Normalization of Power Spectral Density

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Abstract

Normalizations come in many forms, but here we outline the normalization used by rlpSpec: power spectral density.

Contents

1	Bac	Background		
	1.1	stats::spectrum	1	
	1.2	multitaper::spec.mtm	3	
	1.3	SDF··sana	3	

1 Background

There can often be confusion about the different quantities used in spectral analysis ¹, partly due to myriad nomenclature within the incredibly vast literature on the subject. Regarding nomenclature: Phrases including "amplitude spectrum", "energy spectral density", "power", "power spectra", and even "spectra" all mean something, but are rarely equivalent.

Let us, for the sake of brevity, assume we have a stationary signal f(x) having Fourier Transform (FT) F(s), which is complex. The of this signal is simply the real component of the FT, $\arg\{F\}$, which we'll denote as ${}^{(A)}S$. The corresponding is thus $\operatorname{mod} F$, or ${}^{(\phi)}S$. These quantities are, however, meaningless.

The latter assumes that the spectrum has the normalization used in power – which is that the Nyquist is assumed to be 1,0, always (power doesn't ask for a sample interval). But psd assumes that the interval is 1 (unless you specify otherwise) so that the Nyquist is 0.5. So say you have a white noise with variance 1: power will return a flat spectrum with level 0 db (=1) but psd will return a level of 2 (=3 db), both so that the level times the Nyquist will be 1. But logsmoo, told that the interval is 1, will multiply the spectrum by 2 to covert from lc**2/Nyquist to lc**2/Hz: giving 3 dB for the spectrum from power, but 6 dB (=4) for the spectrum from psd.

First load the package into the namespace:

> library(rlpSpec)

1.1 stats::spectrum

Included in the core distribution of R is stats::spec.rum, which accesses stats::spec.ar or stats::spec.pgram for either parametric and non-parametric estimation, respectively. The user can optionally apply a single cosine taper, and/or a smoothing kernel. Our method is non-parametric; hence, we will compare to the latter.

¹ This post to R-help very eloquently describes the problem, and provides some guidance: http://r.789695.n4.nabble.com/Re-How-do-I-normalize-a-PSD-td792902.html

However, the logical arguments detrend and demean to psdcore are passed to spec.pgram; they are, by default, both TRUE.

As a matter of bookkeeping, we must deal with the working environment accessed by rlpSpec functions. Specifically, we should ensure psdcore does not access any inappropriate information by setting refresh=TRUE. We can then re-calculate the multitaper PSD and the raw periodogram with plotpsd=TRUE. The results are shown in Figure 1.1.

- > data(magsat)
- > psdcore(magsat\$clean, ntaper=10, refresh=TRUE, plotpsd=TRUE)

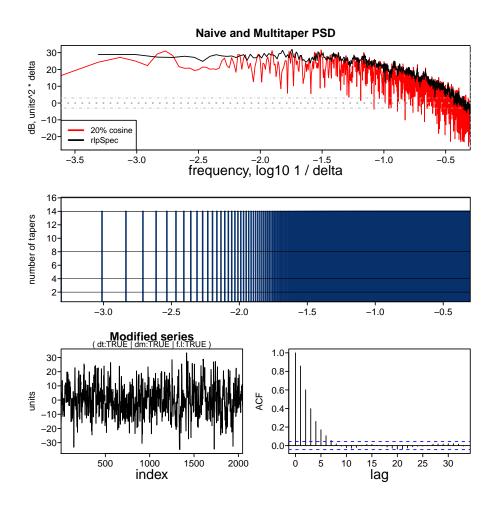


Figure 1: Top: Comparison between naïve and multitaper PSD estimators for the clean MAGSAT data. The frequency axis is in units of \log_{10} km⁻¹, and power axis is in decibels. Bottom: The spatial series used to estimate the PSDs.

- $1.2 \quad \texttt{multitaper::spec.mtm}$
- 1.3 SDF::sapa

\mathbf{Index}

amplitude spectrum, 1

phase spectrum, 1