

Outline

- Free Theory in Curved Spacetime
- The Graviton
- Gauge Connection
- Topological Gravity

FT in curved spacetime

Free source \rightarrow Curved source

$$n_{\mu\nu} \rightarrow g_{\mu\nu}(x)$$

invariant source time interval
becomes:

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu$$

using, $x \rightarrow x'(x)$

$$ds^2 = g'_{\alpha\beta} dx'^\alpha dx'^\beta$$

$$= g'_{\alpha\beta} \frac{\partial x'^\alpha}{\partial x^\mu} \frac{\partial x'^\beta}{\partial x^\nu} dx^\mu dx^\nu$$

$$\Rightarrow g_{\mu\nu} = g'_{\alpha\beta} \frac{\partial x'^\alpha}{\partial x^\mu} \frac{\partial x'^\beta}{\partial x^\nu}$$

Lagrangian is Lorentz invariant
but action integral transforms by

$$d^4x \rightarrow d^4x' = d^4x \det \left(\frac{\partial x'}{\partial x} \right)$$

$$\begin{aligned} \text{since } g &= \text{det} g_{\mu\nu} = \det \left(g'_{\lambda 0} \frac{\partial x'^\lambda}{\partial x^m} \frac{\partial x'^0}{\partial x^\nu} \right) \\ &= g' \left[\det \left(\frac{\partial x'}{\partial x} \right) \right]^2 \end{aligned}$$

$$\begin{aligned} \Rightarrow d^4x \sqrt{-g} &= d^4x \det \left(\frac{\partial x'}{\partial x} \right) \sqrt{-g'} \\ &= d^4x' \sqrt{-g'} \end{aligned}$$

$$\Rightarrow S = \int d^4x \sqrt{-g'} \mathcal{L}(x, \dot{x}, g)$$

is invariant under coordinate
transformations

Quantizing Gravity: The Graviton

Namely we could write
the action for the universe as

$$S = S_g + S_m$$

and quantize gravity by
calculating

$$\int Dg D\varphi e^{iS}$$

But this is not renormalizable
and is not easy to work with !

Einsteini showed us that
gravity is linked to energy/
momentum

the graviton is then the particle
associated with the field $g_{\mu\nu}$.

Defining the stress energy tensor

$$T^{\mu\nu}(x) = \frac{-2}{\sqrt{-g}} \frac{\delta S_m}{\delta g_{\mu\nu}}(x)$$

and expanding around flat spacetime

by writing $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$

$$\begin{aligned}\Rightarrow S_m(h) &= S_m(h=0) + \int d^4x \sqrt{-g} h_{\mu\nu} \frac{\delta S_m}{\delta g_{\mu\nu}} \\ &= S_m - \int d^4x \frac{1}{2} h_{\mu\nu} T^{\mu\nu} \\ &\quad + O(h^2)\end{aligned}$$

$$\begin{aligned}
 g &= \det(g_{\mu\nu}) = \det(\eta_{\mu\nu} + h_{\mu\nu}) \\
 &= -\det(1 + \eta^{\mu\nu} h_{\mu\nu}) \\
 &= -(1 + \eta^{\mu\nu} h_{\mu\nu} + O(h^2))
 \end{aligned}$$

$$g^{\mu\nu} = \eta^{\mu\nu} - h^{\mu\nu} + O(h^2)$$

The Einstein-Hilbert Action

$$S \equiv \frac{1}{16\pi G} \int d^4x \sqrt{-g} R$$

$$= M_p^2 \int d^4x \sqrt{-g} g^{\mu\nu} R_{\mu\nu}$$



Ricci Tensor

The Weak Field Action

turning $x^m \rightarrow x^m - \epsilon^m(x)$

$$\Rightarrow h_{\mu\nu} \rightarrow h_{\mu\nu} + \partial_\mu \epsilon_\nu + \partial_\nu \epsilon_\mu$$

