

Background Independent Quantum Gravity

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Why Quantum Gravity?

General Relativity and Quantum Mechanics are Fundamentally Incompatible

- GR violates the Uncertainty Principle
- QM violates Causality

Why Quantum Gravity?

General Relativity and Quantum Field Theory give different answers

- Black Hole Information Paradox
- Wormholes, Time Machines, and Warp Drives
- Origins and structure of the universe

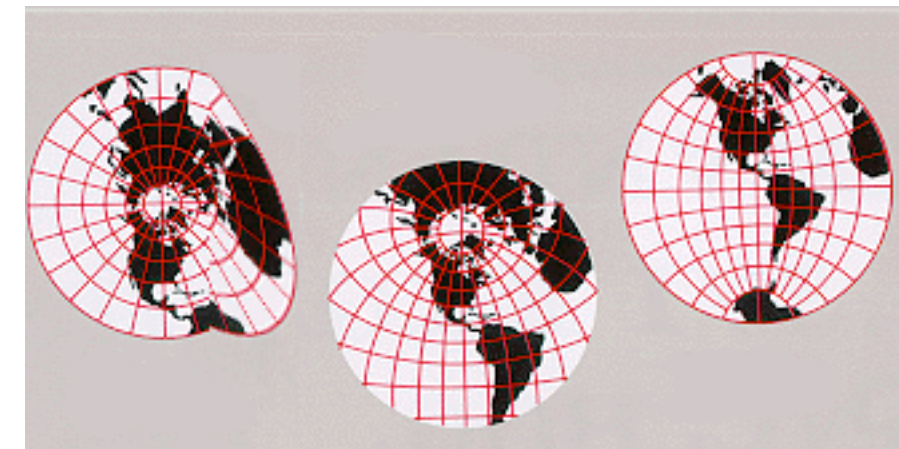
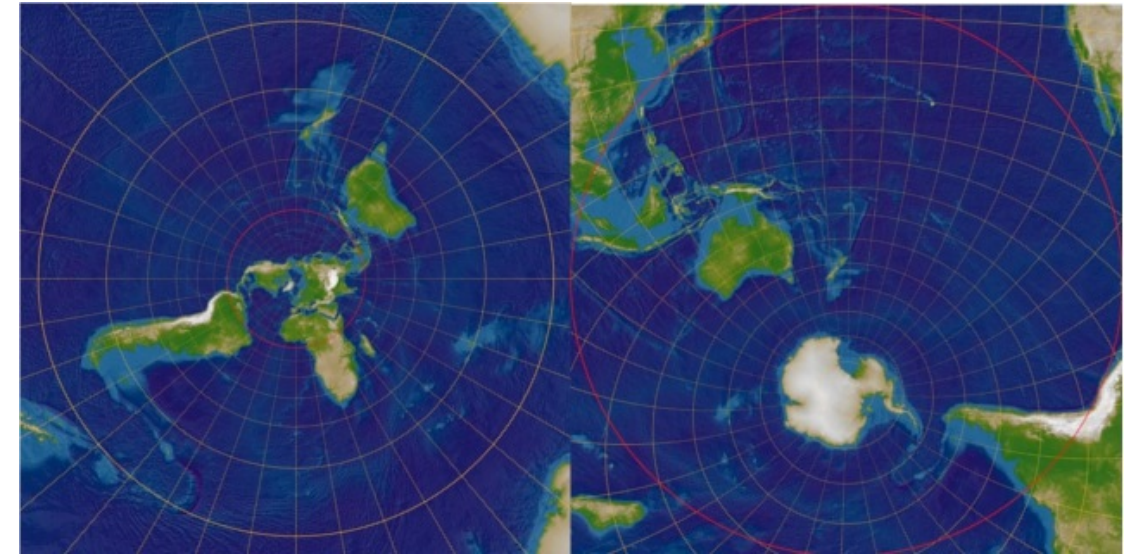
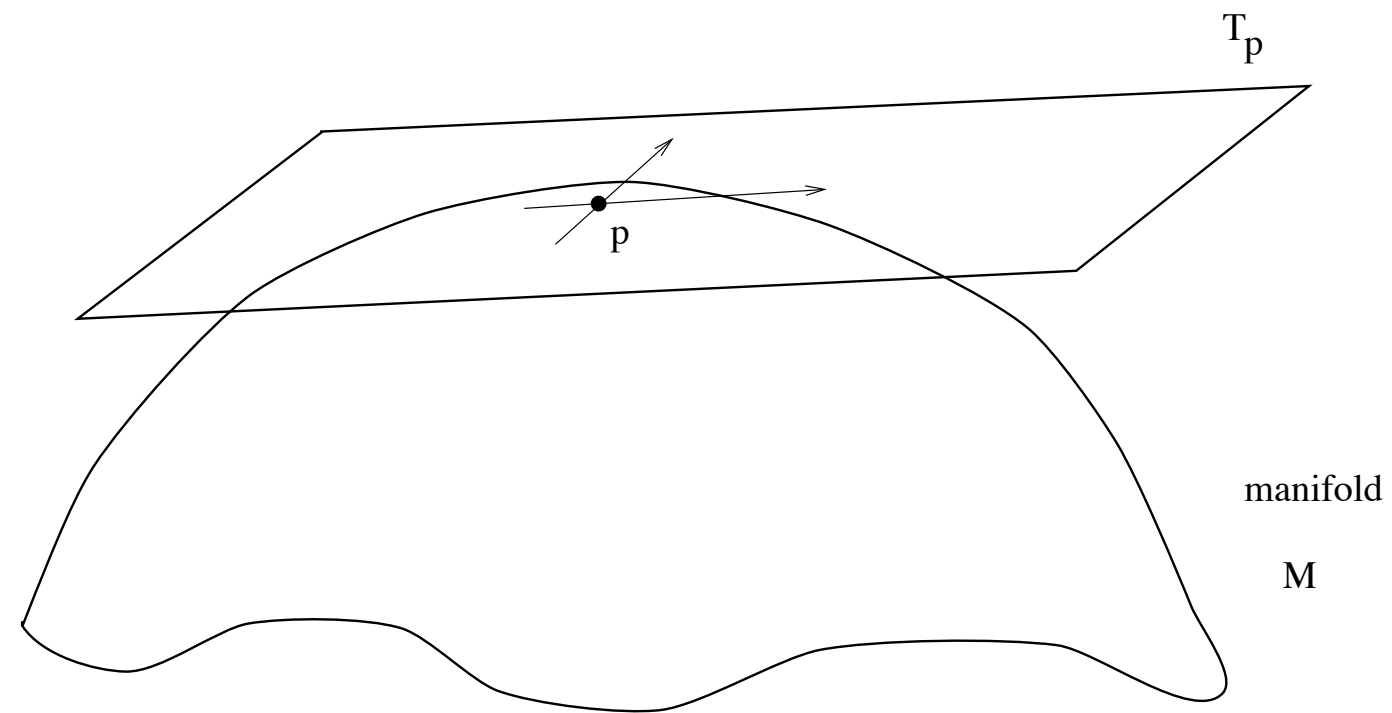
A brief history of Quantum Gravity

