

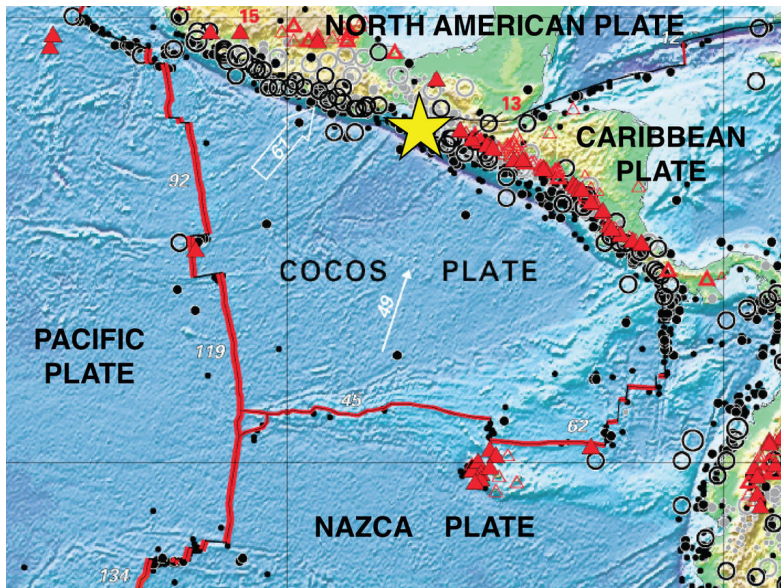
# Infinite-dimensional Bayesian inversion for fault slip from surface measurements

Abhijit Chowdhary and Alen Alexanderian

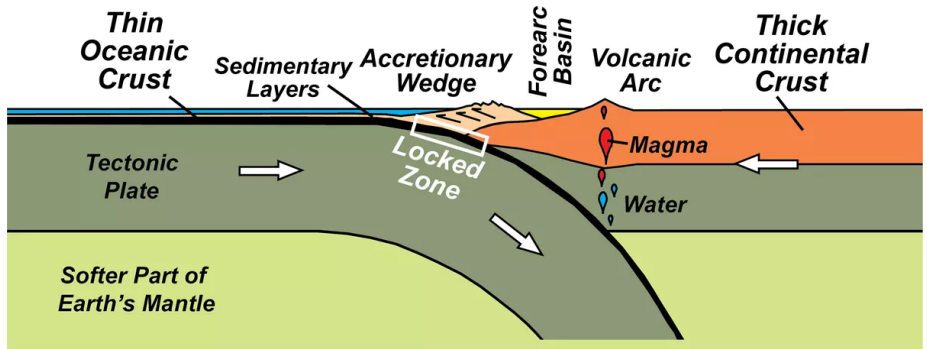
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North Carolina State University

April 22, 2022

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Tom Simkin et al. *This dynamic planet: World map of volcanoes, earthquakes, impact craters and plate tectonics*. Tech. rep. 2006. DOI: [10.3133/i2800](https://doi.org/10.3133/i2800)



Robert J Lillie. *Oregon's Island In The Sky: Geology Road Guide to Marys Peak*. [English](#).  
 OCLC: 979996650. 2017. ISBN: 9781540611963

Understand the subduction zone from collected observations.

# Goals

Understand the subduction zone from collected observations.

Do so while quantifying measurement uncertainties.

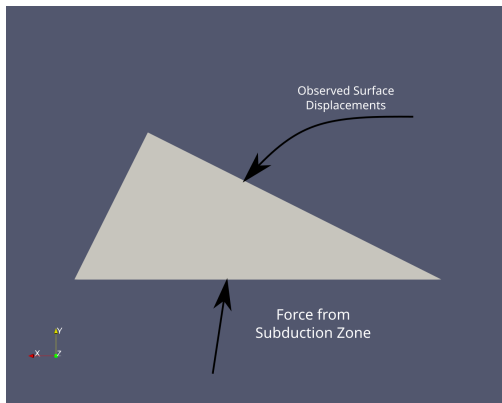
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# Model Assumptions

