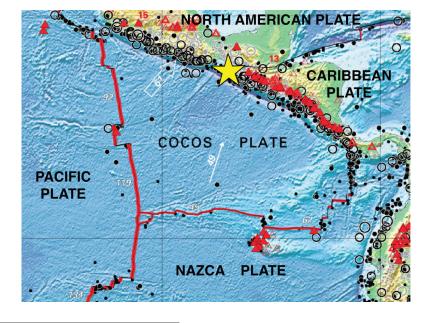
Infinite-dimensional Bayesian inversion for fault slip from surface measurements

Abhijit Chowdhary and Alen Alexanderian

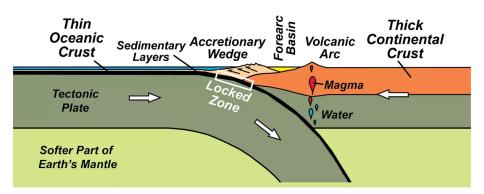
Department of Mathematics North Carolina State University

April 22, 2022

Work done through NSF-DMS-2111044



Tom Simkin et al. This dynamic planet: World map of volcanoes, earthquakes, impact craters and plate tectonics. Tech. rep. 2006. DOI: 10.3133/i2800



Robert J Lillie. *Oregon's Island In The Sky: Geology Road Guide to Marys Peak*. English. OCLC: 979996650. 2017. ISBN: 9781540611963

Goals

Understand the subduction zone from collected observations.

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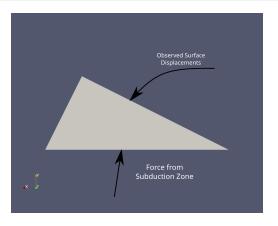
Understand the subduction zone from collected observations.

Do so while quantifying measurement uncertainties.

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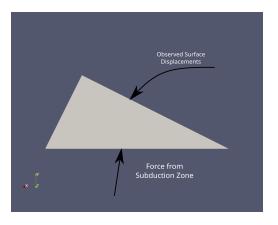
Model Assumptions



For the sake of modeling convienience, assume:

• Governing PDE (forward model): Linear elasticity

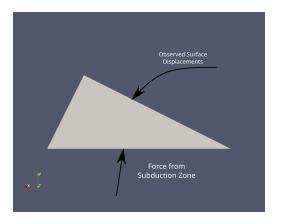
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Model Assumptions



For the sake of modeling convienience, assume:

- Governing PDE (forward model): Linear elasticity
- Uncertain parameter: Displacement on fault plane
- Inverse Problem: Given measurements of surface deformation u^{obs} reconstruct fault plane displacement.

$$-\nabla \cdot \boldsymbol{\sigma}(\boldsymbol{u}) = \mathbf{0} \text{ in } \Omega, \tag{1}$$

where:

- $\sigma(u) = \mathbb{C}\varepsilon(u)$ with
 - $\mathbb{C}[\varepsilon] = 2\mu\varepsilon + \lambda\operatorname{tr}(\varepsilon)$ I the fourth-order linear elasticity tensor:
 - ullet $egin{aligned} ullet$ $egin{aligned} ullet$ ullet ullet
- ullet μ and λ are known as the Láme constants.

Kimberly Alison McCormack. Earthquakes, groundwater and surface deformation: Exploring the poroelastic response to megathrust earthquakes. Aug. 2018. URL: hdl.handle.net/2152/68892

$$-\nabla \left[\mu(\nabla \boldsymbol{u} + (\nabla \boldsymbol{u})^{\mathrm{T}}) + \lambda \nabla \cdot \boldsymbol{u} \mathbf{I} \right] = \mathbf{0} \quad \text{in } \Omega,$$
 (2a)

$$\sigma(\mathbf{u})\mathbf{n} = \mathbf{0}$$
 on Γ_t (2b)

$$\mathbf{u} + \beta \mathbf{\sigma}(\mathbf{u})\mathbf{n} = \mathbf{h} \quad \text{on } \Gamma_{s}$$
 (2c)

$$\mathbf{u} \cdot \mathbf{n} = 0 \quad \text{on } \Gamma_b$$
 (2d)

$$\delta \mathbf{T}(\boldsymbol{\sigma}(\boldsymbol{u})\boldsymbol{n}) + \mathbf{T}\boldsymbol{u} = \boldsymbol{m} \quad \text{on } \Gamma_b$$
 (2e)

• **T** is the tangential operator $\mathbf{T} \boldsymbol{u} = (\mathbf{I} - \boldsymbol{n} \otimes \boldsymbol{n}) \boldsymbol{u} = \boldsymbol{u} - (\boldsymbol{n}^{\mathrm{T}} \boldsymbol{u}) \boldsymbol{n}$.

