

An overview of `psd`: Adaptive sine multitaper power spectral density estimation in R

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Abstract

This vignette provides an overview of some features included in the package `psd`, designed to compute estimates of power spectral density (PSD) for a univariate series in a sophisticated manner, with very little tuning effort. The sine multitapers are used, and the number of tapers varies with spectral shape, according to the optimal value proposed by Riedel and Sidorenko (1995). The adaptive procedure iteratively refines the optimal number of tapers at each frequency based on the spectrum from the previous iteration. Assuming the adaptive procedure converges, this produces power spectra with significantly lower spectral variance relative to results from less-sophisticated estimators. Sine tapers exhibit excellent leakage suppression characteristics, so bias effects are also reduced. Resolution and uncertainty vary with the number of tapers, which means we do not need to resort to either (1) windowing methods, which inherently degrade resolution at low-frequency (e.g. Welch's method); or (2) smoothing kernels, which can badly distort important features without careful tuning (e.g. the Daniell kernel in `stats::spectrum`). In this regards `psd` is best suited for data having large dynamic range and some mix of narrow and wide-band structure, features typically found in geophysical datasets.

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1 Quick start: A minimal example.

First, we load the package into the namespace:

```
library(psd)

## Loading required package: fftw
## Loaded psd (0.5.0) - Adaptive multitaper spectrum estimation
```

For a series to analyze, we can use `magnet`, included in `psd`, which represents along-track measurements of horizontal magnetic-field strength from a gimbaled, airborne magnetometer. These data are a small subset of the full Project MAGNET series (Coleman, 1992), which has provided insight into the history of the Earth's oceanic crust (Parker and O'Brien, 1997; O'Brien et al., 1999; Korte et al., 2002). The sampling interval is once every kilometer (km), so the data will represent crustal magnetization with wavelengths longer than 2 km.

```
data(magnet)
```

The format of the data set is a `data.frame` with four sets of information:

```
names(magnet)

## [1] "km"      "raw"     "clean"   "mdiff"
```

The `raw` and `clean` names represent raw and edited intensities respectively, expressed in units of nanotesla; `mdiff` is the difference between them. The difference between them is a matter of just a few points attributable to instrumental malfunction.

```
subset(magnet, abs(mdiff) > 0)

##      km      raw      clean      mdiff
## 403 402  209.1 -3.6355 -212.7355
## 717 716 -248.7 -9.7775  238.9225
```

These deviations can, as we will see, adversely affect the accuracy of any PSD estimate, multitaper or otherwise.

We can find power spectral density (PSD) estimates for the two series quite simply with `pspectrum`:

```
psdr <- pspectrum(magnet$raw)

## Stage 0 est. (pilot)
## environment **.psdEnv** refreshed
## detrending (and demeaning)
## Stage 1 est. (Ave. S.V.R. 3.0 dB)
## Stage 2 est. (Ave. S.V.R. 6.2 dB)
## Stage 3 est. (Ave. S.V.R. 6.7 dB)
## Stage 4 est. (Ave. S.V.R. 7.2 dB)
## Stage 5 est. (Ave. S.V.R. 7.2 dB)
## Normalized single-sided PSD (PSD) to single-sided PSD for sampling-freq. 1

psdc <- pspectrum(magnet$clean)
```

```
## Stage 0 est. (pilot)
## environment ** .psdEnv ** refreshed
## detrending (and demeaning)
## Stage 1 est. (Ave. S.V.R. 3.6 dB)
## Stage 2 est. (Ave. S.V.R. 7.3 dB)
## Stage 3 est. (Ave. S.V.R. 7.2 dB)
## Stage 4 est. (Ave. S.V.R. 6.6 dB)
## Stage 5 est. (Ave. S.V.R. 6.8 dB)
## Normalized single-sided PSD (PSD) to single-sided PSD for sampling-freq. 1
```

Each application of `pspectrum` calculates a pilot PSD, followed by `niter` iterations of refinement. With each iteration the number of tapers is adjusted based on the proposed optimal number from Riedel and Sidorenko (1995), which depends on spectral shape; we use quadratically weighted spectral derivatives (Prieto et al., 2007) to estimate this shape. By default, a multipanel summary plot of the final PSD compared to the raw periodogram estimate is shown after the final iterative stage. Note that if the user forgets to assign the results of `pspectrum` to the global environment, this can be done with the `psd_envGet` function:

```
psdc_recovered <- psd_envGet("final_psd")
all.equal(psdc, psdc_recovered)

## [1] TRUE
```

In general, spectral variance is reduced with sequential refinements¹, but is not necessarily guaranteed to converge. Note that in the example the sampling frequency of both series is 1 km^{-1} , the assumed value.

Figure 1 compares the power spectra for the `raw` and `clean` series². We expect the Project MAGNET data to be linear in the space of linear-frequencies and logarithmic-power; we see a clear improvement in spectral shape between the two series, simply because the large outliers have been removed. The PSD of the clean series shows a very “red” spectrum typical of geophysical processes (Agnew, 1992), and a rolloff in signal for 10 kilometer wavelengths and longer; whereas, the PSD for the raw series looks somewhat unrealistic at higher wavelengths—features which could be difficult to judge if the spectral variance was higher.

2 Comparisons with other methods

As we have shown in the Project MAGNET example, improved understanding of the physics behind the signals in the data is of great concern. Assuming a sample is free of non-physical points, how do PSD estimates from `psd` compare with other methods? Unfortunately the suite of extensions with similar functionality is relatively limited, but hopefully we have summarized most, if not all, the available functions in Table 1.

We compare results from `psd` with those from a few of the methods in Table 1, using the same data: the cleaned Project MAGNET series.

2.1 `stats::spectrum`

Included in the core distribution of R is `stats::spectrum`, which accesses `stats::spec.ar` or `stats::spec.pgram` for either parametric and non-parametric estimation, respectively. The user can optionally apply a single cosine taper, and/or a smoothing kernel. Our method is non-parametric; hence, we will compare to the latter.

¹ Messages are given by default; ones with “Ave. S.V.R.” are in reference to “average spectral-variance reduction”, which is the variance of the double-differenced spectra at each stage, relative to the pilot estimate’s variance.

² Note that `pspectrum` returns an object with class `spec`, so we have access to methods within `stats`, including `plot.spec`.

Table 1: A comparison of power spectral density estimators in R, excluding extensions which only estimate raw-periodograms. Normalizations are shown as either “single” or “double” for either single- or double-sided spectra, and “various” if there are multiple, optional normalizations. A (*) denotes the default for a function having an option for either single or double.

FUNCTION	NAMESPACE	SINE M.T.?	ADAPTIVE?	NORM.	REFERENCE
<code>bspec</code>	<code>bspec</code>	NO	NO	single*	Röver et al. (2011)
<code>mtapspec</code>	<code>RSEIS</code>	YES	NO	various	Lees and Park (1995)
<code>pspectrum</code>	<code>psd</code>	YES	YES	single	Barbour and Parker (2014, 2015)
<code>spectrum</code>	<code>stats</code>	NO	NO	double	R Core Team (2013)
<code>spec.mtm</code>	<code>multitaper</code>	YES	YES	double	Rahim and Burr (2013)
<code>SDF</code>	<code>sapa</code>	YES	NO	single*	Percival and Walden (1993)

Included in `psdcore` is an option to compare the results with a 20% cosine-tapered periodogram, found with the following command:

```
spec.pgram(X, pad = 1, taper = 0.2, detrend = FALSE, demean = FALSE, plot = FALSE)
```

Within `psdcore` the comparison is made with the logical argument `preproc` passed to `spec.pgram`, which is `TRUE` by default.

As a matter of bookkeeping and good practice, we should consider the working environment accessed by `psd` functions. To ensure `psdcore` does not access any inappropriate information leftover from the previous calculations, we can set `refresh=TRUE`. We can then re-calculate the multitaper PSD and the raw periodogram with `plotpsd=TRUE`; these results are shown in Figure 2.

2.2 RSEIS::mtapspec

In `RSEIS` the spectrum estimation tool is `mtapspec`, which calls the program of Lees and Park (1995). There are numerous optional tuning parameters, including flags for normalization and taper averaging. For our purpose the correct normalization for `mtapspec` is found by using `MTP=list(kind=2, inorm=3)` and scaling the results by 2 (to convert double-sided spectra to single-sided spectra).

