

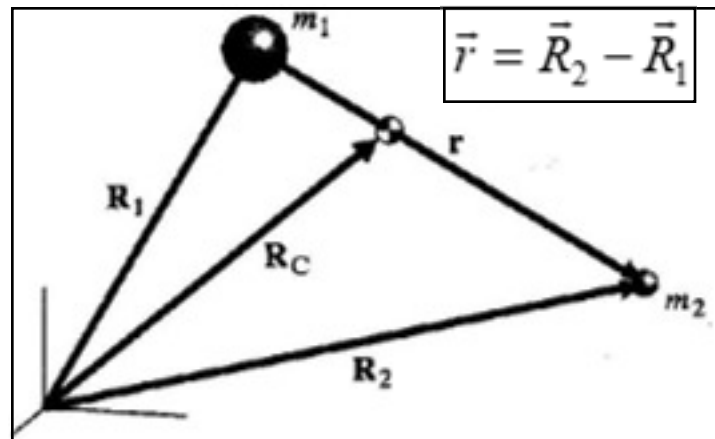


## *Two-body Motion Scenario*



## *Basics of 2-Body Formulation*

Consider the **two-body** system given below.



As **centre of mass** has no **acceleration**, it can be treated as the **inertial** frame origin and all position and velocity **vectors** can be defined as **relative** quantities, as above.



## *Basics of 2-Body Formulation*

**Basic equations for a 2-body system** are as given below.

$$m_1 \ddot{\vec{R}}_1 = - \frac{Gm_1 m_2 (\vec{R}_2 - \vec{R}_1)}{|\vec{R}_2 - \vec{R}_1|^3}, \quad m_2 \ddot{\vec{R}}_2 = + \frac{Gm_1 m_2 (\vec{R}_2 - \vec{R}_1)}{|\vec{R}_2 - \vec{R}_1|^3}$$

$$\vec{R}_c = \frac{m_1 \vec{R}_1 + m_2 \vec{R}_2}{M}; \quad \vec{r} = \vec{R}_2 - \vec{R}_1; \quad M = m_1 + m_2$$

$$\vec{R}_1 = \vec{R}_c + \frac{m_2}{M} \vec{r}; \quad \vec{R}_2 = \vec{r} + \vec{R}_1 = \vec{R}_c + \left(1 - \frac{m_2}{M}\right) \vec{r}$$

$$m_1 \vec{R}_1 \times \dot{\vec{R}}_1 + m_2 \vec{R}_2 \times \dot{\vec{R}}_2 = \vec{H}_0, \quad \frac{1}{2} m_1 \dot{\vec{R}}_1^2 + \frac{1}{2} m_2 \dot{\vec{R}}_2^2 - \frac{Gm_1 m_2}{|\vec{r}|} = E_0$$

This system has **12 unknowns**.



## *2 – Body Simplification*

**2-body** system can be solved if we **drop** 6 variables. An **acceptable** approximation is to assume that **one** body is much **larger**, leading to following **simplification**.

$$M = m_1 + m_2 \approx m_1; \quad \left( \frac{m_2}{m_1} \right) \approx 0$$
$$\bar{R}_c = \frac{m_1 \bar{R}_1 + m_2 \bar{R}_2}{m_1 + m_2} = \frac{\bar{R}_1 + \left( \frac{m_2}{m_1} \right) \bar{R}_2}{1 + \left( \frac{m_2}{m_1} \right)} \approx \bar{R}_1$$

Further, if **origin** is at the **centre of mass**, then we get  **$\mathbf{R}_1 = \mathbf{0}$** , and  **$(\mathbf{R}_2 - \mathbf{R}_1) = \mathbf{r}$** .



## *1 – Body Problem Definition*

**2-body** simplification leads to the **case** that centre of mass of the **system** is at the **centre** of larger body.

Without loss of **generality** we can assume **larger** body to be **stationary**, so that its **centre** is a point fixed in **space**.