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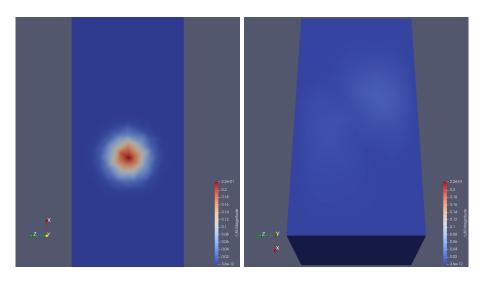
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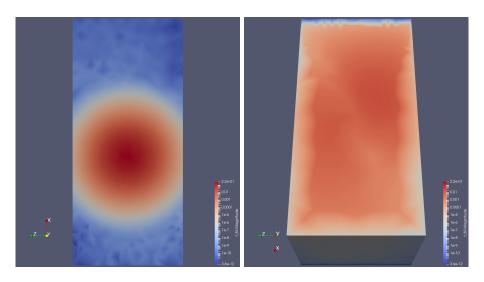
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- The right hand side of (2e) is the displacement on the fault plane that is being inverted for.
- (2e) can be understood as a regularized Dirichlet condition.

Forward Solution



Forward Solution



Weak Formulation

Define:

$$\mathbf{V} := \{ \mathbf{u} \in H^1(\Omega)^3 : \mathbf{u} \cdot \mathbf{n} = 0 \text{ on } \Gamma_b \}$$

Then the weak formulation of the forward model is given by:

$$\int_{\Gamma_{s}} \beta^{-1}(\boldsymbol{u} - \boldsymbol{h}) \cdot \boldsymbol{v} \, \mathrm{d}s + \int_{\Gamma_{b}} \delta^{-1}(\mathbf{T}\boldsymbol{u} - \boldsymbol{m}) \cdot \boldsymbol{v} \, \mathrm{d}s + \int_{\Omega} \varepsilon(\boldsymbol{u}) : \mathbb{C}[\varepsilon(\boldsymbol{v})] \, \mathrm{d}\boldsymbol{x} = 0, \quad \forall \boldsymbol{v} \in \boldsymbol{V} \quad (3)$$

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The Deterministic Inverse Problem

To reconstruct the fault displacement we construct the PDE-constrained optimization problem:

$$\mathcal{J}(\boldsymbol{\textit{m}}) = \frac{1}{2}\|\mathcal{B}\boldsymbol{\textit{u}}(\boldsymbol{\textit{m}}) - \boldsymbol{\textit{u}}^{\mathrm{obs}}\|^2 + \frac{1}{2}\|\mathcal{A}\boldsymbol{\textit{m}}\|^2$$

where \boldsymbol{u} is given by the solution of the linear elasticity equation.

- u(m) is given by the forward model.
- $\mathcal{B}:(L^2(\Omega))^3 \to \mathbb{R}^N$ is an observation operator.
- $\mathbf{u}^{\text{obs}} \in \mathbb{R}^N$ where N is the number of data points.

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If we let S be the forward model operator, i.e. u = Sm, and let F = BS, then:

$$\mathcal{J}(\mathbf{m}) = \frac{1}{2} \|\mathcal{F}\mathbf{m} - \mathbf{u}^{\text{obs}}\|^2 + \frac{1}{2} \|\mathcal{A}\mathbf{m}\|^2$$

¹In literature, called the *parameter-to-observable operator*.

