

Analysis of Gravitational Wave Event GW150914

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PHYS 450

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1. Libraries

The code relies on `gwpy` for data handling, `scipy` for signal analysis, `numpy` for numerical calculations, and `matplotlib` for plotting.

```
1 import numpy as np
2 from numpy . fft import rfft , rfftfreq
3 import matplotlib.pyplot as plt
4 from scipy.signal import butter, filtfilt
5 from gwosc.datasets import event_gps
6 from gwpy.timeseries import TimeSeries
7 from gwpy.plot import Plot
8 from astropy.time import Time
```

2. Data Fetching

We first find the GPS time associated with the event.

```
1 gps = event_gps("GW150914")
```

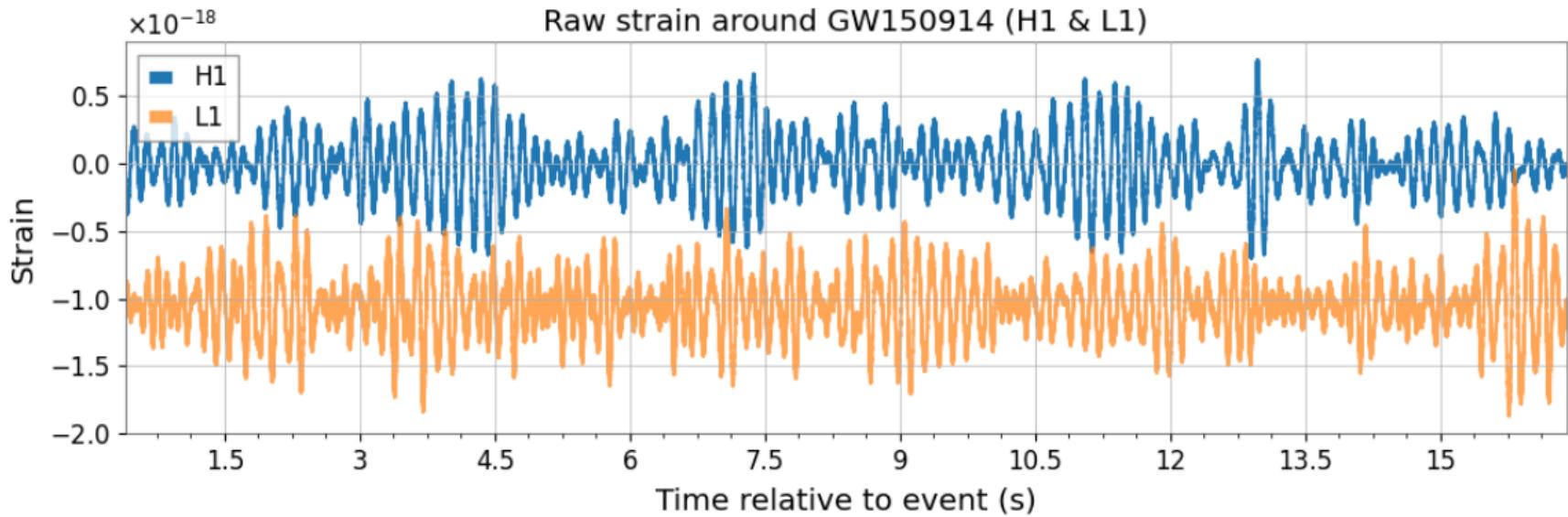
And set up the duration around it

```
1 duration = 16
2 start = gps - duration//2
3 end = gps + duration//2
```

We fetch open strain data from the Hanford (H1) and Livingston (L1) detectors.

```
1 h1 = TimeSeries.fetch_open_data('H1', start, end, cache=True)
2 l1 = TimeSeries.fetch_open_data('L1', start, end, cache=True)
```

3. Raw Signal



Observation: The signal is invisible. The amplitude is dominated by low-frequency seismic drift and instrumental noise.

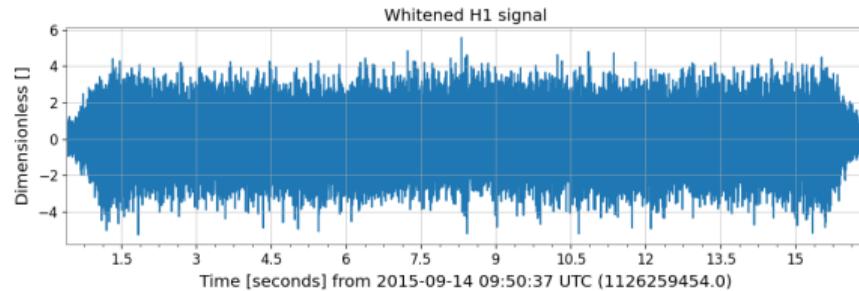
4. Spectral Whitening

To normalize the noise floor, we "whiten" the data.