We assume `mtapspec` doesn’t remove a mean and trend from the input series. We can do this easily with the `prewhiten` methods:

```
require(RSEIS)

## Loading required package: RSEIS

dt = 1 # km
# prewhiten the data after adding a linear trend + offset
summary(prewhiten(mc <- (ts(magnet$clean + 1000, deltat = dt) + seq_along(magnet$clean)),
  plot = FALSE))

## detrending (and demeaning)

##          Length Class  Mode
## lmdfit      12    lm     list
## ardfit       0 -none-  NULL
```

```
## prew_lm 2048    ts      numeric
## prew_ar      0  -none-  NULL
## imputed      1  -none-  logical
```

Although the default operation of `prewhiten` is to fit a linear model of the form $f(x) = \alpha x + \beta + \epsilon$ using ordinary linear least squares, setting `AR.max` higher than zero to fit an auto-regressive (AR) model to the data³. This fit uses the Akaike information criterion (AIC) to select the highest order appropriate for the data.

```
summary(atsar <- prewhiten(mc, AR.max = 100, plot = FALSE))

## detrending (and demeaning)
## autoregressive model fit (returning innovations)

##      Length Class  Mode
## lmdfit    12    lm    list
## ardfit     14    ar    list
## prew_lm 2048    ts      numeric
## prew_ar 2048    ts      numeric
## imputed    1  -none-  logical

print(atsar$ardfit)

##
## Call:
## ar.yw.default(x = tser_prew_lm, aic = TRUE, order.max = AR.max,      demean = TRUE)
##
## Coefficients:
##      1      2      3      4      5      6
## 1.5134 -1.1037  0.6723 -0.3880  0.2108 -0.0786
##
## Order selected 6  sigma^2 estimated as 19.46

ats_lm <- atsar$prew_lm
ats_ar <- atsar$prew_ar
```

We didn't necessarily need to deal with the sampling information since it is just 1 per km; but, supposing the sampling information was based on an interval, we could have used a negative value for `X.frq`, with which `psdcore` would interpret as an interval (instead of a frequency). A quick example highlights the equivalency:

```
a <- rnorm(32)
all.equal(psdcore(a, 1, 1, first.last = FALSE)$spec, psdcore(a, -1, 1,
  first.last = FALSE)$spec)

## [1] TRUE
```

Returning the the RSEIS comparison, we first estimate the PSD from `mtapspec` with 10 tapers:

³Note that the linear trend fitting is removed from the series prior to AR estimation, and the residuals from this fit are also returned.

```
tapinit <- 10
Mspec <- mtapspec(ats_lm, deltat(ats_lm), MTP = list(kind = 2, inorm = 3,
  nwin = tapinit, npf = 0))
```

where `nwin` is the number of tapers taken and `npf` is, from the documentation, the “number of Pi-prolate functions” (we leave it out for the sake of comparison). Note that the object returned is not of class `spec`:

```
str(Mspec)

## List of 12
## $ dat      : ts [1:2048, 1] -16.23 -14.56 -12.02 -7.21 -3.13 ...
## .. attr(*, "dimnames")=List of 2
## .. ..$ : NULL
## .. ..$ : NULL
## .. attr(*, "tsp")= num [1:3] 1 2048 1
## $ dt       : num 1
## $ spec     : num [1:4096] 528 557 600 595 615 ...
## $ dof      : num [1:4096] 20 20 20 20 20 20 20 20 20 20 ...
## $ Fv       : num [1:4096] 4.45e-20 4.78e-02 5.36e-01 1.54 1.15 ...
## $ Rspec    : num [1:2049, 1:10] 1.86e-07 -9.32e+01 6.05e+02 1.16e+03 -2.97e+02 ...
## $ Ispec    : num [1:2049, 1:10] 0 -227 -569 665 1157 ...
## $ freq     : num [1:2049] 0 0.000244 0.000488 0.000732 0.000977 ...
## $ df       : num 0.000244
## $ numfreqs : num 2049
## $ klen     : num 4096
## $ mtm      :List of 4
## ..$ kind : num 2
## ..$ nwin : num 10
## ..$ npf  : num 0
## ..$ inorm: num 3
```

We will calculate the comparative spectra from

1. `spectrum` (20% cosine taper),
2. `psdcore` (with fixed tapers), and
3. `pspectrum` (allowing adaptive taper refinement)

and we will need to correct for normalization factors, as necessary, with `normalize`. Note that by default the normalization is set within `pspectrum` (with `normalize`) once the adaptive procedure is finished.

```
Xspec <- spec.pgram(ats_lm, pad = 1, taper = 0.2, detr = TRUE, dem = TRUE,
  plot = FALSE)
Pspec <- psdcore(ats_lm, dt, tapinit)
Aspec <- pspectrum(ats_lm, dt, tapinit, plot = FALSE)

## Stage 0 est. (pilot)
## environment ** .psdEnv ** refreshed
## detrending (and demeaning)
## Stage 1 est. (Ave. S.V.R. -9.3 dB)
## Stage 2 est. (Ave. S.V.R. -12.3 dB)
```

```
## Stage 3 est. (Ave. S.V.R. -14.3 dB)
## Stage 4 est. (Ave. S.V.R. -15.2 dB)
## Stage 5 est. (Ave. S.V.R. -16.2 dB)
## Stage 6 est. (Ave. S.V.R. -16.2 dB)
## Stage 7 est. (Ave. S.V.R. -16.1 dB)
## Stage 8 est. (Ave. S.V.R. -16.1 dB)
## Stage 9 est. (Ave. S.V.R. -16.1 dB)
## Stage 10 est. (Ave. S.V.R. -16.1 dB)
## Normalized single-sided PSD (PSD) to single-sided PSD for sampling-freq. 1

# Correct for double-sidedness of spectrum and mtapspec results
class(Mspec)

Mspec <- normalize(Mspec, dt, "spectrum")

## Normalized double-sided PSD (SPECTRUM) to single-sided PSD for sampling-freq. 1

nt <- 1:Mspec$numfreqs
mspec <- Mspec$spec[nt]
class(Xspec)

Xspec <- normalize(Xspec, dt, "spectrum")

## Normalized double-sided PSD (SPECTRUM) to single-sided PSD for sampling-freq. 1
```

These estimates are shown on the same scale in Figure 4.

Because we did not specify the length of the FFT in `mtapspec` we end up with different length spectra. So, to form some statistical measure of the results, we can interpolate PSD levels onto the `psd`-based frequencies (or reciprocally):

```
library(signal, warn.conflicts = FALSE)
pltpi <- interp1(pltf, pltp, Pspec$freq)
```

We regress the spectral values from `mtapspec` against the `psdcore` results because we have used them to produce uniformly tapered spectra with an equal number of sine tapers.

```
df <- data.frame(x = dB(Pspec$spec), y = pltpi, tap = unclass(Aspec$taper))
summary(dflm <- lm(y ~ x + 0, df))

##
## Call:
## lm(formula = y ~ x + 0, data = df)
##
## Residuals:
##      Min       1Q   Median       3Q      Max
## -2.2580 -0.3454  0.1711  0.9643  5.5561
##
## Coefficients:
```

```
## Estimate Std. Error t value Pr(>|t|)
## x 0.992045 0.002035 487.6 <2e-16 ***
## ---
## Signif. codes: 0 '***' 0.001 '**' 0.01 '*' 0.05 '.' 0.1 ' ' 1
##
## Residual standard error: 1.189 on 1023 degrees of freedom
## Multiple R-squared: 0.9957, Adjusted R-squared: 0.9957
## F-statistic: 2.377e+05 on 1 and 1023 DF, p-value: < 2.2e-16

df$res <- residuals(dflm)
```

We show the regression residuals in Figure 5. The structure visible at low power levels might be from curvature bias in the `mtapspec` results, which manifests at short wavelengths in Figure 4.

2.3 `multitaper::spec.mtm`

The function with the highest similarity to `psd` is `spec.mtm` in the `multitaper` package: it uses the sine multitapers, and can adaptively refine the spectrum. In fact, this function calls source code of a Fortran equivalent to `psd` authored by R.L. Parker (2013) to do these operations.

There are some notable differences, though. By default `spec.mtm` uses the Discrete Prolate Spheroidal Sequences (dpss) of Thomson (1982), which can have very good spectral leakage suppression (assuming the number of tapers used is appropriate for the desired resolution, which varies inversely with the time-bandwidth product). Spectral analyses using dpss can have superior results if the series is relatively short (e.g. $N < 1000$), or has inherent spectra with sharply changing features or deep wells. Improper usage of the dpss, however, can lead to severe bias. Thus, considerable care should be given to parameter choices, which translates practicably to having many more knobs to turn.

2.4 `sapa::SDF`

As of this writing, the package has no maintainer; lest we end up discussing deprecated and archived functions, we will not compare it to `psd`.

2.5 `bspec::bspec`

An intriguing method for producing power spectral density estimates using Bayesian inference is presented by Röver et al. (2011) and included in the `bspec` package. Simplistically, the method uses a *Student's t* likelihood function to estimate the distribution of spectral densities at a given frequency. We will use the spectra from the previous calculation to compare with `bspec` results. For this comparison we use the default settings for the *a priori* distribution scale and degrees of freedom. In Figure 6 we have used the `plot.bspect` method and overlain the results found previously by `psdcore`.

```
library(bspec)

##
## Attaching package: 'bspec'
##
## The following object is masked from 'package:stats':
##
## acf
```



```
##
## The following object is masked from 'package:base':
##
##      sample
print(Bspec <- bspec(ts(magnet$clean)))
```

3 Can AR prewhitening improve the spectrum?

This question must be addressed on a case-by-case basis; but, if there is significant auto-regressive structure in the series then the answer is likely YES. The MAGNET dataset is an example where the structure of the series is nicely represented by an AR model with a random noise component.