Bayesian Inversion in Finite Dimensions

Theorem (Bayes Theorem in Finite Dimensions)

$$\pi_{
m post}(\mathbf{m}|\mathbf{u}^{
m obs}) \propto \pi_{
m like}(\mathbf{u}^{
m obs}|\mathbf{m})\pi_{
m prior}(\mathbf{m})$$

- Gaussian Prior $\mathbf{m} \sim \mathcal{N}(\mathbf{m}_{\mathrm{pr}}, \mathbf{\Gamma}_{\mathrm{pr}})$.
- Additive Gaussian noise

$$\mathbf{u}^{\mathrm{obs}} = \mathbf{Fm} + \boldsymbol{\eta}, \quad \boldsymbol{\eta} \sim \mathcal{N}(\mathbf{0}, \boldsymbol{\Gamma}_{\mathrm{noise}})$$

Posterior is therefore Gaussian with:

$$\boldsymbol{m}|\boldsymbol{u}^{\mathrm{obs}} \sim \mathcal{N}(\boldsymbol{m}_{\mathrm{post}}, \boldsymbol{\Gamma}_{\mathrm{post}})$$

where:

$$\begin{split} & m_{\mathrm{post}} = \Gamma_{\mathrm{post}} \left(\textbf{F}^{\mathrm{T}} \boldsymbol{\Gamma}_{\mathrm{noise}}^{-1} \textbf{u}^{\mathrm{obs}} + \boldsymbol{\Gamma}_{\mathrm{pr}}^{-1} \textbf{m}_{\mathrm{pr}} \right) \\ & \boldsymbol{\Gamma}_{\mathrm{post}} = \left(\textbf{F}^{\mathrm{T}} \boldsymbol{\Gamma}_{\mathrm{noise}}^{-1} \textbf{F} + \boldsymbol{\Gamma}_{\mathrm{pr}}^{-1} \right)^{-1} \end{split}$$

Corresponds to
$$\mathcal{J}(\mathbf{m}) = \frac{1}{2} \|\mathcal{F}\mathbf{m} - \mathbf{u}^{\mathrm{obs}}\|_{\mathbf{\Gamma}_{\mathrm{noise}}^{-1}}^2 + \frac{1}{2} \|\mathbf{m} - \mathbf{m}_{\mathrm{pr}}\|_{\mathbf{\Gamma}_{\mathrm{pr}}^{-1}}^2$$
.

Bayesian Inversion in Infinite Dimensions

Theorem (Bayes Theorem in Infinite Dimensions)

$$rac{\mathrm{d} \mu_\mathrm{post}^{oldsymbol{u}^\mathrm{obs}}}{\mathrm{d} \mu_\mathrm{pr}} \propto \pi_\mathrm{like}(oldsymbol{u}^\mathrm{obs} | oldsymbol{m})$$

- Gaussian Prior $\mathbf{m} \sim \mathcal{N}(\mathbf{m}_{\mathrm{pr}}, \mathcal{C}_0)$.
- Additive Gaussian noise

$$\mathbf{u}^{ ext{obs}} = \mathcal{F} m{m} + m{\eta}, \quad m{\eta} \sim \mathcal{N}(m{0}, m{\Gamma}_{ ext{noise}})$$

• The pair (m, u^{obs}) is jointly Gaussian with:

$$\mu_{\mathrm{post}}^{\mathbf{u}^{\mathrm{obs}}} = \mathcal{N}(\mathbf{m}_{\mathrm{post}}, \mathcal{C}_{\mathrm{post}})$$

where:

$$\begin{split} & \textbf{\textit{m}}_{\mathrm{post}} = \mathcal{C}_{\mathrm{post}} \left(\mathcal{F}^* \textbf{\textit{\Gamma}}_{\mathrm{noise}}^{-1} \textbf{\textit{u}}^{\mathrm{obs}} + \mathcal{C}_{\mathrm{pr}}^{-1} \textbf{\textit{m}}_{\mathrm{pr}} \right) \\ & \mathcal{C}_{\mathrm{post}} = \left(\mathcal{F}^* \textbf{\textit{\Gamma}}_{\mathrm{noise}}^{-1} \mathcal{F} + \mathcal{C}_{\mathrm{pr}}^{-1} \right)^{-1} \end{split}$$

Corresponds to $\mathcal{J}(\mathbf{m}) = \frac{1}{2} \|\mathcal{F}\mathbf{m} - \mathbf{u}^{\mathrm{obs}}\|_{\Gamma_{\mathrm{prior}}^{-1}}^2 + \frac{1}{2} \|\mathcal{C}^{-1/2}(\mathbf{m} - \mathbf{m}_{\mathrm{pr}})\|^2$.