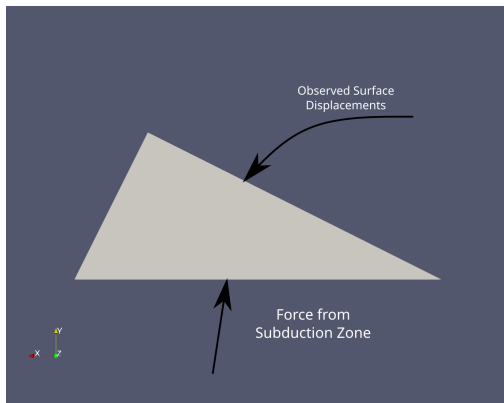


For the sake of modeling convenience, assume:

- **Governing PDE** (forward model): Linear elasticity



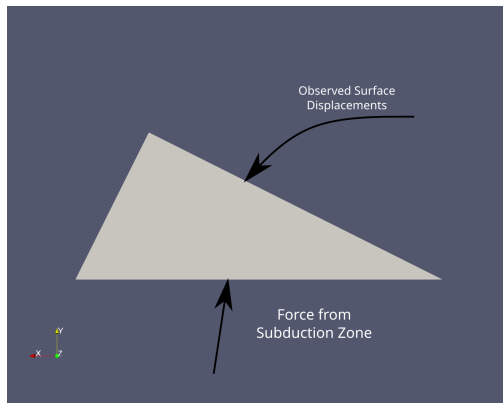
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For the sake of modeling convenience, assume:

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- **Uncertain parameter**: Displacement on fault plane

# Model Assumptions



For the sake of modeling convenience, assume:

- **Governing PDE** (forward model): Linear elasticity
- **Uncertain parameter**: Displacement on fault plane
- **Inverse Problem**: Given measurements of surface deformation  $\mathbf{u}^{\text{obs}}$  reconstruct fault plane displacement.

# Forward Model

$$-\nabla \cdot \boldsymbol{\sigma}(\mathbf{u}) = \mathbf{0} \text{ in } \Omega, \quad (1)$$

where:

- $\boldsymbol{\sigma}(\mathbf{u}) = \mathbb{C}\boldsymbol{\varepsilon}(\mathbf{u})$  with
  - $\mathbb{C}[\boldsymbol{\varepsilon}] = 2\mu\boldsymbol{\varepsilon} + \lambda \text{tr}(\boldsymbol{\varepsilon})\mathbf{I}$  the fourth-order linear elasticity tensor:
  - $\boldsymbol{\varepsilon}(\mathbf{u}) = \frac{1}{2} \left[ \nabla \mathbf{u} + (\nabla \mathbf{u})^T \right]$  the strain tensor.
- $\mu$  and  $\lambda$  are known as the Lame constants.

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Kimberly Alison McCormack. *Earthquakes, groundwater and surface deformation : Exploring the poroelastic response to megathrust earthquakes*. Aug. 2018. URL: [hdl.handle.net/2152/68892](https://hdl.handle.net/2152/68892)

# Forward Model

$$-\nabla \left[ \mu(\nabla \mathbf{u} + (\nabla \mathbf{u})^T) + \lambda \nabla \cdot \mathbf{u} \mathbf{I} \right] = \mathbf{0} \quad \text{in } \Omega, \quad (2a)$$

$$\boldsymbol{\sigma}(\mathbf{u})\mathbf{n} = \mathbf{0} \quad \text{on } \Gamma_t \quad (2b)$$

$$\mathbf{u} + \beta \boldsymbol{\sigma}(\mathbf{u})\mathbf{n} = \mathbf{h} \quad \text{on } \Gamma_s \quad (2c)$$

$$\mathbf{u} \cdot \mathbf{n} = 0 \quad \text{on } \Gamma_b \quad (2d)$$

$$\delta \mathbf{T}(\boldsymbol{\sigma}(\mathbf{u})\mathbf{n}) + \mathbf{T}\mathbf{u} = \mathbf{m} \quad \text{on } \Gamma_b \quad (2e)$$

- $\mathbf{T}$  is the tangential operator  $\mathbf{T}\mathbf{u} = (\mathbf{I} - \mathbf{n} \otimes \mathbf{n})\mathbf{u} = \mathbf{u} - (\mathbf{n}^T \mathbf{u})\mathbf{n}$ .

