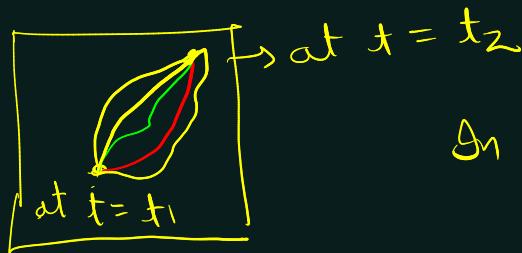


(End) Lagrange equations of motion: $\frac{\partial L}{\partial q_i} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_i} \right) = 0$



In the intermediate region $q = q(t)$
along the main path

For any other path: $q(t) + \delta q$

Actual solution
Variation on top of the actual solution

Free particle:

Not in a potential / field
or not subject to any force.

In 3D: $T = \frac{1}{2} m(\dot{x}^2 + \dot{y}^2 + \dot{z}^2)$; $V = 0$

How many coordinates: $(q_i) = x, y$ and z

$$\textcircled{1} \quad \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{x}} \right) - \frac{\partial L}{\partial x} = 0 \quad \Rightarrow$$

$$\frac{d}{dt} \left(m \dot{x} \right) = 0 \quad \Rightarrow m \ddot{x} (\text{acceleration}) =$$

$$\textcircled{2} \quad m \ddot{y} = 0$$

$$\textcircled{3} \quad m \ddot{z} = 0$$

consider a particle in a potential $V \equiv V(x, y, z)$ and $L = T - V$

$$\textcircled{1} \quad m \ddot{x} + \frac{\partial V}{\partial x} = 0 \quad \Rightarrow m \ddot{x} = - \frac{\partial V}{\partial x}$$

$$\textcircled{2} \quad m \ddot{y} + \frac{\partial V}{\partial y} = 0 \quad \Rightarrow m \ddot{y} = - \frac{\partial V}{\partial y}$$

$$\textcircled{3} \quad m \ddot{z} + \frac{\partial V}{\partial z} = 0 \quad \Rightarrow m \ddot{z} = - \frac{\partial V}{\partial z}$$

$$(m \ddot{x}) \hat{x} + (m \ddot{y}) \hat{y} + (m \ddot{z}) \hat{z} = - \left(\frac{\partial V}{\partial x} \hat{x} + \frac{\partial V}{\partial y} \hat{y} + \frac{\partial V}{\partial z} \hat{z} \right)$$

$$\Rightarrow \vec{F} = -\vec{\nabla} V$$

Let's see the Lagrange's equation

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}} \right) - \frac{\partial L}{\partial q} = 0 \quad ; \quad L(q, \dot{q}, t)$$

For free particle $L = \frac{1}{2} m(\dot{x}^2 + \dot{y}^2 + \dot{z}^2)$

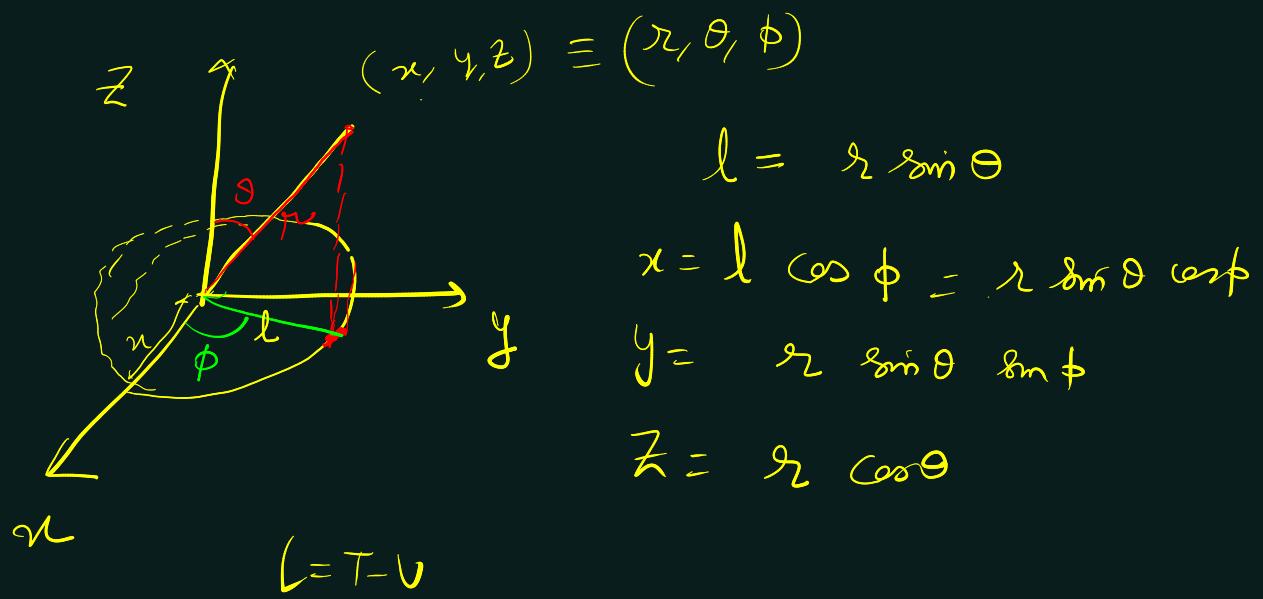
If L does not explicitly depend on q

then $\frac{\partial L}{\partial q} = 0$; q is called a cyclic coordinate

$$\Rightarrow \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}} \right) = 0 \Rightarrow \frac{\partial L}{\partial \dot{q}} = \text{constant} \quad \left\{ \begin{array}{l} \text{Conservation} \\ \text{Law} \end{array} \right.$$