Recall the results of the prewhitening in Section 2.2. While `AR.max` was set relatively high, only an AR(6) model was fit significantly, according to the AIC requirements. The estimated variance of the innovations is about 20 nT². If the innovation spectrum is flat (as we expect), this variance translates to power levels of about 16 decibels for a 1 km sampling interval.

```
ntap <- 7
psd_ar <- psdcore(ats_ar, ntaper = ntap, refresh = TRUE)
dB(mean(psd_ar$spec))

## [1] 15.82824
```

In Figure 7 we have used `pilot_spec` to model the spectral response of the AR component of the series (solid black line). The non-AR component (labelled "AR-innovations") contributes approximately ± 3 dB to the original spectrum. Overlain on these series is the adaptive spectrum found previously.

4 Assessing spectral properties

4.1 Spectral uncertainties

It is important to place bounds on the uncertainties associated with a spectral estimate. In a multitaper algorithm the uncertainty is distributed as a χ^2_ν variate where ν is the number of degrees of freedom, which is twice the number of tapers applied. A proxy for this is simply $1/\sqrt{\nu-1}$. Using $\nu = 2 * K$ we can approximate the distribution of uncertainties from the tapers alone; however, a more rigorous estimate comes from evaluating the appropriate distribution for a coverage probability (e.g. $p = 0.95$). Among other calculations, `spectral_properties` returns the χ^2_ν based confidence intervals for $p = 0.95$, as well as the approximate uncertainties.

To illustrate, we plot the uncertainties for an integer sequence⁴ of tapers [0,50], shown in Figure 8. The benefits of having more than just a few tapers becomes obvious, though the spectral uncertainty is

⁴ Note the χ^2_ν distribution is defined for non-negative, non-integer degrees of freedom, but we cannot apply fractions of tapers.

asymptotically decreasing with taper numbers and yields only slight improvements with logarithmic number of tapers.

Returning to the Project MAGNET spectra, we will compare the spectral uncertainties from `psd` to the those from `bspec`, the Bayesian method, for a coverage probability of 95%. Figure 9 shows the uncertainties as bounded polygons, which we calculate here:

```
spp <- spectral_properties(Pspec$taper, db.ci = TRUE)
spa <- spectral_properties(Aspec$taper, db.ci = TRUE)
str(spa)

## 'data.frame': 1024 obs. of 8 variables:
## $ taper : num 4 5 4 3 2 3 4 5 6 7 ...
## $ stderr.chi.lower : num -3.51 -3.2 -3.51 -3.95 -4.64 ...
## $ stderr.chi.upper : num 5.47 4.75 5.47 6.61 8.76 ...
## $ stderr.chi.median : num 1.73 1.52 1.73 2.05 2.64 ...
## $ stderr.chi.approx : num 1.39 1.25 1.39 1.61 1.98 ...
## $ resolution : num 0.00977 0.01172 0.00977 0.00781 0.00586 ...
## $ dof : num 8 10 8 6 4 6 8 10 12 14 ...
## $ bw : num 0.00488 0.00586 0.00488 0.00391 0.00293 ...

create_poly <- function(x, y, dy, from.lower = FALSE) {
  xx <- c(x, rev(x))
  if (from.lower) {
    yy <- c(y, rev(y + dy))
  } else {
    yy <- c(y + dy, rev(y - dy))
  }
  return(data.frame(xx = xx, yy = yy))
}

psppu <- create_poly(Pspec$freq, dB(Pspec$spec), spp$stderr.chi.upper)
pspau <- create_poly(Aspec$freq, dB(Aspec$spec), spa$stderr.chi.upper)
# and the Bayesian spectrum 95% limits
pspb <- create_poly(Bspec_plt$freq, Bspec_plt$spectrum[, 1], Bspec_plt$spectrum[,
3], from.lower = TRUE)
```

4.2 Spectral resolution

There is an inherent tradeoff between the number of tapers applied and the spectral resolution (effectively, the spectral bandwidth). In general, the greater the number of tapers applied, the lower the spectral resolution. We can use the information returned from `spectral_properties` to visualize the actual differences in resolution for the Project MAGNET PSD estimates; these are shown in Figure 10.

4.3 Visualizing the adaptive history

One might be curious to study how the uncertainties change with each iteration. `pspectrum` saves an array of “historical” data in its working environment. Specifically, it saves the frequencies, spectral values, and number of tapers at each stage of the adaptive procedure, accessible with `get_adapt_history`. To ensure a fresh calculation and to add a few more iterations to visualize, we repeat the adaptive spectral analysis, and then bring the stage history into the `.GlobalEnv` environment:

```

pspectrum(ats_lm, niter = 4, plot = FALSE)

## Stage 0 est. (pilot)
## environment ** .psdEnv ** refreshed
## detrending (and demeaning)
## Stage 1 est. (Ave. S.V.R. 3.6 dB)
## Stage 2 est. (Ave. S.V.R. 7.3 dB)
## Stage 3 est. (Ave. S.V.R. 7.2 dB)
## Stage 4 est. (Ave. S.V.R. 6.6 dB)
## Normalized single-sided PSD (PSD) to single-sided PSD for sampling-freq. 1

str(AH <- get_adapt_history())

## List of 3
## $ freq : num [1:1024] 0 0.000489 0.000978 0.001466 0.001955 ...
## $ stg_kopt:List of 5
## ..$ : atomic [1:1024] 7 7 7 7 7 7 7 7 7 ...
## ..$ : attr(*, "span_was_set")= logi TRUE
## ..$ : attr(*, "n_taper_limits_orig")= num [1:2] 1 7
## ..$ : num [1:1024] 4 5 6 7 6 5 4 5 6 7 ...
## ..$ : num [1:1024] 4 5 6 5 4 3 4 5 6 7 ...
## ..$ : num [1:1024] 4 5 4 3 2 3 4 5 6 7 ...
## ..$ : num [1:1024] 4 5 4 3 2 3 4 5 6 7 ...
## $ stg_psd :List of 5
## ..$ : num [1:1024] 654 1306 1337 1381 1335 ...
## ..$ : num [1:1024] 640 1274 1532 1411 1444 ...
## ..$ : num [1:1024] 638 1273 1530 1746 1563 ...
## ..$ : num [1:1024] 631 1259 1977 2290 1885 ...
## ..$ : num [1:1024] 629 1254 1970 2282 1879 ...

```

Followed by some trivial manipulation:

```

Freqs <- AH[["freq"]]
Dat <- AH[["stg_psd"]]
numd <- length(Freqs)
numit <- length(Dat)
StgPsd <- dB(matrix(unlist(Dat), ncol = numit))
Dat <- AH[["stg_kopt"]]
StgTap <- matrix(unlist(Dat), ncol = numit)

```

We can plot these easily with `matplot` or other tools. We show the adaptive history in Figure 11.

It may be informative to investigate cross correlation coefficients between the stages; but, in this case, only the PSD estimates are significantly correlated:

```

suppressWarnings(symnum(cT <- cor(StgTap)))

##
## [1,] 1
## [2,] ? 1
## [3,] ? 1
## [4,] ? + 1

```

```
## [5,] ?   + + 1
## attr("legend")
## [1] 0 ' ' 0.3 '.' 0.6 ',' 0.8 '+' 0.9 '*' 0.95 'B' 1 \t    ## NA: '?'
```

```
suppressWarnings(symnum(cP <- cor(StgPsd)))

##
## [1,] 1
## [2,] B 1
## [3,] B B 1
## [4,] B B B 1
## [5,] B B B B 1
## attr("legend")
## [1] 0 ' ' 0.3 '.' 0.6 ',' 0.8 '+' 0.9 '*' 0.95 'B' 1
```

5 Call overview

Shown in Figure 12 is a flow chart highlighting the essential functions involved in the adaptive estimation process. The primary function is `pspectrum`.