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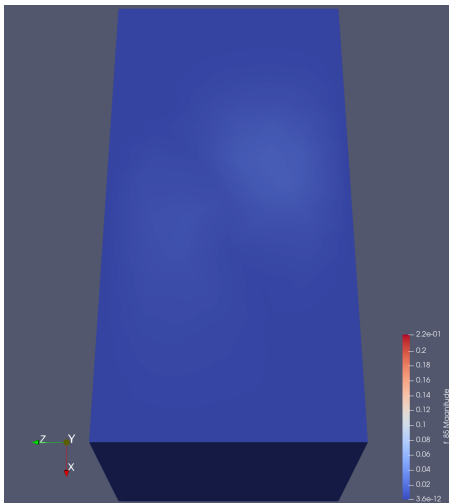
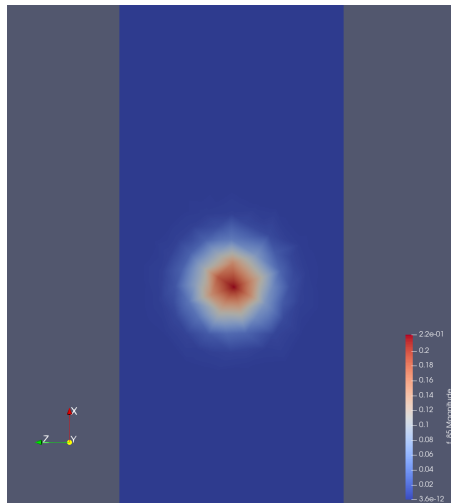
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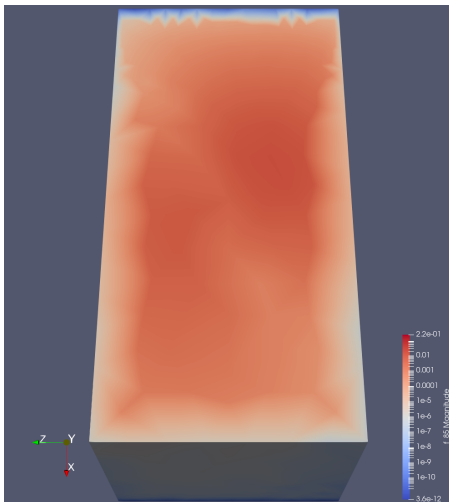
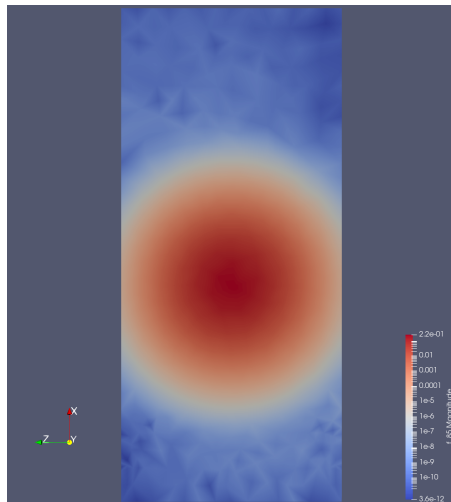
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- The right hand side of (2e) is the displacement on the fault plane that is being inverted for.
- (2e) can be understood as a regularized Dirichlet condition.

