japan Physics Olympiad Japan. Physics Olympiad Basic To Advanced Exercises. Wspc (January 15, 2014).