Session Info

```
utils::sessionInfo()

## R version 3.1.2 (2014-10-31)
## Platform: x86_64-apple-darwin13.4.0 (64-bit)
##
## locale:
## [1] en_US.UTF-8/en_US.UTF-8/en_US.UTF-8/C/en_US.UTF-8/en_US.UTF-8
##
## attached base packages:
## [1] stats      graphics  grDevices  utils      datasets  methods
## [7] base
##
## other attached packages:
## [1] bspec_1.4          ggplot2_1.0.0      signal_0.7-4
## [4] RColorBrewer_1.1-2 RSEIS_3.3-3        psd_0.5-0
## [7] fftw_1.0-3         knitr_1.9
##
## loaded via a namespace (and not attached):
## [1] colorspace_1.2-4 digest_0.6.8      evaluate_0.5.5
## [4] formatR_1.0       grid_3.1.2       gtable_0.1.2
## [7] highr_0.4         labeling_0.3      lattice_0.20-29
## [10] MASS_7.3-37       munsell_0.4.2     plyr_1.8.1
## [13] proto_0.3-10      Rcpp_0.11.4       reshape2_1.4.1
## [16] RPMG_2.1-5        Rwave_2.2         scales_0.2.4
## [19] stringr_0.6.2     tools_3.1.2       zoo_1.7-11
```

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```
plot(psd, log = "dB", main = "Raw and Clean Project MAGNET power spectral density",
     lwd = 3, ci.col = NA, ylim = c(0, 32), yaxs = "i")
plot(psd, log = "dB", add = TRUE, lwd = 3, lty = 5)
text(c(0.25, 0.34), c(11, 24), c("Clean", "Raw"), cex = 1)
```

Raw and Clean Project MAGNET power spectral density

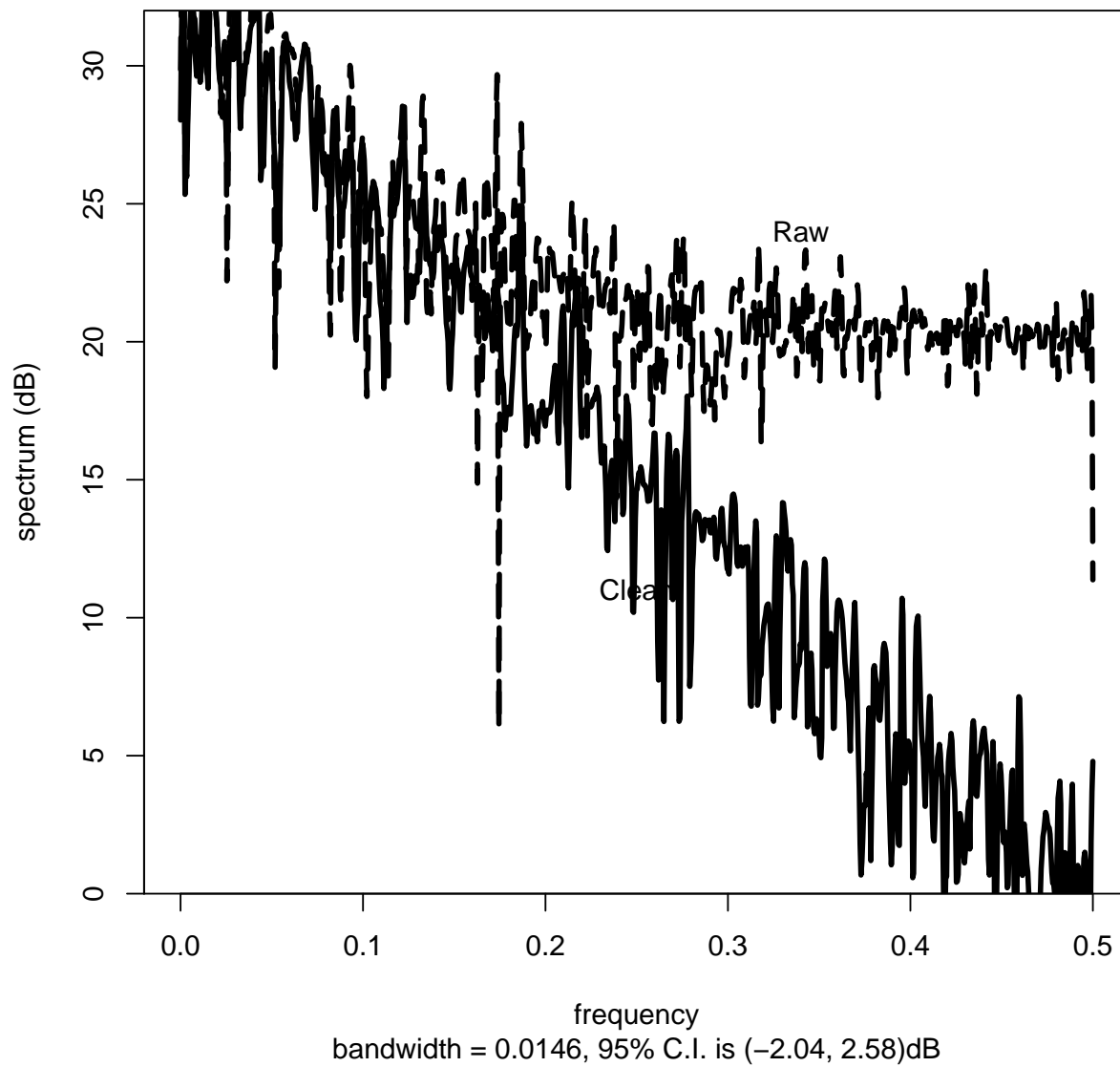


Figure 1: Power spectral density estimates for the raw and cleaned Project MAGNET data bundled with [psd](#). Note that because the class is 'spec' we have utilized existing methods in the stats namespace.

```
ntap <- psdc$taper
psdcore(magnet$clean, ntaper = ntap, refresh = TRUE, plotpsd = TRUE)
```

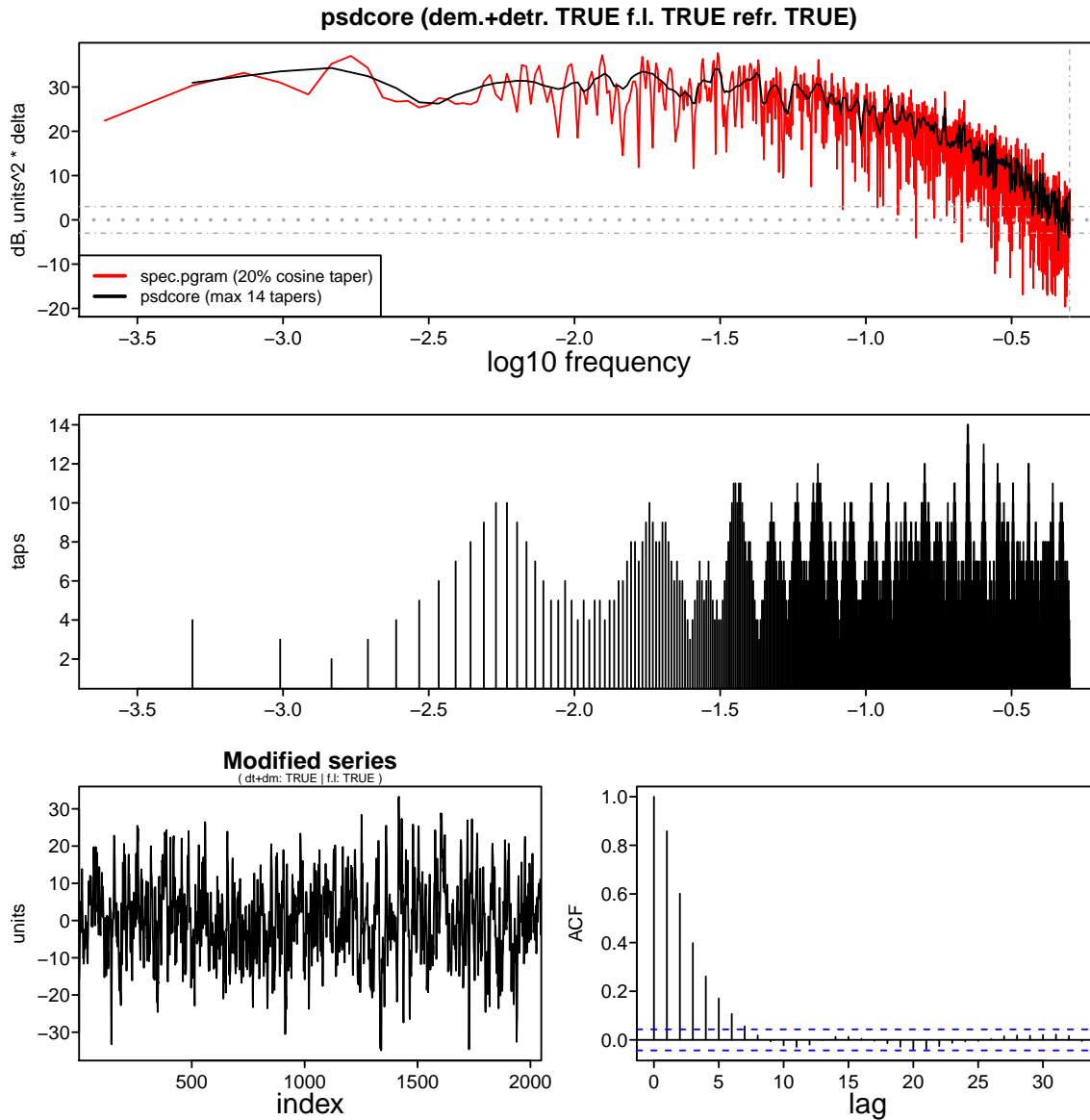


Figure 2: A summary plot produced by `psdcore` when `plotpsd=TRUE`. Top: Comparison between PSD estimators for the clean Project MAGNET data. The frequency axis is in units of $\log_{10} \text{ km}^{-1}$, and power axis is in decibels. Middle: The number of tapers applied as a function of frequency from the `plot.tapers` method. Bottom: The spatial series used to estimate the PSDs and a subset of the full autocorrelation function.


