

Simulating Gaussian tracers correlated with CMB lensing

Anton Baleato Lizancos
ab2368@cam.ac.uk

The goal is to generate simulations of large-scale-structure tracers which are appropriately correlated with a given simulation of the CMB lensing potential. The basic idea is to set up a linear system of equations from which it will be possible to solve for the coefficients involved granted we know the theoretical auto- and cross- spectra of the tracers.

Suppose we are to work with just three tracers: the CIB, LSST galaxies and an internal reconstruction, with spherical harmonic coefficients expressed as $I_{lm}, g_{lm}, \kappa_{lm}^{rec}$, respectively. Given a map of the true convergence with spherical harmonic coefficients κ_{lm} , the different tracers are correlated as

$$\kappa_{lm}^{rec} = \kappa_{lm} + n_{lm}, \quad (1)$$

$$g_{lm} = A_l^{g\kappa} \kappa_{lm} + u_{lm}, \quad (2)$$

$$I_{lm} = A_l^{I\kappa} \kappa_{lm} + A_l^{gI} u_{lm} + e_{lm}, \quad (3)$$

$$(4)$$

where n_{lm}, u_{lm}, e_{lm} are coefficients of the noise – and as such are presumed to each be uncorrelated with everything else – and κ_{lm} are the coefficients of the true convergence for the particular realisation that we wish to generate correlated tracers for. Solving for the coefficients

$$A_l^{g\kappa} = \frac{C_l^{g\kappa}}{C_l^{\kappa\kappa}} \quad (5)$$

$$A_l^{I\kappa} = \frac{C_l^{I\kappa}}{C_l^{\kappa\kappa}} \quad (6)$$

$$A_l^{gI} = \frac{C_l^{gI} - A_l^{g\kappa} A_l^{I\kappa} C_l^{\kappa\kappa}}{C_l^{uu}} \quad (7)$$

and the noise spectra

$$C_l^{nn} = N_l^{\kappa\kappa} \quad (8)$$

$$C_l^{uu} = C_l^{gg} - (A_l^{g\kappa})^2 C_l^{\kappa\kappa} \quad (9)$$

$$C_l^{ee} = C_l^{II} - (A_l^{I\kappa})^2 C_l^{\kappa\kappa} - (A_l^{gI})^2 C_l^{uu}, \quad (10)$$

where $N_l^{\kappa\kappa}$ is the internal reconstruction noise level.

From the above, it is clear that the problem is soluble if we know all the auto- and cross- spectra, in which case all we have to do is draw the coefficients n_{lm}, u_{lm} and e_{lm} from probability distributions with angular power spectra $N_l^{\kappa\kappa}, C_l^{uu}$ and C_l^{ee} .

The solution above can be generalised to any number of tracers as follows. Consider tracer i with auto- spectrum C^{ii} and cross-spectrum C^{ij} with tracer j . Let it be described as a linear combination $\sum_p \sum_{lm} A_l^{ij} a^p(l, m)$ of harmonic coefficients a^p with angular power spectra $C_l^{a^p a^p}$ and scaled with weights given by the square, off-diagonal matrix A_l^{ij} . In the case above, $a^p = \{\kappa_{lm}, u_{lm}, e_{lm}\}$, $C_l^{a^p a^p} = \{C_l^{\kappa\kappa}, C_l^{uu}, C_l^{ee}\}$. The general formula for the weights is

$$A_l^{ij} = \frac{1}{C_l^{a^j a^j}} \left(C_l^{ij} - \sum_{p=0}^{j-1} A_l^{jp} A_l^{ip} C_l^{a^p a^p} \right), \quad (11)$$

with auxiliary spectra given by:

$$C_l^{a^j a^j} = C_l^{jj} - \sum_{p=0}^{j-1} (A_l^{jp})^2 C_l^{a^p a^p}. \quad (12)$$