\Rightarrow for free particle along x : $\frac{d}{dt}(m\dot{x}) = 0 \Rightarrow m\dot{x} = \text{constant}$
 $m\ddot{x} = 0 \Rightarrow m = \text{constant}$

Conservation of momentum



$$\text{For a free particle } L = T = 0 = \frac{1}{2} m (\dot{x}^2 + \dot{y}^2 + \dot{z}^2)$$

$$L = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2 + r^2 \sin^2 \theta \dot{\phi}^2)$$

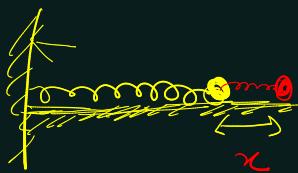
L does not depend on ϕ explicitly therefore ϕ is cyclic

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{\phi}} \right) - \cancel{\frac{\partial L}{\partial \phi}} = 0$$

$$\Rightarrow \frac{\partial L}{\partial \dot{\phi}} = \text{constant} \Rightarrow m r^2 \sin^2 \theta \dot{\phi} = \text{constant}$$

$$\Rightarrow m \dot{\ell}^2 \dot{\phi} = \text{constant} = L_z \Rightarrow \text{conservation of angular momentum}$$

What if $V \neq 0$ but $V = V_{\text{Coulomb}}$



Scheibenoberfläche

$$T = \frac{1}{2} m \dot{x}^2$$

$$V = \frac{1}{2} k x^2$$

$$L = \frac{1}{2} m \dot{x}^2 - \frac{1}{2} k x^2$$

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{x}} \right) - \frac{\partial L}{\partial x} = 0 \Rightarrow m \ddot{x} + k x = 0$$

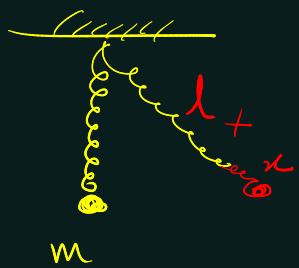
$$\Rightarrow \ddot{x} + \left(\frac{k}{m} \right) x = 0$$

$$\frac{d^2 x}{dt^2} + \omega^2 x = 0$$

$$x_2 \propto e^{i \omega t} + \beta e^{-i \omega t} = C \sin \omega t + D \cos \omega t$$

Spring

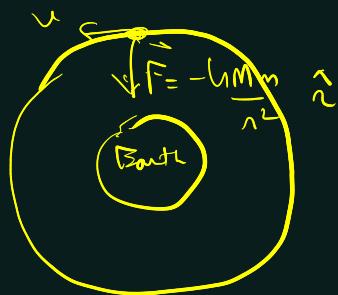
Pendulum



Spring force constant k
mass of bob m
acc. due to gravity g

- (1) Non-inertial frames of reference }
 (2) Dumbitative forces: Fictitious }
 How do we deal with
 those in Lagrangian Mech

(b) Non-inertial frames



Student A: There is some other pseudo-force

Student B: It is always falling but not falling towards the Earth

Student C: In the frame of ref. of satellite
centrifugal (?)

To calculate velocity $m \frac{v^2}{r} = \frac{Gm}{r^2}$

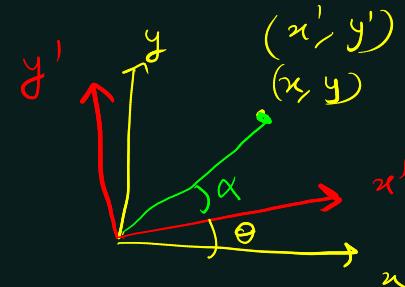
Centrifugal force

$$\text{Lagrangian for a free particle: } L = T - V = \frac{1}{2} m(\dot{x}^2 + \dot{y}^2 + \dot{z}^2) - 0$$

Rotating frame of reference:

$$x = r \cos(\theta + \alpha)$$

$$y = r \sin(\theta + \alpha)$$



$$x = r \cos \theta \cos \alpha - r \sin \theta \sin \alpha = (r \cos \alpha) \cos \theta - (r \sin \alpha) \sin \theta$$

$$y = r \sin \theta \cos \alpha + r \cos \theta \sin \alpha = (r \cos \alpha) \sin \theta + (r \sin \alpha) \cos \theta$$

$$\begin{aligned} x &= x' \cos \theta - y' \sin \theta \\ y &= x' \sin \theta + y' \cos \theta \end{aligned}$$

$$z = z$$

$$\dot{x} = \dot{x}' \cos \theta - \dot{y}' \sin \theta - (x' \sin \theta) \omega - (y' \cos \theta) \omega$$

$$\dot{y} = \dot{x}' \sin \theta + \dot{y}' \cos \theta + (x' \cos \theta) \omega - (y' \sin \theta) \omega$$

$$\Rightarrow \dot{z} = (\dot{x}' - y' \omega) \cos \theta - (\dot{y}' + x' \omega) \sin \theta = A \cos \theta - B \sin \theta$$