```
plot(ts.union(orig.plus.trend = mc, linear = ats_lm, ar = ats_ar), yax.flip = TRUE,
     main = sprintf("Prewhitened Project MAGNET series"))
mtext(sprintf("linear and linear+AR(%s)", atsar$ardfit$order), line = 1.1)
```

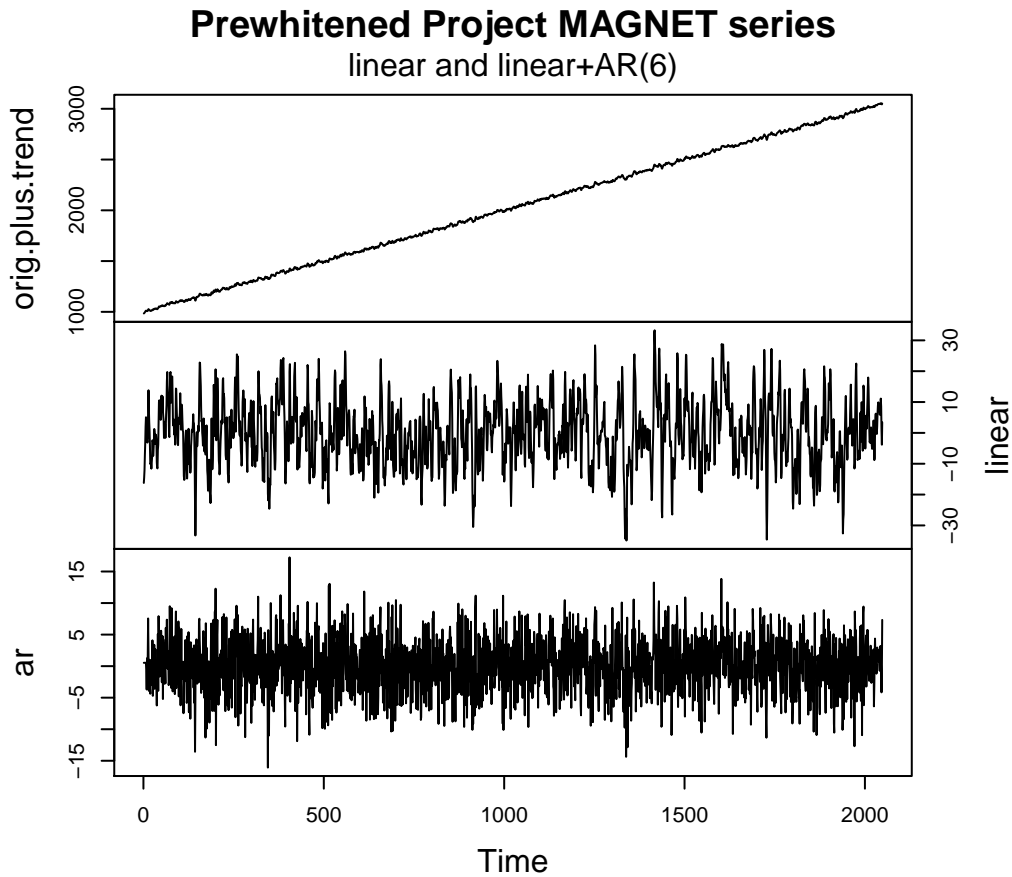


Figure 3: Pre-whitening of the Project MAGNET series (with a synthetic linear model superimposed on it) assuming linear and linear-with-AR models.

```

library(RColorBrewer)
cols <- c("dark grey", brewer.pal(8, "Set1")[c(5:4, 2)])
lwds <- c(1, 2, 2, 5)
par(las = 1)
plot(Xspec, log = "dB", ylim = 40 * c(-0.4, 1), ci.col = NA, col = cols[1],
     lwd = lwds[1], main = "PSD Comparisons")
pltf <- Mspec$freq
lines(pltf, pltf <- dB(mspec), col = cols[2], lwd = lwds[2])
plot(Pspec, log = "dB", add = TRUE, col = cols[3], lwd = lwds[3])
plot(Aspec, log = "dB", add = TRUE, col = cols[4], lwd = lwds[4])
legend("topright", c("spec.pgram", "RSEIS::mtapspec", "psdcore", "pspectrum"),
     title = "Estimator", lwd = 3, cex = 1.1, col = cols)

```

PSD Comparisons

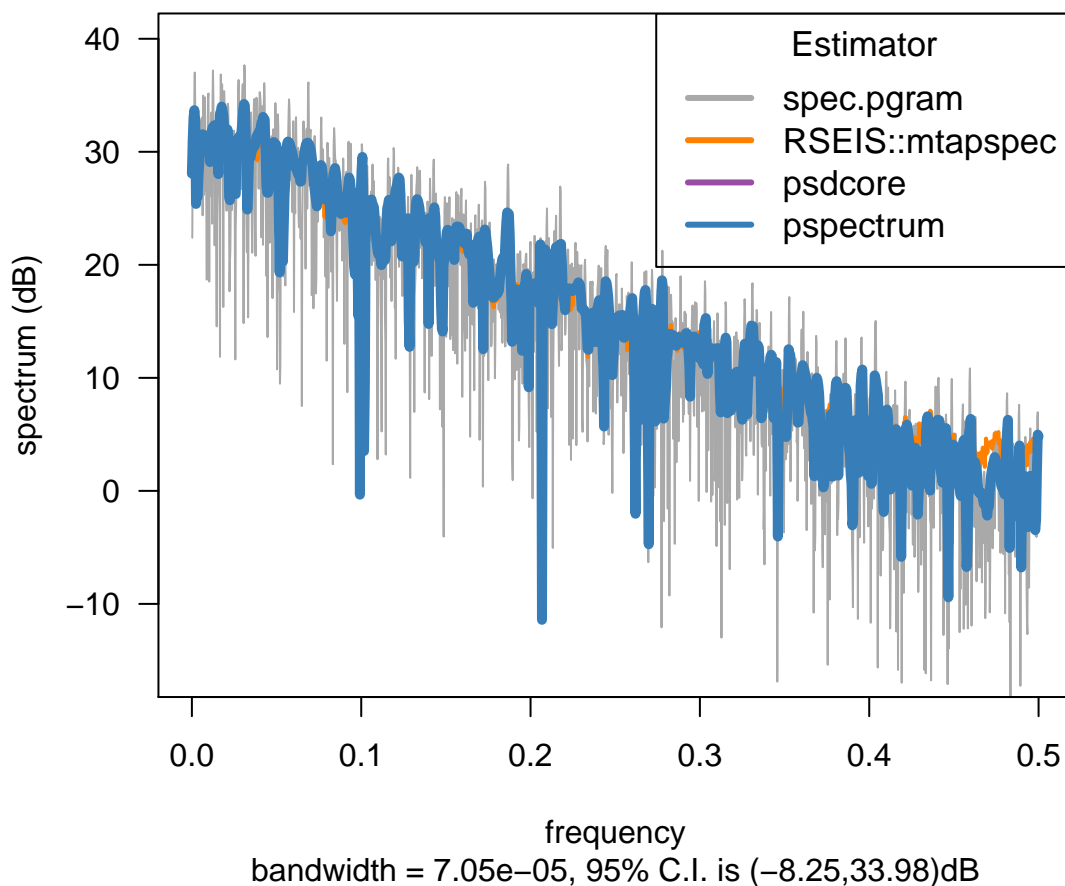


Figure 4: Comparisons of estimations of Project MAGNET power spectral densities.

