

(Note that the handwritten solutions for problems 2,3,4 are numbered 3,4,5 (from last year))

1) Starting with

$$B_\lambda = \frac{2hc^2\lambda^{-5}}{\exp\left(\frac{hc}{k_b\lambda T}\right) - 1} \quad (1)$$

derive the equivalent relationship for frequency $\hat{\nu}$:

$$B_{\hat{\nu}}(T) = \frac{2h\hat{\nu}^3}{c^2} \frac{1}{e^{\frac{h\hat{\nu}}{kT}} - 1} \quad (2)$$

and show that it has the correct unit of $\text{W m}^{-2} \text{sr}^{-1} \text{s}$

solution: From the notes we know that

$$\int_{\lambda_1}^{\lambda_2} B_\lambda d\lambda = \int_{\hat{\nu}_2}^{\hat{\nu}_1} B_{\hat{\nu}} d\hat{\nu} = - \int_{\hat{\nu}_1}^{\hat{\nu}_2} B_{\hat{\nu}} d\hat{\nu} \quad (3)$$

where $\lambda = c\hat{\nu}^{-1}$.

By the rules for substitution of variables we have:

$$B_\lambda \frac{d\lambda}{d\hat{\nu}} d\hat{\nu} = B_{\hat{\nu}} d\hat{\nu} \quad (4)$$

where

$$\frac{d\lambda}{d\hat{\nu}} = -\frac{c}{\hat{\nu}^2} = -c\hat{\nu}^{-2} \quad (5)$$

So

$$B_{\hat{\nu}} = -B_\lambda \frac{d\lambda}{d\hat{\nu}} \quad (6)$$

with the limits of integration taken from large $\hat{\nu}_1$ (small λ_1) to smaller $\hat{\nu}_2$ (larger λ_2).

Replace λ by the equivalent $c\hat{\nu}^{-1}$ in (1):

$$\frac{2hc^2c^{-5}\hat{\nu}^5}{\exp\left(\frac{h\hat{\nu}}{k_b}\right) - 1} \quad (7)$$

multiply (8) by (5)

$$-B_{\hat{\nu}} = -c\hat{\nu}^{-2} \frac{2hc^2c^{-5}\hat{\nu}^5}{\exp\left(\frac{h\hat{\nu}}{k_b}\right) - 1} = -\frac{2hc^{-2}\hat{\nu}^3}{\exp\left(\frac{h\hat{\nu}}{k_b}\right) - 1} \quad (8)$$

Units: $(\text{Js})(\text{ms}^{-1})^{-2}(\text{s}^{-1})^3 = \text{Jm}^{-2}\text{s}^{-1} = \text{Wm}^{-2} = \text{W m}^{-2} \text{sr}^{-1}/\text{s}^{-1}$

$$\Rightarrow \int_0^\infty B_\lambda d\lambda = \int_0^\infty \frac{2hc^2}{\lambda^5 \exp\left(\frac{hc}{k_b \lambda T}\right) - 1} d\lambda$$

Only λ varies, T is constant.

Take the hint and write

$$x = \frac{hc}{k_b \lambda T}, \quad dx = -\frac{hc}{k_b \lambda^2 T} d\lambda$$

so $dx = -x \frac{d\lambda}{\lambda}$ and $\frac{1}{\lambda} = \frac{x k_b T}{hc}$

$$d\lambda = -\frac{hc}{k_b T} x^{-2} dx$$

Use this to rewrite the integration in terms of x : for the limits note that as $\lambda \rightarrow \infty$, $x \rightarrow 0$.

$$\int_0^\infty B_\lambda d\lambda = - \int_\infty^0 \frac{2hc^2 \left(\frac{hc}{k_b \lambda T}\right)^{-5} \left(\frac{hc}{k_b T} x^{-2}\right) dx}{\exp(x) - 1}$$

Rearrange, flipping the limits of integration

$$\begin{aligned} \int_0^\infty B_\lambda d\lambda &= 2hc^2 \left(\frac{hc}{k_b T}\right)^{-5} \left(\frac{hc}{k_b T}\right) \int_0^\infty \frac{x^5 x^{-2}}{\exp(x) - 1} dx \\ &= 2hc^2 \left(\frac{hc}{k_b}\right)^{-4} T^4 \int_0^\infty \frac{x^3}{\exp(x) - 1} dx \propto T^4 \end{aligned}$$