$$\dot{y} = (\dot{x}' - y' \omega) \sin \theta + (\dot{y}' + x' \omega) \cos \theta = A \sin \theta + B \cos \theta$$

$$\dot{x}^2 + \dot{y}^2 + \dot{z}^2 = (\dot{x}' - y' \omega)^2 + (\dot{y}' + x' \omega)^2 + \dot{z}'^2$$

$$L = \frac{m}{2} \left[(\dot{x}' - y' \omega)^2 + (\dot{y}' + x' \omega)^2 + \dot{z}'^2 \right]$$

$$\omega = \dot{\theta} \quad \text{But } \omega \text{ is a vector} \quad \vec{\omega} = \dot{\theta} \hat{z}$$

$$\vec{r} = x \hat{x} + y \hat{y} + z \hat{z}$$

$$\text{Calculate} \quad \vec{\omega} \times \vec{r} = \omega x \hat{y} - \omega y \hat{x}$$

$$\begin{aligned} \dot{\vec{r}} + \vec{\omega} \times \vec{r}' &= \dot{x} \hat{x} + \dot{y} \hat{y} + \dot{z} \hat{z} + \omega x \hat{y} - \omega y \hat{x} \\ &= (\dot{x} - \omega y) \hat{x} + (\dot{y} + \omega x) \hat{y} + \dot{z} \hat{z} \end{aligned}$$

$$|\dot{\vec{r}} + \vec{\omega} \times \vec{r}'|^2 = (\dot{x} - \omega y)^2 + (\dot{y} + \omega x)^2 + \dot{z}^2$$

$$L = \frac{m}{2} |\dot{\vec{x}}' + \omega \vec{x}'|^2$$

Find out the Equations of motion and make for yourself that there are two additional force there : Centrifugal & Coriolis

$$\frac{\partial L}{\partial q} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}} \right) = 0$$

Dissipative
Forces

In case of dampener: the above equation has to be modified -

If dampener force has the form $F = -\alpha \dot{q}$

$$\text{then } R = -\frac{1}{2} \alpha \dot{q}^2$$

And the equations are

$$\frac{\partial L}{\partial q} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}} \right) = \frac{\partial R}{\partial \dot{q}}$$

Formal

$$F_{\text{ext}} = -kx$$

$$F_{\text{int}} = -\alpha \dot{x} = -\alpha v$$

$$T = \frac{1}{2} m \dot{x}^2$$

$$V = \frac{1}{2} k x^2$$

$$R = \frac{1}{2} \alpha \dot{x}^2$$

$$\frac{\partial L}{\partial x} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{x}} \right) = \frac{dR}{dx} \Rightarrow -kx - m\ddot{x} = \alpha \ddot{x}$$

$$\Rightarrow m\ddot{x} + \alpha \dot{x} + kx = 0$$

$$\Rightarrow \frac{d^2x}{dt^2} + \left(\frac{\alpha}{m}\right) \frac{dx}{dt} + \left(\frac{k}{m}\right)x = 0$$

Let's say $\frac{\alpha}{m} = 2\gamma$ and $\frac{k}{m} = \omega^2$

$$\boxed{\frac{d^2x}{dt^2} + 2\gamma \frac{dx}{dt} + \omega^2 x = 0}$$

For no-friction $\alpha \rightarrow \gamma = 0$

$$\frac{d^2x}{dt^2} + \omega^2 x = 0$$

we get back S.H.M. equation

Let's consider a trial function : $x(t) = A e^{i\omega t}$

$$\dot{x} = \frac{dx}{dt} = i\omega A e^{i\omega t}$$

$$\ddot{x} = \frac{d^2x}{dt^2} = (i\omega)^2 A e^{i\omega t} = -\omega^2 A e^{i\omega t}$$

$$\Rightarrow -\omega^2 A e^{i\omega t} + \omega^2 A e^{i\omega t} = 0$$

Try with $x(t) = B e^{-i\omega t}$

This is also a solution (you can verify)

Then a linear combination $x(t) = A e^{i\omega t} + B e^{-i\omega t}$ is also a solution

$$\Rightarrow x(t) = A (\cos \omega t + i \sin \omega t) + B (\cos \omega t - i \sin \omega t)$$

$$\Rightarrow x(t) = \underbrace{(A+B)}_{D} \cos \omega t + \underbrace{i(A-B)}_{C} \sin \omega t$$

$$x(t) = C \sin \omega t + D \cos \omega t$$

$\left\{ \begin{array}{l} \text{Fix the constants, say} \\ x = 0 \text{ at } t = 0 \end{array} \right.$

Let's say one more $x(t) = A \cos(\omega t + \phi)$

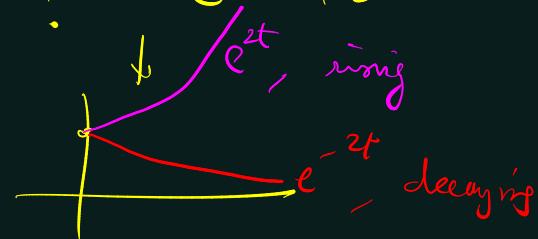
$$\frac{d^2x}{dt^2} + \omega^2 x = 0 \Rightarrow D^2 + \omega^2 = 0$$

$$D \begin{cases} \alpha_1 \rightarrow i\omega \\ \alpha_2 \rightarrow -i\omega \end{cases}$$

$$\Rightarrow x(t) = A e^{\alpha_1 t} + B e^{\alpha_2 t}$$

$$= A e^{i\omega t} + B e^{-i\omega t} \rightarrow \text{A sinusoidal wave}$$

Say $\frac{d^2x}{dt^2} - 4x = 0 \Rightarrow x(t) = ?$ $= A e^{2t} + B e^{-2t}$



Let's take up the spring problem with friction.