Returning to Fault Inversion

Slight Wrinkle: Our parameter-to-observable map is affine.

Returning to Fault Inversion

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We want to minimize

$$\mathcal{J}(\mathbf{m}) = \frac{1}{2} \|\mathcal{B}\mathbf{u} - \mathbf{u}^{\text{obs}}\|_{\mathbf{\Gamma}^{-1}_{\text{noise}}}^2 + \frac{1}{2} \|\mathcal{C}^{-1/2}(\mathbf{m} - \mathbf{m}_{\text{pr}})\|^2$$

Returning to Fault Inversion

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We want to minimize

$$\mathcal{J}(\textbf{\textit{m}}) = \frac{1}{2}\|\mathcal{B}\textbf{\textit{u}} - \textbf{\textit{u}}^{\rm obs}\|_{\textbf{\textit{\Gamma}}_{\rm noise}^{-1}}^2 + \frac{1}{2}\|\mathcal{C}^{-1/2}(\textbf{\textit{m}} - \textbf{\textit{m}}_{\rm pr})\|^2$$

Define:

$$\begin{split} \mathcal{L}(\boldsymbol{m}, \boldsymbol{u}, \boldsymbol{v}) &= \frac{1}{2} \| \mathcal{B} \boldsymbol{u} - \boldsymbol{u}^{\mathrm{obs}} \|_{\boldsymbol{\Gamma}_{\mathrm{noise}}^{-1}}^2 + \frac{1}{2} \| \mathcal{C}^{-1/2} (\boldsymbol{m} - \boldsymbol{m}_{\mathrm{pr}}) \|^2 + \int_{\Gamma_s} \beta^{-1} (\boldsymbol{u} - \boldsymbol{h}) \cdot \boldsymbol{v} \, \mathrm{d}s \\ &+ \int_{\Gamma_b} \delta^{-1} (\boldsymbol{\mathsf{T}} \boldsymbol{u} - \boldsymbol{m}) \cdot \boldsymbol{v} \, \mathrm{d}s + \int_{\Omega} \varepsilon(\boldsymbol{u}) : \mathbb{C}[\varepsilon(\boldsymbol{v})] \, \mathrm{d}\boldsymbol{x} \end{split}$$

Procedure: Obtain m by enforcing $\mathcal{L}_m = \mathcal{L}_u = \mathcal{L}_v = 0$.

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FEM Discretization

- Consider finite element space $V_h = \operatorname{Span}(\phi_1, \dots, \phi_n)$.
- ϕ_k : Lagrange basis functions.
- For $\mathbf{f} \in \mathcal{V}_h$:

$$\mathbf{f} = \sum_{k=1}^{n} f_k \phi_k$$

In particular, construct $\mathbf{u}, \mathbf{v}, \mathbf{m}, \mathbf{m}_{\mathrm{pr}}, \mathbf{h} \in \mathcal{V}_h$ from $\mathbf{u}, \mathbf{v}, \mathbf{m}, \mathbf{m}_{\mathrm{pr}}$ and \mathbf{h} .

• Finite dimensional Hilbert space

$$(\mathcal{V}_h, \langle \cdot, \cdot \rangle_{L^2}) \cong (\mathbb{R}^n, \langle \cdot, \cdot \rangle_M), \qquad \langle \mathbf{u}, \mathbf{v} \rangle = \mathbf{u}^{\mathrm{T}} \mathbf{M} \mathbf{v}$$

• From the Lagrangian we find

$$\begin{bmatrix} \mathbf{C}_{\mathrm{pr}}^{-1} & & -\mathbf{C}^* \\ & \mathbf{B}^*\boldsymbol{\Gamma}_{\mathrm{noise}}^{-1}\mathbf{B} & \mathbf{A}^* \\ -\mathbf{C} & \mathbf{A} & \end{bmatrix} \begin{bmatrix} \mathbf{m} \\ \mathbf{u} \\ \mathbf{v} \end{bmatrix} = \begin{bmatrix} \mathbf{C}_{\mathrm{pr}}^{-1}\mathbf{m}_0 \\ \mathbf{B}^*\boldsymbol{\Gamma}_{\mathrm{noise}}^{-1}\mathbf{u}^{\mathrm{obs}} \\ \mathbf{D}\mathbf{h} \end{bmatrix}$$

where Au = Cm + Dh is the discretized weak form.

