

A.1) $\phi(r) = \frac{k}{r^4}$ (No step marking)

$$\vec{E} = -\vec{\nabla}\phi$$

$$\Rightarrow \vec{E} = -\frac{\partial \phi}{\partial r} \hat{r}$$

$$\vec{E} = \frac{4k}{r^5} \hat{r} \rightarrow \text{It can be written as } E = \frac{4k}{r^5}$$

but $\vec{E} = \frac{4k}{r^5}$ (one side without vector sign will result in -1).

$$\begin{aligned} \text{Energy Density} &= \frac{1}{2} \epsilon_0 E^2 \\ &= \frac{1}{2} \epsilon_0 \left(\frac{4k}{r^5} \right)^2 \quad (3) \\ &= \frac{8\epsilon_0 k^2}{r^{10}} \end{aligned}$$

A.2) $\therefore \vec{\nabla} \cdot \left(\frac{\hat{r}}{r^2} \right) = 4\pi \delta^3(r)$

and $\int_V f(r) \delta^3(r-a) dV = f(a)$, if $r=a$ is in V and is 0 otherwise.

(a) $\therefore \int_{-1}^{+1} (r^2+2) \vec{\nabla} \cdot \left(\frac{\hat{r}}{r^2} \right) dV = \int_{-1}^{+1} (r^2+2) 4\pi \delta^3(r) dV$
 No step marking (1.5) $= 4\pi \int_{-1}^{+1} (r^2+2) \delta^3(r) dV$
 $= 4\pi (0+2) = 8\pi$

(b) $\int_1^\infty (r^2+2) \vec{\nabla} \cdot \left(\frac{\hat{r}}{r^2} \right) dV = \int_1^\infty (r^2+2) 4\pi \delta^3(r) dV$
 No step marking (1.5) $= 4\pi \int_1^\infty (r^2+2) \delta^3(r) dV = 0$

A.3) (a) (i) $\int_V (\vec{\nabla} \cdot \vec{E}) dV = \oint_S \vec{E} \cdot d\vec{S}$ (1)

(b) (i) $E_{t1} = E_{t2}$ & (ii) $\epsilon_1 E_{n1} = \epsilon_2 E_{n2}$ (1)

↑ Award marks only when both options are written or otherwise.

A.4) (a) Inside $E=0$ ($r < R$)

$$\therefore \text{Energy density} = \frac{1}{2} \epsilon_0 E^2 = 0$$



Outside, $E = \frac{1}{4\pi\epsilon_0} \frac{Q}{r^2}$ ($r > R$)

① mark



$$\begin{aligned} \therefore \text{Energy density} &= \frac{1}{2} \epsilon_0 E^2 \\ &= \frac{1}{2} \epsilon_0 \left(\frac{Q}{4\pi\epsilon_0 r^2} \right)^2 \\ &= \frac{Q^2}{32\pi^2 \epsilon_0 r^4} \end{aligned}$$

① mark

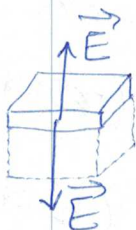
(b) Total Energy

No part marking

$$\begin{aligned} &= \frac{Q^2}{32\pi^2 \epsilon_0} \int_{r=R}^{\infty} \int_{\theta=0}^{\pi} \int_{\phi=0}^{2\pi} \frac{1}{r^4} r^2 \sin\theta d\theta d\phi dr \\ &= \frac{Q^2}{32\pi^2 \epsilon_0} \int_R^{\infty} \frac{1}{r^2} \int_0^{\pi} \sin\theta d\theta \int_0^{2\pi} d\phi dr \\ &= \frac{Q^2}{32\pi^2 \epsilon_0} \left[-\frac{1}{r} \right]_R^{\infty} \left[-\cos\theta \right]_0^{\pi} \left[\phi \right]_0^{2\pi} \\ &= \frac{Q^2}{32\pi^2 \epsilon_0} \left[\frac{1}{r} \right]_{\infty}^R \left[(-1) + 1 \right] 2\pi \\ &= \frac{Q^2}{32\pi^2 \epsilon_0} \frac{1}{R} \times 4\pi = \frac{Q^2}{8\pi \epsilon_0 R} \end{aligned}$$

③ mark

A.5)



$$EA + EA = \frac{Q_{enc}}{\epsilon_0}$$

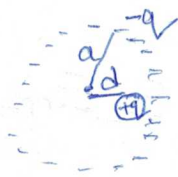
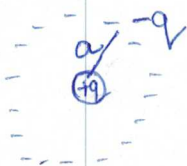
$$\Rightarrow 2EA = \frac{\sigma A}{\epsilon_0}$$

$$\Rightarrow \vec{E} = \frac{\sigma}{2\epsilon_0} \Rightarrow \vec{E} = \frac{\sigma}{2\epsilon_0} \hat{n}$$

③

No part marking

A.6)



In presence of an external field \vec{E} , the nucleus will shift slightly to the right, creating a dipole moment $\vec{P} = q\vec{d}$

The shift stop increasing when the restoring electric force from the displaced cloud balance the external force on the nucleus.

For a uniformly charged sphere of radius a and total charge $-q$. The electric field inside at a distance r from centre is

$$E_{\text{cloud}}(r) = \frac{1}{4\pi\epsilon_0} \frac{-q(r)}{a^3}$$

(valid for $r < a$)
(Using Gauss's law)

~~At equilibrium~~

← Award

② 'if correct till here'

So, if the nucleus is displaced by \vec{d} , it feels a restoring force from this field,

$$\vec{F}_{\text{restoring}} = q \cdot \vec{E}_{\text{cloud}}(d)$$

$$= q \cdot \frac{1}{4\pi\epsilon_0} \frac{-q\vec{d}}{a^3} = \frac{-q^2}{4\pi\epsilon_0 a^3} \vec{d}$$

In the external field, the nucleus also experience a force:

$$\vec{F}_{\text{ext}} = q \cdot \vec{E}$$

At equilibrium, ~~the~~

$$\vec{F}_{\text{restoring}} + \vec{F}_{\text{ext}} = 0$$

$$\Rightarrow \frac{-q^2}{4\pi\epsilon_0 a^3} \vec{d} + q\vec{E} = 0$$

$$\Rightarrow q\vec{E} = \frac{q^2}{4\pi\epsilon_0 a^3} \vec{d}$$

$$\Rightarrow \vec{d} = (4\pi\epsilon_0 a^3) \frac{\vec{E}}{q}$$

$$\text{Now, } \therefore \vec{p} = q\vec{d}$$

$$\text{Substituting } \vec{d}, \vec{p} = (4\pi\epsilon_0 a^3) \vec{E}$$

$$\text{Also, it is given that } \vec{p} = \alpha \vec{E}$$

where α is polarizability.

Comparing both eqⁿs for \vec{p} , we get

$$\alpha = 4\pi\epsilon_0 a^3 \quad \text{④}$$

~~No part marking~~

← Award 3 if correct till here.