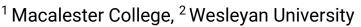
Constraining Dust Structure in Three Protoplanetary and Transitional Disks



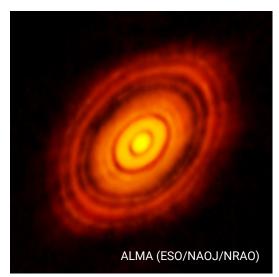
Macalester College

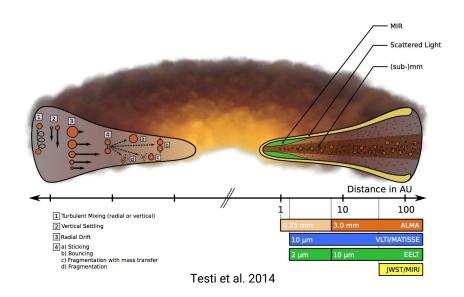
Alyssa M. Bulatek ¹, Dr. Kevin Flaherty ²

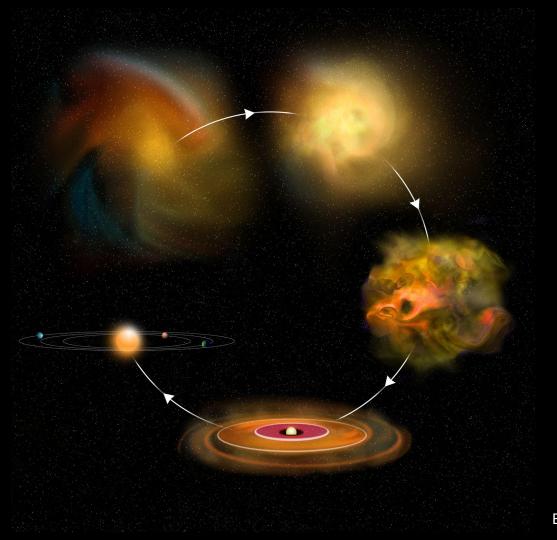


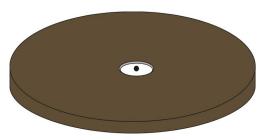




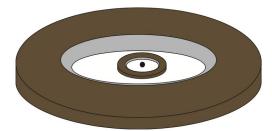




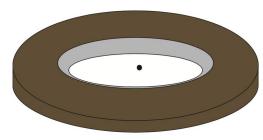




Full Disk



Pre-Transitional Disk

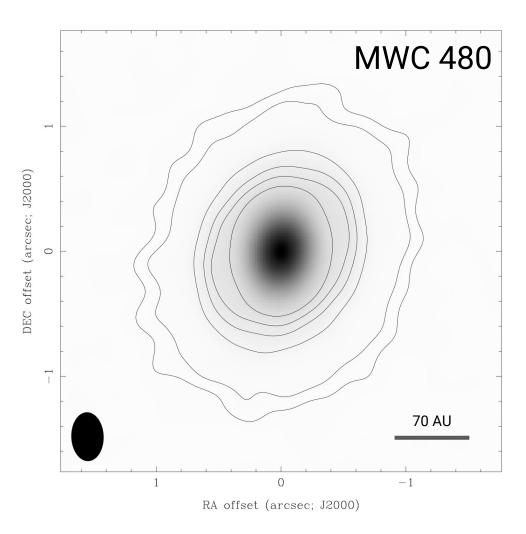


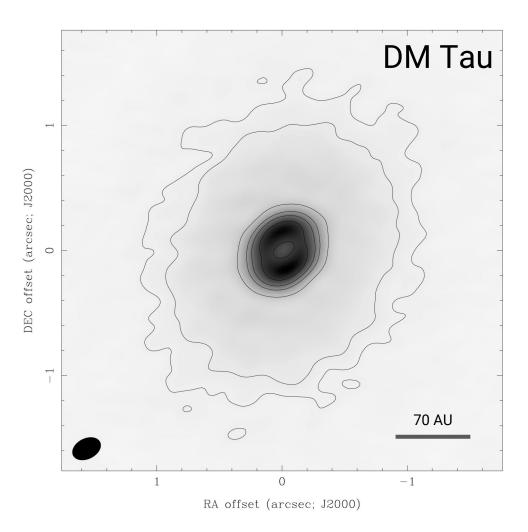
Transitional Disk

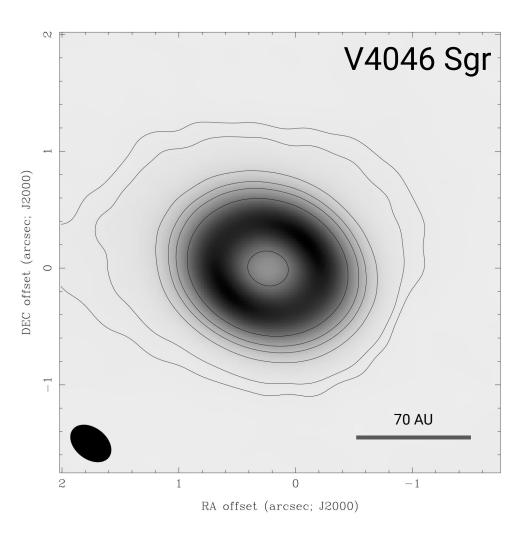
Figure from Espaillat et. al. 2014











MWC 480