Reduced Linear System

Parameter-to-Observable Map

From the discretized weak form we can define:

$$Au = Cm + Dh \implies \mathcal{F}(m) = BA^{-1}[Cm + Dh] = \mathcal{G}m + g$$

Reduced Linear System

Parameter-to-Observable Map

From the discretized weak form we can define:

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With this:

$$\begin{bmatrix} \textbf{C}_{\mathrm{pr}}^{-1} & & -\textbf{C}^* \\ & \textbf{B}^* \textbf{\Gamma}_{\mathrm{noise}}^{-1} \textbf{B} & \textbf{A}^* \\ -\textbf{C} & \textbf{A} & \end{bmatrix} \begin{bmatrix} \textbf{m} \\ \textbf{u} \\ \textbf{v} \end{bmatrix} = \begin{bmatrix} \textbf{C}_{\mathrm{pr}}^{-1} \textbf{m}_0 \\ \textbf{B}^* \textbf{\Gamma}_{\mathrm{noise}}^{-1} \textbf{u}^{\mathrm{obs}} \\ \textbf{Dh} \end{bmatrix}$$

can be reduced to

$$\left(\boldsymbol{\mathsf{C}}_{\mathrm{pr}}^{-1} + \mathcal{G}^*\boldsymbol{\mathsf{\Gamma}}_{\mathrm{noise}}^{-1}\mathcal{G}\right)\boldsymbol{\mathsf{m}} = \mathcal{G}^*\boldsymbol{\mathsf{\Gamma}}_{\mathrm{noise}}^{-1}\left(\boldsymbol{\mathsf{g}} + \boldsymbol{\mathsf{u}}^{\mathrm{obs}}\right) + \boldsymbol{\mathsf{C}}_{\mathrm{pr}}^{-1}\boldsymbol{\mathsf{m}}_0$$

Tractability of Computation

$$\underbrace{\left[\boldsymbol{C}_{\mathrm{pr}}^{-1} + \mathcal{G}^* \boldsymbol{\Gamma}_{\mathrm{noise}}^{-1} \mathcal{G} \right]}_{\boldsymbol{C}_{\mathrm{post}}^{-1}} \boldsymbol{m}_{\mathrm{post}} = \left[\mathcal{G}^* \boldsymbol{\Gamma}_{\mathrm{noise}}^{-1} \left(\boldsymbol{g} + \boldsymbol{u}^{\mathrm{obs}} \right) + \boldsymbol{C}_{\mathrm{pr}}^{-1} \boldsymbol{m}_0 \right]$$

How do we make this computation tractable?

Tractability of Computation

$$\underbrace{\left[\boldsymbol{C}_{\mathrm{pr}}^{-1} + \mathcal{G}^* \boldsymbol{\Gamma}_{\mathrm{noise}}^{-1} \mathcal{G} \right]}_{\boldsymbol{C}_{\mathrm{post}}^{-1}} \boldsymbol{m}_{\mathrm{post}} = \left[\mathcal{G}^* \boldsymbol{\Gamma}_{\mathrm{noise}}^{-1} \left(\boldsymbol{g} + \boldsymbol{u}^{\mathrm{obs}} \right) + \boldsymbol{C}_{\mathrm{pr}}^{-1} \boldsymbol{m}_0 \right]$$

How do we make this computation tractable?

- Exploit the dimensionality reduction in the parameter space
- Use symmetry aware solvers
- Leverage sparsity of
 - matrices induced by differential operators
 - observations

Parameter space dimensionality reduction

Symmetry: Adjoints want to kill you

Sparsity: Matrices induced by differential operators

Sparsity: Observations

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