```
library(ggplot2)
gr <- ggplot(df, aes(x = x, y = res)) + geom_abline(intercept = 0, slope = 0,
  size = 2, color = "salmon") + geom_point(aes(color = tap))
print(gr + theme_bw() + ggtitle("Regression residuals, colored by optimized tapers") +
  xlab("Power levels, dB") + ylab(""))
```

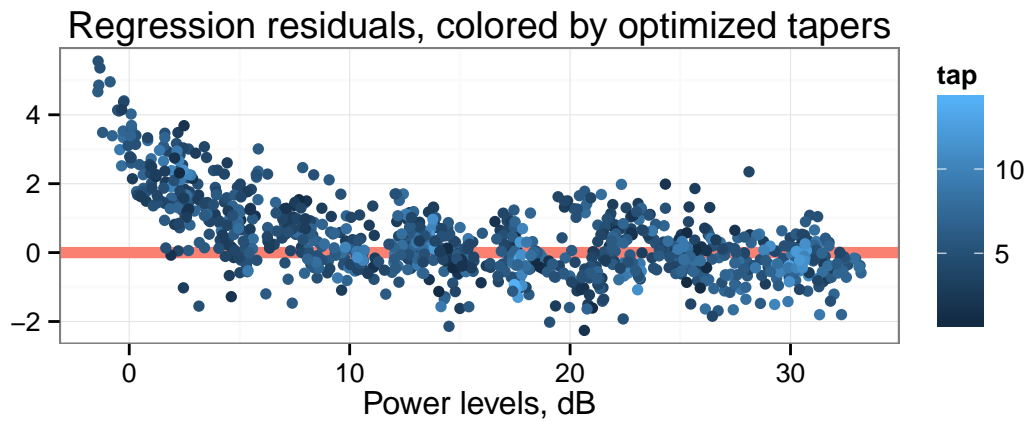


Figure 5: Linear regression residuals of `mtapspec` against `psdcore` for Project MAG-NET PSD estimates.

```
Bspec_plt <- plot(Bspec)
lines(Pspec$freq, Pspec$spec, col = "red", lwd = 2)
```

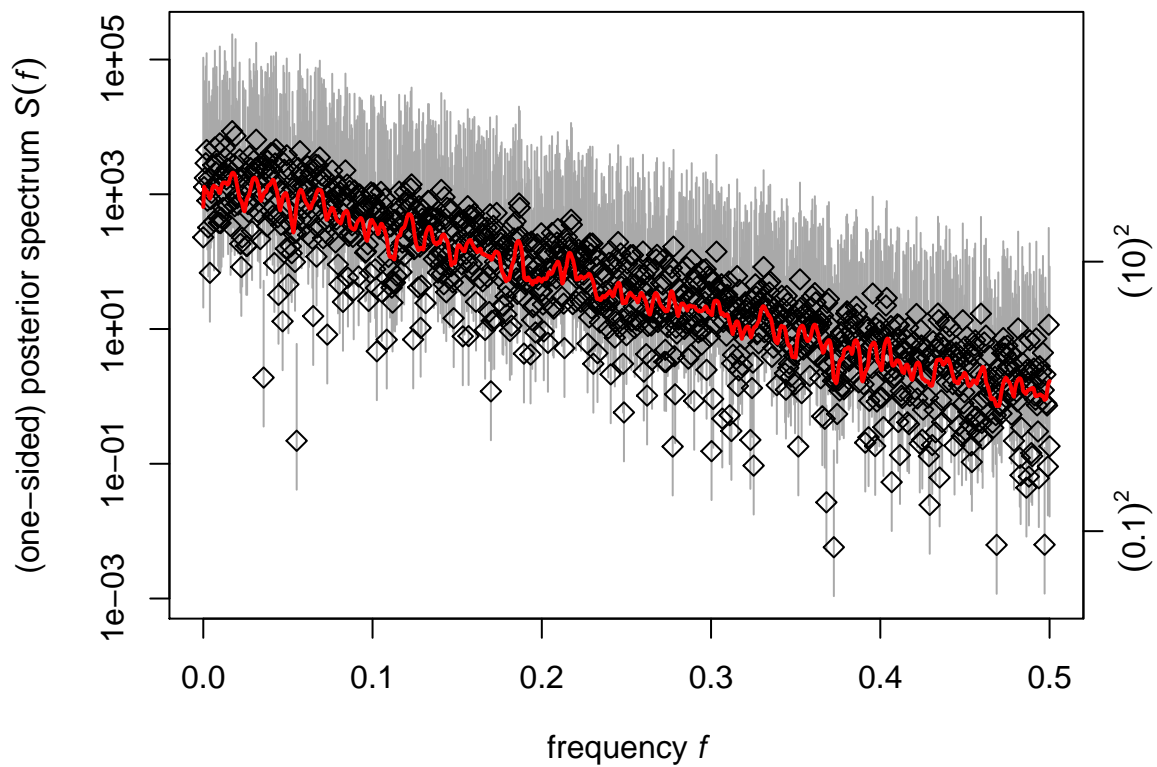


Figure 6: Project MAGNET PSD estimates from `bspec`, a Bayesian method, compared to the `psdcore` results shown in Figure 4.

```

pilot_spec(ats_lm, ntap = ntap, remove.AR = 100, plot = TRUE)
plot(Aspec, log = "dB", add = TRUE, col = "grey", lwd = 4)
plot(Aspec, log = "dB", add = TRUE, lwd = 3, lty = 3)
spec.ar(ats_lm, log = "dB", add = TRUE, lwd = 2, col = "grey40")

```

Pilot spectrum estimation (with AR(6) response)

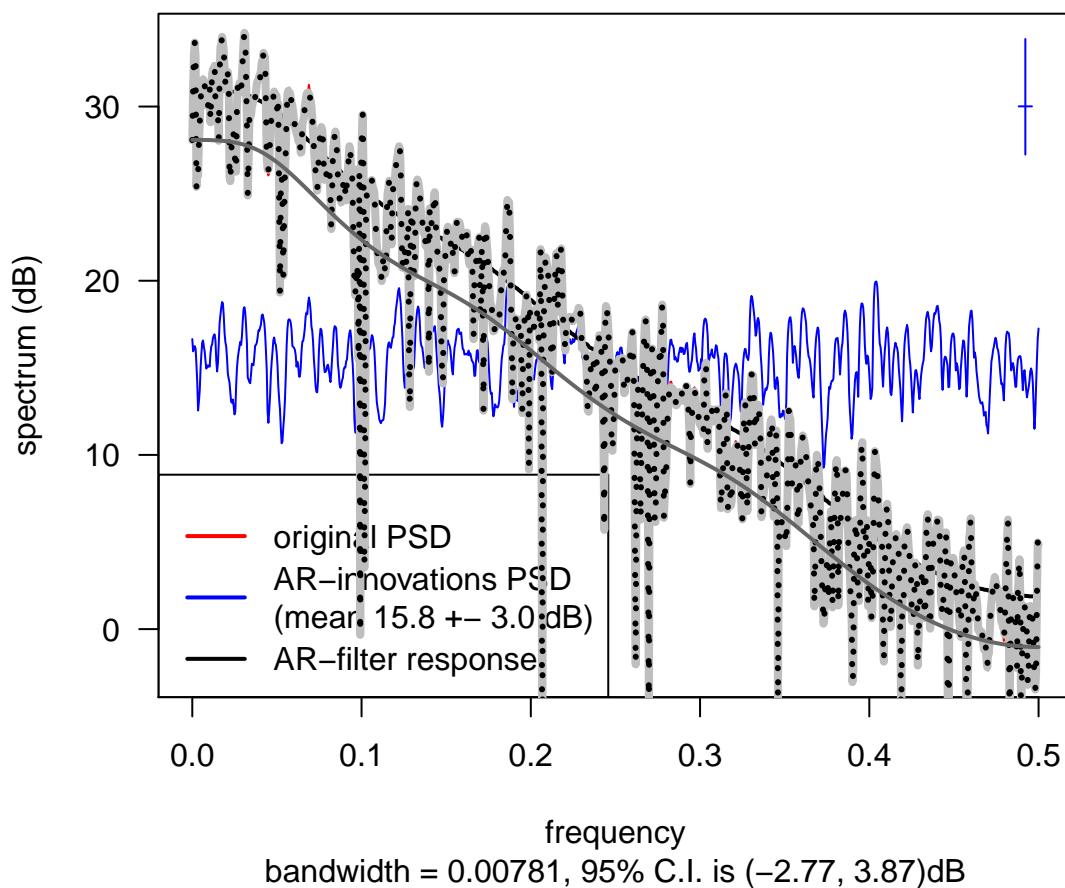


Figure 7: AR response spectrum for the MAGNET dataset produced by `pilot_spec`. Overlain on the figure is the adaptive estimation from Figure 4 (dotted line), and the results from `spec.ar` in dark grey; the shift is due to a normalization difference.