↘ * Similar
to gauge
transformation
of

$$A_m \rightarrow A_m - \partial_m A$$

Since S must be invariant
under coordinate transforms and
if we expand out to $O(h^2)$ with

$$\sqrt{-g} g^{\mu\nu} = \eta^{\mu\nu} - h^{\mu\nu} + \frac{1}{2} \eta^{\mu\nu} h + O(h^2)$$

lead to

$$S_{Wfj} = \int d^4x \left(\frac{1}{2} M_p^2 \mathcal{I} - \frac{1}{2} h_{\mu\nu} T^{\mu\nu} \right)$$

where

$$\begin{aligned} \mathcal{I} = & \frac{1}{2} \partial_\lambda h^{\mu\nu} \partial^\lambda h_{\mu\nu} - \frac{1}{2} \partial_\lambda h^\mu_m \partial^\lambda h^\nu_n \\ & - \partial_\lambda h^{\lambda\nu} \partial^\mu h_{\mu\nu} + \partial^\nu h^\lambda_\lambda \partial^\mu h_{\mu\nu} \end{aligned}$$

The Graviton Propagator

trick 1 : add $(\partial^\mu h_{\mu\nu} - \frac{1}{2} \partial_\nu h^\lambda_\lambda)^2$ to action

$$\Rightarrow S_{Wfj} = \int d^4x \frac{1}{2} \left[\frac{M_p^2}{2} (\partial_\lambda h^{\mu\nu} \partial^\lambda h_{\mu\nu} - \frac{1}{2} \partial_\lambda h^\mu_m \partial^\lambda h^\nu_n) - h_{\mu\nu} T^{\mu\nu} \right]$$

this is also known as Harmonic
Gauge

as we are free to

choose $h_{\mu\nu}$ s.t. $\boxed{\partial^\mu h_{\mu\nu} = \frac{1}{2} \partial_\nu h^\lambda_\lambda}$

we can then write

$$S = -\frac{1}{2} M_p^2 \int d^4x \left[h^{\mu\nu} K_{\mu\nu;20} \square h^{20} + \mathcal{O}(h^3) \right]$$

with $K_{\mu\nu;20} = \frac{1}{2} \left(\gamma_{\mu 2} \gamma_{\nu 0} + \gamma_{\mu 0} \gamma_{\nu 2} - \gamma_{\mu\nu} \gamma_{20} \right)$

because $h_{\mu\nu}$ is symmetric, in harmonic gauge

$$K^{-1} = K ! \quad \text{so}$$

$$D_{\mu\nu,20}(k) = \frac{K_{\mu\nu;20}}{k^2 + i\epsilon}$$

Differential Geometry

a simple k -form:

$$\varphi = \varphi_{m_1 m_2 \dots m_k} dx^{m_1} dx^{m_2} \dots dx^{m_k}$$

$$d\varphi = \frac{1}{k!} \partial_\nu \varphi_{m_1 m_2 \dots m_k} dx^\nu dx^{m_1} dx^{m_2} \dots dx^{m_k}$$

$$* dd = 0 *$$

locally Euclidean

for a ^V Riemannian Manifold

$$g_{\mu\nu}(x) = e_m^a \delta_{ab} e_\nu^b \quad \text{(a similarity transformation of series)}$$

$$a = 1, \dots, D$$

"vielbeins" or "World Vectors"

↳ many legs

↳ "same role of metric"

on a curved manifold, parallel transport recorded infinitesimally rotates the vielbein:

$$de^a = -\omega^{ab} e^b$$

also

$$e^a = e_m^a dx^m$$

vielbein transform by rotation:

$$e^a(x) \rightarrow e'^a(x) = O_b^a(x) e^b(x)$$

$$\text{and } w \rightarrow w'$$

$$\text{defined by } de'^a = -\omega'^{ab} e^b$$

An instructive calculation:

$$de^a = d(O_b^a e^b)$$

$$= d(O_b^a e_m^b dx^m)$$

$$= \partial_\nu (O_b^a e_m^b) dx^\nu dx^m$$

$$= \partial_\nu O_b^a dx^\nu e^b + O_b^a \partial_\nu e_m^b dx^\nu dx^m$$

$$= dO_b^a e^b + O_b^a de^b$$

* Leibniz Property! *

$$= dO_b^a e^b - O_b^a \omega^{bc} e^c$$

$$= (dO_b^a O_d^b - O_b^a \omega^{bc} O_d^c) e^d$$

$$\Rightarrow \omega' = \Omega \omega \Omega^\top - (\delta \Omega) \Omega^\top$$

the 1-form ω transforms
just like the non-abelian gauge
potentials A_m^a

But (!), no analog of $e \dots$

ω is the connection between
local Euclidean frames, and
varies in a curved manifold:

$$d\omega + \omega^2 = R$$

& for completeness

$$de + \omega e = 0$$

Just like $F = dA + A^2$

field strength in NAGT

In Einstein Gravity in order
for fields and their derivatives
to transform in the same way
we must introduce a covariant
derivative, D to act on a
vector field W :

$$D_1 W^m = \partial_1 W^m + \sum_{\lambda} \Gamma_{\lambda}^m W^{\lambda}$$



Christoffel Symbol

* for spin $\frac{1}{2}$ fields *

$$\mathcal{L} = \bar{\psi} (\not{D}_m - m) \psi$$

$$\text{where } D_m = \partial_m - \frac{i}{4} \omega_{mab} \sigma^{ab}$$

because ψ is defined in local
Lorentz frame (spins, not
4-vectors?)

Topological Field Theory

in $(2+1)$ -D spacetime

anyons emerge, which acquire
a phase $e^{i \frac{\Theta}{\pi} \Delta q}$ after being rotated
 Δq around each other CC

this can be accounted for by
with

$$\mathcal{L} = \mathcal{L}_0 + \gamma \epsilon^{\mu\nu\lambda} u_\mu \partial_\nu u_\lambda + u_\mu i^\mu$$

Chern-Simons term

$$a_m \rightarrow a_m + \partial_m \lambda$$

$$\varepsilon^{\mu\nu\lambda} a_m \partial_\nu a_\lambda \rightarrow \varepsilon^{\mu\nu\lambda} (a_m + \partial_m \lambda) \\ \partial_\nu (a_\lambda + \partial_\lambda \lambda)$$

$$= \varepsilon^{\mu\nu\lambda} (a_m \partial_\nu a_\lambda + \partial_m \lambda \partial_\nu a_\lambda)$$

$$a_m \partial_\nu \partial_\lambda \lambda + \partial_m \lambda \partial_\nu \partial_\lambda \lambda)$$

$$= \varepsilon^{\mu\nu\lambda} (a_m \partial_\nu a_\lambda + \partial_m \lambda \partial_\nu a_\lambda)$$

$$JS = 2 \int d^3x \varepsilon^{\mu\nu\lambda} \partial_m (\lambda \partial_\nu a_\lambda)$$

\hookrightarrow dropped

$\Rightarrow JS$ Gauge invariant

C.S. action

$$S = \gamma \int_M d^3x \epsilon^{\mu\nu\lambda} a_\mu \partial_\nu a_\lambda$$

$$= \gamma \int_M d^3x a \wedge da$$

now under coordinate transformations
a vector:

$$a_m(x) = \frac{\partial x'^\lambda}{\partial x^m} a'_\lambda(x')$$

so

$$\epsilon^{\mu\nu\lambda} a_\mu(x) b_\nu(x) c_\lambda(x)$$

$$= \epsilon^{\mu\nu\lambda} \frac{\partial x'^\sigma}{\partial x^m} \frac{\partial x'^\tau}{\partial x^\nu} \frac{\partial x'^\rho}{\partial x^\lambda} a_\sigma(x') \dots$$

$$= \det \left(\frac{\partial x'}{\partial x} \right) \epsilon^{\sigma\tau\rho} a'_\sigma(x') b'_\tau(x') c'_\rho(x')$$

$$\Rightarrow d^3x \varepsilon^{m\nu\lambda} a_\mu(x) b_\nu(x) c_\lambda(x)$$

$$= d^3x' \varepsilon^{\sigma\tau\rho} a'_\sigma(x') b'_\tau(x') c'_\rho(x')$$

with $d^3x' = d^3x \det \left(\frac{\partial x'}{\partial x} \right)$

\Rightarrow no metric,

$$\int D\alpha e^{i \int_M a da}$$

\hookrightarrow closed manifold

depends only on topology

but what about $T^{\mu\nu}$

$$\text{no } g_{\mu\nu} \Rightarrow T^{\mu\nu} = 0$$

$$\Rightarrow H = 0$$

\Rightarrow Ground state degeneracy!