1 ring dust-to-gas: $0.0147^{+3.2e-5}_{-3.1e-5}$ inner radius: $9.25^{+0.0074}_{-0.0008}$ outer radius: $84^{+0.02}_{-0.2}$

DM Tau

2 rings dust-to-

dust-to-gas: 0.0048 +4.2e-5 -5.6e-5 inner radius: 73.06 +0.10 outer radius: 194.36 +0.41 -0.07

dust-to-gas: $0.0066^{+1.0e-5}_{-2.3e-5}$ inner radius: $18.92^{+0.07}_{-0.006}$ outer radius: $46.58^{+0.08}_{-0.009}$

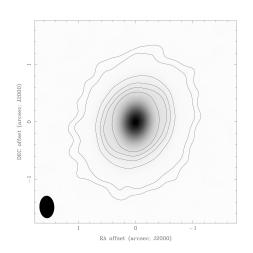


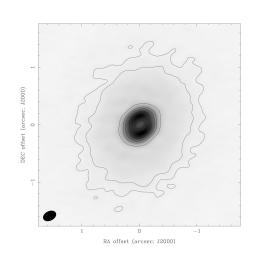
3 rings

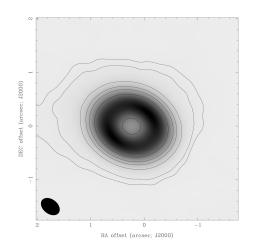
dust-to-gas: $8.8e-5 \, {}^{+1.5e-5}_{-1.1e-5}$ inner radius: $73.45 \, {}^{+3.48}_{-0.75}$ outer radius: $90.58 \, {}^{+0.94}_{-1.19}$

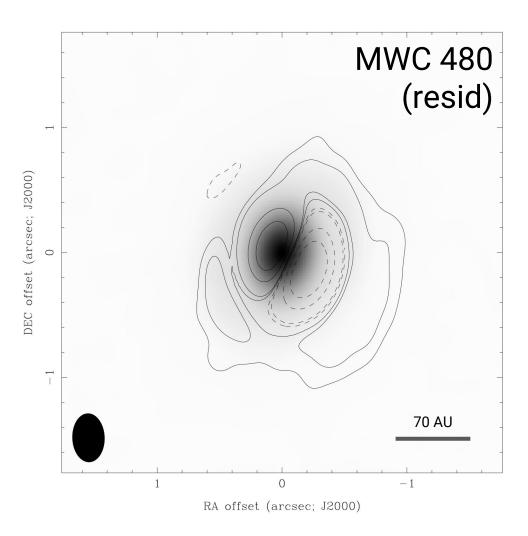
dust-to-gas: $0.01^{+1.1e-4}_{-6.0e-5}$ inner radius: $24.74^{+0.12}_{-0.17}$ outer radius: $62.01^{+0.11}_{-0.24}$