```

sp <- spectral_properties(as.tapers(1:50), p = 0.95, db.ci = TRUE)
par(las = 1)
plot(stderr.chi.upper ~ taper, sp, type = "s", ylim = c(-10, 20), yaxs = "i",
     xaxs = "i", xlab = expression("number of tapers (" * nu/2 * ")"),
     ylab = "dB", main = "Spectral uncertainties")
mtext("(additive factor)", line = 0.3)
lines(stderr.chi.lower ~ taper, sp, type = "s")
lines(stderr.chi.median ~ taper, sp, type = "s", lwd = 2)
lines(stderr.chi.approx ~ taper, sp, type = "s", col = "red", lwd = 2)
# to reach 3 db width confidence interval at p=.95
abline(v = 33, lty = 3)
legend("topright", c(expression("Based on " * chi^2 * "(p," * nu * ") and (1-p," *
  nu * ")"), expression(" " * chi^2 * "(p=0.5," * nu * ")"), "approximation"),
     lwd = c(1, 3, 3), col = c("black", "black", "red"), bg = "white")

```

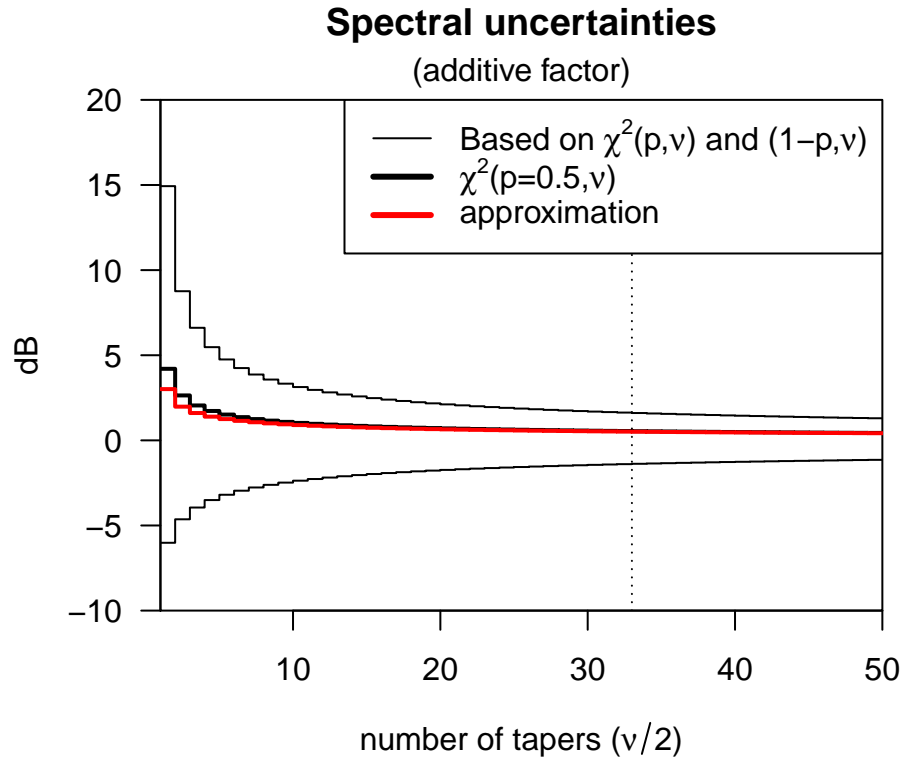


Figure 8: Additive spectral uncertainties by number of tapers needed to create 95% confidence intervals. These quantized curves are found by evaluating the χ^2_ν distribution, where ν is the number of degrees of freedom (two per taper). The thick, red line shows an approximation to these uncertainties based on $1/\sqrt{\nu-1}$, which is accurate to within a few percent in most cases. The vertical dotted-line shows the number of tapers need to make the width less than 3 decibels.

```

plot(c(0, 0.5), c(-5, 40), col = "white", main = "Project MAGNET Spectral Uncertainty (p > 0.95)",
     ylab = "", xlab = "spatial frequency, 1/km", yaxt = "n", frame.plot = FALSE)
lines(c(2, 1, 1, 2) * 0.01, c(0, 0, 7, 7))
text(0.04, 3.5, "7 dB")
polygon(pspb$xx, dB(pspb$yy), col = "light blue", border = NA)
text(0.26, 37, "Bayesian (bspec)", col = "#0099FF", cex = cx <- 0.9)
polygon(pspu$xx, pspu$yy, col = "dark grey", border = "black", lwd = 0.2)
text(0.15, 6, "Light: adaptive\ntaper refinement\n(pspectrum)", cex = cx)
polygon(pspau$xx, pspau$yy, col = "light grey", border = "black", lwd = 0.2)
text(0.4, 22, "Dark: Uniform\ntapering (psdcore)", cex = cx)

```

Project MAGNET Spectral Uncertainty (p > 0.95)

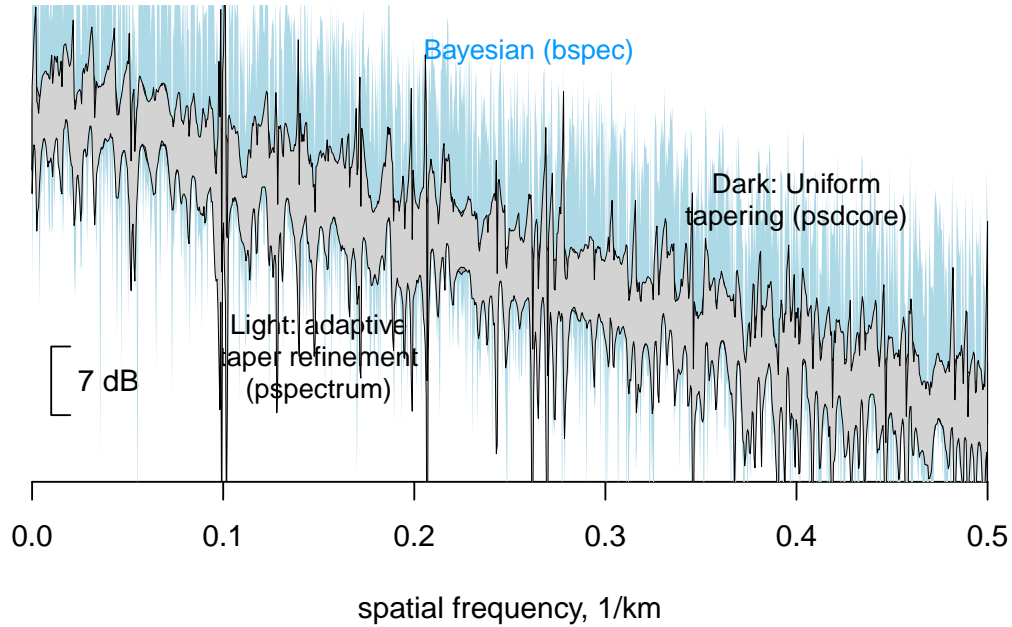


Figure 9: Project MAGNET spectral uncertainties for 95% coverage probability. The filled regions encompass the spectral uncertainties values based on the upper χ^2_L curve shown in Figure 8, light and dark for PSDs with and without adaptive taper optimization, respectively. The results from Figure 6 (Bayesian method) are shown in blue.