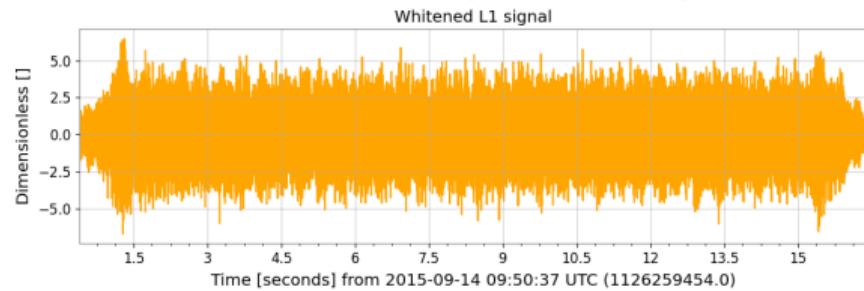
```
1 # Whiten the data using a 4-second FFT window with a 2 second overlap to
  estimate noise
2 h1_white = h1.whiten(fftlength=4, overlap=2)
3 l1_white = l1.whiten(fftlength=4, overlap=2)
```

5. Spectral Whitening (Output)

Whitened H1 (Hanford)



Whitened L1 (Livingston)



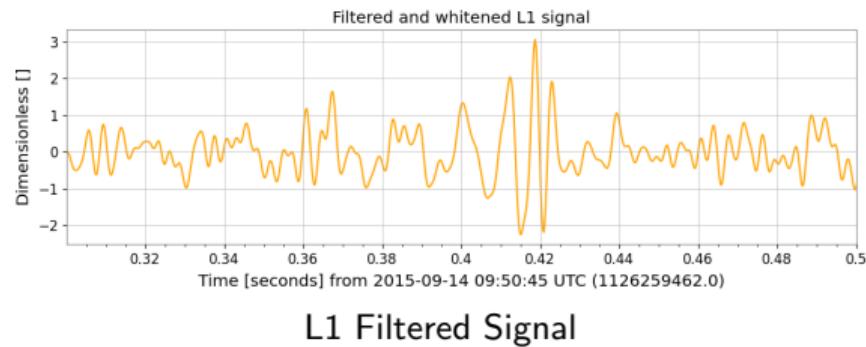
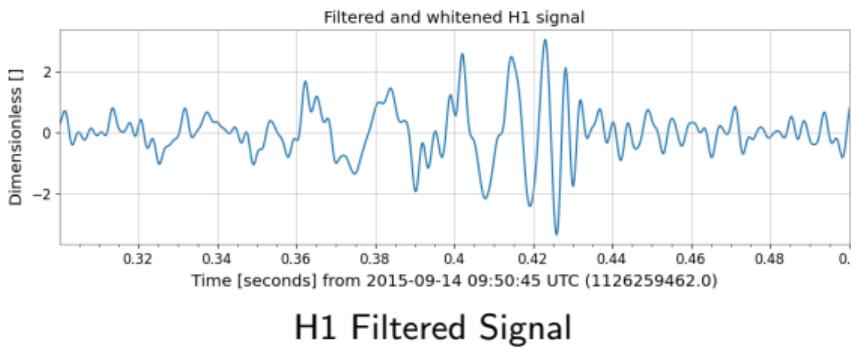
6. Bandpass Filtering

We design a 4th-order Butterworth filter to isolate the 35-350 Hz band.

```
1 def bandpass(data, f_s, f_l, f_h, order=4):
2     f_nyq = 0.5 * f_s # Nyquist Frequency
3     # Design filter (normalized by Nyquist)
4     b, a = butter(order, [f_l/f_nyq, f_h/f_nyq], btype='band')
5     return filtfilt(b, a, data)
6
7 # Apply to whitened data
8 h1_bp=bandpass(h1_white , f_s_h , 35 , 350 , 4)
9 l1_bp=bandpass(l1_white , f_s_l , 35 , 350 , 4)
```

Why filtfilt? It processes the data forward and backward, canceling out phase delays. This ensures the peak time corresponds exactly to the physical event.

7. Bandpass Filtering (Output)

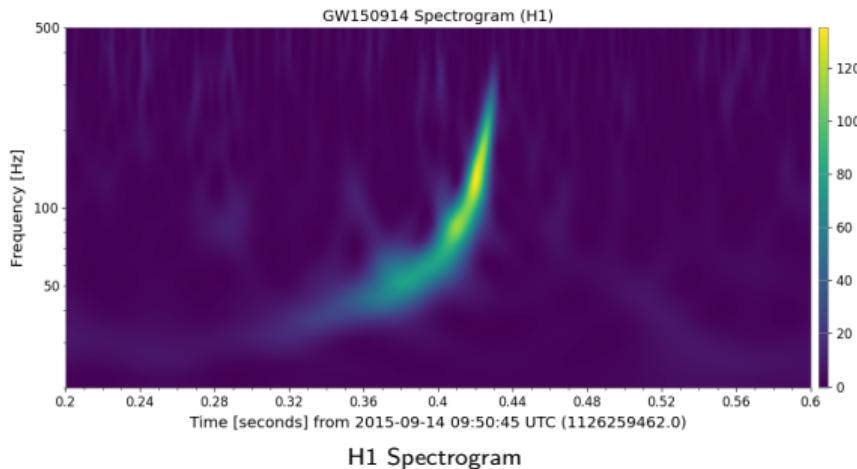


The characteristic "ringdown" of the black hole merger is now clearly visible in both detectors.