Also: $\int_0^\infty \frac{x^3}{\exp(x) - 1} dx = \frac{\pi^4}{15}$

$$\begin{aligned}
 \text{So } \int_0^{\infty} B_{\lambda} d\lambda &= \frac{2hc^2 k_b^4}{h^4 c^4} \frac{\pi^4}{15} T^4 \quad \text{p5 1 p3 2/2} \\
 &= \frac{2\pi^4 k_b^4}{15 h^3 c^2} T^4 = \frac{5.67 \times 10^{-8}}{\pi} T^4
 \end{aligned}$$

as expected.

$$4) B_{\lambda} = \frac{2hc^2 \lambda^{-5}}{\exp\left(\frac{hc}{k_b \lambda T}\right) - 1}$$

ps1 p4 and ps2 p5

If $x \ll 1$, $\exp(x) \sim 1+x$

$$\text{So } B_{\lambda} \sim \frac{2hc^2 \lambda^{-5}}{1 + \frac{hc}{k_b \lambda T} - 1} = 2k_b \lambda T c \lambda^{-5} = \frac{2ck_b T}{\lambda^4}$$

with $\lambda = 0.01 \text{ m} = 10000 \mu\text{m}$

and $T = 300 \text{ K}$, $\frac{hc}{k_b \lambda T} \sim 0.0048 \ll 1$

Spreadsheet says: exact 2.477×10^{-10}

$\text{W m}^{-2} \mu\text{m}^{-1} \text{sr}^{-1}$

Rayleigh-Jeans =

2.4834×10^{-10}

$\text{W m}^{-2} \mu\text{m}^{-1} \text{sr}^{-1}$

⇒ Take the hint to disregard the -1 term:

$$B_{\lambda} = \frac{2hc^2}{\lambda^5 \exp\left(\frac{hc}{k_b \lambda T}\right)} = 2hc^2 \lambda^{-5} \exp\left(-\frac{hc}{k_b \lambda T}\right)$$

As the text says, we want to set $\frac{dB_{\lambda}}{d\lambda} = 0$

and solve for λ . Use the chain

$$\text{rule: } \frac{dB_{\lambda}}{d\lambda} = \frac{2hc^2}{\lambda^6 \exp\left(\frac{hc}{k_b \lambda T}\right)} (-5) + 2hc^2 \lambda^{-5} \left(\frac{hc}{k_b \lambda^2 T}\right) \exp(\dots)$$

ϵ cont.

p51, p5 and p6

set this to zero:

$$\frac{dB_\lambda}{d\lambda} = \frac{2hc^2}{\lambda^6 \exp\left(\frac{hc}{k_B \lambda T}\right)} \left(\frac{hc}{k_B \lambda T} - 5\right) = 0$$

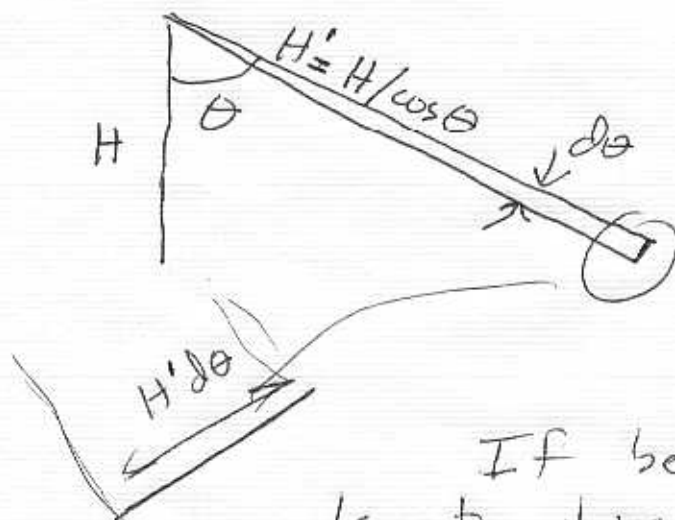
$$\text{so } \frac{hc}{k_B \lambda_{\max} T} = 5 = 0 \quad \text{and}$$

$$\lambda_{\max} = \frac{hc}{k_B 5T} = \frac{2.877 \times 10^{-3} \text{ (metres)}}{T}$$

$$b) \text{ i) } \delta\omega = \frac{\text{area}}{r^2} = \frac{1 \text{ km}^2}{(830 \text{ km})^2} = 1.45 \times 10^{-6} \text{ sr}$$

ii) Changing the "look angle" doesn't change $\delta\omega$, because it is a property of the telescope.

iii) If the pixel is square, then as θ increases, the distance from ground to satellite increases



This affects both the width and the ~~height~~ length

If before, the length was $H d\theta = 1 \text{ km}$