```

frq <- Aspec$freq
relp <- (spa$resolution - spp$resolution)/spp$resolution
par(las = 1, oma = rep(0, 4), omi = rep(0, 4), mar = c(4, 3, 2, 0))
layout(matrix(c(1, 2), 1, 2, byrow = TRUE), heights = c(2, 2), widths = c(3,
0.5), respect = TRUE)
plot(frq, relp, main = "Percent change in spectral resolution", col = "light grey",
ylim = yl <- range(pretty(relp)), type = "h", ylab = "dB", xlab = "frequency, 1/km")
lines(frq, relp)
text(0.25, 45, "Adaptive relative to fixed", cex = 0.9)
par(mar = c(4, 0, 2, 2))
# empirical distribution of values
boxplot(relp, range = 0, main = sprintf("%.01f", median(relp)), axes = FALSE,
ylim = yl, yaxs = "i", notch = TRUE)
axis(4)

```

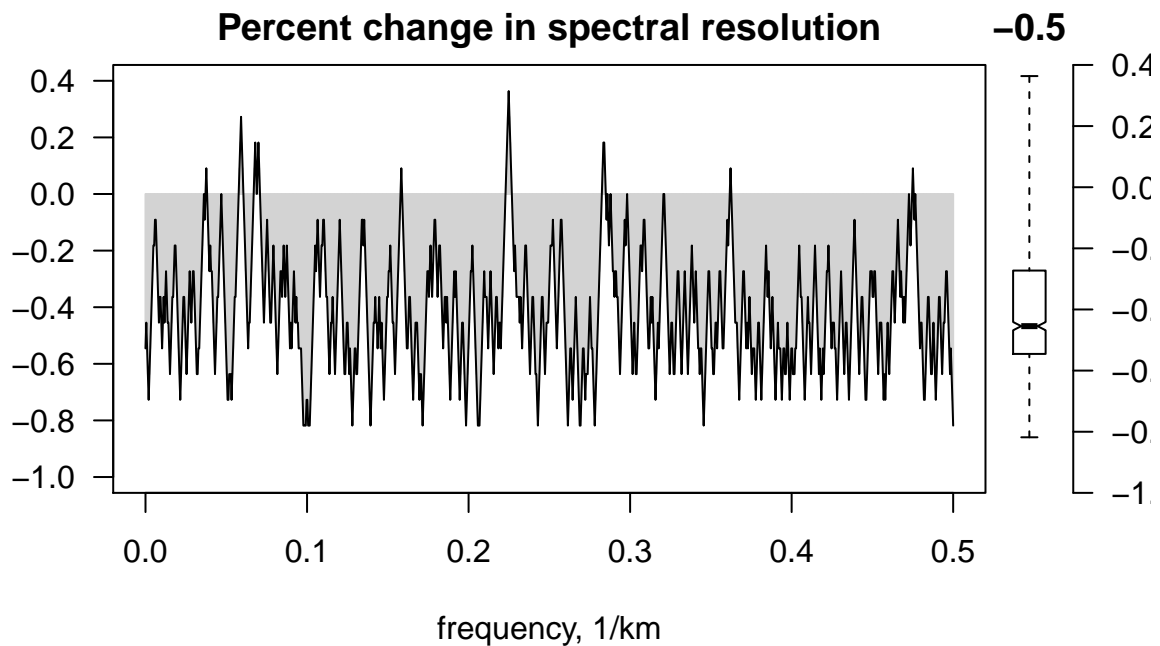


Figure 10: Relative changes in resolution of the adaptive method relative to the fixed multitaper method, plotted as a function of spatial frequency in units of percent. The non-zero median value implies the pilot spectrum was found using too-few tapers, according to the optimization algorithm. Positive values indicate broadening resolution bandwidth.

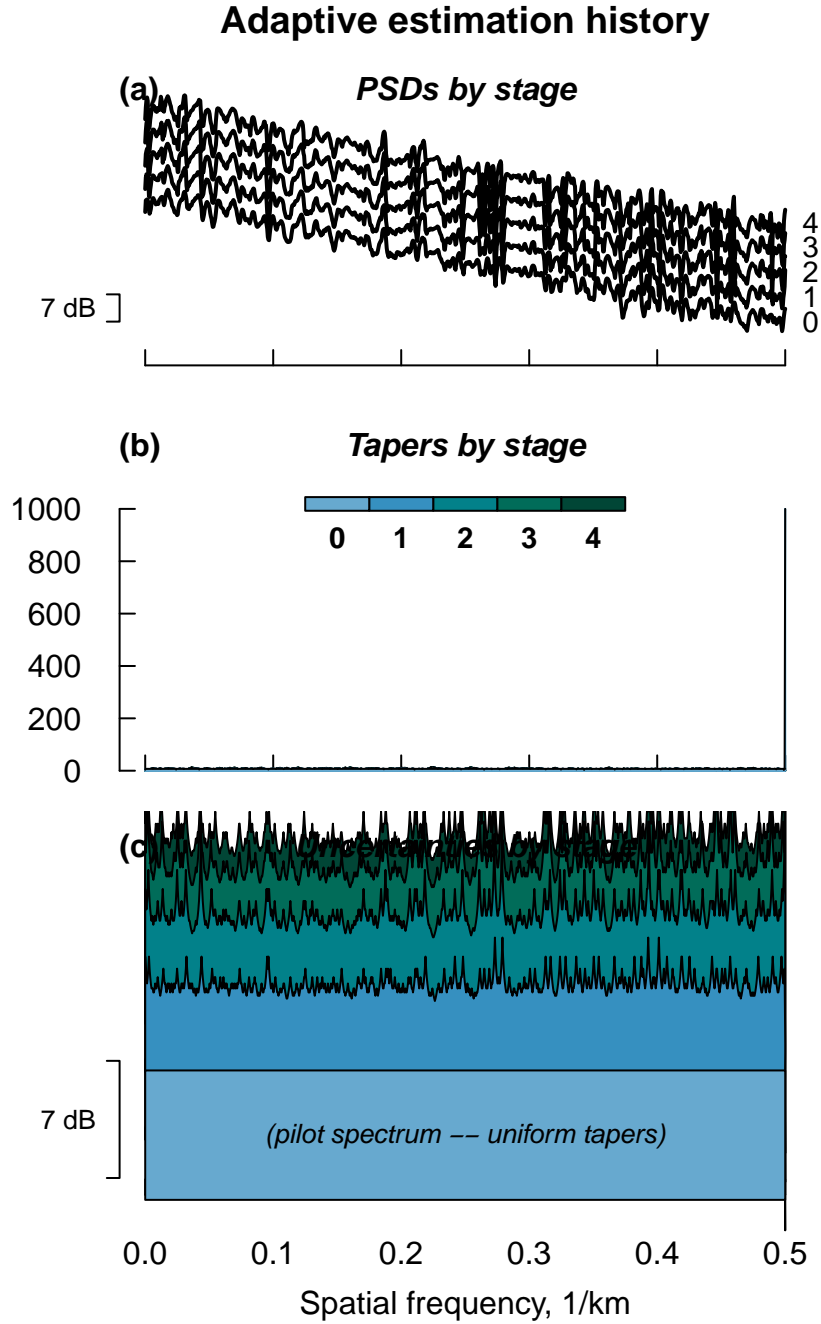


Figure 11: Adaptive spectral estimation history. (A) PSD series for each stage of the adaptive method, offset by a few decibels for visualization purposes. Filled polygons are shown in (B) for the number of tapers at each stage, and (C) the relative uncertainties of the PSDs.

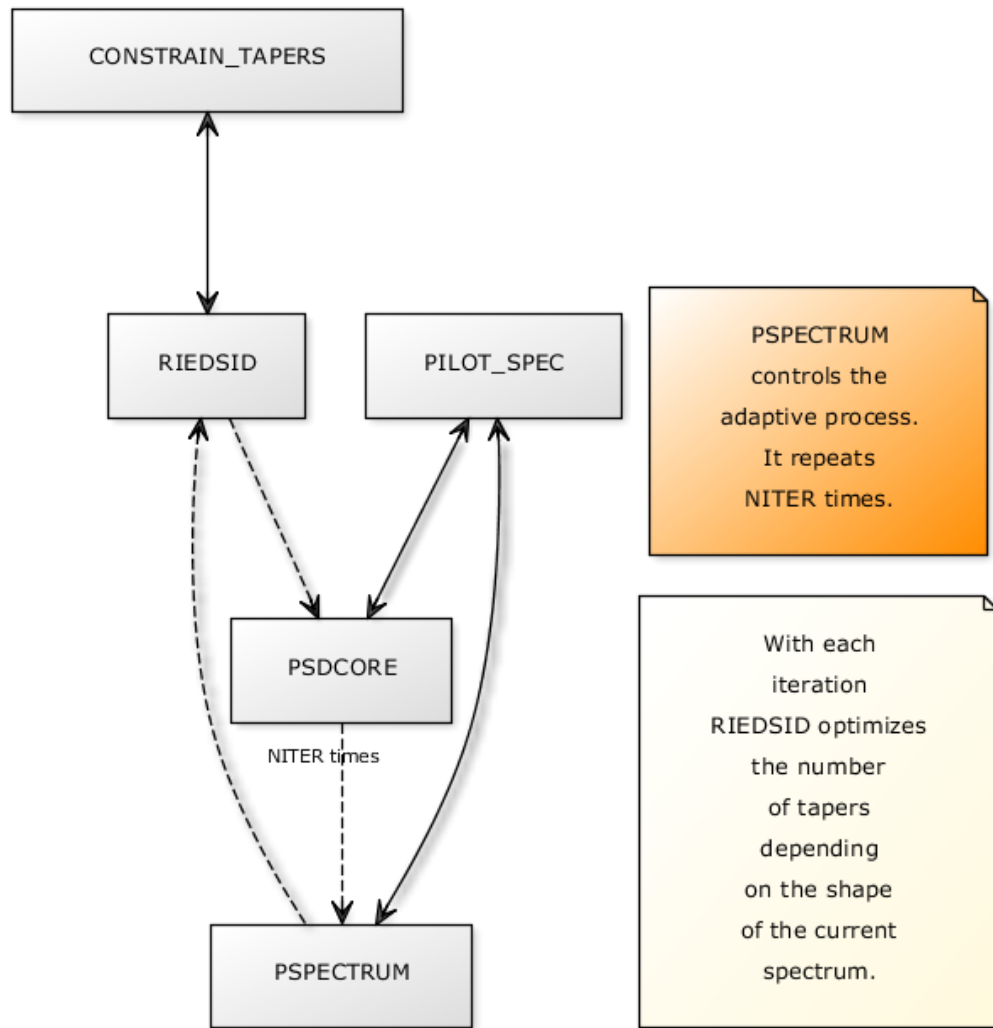


Figure 12: Simplified call graph for `psd`. The dashed lines show a simplified circuit which the spectra and its tapers make during the iterative process.