# Forward Solution



# Forward Solution





# Weak Formulation

Define:

$$\mathbf{V} := \{\mathbf{u} \in H^1(\Omega)^3 : \mathbf{u} \cdot \mathbf{n} = 0 \text{ on } \Gamma_b\}$$

Then the weak formulation of the forward model is given by:

$$\begin{aligned} \int_{\Gamma_s} \beta^{-1}(\mathbf{u} - \mathbf{h}) \cdot \mathbf{v} \, ds + \int_{\Gamma_b} \delta^{-1}(\mathbf{T}\mathbf{u} - \mathbf{m}) \cdot \mathbf{v} \, ds \\ + \int_{\Omega} \boldsymbol{\varepsilon}(\mathbf{u}) : \mathbb{C}[\boldsymbol{\varepsilon}(\mathbf{v})] \, d\mathbf{x} = 0, \quad \forall \mathbf{v} \in \mathbf{V} \quad (3) \end{aligned}$$

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# The Deterministic Inverse Problem

To reconstruct the fault displacement we construct the PDE-constrained optimization problem:

$$\mathcal{J}(\mathbf{m}) = \frac{1}{2} \|\mathcal{B}\mathbf{u}(\mathbf{m}) - \mathbf{u}^{\text{obs}}\|^2 + \frac{1}{2} \|\mathcal{A}\mathbf{m}\|^2$$

where  $\mathbf{u}$  is given by the solution of the linear elasticity equation.

- $\mathbf{u}(\mathbf{m})$  is given by the forward model.
- $\mathcal{B} : (L^2(\Omega))^3 \rightarrow \mathbb{R}^N$  is an observation operator.
- $\mathbf{u}^{\text{obs}} \in \mathbb{R}^N$  where  $N$  is the number of data points.

---

<sup>1</sup>In literature, called the *parameter-to-observable operator*.

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If we let  $\mathcal{S}$  be the forward model operator, i.e.  $\mathbf{u} = \mathcal{S}\mathbf{m}$ , and let<sup>1</sup>  $\mathcal{F} = \mathcal{B}\mathcal{S}$ , then:

$$\mathcal{J}(\mathbf{m}) = \frac{1}{2} \|\mathcal{F}\mathbf{m} - \mathbf{u}^{\text{obs}}\|^2 + \frac{1}{2} \|\mathcal{A}\mathbf{m}\|^2$$

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# Bayesian Inversion in Finite Dimensions

## Theorem (Bayes Theorem in Finite Dimensions)

$$\pi_{\text{post}}(\mathbf{m}|\mathbf{u}^{\text{obs}}) \propto \pi_{\text{like}}(\mathbf{u}^{\text{obs}}|\mathbf{m})\pi_{\text{prior}}(\mathbf{m})$$

- Gaussian Prior  $\mathbf{m} \sim \mathcal{N}(\mathbf{m}_{\text{pr}}, \mathbf{\Gamma}_{\text{pr}})$ .
- Additive Gaussian noise

$$\mathbf{u}^{\text{obs}} = \mathbf{F}\mathbf{m} + \boldsymbol{\eta}, \quad \boldsymbol{\eta} \sim \mathcal{N}(\mathbf{0}, \mathbf{\Gamma}_{\text{noise}})$$

- Posterior is therefore Gaussian with:

$$\mathbf{m}|\mathbf{u}^{\text{obs}} \sim \mathcal{N}(\mathbf{m}_{\text{post}}, \mathbf{\Gamma}_{\text{post}})$$

where:

$$\begin{aligned}\mathbf{m}_{\text{post}} &= \mathbf{\Gamma}_{\text{post}} \left( \mathbf{F}^T \mathbf{\Gamma}_{\text{noise}}^{-1} \mathbf{u}^{\text{obs}} + \mathbf{\Gamma}_{\text{pr}}^{-1} \mathbf{m}_{\text{pr}} \right) \\ \mathbf{\Gamma}_{\text{post}} &= \left( \mathbf{F}^T \mathbf{\Gamma}_{\text{noise}}^{-1} \mathbf{F} + \mathbf{\Gamma}_{\text{pr}}^{-1} \right)^{-1}\end{aligned}$$

Corresponds to  $\mathcal{J}(\mathbf{m}) = \frac{1}{2} \|\mathcal{F}\mathbf{m} - \mathbf{u}^{\text{obs}}\|_{\mathbf{\Gamma}_{\text{noise}}^{-1}}^2 + \frac{1}{2} \|\mathbf{m} - \mathbf{m}_{\text{pr}}\|_{\mathbf{\Gamma}_{\text{pr}}^{-1}}^2$ .