$$\frac{d^2x}{dt^2} + 2\gamma \frac{dx}{dt} + \omega^2 x = 0 \Rightarrow D^2 + 2\gamma D + \omega^2 = 0$$

Find two roots α_1 & α_2

$$\alpha_1, \alpha_2 : -\frac{2\gamma \pm \sqrt{4\gamma^2 - 4\omega^2}}{2}$$

$$\approx \alpha_1 = -\gamma + \sqrt{\gamma^2 - \omega^2} \quad \text{and} \quad \alpha_2 = -\gamma - \sqrt{\gamma^2 - \omega^2}$$

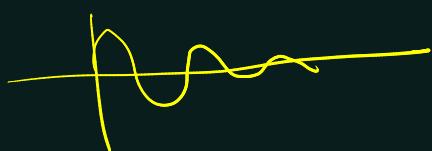
$$x(t) = A e^{\alpha_1 t} + B e^{\alpha_2 t} = A e^{(-\gamma + \sqrt{\gamma^2 - \omega^2})t} + B e^{(-\gamma - \sqrt{\gamma^2 - \omega^2})t}$$

$$= e^{-\gamma t} \left[A e^{\sqrt{\gamma^2 - \omega^2} t} + B e^{-\sqrt{\gamma^2 - \omega^2} t} \right]$$

✓ Decaying Term If $\gamma > \omega$ then rising or decaying

If $\gamma < \omega$ then consider $\gamma^2 - \omega^2 = -k^2$ which is negative

$$x(t) = e^{-\gamma t} \left[A e^{ikt} + B e^{-ikt} \right]$$



oscillates with dissipation / damping and also forcing

$$\frac{d^2x}{dt^2} + 2\gamma \frac{dx}{dt} + \omega^2 x = F \cos \omega_0 t$$
$$= \frac{F}{2} e^{i\omega_0 t} + \underbrace{\frac{F}{2} e^{-i\omega_0 t}}$$

$$\frac{d^2x}{dt^2} + 2\gamma \frac{dx}{dt} + \omega^2 x = \frac{F}{2} e^{i\omega_0 t}$$

Trial solution $x(t) = A e^{i\omega_0 t}$

$$\Rightarrow (\omega_0)^2 A e^{i\omega_0 t} + 2\gamma (\omega_0) A e^{i\omega_0 t} + \omega^2 A e^{i\omega_0 t} = \frac{F}{2} e^{i\omega_0 t}$$

$$\Rightarrow A = \frac{F/2}{\omega^2 + 2i\gamma\omega_0 - \omega_0^2}$$

$$\frac{d^2x}{dt^2} + 2\gamma \frac{dx}{dt} + \omega^2 x = \frac{F_0}{2} e^{-i\omega t}$$

Try with $x(t) = B e^{-i\omega t}$

$$B = \frac{F_0/2}{\omega^2 - 2i\gamma\omega_0 - \omega^2}$$

$$\frac{d^2x}{dt^2} + 2\gamma \frac{dx}{dt} + \omega^2 x = F_0 e^{i\omega t}$$

$$x(t) = A e^{i\omega t} + B e^{-i\omega t} = \frac{F_0/2}{\omega^2 + 2i\gamma\omega_0 - \omega^2} e^{i\omega t} + \frac{F_0/2}{\omega^2 - 2i\gamma\omega_0 - \omega^2} e^{-i\omega t}$$

(Real) + (Imaginary)

Homework =

$$F_0/2 \left[\frac{\alpha + i\beta}{(\omega^2 + \omega_s^2) + i(2\gamma\omega_s)} + \frac{\alpha - i\beta}{(\omega^2 - \omega_s^2) - i(2\gamma\omega_s)} \right]$$

Astrophysics:

we will use the Lagrangian dynamics to solve the central force problem: Planetary Motion: (Atomic Model)