8. Time-Frequency Analysis of H

We use the Q-transform to visualize frequency evolution over time.

```
1 dt_zoom = 0.2
2 outseg = (gps - dt_zoom, gps + dt_zoom)
3
4 hq = h1_final.q_transform(outseg=outseg, qrange=(4, 8), frange=(20, 500))
```

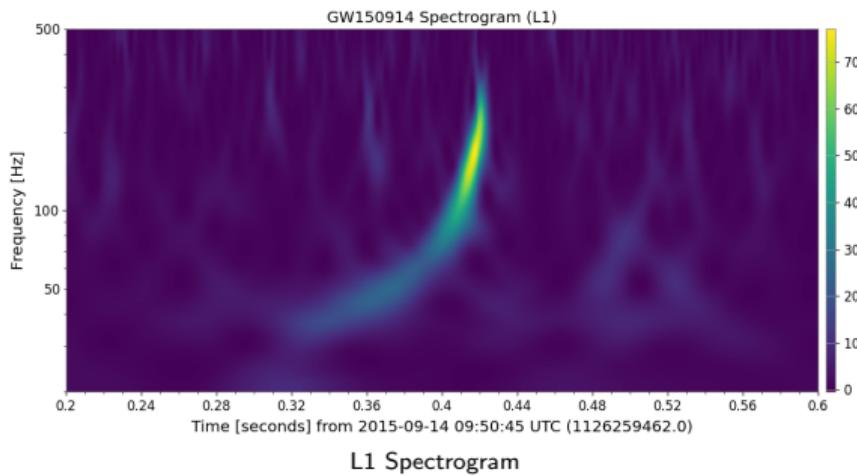


H1 Spectrogram

Physics: The "swoosh" shape (frequency rising with time) confirms compact binary collision.

9. Time-Frequency Analysis of L

```
1 lq = l1_final.q_transform(outseg=outseg, qrange=(4, 8), frange=(20, 500))
```



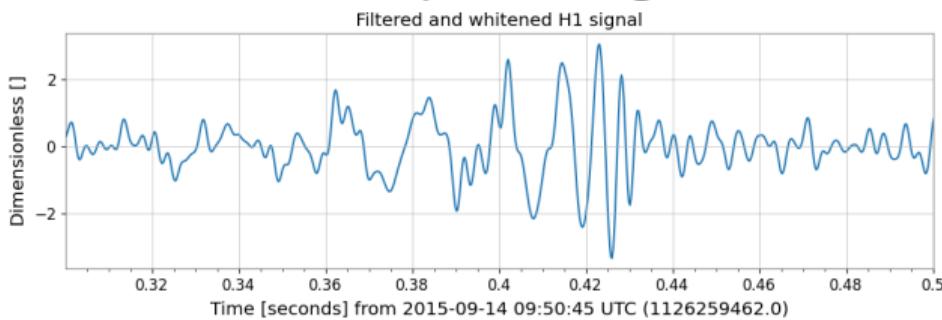
L1 Spectrogram

Physics: The "swoosh" shape (frequency rising with time) confirms compact binary collision.

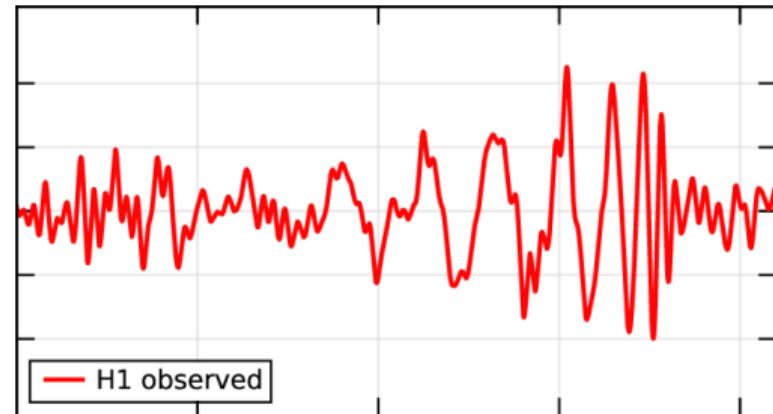
10. Validation vs Published Results

We compare our filtered H1 signal (left) with the official discovery figure (right).

Our Computation figure



LIGO published figure
Hanford, Washington (H1)



Conclusion

The analysis successfully extracts a waveform that matches the frequency, duration, and amplitude envelope of the official GW150914 discovery.