# Bayesian Inversion in Infinite Dimensions

## Theorem (Bayes Theorem in Infinite Dimensions)

$$\frac{d\mu_{\text{post}}^{\mathbf{u}^{\text{obs}}}}{d\mu_{\text{pr}}} \propto \pi_{\text{like}}(\mathbf{u}^{\text{obs}} | \mathbf{m})$$

- Gaussian Prior  $\mathbf{m} \sim \mathcal{N}(\mathbf{m}_{\text{pr}}, \mathcal{C}_0)$ .
- Additive Gaussian noise

$$\mathbf{u}^{\text{obs}} = \mathcal{F}\mathbf{m} + \boldsymbol{\eta}, \quad \boldsymbol{\eta} \sim \mathcal{N}(\mathbf{0}, \boldsymbol{\Gamma}_{\text{noise}})$$

- The pair  $(\mathbf{m}, \mathbf{u}^{\text{obs}})$  is jointly Gaussian with:

$$\mu_{\text{post}}^{\mathbf{u}^{\text{obs}}} = \mathcal{N}(\mathbf{m}_{\text{post}}, \mathcal{C}_{\text{post}})$$

where:

$$\begin{aligned}\mathbf{m}_{\text{post}} &= \mathcal{C}_{\text{post}} \left( \mathcal{F}^* \boldsymbol{\Gamma}_{\text{noise}}^{-1} \mathbf{u}^{\text{obs}} + \mathcal{C}_{\text{pr}}^{-1} \mathbf{m}_{\text{pr}} \right) \\ \mathcal{C}_{\text{post}} &= \left( \mathcal{F}^* \boldsymbol{\Gamma}_{\text{noise}}^{-1} \mathcal{F} + \mathcal{C}_{\text{pr}}^{-1} \right)^{-1}\end{aligned}$$

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# Returning to Fault Inversion

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We want to minimize

$$\mathcal{J}(\mathbf{m}) = \frac{1}{2} \|\mathcal{B}\mathbf{u} - \mathbf{u}^{\text{obs}}\|_{\mathbf{r}_{\text{noise}}^{-1}}^2 + \frac{1}{2} \|\mathcal{C}^{-1/2}(\mathbf{m} - \mathbf{m}_{\text{pr}})\|^2$$



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Define:

$$\begin{aligned} \mathcal{L}(\mathbf{m}, \mathbf{u}, \mathbf{v}) = & \frac{1}{2} \|\mathcal{B}\mathbf{u} - \mathbf{u}^{\text{obs}}\|_{\mathbf{r}_{\text{noise}}^{-1}}^2 + \frac{1}{2} \|\mathcal{C}^{-1/2}(\mathbf{m} - \mathbf{m}_{\text{pr}})\|^2 + \int_{\Gamma_s} \beta^{-1}(\mathbf{u} - \mathbf{h}) \cdot \mathbf{v} \, ds \\ & + \int_{\Gamma_b} \delta^{-1}(\mathbf{T}\mathbf{u} - \mathbf{m}) \cdot \mathbf{v} \, ds + \int_{\Omega} \varepsilon(\mathbf{u}) : \mathbb{C}[\varepsilon(\mathbf{v})] \, d\mathbf{x} \end{aligned}$$

**Procedure:** Obtain  $\mathbf{m}$  by enforcing  $\mathcal{L}_{\mathbf{m}} = \mathcal{L}_{\mathbf{u}} = \mathcal{L}_{\mathbf{v}} = 0$ .

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# FEM Discretization

- Consider finite element space  $\mathcal{V}_h = \text{Span}(\phi_1, \dots, \phi_n)$ .
- $\phi_k$ : Lagrange basis functions.
- For  $\mathbf{f} \in \mathcal{V}_h$ :

$$\mathbf{f} = \sum_{k=1}^n f_k \phi_k$$

In particular, construct  $\mathbf{u}, \mathbf{v}, \mathbf{m}, \mathbf{m}_{\text{pr}}, \mathbf{h} \in \mathcal{V}_h$  from  $\mathbf{u}, \mathbf{v}, \mathbf{m}, \mathbf{m}_{\text{pr}}$  and  $\mathbf{h}$ .