CS Gravity

in terms of connection ω

$$R_{ij}^k = \partial_i \omega_j{}^k - \partial_j \omega_i{}^k + [\omega_i, \omega_j]_k$$

or

$$R = d\omega + \omega \wedge \omega$$

and in $d=4$

$$\int_{\text{Einstein}} = \frac{1}{2} \int_M \epsilon^{ijkl} \epsilon_{abcd} (e_i{}^a e_j{}^b R_{kl}{}^{cd})$$

$$= \frac{1}{2} \int_M e \wedge e \wedge R$$

$$= \frac{1}{2} \int_M e \wedge e \wedge (dw + \omega^2)$$

ω has group structure $SO(3, 1)$

(interpreting e 's as translations)

(e, ω) is in $ISO(3, 1)$

\hookrightarrow GR is ISO gauge theory?

not in $d = 4$

no such gauge action of
form

$$\int A \wedge A \wedge (dA + A^2)$$

in $(2+1) - D$:

$$S_{EH} = \frac{1}{2} \int_M e \wedge (d\omega + \omega^2)$$



leads one S

$$S_{CS} = \frac{1}{2} \int_M Tr \left(A dA + \frac{2}{3} A^3 \right)$$

A a lie-algebra valued 1-form

" Tr " a non-degenerate invariant

bilinear form on Lie Algebra

$$A = A^a T_a \Rightarrow Tr(A dA) = Tr(T_a T_b) \xrightarrow{\text{das "neue" }} \times A^a \wedge dA^b$$

$ISO(2,1)$ Gauge Theory:

J^{ab} Laxez generators

$$[J_a, J_b] = \epsilon_{abc} J^c$$

P^a translation generators

$$[P_a, P_b] = 0$$

let $J^a = \frac{1}{2} \epsilon^{abc} J_{bc}$

Then the gauge field

is the 1-form

$$A_i = e_i^a P_a + \omega_i^a J_a$$

$$A = Pe + J\omega$$

an infinitesimal gauge parameter would be

$$u = \rho^a P_a + \tau^\mu J_\mu$$

$$\Rightarrow \delta A_i = - D_i u$$

with $D_i u = \partial_i u + [A_i, u]$

$$D = d + A$$

$$\Rightarrow D^2 = (d + A)(d + A)$$

$$= d^2 + Ad + dA + A^2$$

$$= (dA) + A^2$$

function:

$$(dA + Ad) \omega = d(A\omega) - Ad\omega$$

$$= (dA)\omega - 2Ad\omega$$

$$\Rightarrow Ad = (dA) - Ad$$

$$F = D^2$$

$$= (dA)^2 + A^2$$

$$= d(Pe + J\omega) + (Pe + J\omega)^2$$

$$= Pde + Jdw + P(e\omega + \omega e)$$

$$+ J\omega^2$$

$$= P(de + \omega e) + J(dw + \omega^2)$$

$$= 0 + J R$$

$$F \propto R$$

$$S_{cs} = \frac{1}{2} \int d^3x \quad Tr (A \wedge D^2)$$

$$\propto \frac{1}{2} \int d^3x \quad A \wedge R$$

$$\propto \frac{1}{2} \int d^3x \quad e \wedge R + S_w$$

↳ S_{EH}

for constant curvature:

$$\Rightarrow \int D\epsilon D\omega e^{iS_{cs}}$$

$$= \left(\int D\omega e^{iS_w} \right) \left(\int D\epsilon e^{iS_{EH}} \right)$$