Gravity

$$F_g = \frac{G m_1 m_2}{r^2}$$

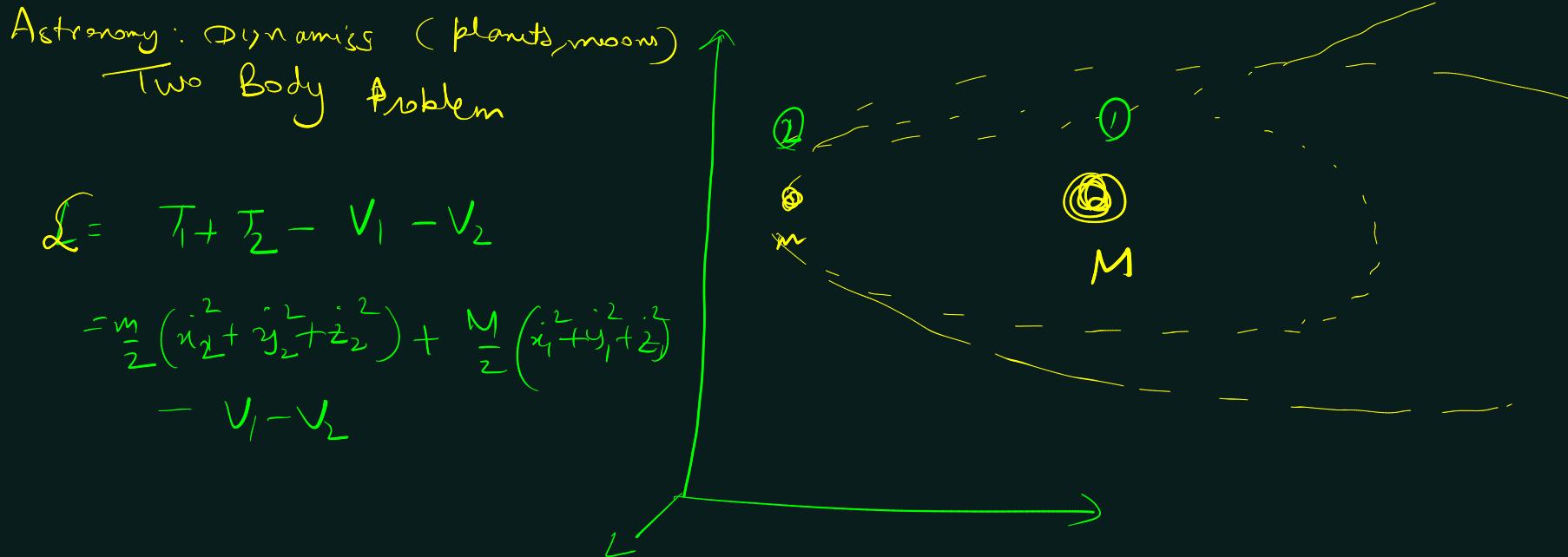
Electrostatic

$$F_E = \frac{1}{4\pi\epsilon_0} \frac{q_1 q_2}{r^2}$$

$$\frac{G}{(4\pi\epsilon_0)} \approx \frac{10^{-11}}{10^{+9}} \approx 10^{-20}$$

(Gravity
is
extremely
weak)

Then why Gravity moves planets ; Not the electromagnetic force



This problem is equivalent to

motion of one body moving in central potential / force field
(centred at centre of mass of 2-body system)

Equivalent to

1-body
in
central

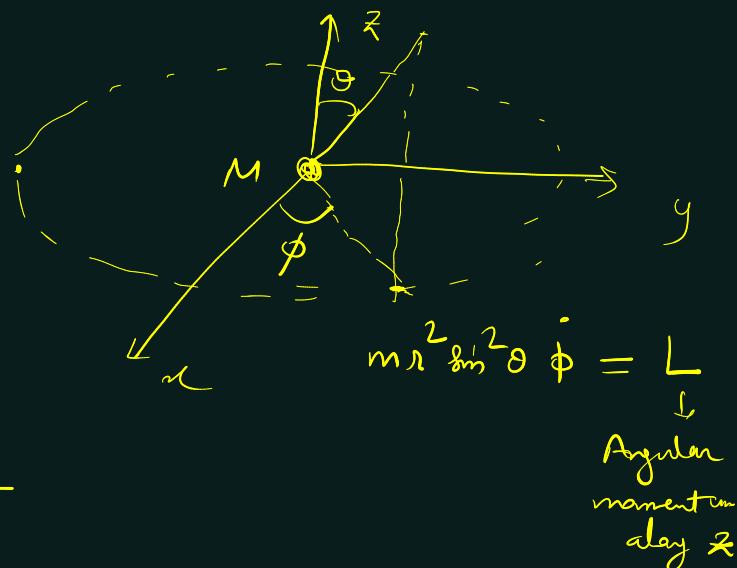
$$\mathcal{L} = \frac{1}{2} \mu (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) + \frac{GM\mu}{r}$$

$$\mu = \frac{mM}{m+M} \approx m \text{ if } m \ll M$$

$$\mathcal{L} = \frac{m}{2} (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) + \frac{GMm}{r}$$

Go to r, θ, ϕ system

$$\mathcal{L} = \frac{m}{2} (\dot{r}^2 + r^2 \dot{\theta}^2 + r^2 \sin^2 \theta \dot{\phi}^2) + \frac{GMm}{r}$$



ϕ is cyclic: So; $\frac{\partial \mathcal{L}}{\partial \dot{\phi}} - \frac{d}{dt} \left(\frac{\partial \mathcal{L}}{\partial \phi} \right) = \frac{d}{dt} (m r^2 \sin^2 \theta \dot{\phi}) = 0$

$$\Rightarrow m r^2 \sin^2 \theta \dot{\phi} = L$$

If L is preserved/closed then $\theta = \frac{\pi}{2}$; motion is in the plane

$$x = r \cos \phi$$

$$y = r \sin \phi$$

$$\phi = \frac{\pi}{2}$$

$$z = 0$$

~~$$\mathcal{L} = \frac{m}{2} (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) - \frac{GMm}{r}$$~~