- Finite dimensional Hilbert space

$$(\mathcal{V}_h, \langle \cdot, \cdot \rangle_{L^2}) \cong (\mathbb{R}^n, \langle \cdot, \cdot \rangle_M), \quad \langle \mathbf{u}, \mathbf{v} \rangle = \mathbf{u}^T \mathbf{M} \mathbf{v}$$

- From the Lagrangian we find

$$\begin{bmatrix} \mathbf{C}_{\text{pr}}^{-1} & & -\mathbf{C}^* \\ & \mathbf{B}^* \boldsymbol{\Gamma}_{\text{noise}}^{-1} \mathbf{B} & \mathbf{A}^* \\ -\mathbf{C} & \mathbf{A} & \end{bmatrix} \begin{bmatrix} \mathbf{m} \\ \mathbf{u} \\ \mathbf{v} \end{bmatrix} = \begin{bmatrix} \mathbf{C}_{\text{pr}}^{-1} \mathbf{m}_0 \\ \mathbf{B}^* \boldsymbol{\Gamma}_{\text{noise}}^{-1} \mathbf{u}^{\text{obs}} \\ \mathbf{D} \mathbf{h} \end{bmatrix}$$

where  $\mathbf{A} \mathbf{u} = \mathbf{C} \mathbf{m} + \mathbf{D} \mathbf{h}$  is the discretized weak form.

# Reduced Linear System

## Parameter-to-Observable Map

From the discretized weak form we can define:

$$\mathbf{A}\mathbf{u} = \mathbf{C}\mathbf{m} + \mathbf{D}\mathbf{h} \implies \mathcal{F}(\mathbf{m}) = \mathbf{B}\mathbf{A}^{-1}[\mathbf{C}\mathbf{m} + \mathbf{D}\mathbf{h}] = \mathcal{G}\mathbf{m} + \mathbf{g}$$

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With this:

$$\begin{bmatrix} \mathbf{C}_{\text{pr}}^{-1} & & -\mathbf{C}^* \\ & \mathbf{B}^* \boldsymbol{\Gamma}_{\text{noise}}^{-1} \mathbf{B} & \mathbf{A}^* \\ -\mathbf{C} & \mathbf{A} & \end{bmatrix} \begin{bmatrix} \mathbf{m} \\ \mathbf{u} \\ \mathbf{v} \end{bmatrix} = \begin{bmatrix} \mathbf{C}_{\text{pr}}^{-1} \mathbf{m}_0 \\ \mathbf{B}^* \boldsymbol{\Gamma}_{\text{noise}}^{-1} \mathbf{u}^{\text{obs}} \\ \mathbf{D}\mathbf{h} \end{bmatrix}$$

can be reduced to

$$\left( \mathbf{C}_{\text{pr}}^{-1} + \mathcal{G}^* \boldsymbol{\Gamma}_{\text{noise}}^{-1} \mathcal{G} \right) \mathbf{m} = \mathcal{G}^* \boldsymbol{\Gamma}_{\text{noise}}^{-1} \left( \mathbf{g} + \mathbf{u}^{\text{obs}} \right) + \mathbf{C}_{\text{pr}}^{-1} \mathbf{m}_0$$

# Tractability of Computation

$$\underbrace{\left[ \mathbf{C}_{\text{pr}}^{-1} + \mathcal{G}^* \mathbf{\Gamma}_{\text{noise}}^{-1} \mathcal{G} \right]}_{\mathbf{C}_{\text{post}}^{-1}} \mathbf{m}_{\text{post}} = \left[ \mathcal{G}^* \mathbf{\Gamma}_{\text{noise}}^{-1} \left( \mathbf{g} + \mathbf{u}^{\text{obs}} \right) + \mathbf{C}_{\text{pr}}^{-1} \mathbf{m}_0 \right]$$

How do we make this computation tractable?

# Tractability of Computation

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How do we make this computation tractable?

- Use symmetry aware solvers
- Leverage sparsity of
  - matrices induced by differential operators
  - observations
- Exploit the dimensionality reduction in the parameter space

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