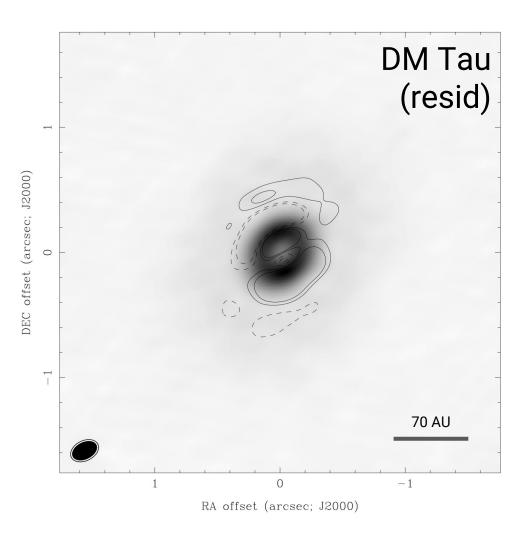
dust-to-gas: 9.5e-4 +1.35e-4 inner radius: 1.5e-4 +1.64e-4 outer radius: 14.90 +1.19

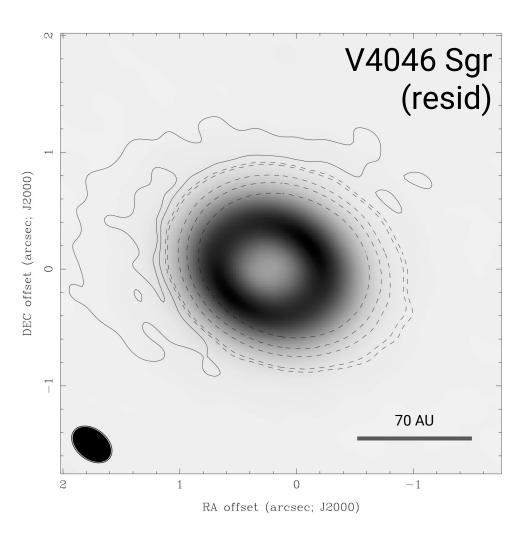












Our models suggest that MWC 480, DM Tau,

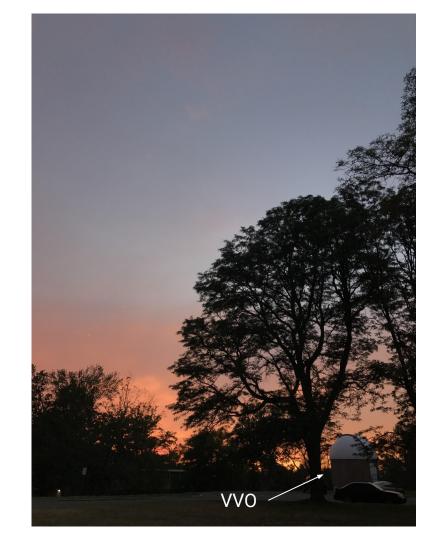
and V4046 Sgr might host asymmetric disks.

We can still make conclusions about

mechanisms at work within the disks!

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- Dr. John Cannon (of Macalester College)



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Images

1. HI Tau:

http://www.eso.org/public/images/eso1436a/

2. Star formation:

https://www.nrao.edu/pr/2012/clumpcores/

- 2. Disk evolution: Espaillat, C., Muzerolle, J., Najita, J., et
- al. 2014, Protostars and Planets VI, 497
- 3. Disk mechanisms: Testi, L., Birnstiel, T., Ricci, L., et al.
- 2014, Protostars and Planets VI, 339
- 4. ALMA:

http://www.almaobservatory.org/en/images/antennas-a t-sunset-in-aos/

5. Interferometry:

http://www.cv.nrao.edu/course/astr534/Interferometer s1.html