$$\mathcal{L} = \frac{m}{2} (\dot{r}^2 + r^2 \dot{\phi}^2) + \frac{GMm}{r}$$

$$\frac{\partial \mathcal{L}}{\partial r} - \frac{d}{dt} \left(\frac{\partial \mathcal{L}}{\partial \dot{r}} \right) = 0$$

$$\Rightarrow m r \dot{\phi}^2 - m \ddot{r} = \frac{GMm}{r^2}$$

$$\Rightarrow \ddot{r} - r \dot{\phi}^2 = - \frac{GM}{r^2}$$

Let's extract all the information about the possible solution physically

$$\mathcal{L} = T - V$$

$$E = T + V = \frac{m(\dot{r}^2 + r^2 \dot{\phi}^2)}{2} - \frac{GMm}{r}$$

$$\text{But } m r^2 \dot{\phi}^2 = L \Rightarrow \dot{\phi} = \frac{L}{mr^2}$$

$$\Rightarrow E = \frac{m \dot{r}^2}{2} + \frac{m}{2} r^2 \frac{L^2}{m^2 r^4} - \frac{GMm}{r}$$

$$= \frac{m \dot{r}^2}{2} + \frac{L^2}{2mr^2} - \frac{GMm}{r}$$

$$E = \frac{1}{2} m \dot{r}^2 + V_{eff} \quad \text{where} \quad V_{eff} = \frac{L^2}{2mr^2} - \frac{GMm}{r} = \frac{A}{r^2} - \frac{B}{r}$$

where $A = \frac{L^2}{2m}$ and $B = GMm$

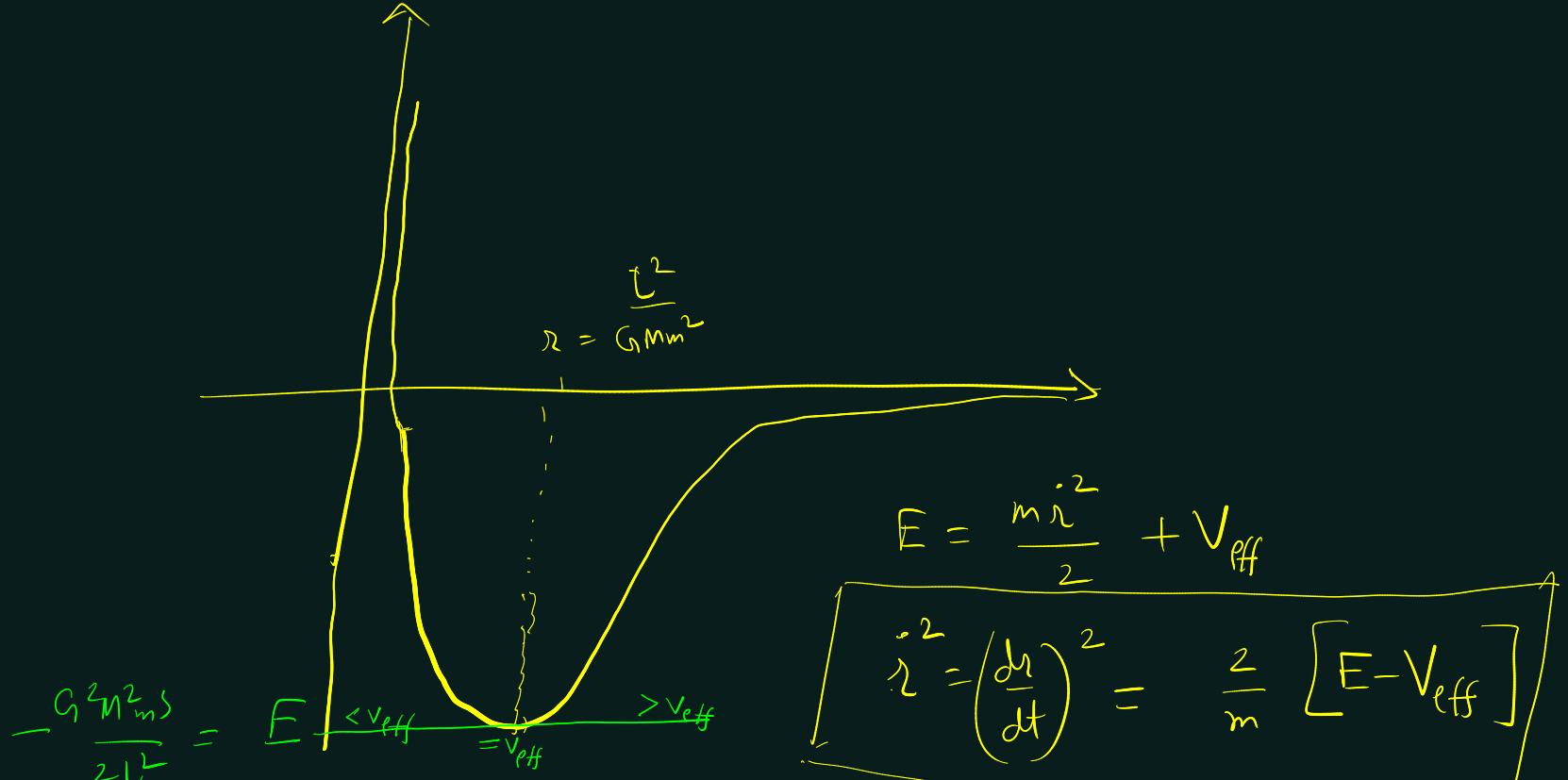
Plot V_{eff} as a function of r :