## *1 – Body Problem Definition*

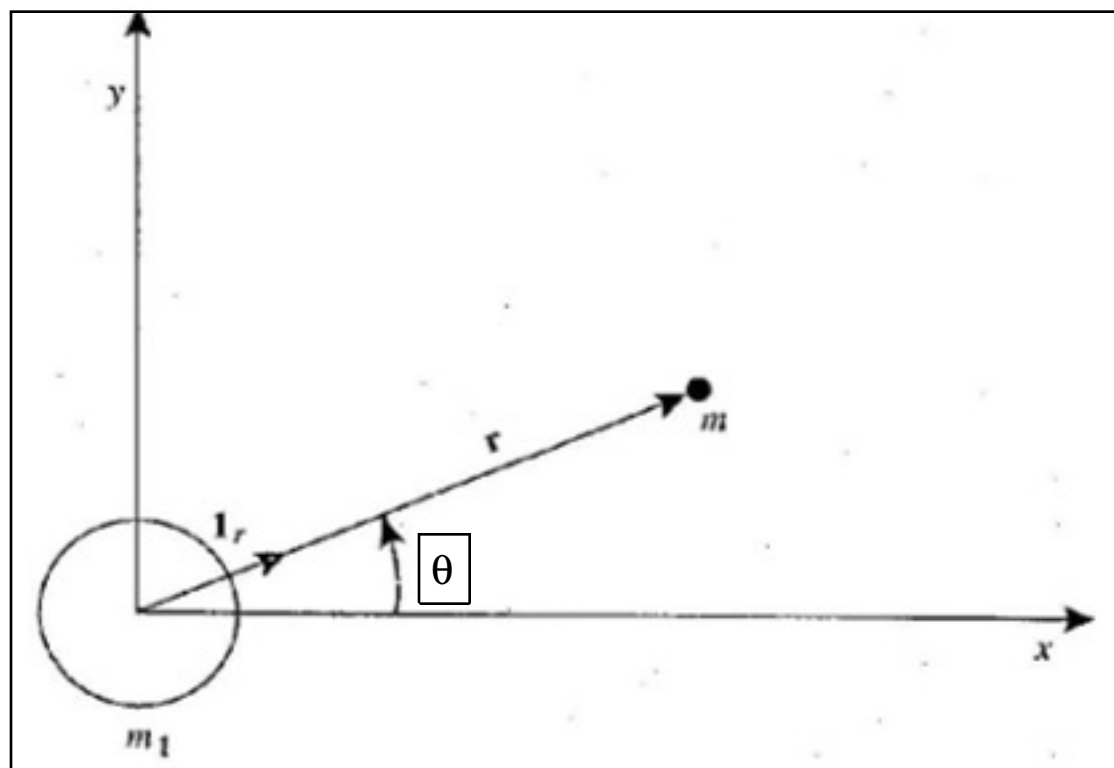
This is the case of **attractive** central force, where **smaller** body moves around a **fixed** point such that **force** is directed **along** the line joining centre and **fixed point**.

**This** results in what is commonly called **1-body** problem or **central force** motion.



## *Central Force Motion Formulation*

Consider the **central** force system, as shown below.





## *Central Force Formulation*

Assuming that  $(\mathbf{m}_1 + \mathbf{m}) \approx \mathbf{m}_1$ , we get the following.

$$m\ddot{\vec{r}} = -G \frac{m_1 m}{\vec{r}^3} \vec{r}; \quad \ddot{\vec{r}} = -G m_1 \frac{\vec{r}}{\vec{r}^3} \rightarrow \text{Acceration of small body}$$

Further, we drop equation involving acceleration of large body,

$$\text{which is small, as follws. } m_1 \ddot{\vec{r}} = G \frac{m m_1}{\vec{r}^3} \vec{r} \rightarrow \ddot{\vec{r}} = G m \frac{\vec{r}}{\vec{r}^3} \approx 0$$

Thus, we now have one vector equation in one variable,  $\vec{r}$  for a spacecraft moving in Earth's gravitational field.

$$\ddot{\vec{r}} + \frac{\mu}{r^3} \vec{r} = 0; \quad \mu = G m_1; \quad G = 6.670 \times 10^{-11} \text{ Nm}^2 / \text{kg}^2$$

$$M_{Earth} = 5.977 \times 10^{24} \text{ kg}; \quad \mu_E = G M_E = 3.986 \times 10^{14} \text{ m}^3 / \text{s}^2$$





## *Central Force Model Features*

We see that we have **one** 2<sup>nd</sup> order vector differential **equation** in one variable '**r**', so that we have **6 unknowns** and **6 initial conditions** on the smaller body.

These **6 initial conditions** are nothing but the **solution** at the end of the **ascent mission**, in terms of position and velocity **vectors** of the spacecraft.



## *Central Force Solution Basics*

There are **many** ways in which the **central** force motion equation **can** be solved.

**One** way is to obtain the **closed** form time **solution** by integrating the corresponding **scalar** differential equations in  $(x, y, z)$ , for specified **initial** conditions.



## *Central Force Solution Basics*

**Another** way is to **numerically** integrate the **scalar** differential equations to **generate** time histories.

We can also use **vector** & scalar **products** (see earlier solutions) to **integrate** equations in terms of momentum & **energy**.



## *Summary*

**Two-body** simplification is an **important** step in understanding the **nature** of motion any spacecraft.

It is to be **noted** that while, the proposed **model** is not exact, the **error** for most cases is quite **negligible**.