Minima: $\frac{dV_{eff}}{dr} = 0 \Rightarrow -\frac{2A}{r^3} + \frac{B}{r^2} = 0$

$$\Rightarrow r = \frac{2A}{B} = \frac{2L^2}{2GMm^2}$$

Minima at $r = \frac{L^2}{GMm^2}$ and V_{eff} at minimum = $\frac{1}{r} \left[\frac{A GMm^2}{L^2} - B \right]$

$$V_{eff, min} = \frac{GMm^2}{L^2} \left[\frac{L^2}{2m} \frac{GMm^2}{L^2} - B \right] = -\frac{GMm^3}{2L^2}$$



E is a constant of motion like L

Consider ; say
different values
we consider

$$E = \frac{GMm}{r^2} \equiv V_{\text{eff}, \min}$$

of E

$$\Rightarrow \frac{dr}{dt} = \sqrt{\frac{2}{m} (E - V_{\text{eff}})} = 0 \text{ at min. point}$$

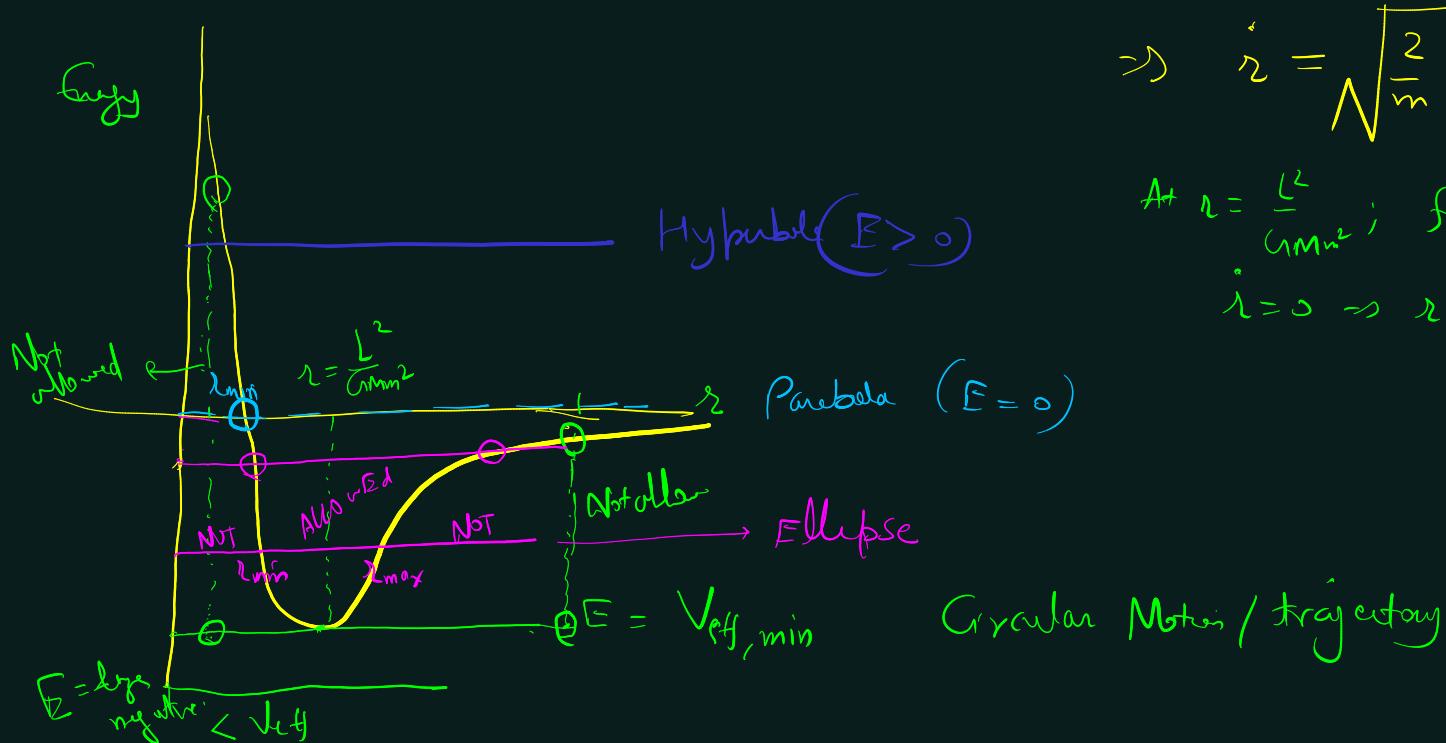
$$E = \frac{m\dot{r}^2}{2} + m\frac{\dot{r}^2\phi^2}{2} - \frac{GMm}{r}$$

$$E = \frac{m\dot{r}^2}{2} + \underbrace{\frac{L^2}{2mr^2}}_{V_{\text{eff}}(r)} - \frac{GMm}{r}$$

where $\phi = \frac{L}{mr^2}$

} of
Angular
is
Cyclic

$$\Rightarrow E = m\frac{\dot{r}^2}{2} + V_{\text{eff}}(r)$$



At $r = \frac{L^2}{Gm r^2}$; for $E = V_{\text{eff}, \min}$

$$r = \text{constant} = \frac{L^2}{GMm^2}$$

What are the r_{\min} & r_{\max} for Elliptical orbit. Let's calculate below

$$E = \frac{m\dot{r}^2}{2} + \frac{L^2}{2mr^2} - \frac{GM_m}{r}$$

Elliptic: At $r=r_{min}$ or at $r=r_{max}$; $\dot{r}=0$

$$\Rightarrow E = \frac{L^2}{2m} u^2 - GM_m u \quad \text{when we defined } u = \frac{1}{r}$$

$$\Rightarrow u^2 - 2 \frac{GM_m^2}{L^2} u - \frac{2mE}{L^2} = 0$$

$$\Rightarrow u = \frac{2 \frac{GM_m^2}{L^2} \pm \sqrt{4 \frac{G^2 M_m^4}{L^4} + \frac{8mE}{L^2}}}{2} =$$

$$= \frac{GM_m^2}{L^2} \left[1 \pm \sqrt{1 + \frac{2EL^2}{G^2 M_m^2 m^3}} \right]$$

$$\Rightarrow r = \frac{L^2}{GM_m^2} \left[1 \pm \varepsilon \right]^{-1} \quad \text{where } \varepsilon = \sqrt{1 + \frac{2EL^2}{G^2 M_m^2 m^3}} = \text{Eccentricity}$$

$$\left. \begin{array}{l} r_{\min} = \frac{L^2}{GMm^2(1+\varepsilon)} \\ r_{\max} = \frac{L^2}{GMm^2(1-\varepsilon)} \end{array} \right\} \quad \begin{array}{l} \text{for } E=0; \quad \varepsilon = 1 \\ \rightarrow r_{\min} = \frac{L^2}{2GMm^2}; \quad r_{\max} = \infty \end{array}$$

what if eccentricity = 0 $\Rightarrow \varepsilon = 0 \Rightarrow r_{\min} = \frac{L^2}{GMm^2}$

Entire trajectory
understood.

which is a circle.

$$r_{\max} = \frac{L^2}{GMm^2}$$

We will solve the Lagrangian Equations now

$$\text{Eqn is : } \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{r}} \right) - \frac{\partial L}{\partial r} = 0$$

$$\Rightarrow L = \frac{m\dot{r}^2}{2} + m\frac{r^2\dot{\phi}^2}{2} + \frac{GMm}{r}$$

$$\Rightarrow \frac{\partial L}{\partial \dot{r}} = m\dot{r}; \quad \frac{\partial L}{\partial r} = m r \dot{\phi}^2 - \frac{GMm}{r^2}$$

$$\Rightarrow m \ddot{r} - m r \dot{\phi}^2 + \frac{GMm}{r^2} = 0$$

$$\Rightarrow \ddot{r} - r\dot{\phi}^2 = -\frac{GM}{r^2}$$

Also from $\dot{\phi}$ eqn we have

$$mr^2\dot{\phi} = L \Rightarrow \dot{\phi} = \frac{L}{mr^2}$$

$$\Rightarrow \ddot{r} - \frac{L^2}{m^2 r^3} = -\frac{GM}{r^2}$$

Define $u = \frac{1}{r} \Rightarrow r = \frac{1}{u}$

$$\frac{dr}{dt} = -\frac{1}{u^2} \frac{du}{dt}$$

$$\boxed{\dot{\phi} = \frac{L}{mr^2}}$$

$$\ddot{r} = \frac{d}{dt} \left(\frac{dr}{dt} \right) = \frac{d}{d\phi} \left[\frac{d\phi}{dt} \frac{du}{d\phi} \right] \frac{d\phi}{dt} = \frac{d}{d\phi} \left[\dot{\phi} \frac{du}{d\phi} \right] \dot{\phi}$$

$$\Rightarrow \ddot{r} = \frac{L}{mr^2} \frac{d}{d\phi} \left[\frac{L}{mr^2} \left(-\frac{1}{u^2} \frac{du}{d\phi} \right) \right]$$

$$= -\frac{L^2}{m^2} u^2 \frac{d^2 u}{d\phi^2}$$

$$\Rightarrow -\frac{L^2}{m^2} u^2 \frac{d^2 u}{d\phi^2} - \frac{L^2}{m^2} u^3 = -GMu^2$$

⇒

$$\left[\frac{d^2u}{d\phi^2} + u = \frac{GMm^2}{L^2} \right] \quad (1)$$

$$\frac{d^2u}{d\phi^2} + u = 0 + \frac{GMm^2}{L^2}$$

first solve $\frac{d^2u}{d\phi^2} + u = 0$

then add to ~~it~~ the sol. of
with trial method

$$\frac{d^2u}{d\phi^2} + u = \frac{GMm^2}{L^2}$$

$$u = A \cos(\phi + \phi_0)$$

Let's say now the solution of (1) is $u = A \cos(\phi + \phi_0) + C$

$$\frac{d^2u}{d\phi^2} = -A \cos(\phi + \phi_0)$$

$$\frac{d^2u}{d\phi^2} + u = -A \cos(\phi + \phi_0) + A \cos(\phi + \phi_0) + C = \frac{GMm^2}{L^2}$$

$$\therefore C = \frac{GMm^2}{L^2}$$

Therefore $u = A \cos(\phi + \phi_0) + \frac{GMm^2}{L^2}$