- 1916: Einstein points out that quantum effects must modify GR
- 1927: Oskar Klein suggests quantum gravity should modify concepts of spacetime
- 1930s: Rosenfeld applies Pauli method to linearized Einstein field equations
- 1934: Blokhintsev and Gal'perin discuss the spin-two quantum of the gravitational field, naming it the graviton
- 1936: Bronstein realizes need for background independence, issues with Riemannian geometry
- 1938: Heisenburg points out issues with the dimensionality of the gravitational coupling constant for a quantum theory of gravity
- 1949: Peter Bergmann and group begin the canonical approach to quantum gravity. Bryce Dewitt applies Schwinger's covariant quantization method to the gravitational field for his thesis.
- 1952: Gupta develops the covariant approach to quantum gravity.


A brief history of Quantum Gravity


- 1957: Charles Misner introduces the “Feynman quantization of general relativity”, or sum over histories approach. His paper outlines the three approaches, canonical, covariant, and sum over history in modern terms
- 1961: Arnowit, Deser, and Misner complete the ADM formulation of GR
- 1962: Using ADM, Peres writes the Hamilton-Jacobi formulation of GR
- 1964: Feynman starts loop corrections to GR, but notices unitarity was lost for naive Feynman rules. DeWitt adds correction terms involving loops of fictitious fermionic particles, solving the problem.
- 1966: Faddeev and Popov independently come up with DeWitt’s “Faddeev-Popov” ghosts.
- 1967: DeWitt’s seminal paper, completed in 1966 but not published until 1967 due to publishing charges, develops canonical quantum gravity and the Wheeler-DeWitt equation. DeWitt and Feynman complete the Feynman rules for GR.
- 1969: Charles Misner starts quantum cosmology by truncating the Wheeler-DeWitt equation to a finite number of degrees of freedom

Quick intro to GR




Quick intro to GR

Metric $ds^2 = e^{2\lambda} dt^2 - e^{2(\nu-\lambda)} (dr^2 + dz^2) - r^2 e^{-2\lambda} d\phi^2$  **Line Element**

 $g_{\mu\nu} = \begin{pmatrix} e^{2\lambda} & 0 & 0 & 0 \\ 0 & -e^{2(\nu-\lambda)} & 0 & 0 \\ 0 & 0 & -e^{2(\nu-\lambda)} & 0 \\ 0 & 0 & 0 & -r^2 e^{-2\lambda} \end{pmatrix}$


$$R_{\mu\nu} = R^\lambda_{\mu\lambda\nu}$$

$$R = R^\mu_\mu = g^{\mu\nu} R_{\mu\nu}$$

 **Connection**

$$\Gamma^\lambda_{\mu\nu} = \frac{1}{2} g^{\lambda\sigma} (\partial_\mu g_{\nu\sigma} + \partial_\nu g_{\sigma\mu} - \partial_\sigma g_{\mu\nu})$$

$$G_{\mu\nu} \equiv R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = 8\pi G T_{\mu\nu}$$

 **Curvature**

 **Matter**

$$R^\rho_{\sigma\mu\nu} = \partial_\mu \Gamma^\rho_{\nu\sigma} - \partial_\nu \Gamma^\rho_{\mu\sigma} + \Gamma^\rho_{\mu\lambda} \Gamma^\lambda_{\nu\sigma} - \Gamma^\rho_{\nu\lambda} \Gamma^\lambda_{\mu\sigma}$$

Parallel Transport

$$\partial_\mu T_{\mu\nu} = ??$$

Not a tensor

Tensor

$$\nabla_\mu V^\nu = \partial_\mu V^\nu + \Gamma_{\mu\lambda}^\nu V^\lambda$$

Parallel Transport $\rightarrow \frac{D}{d\lambda} = \frac{dx^\mu}{d\lambda} \nabla_\mu = 0$ along $x^\mu(\lambda)$

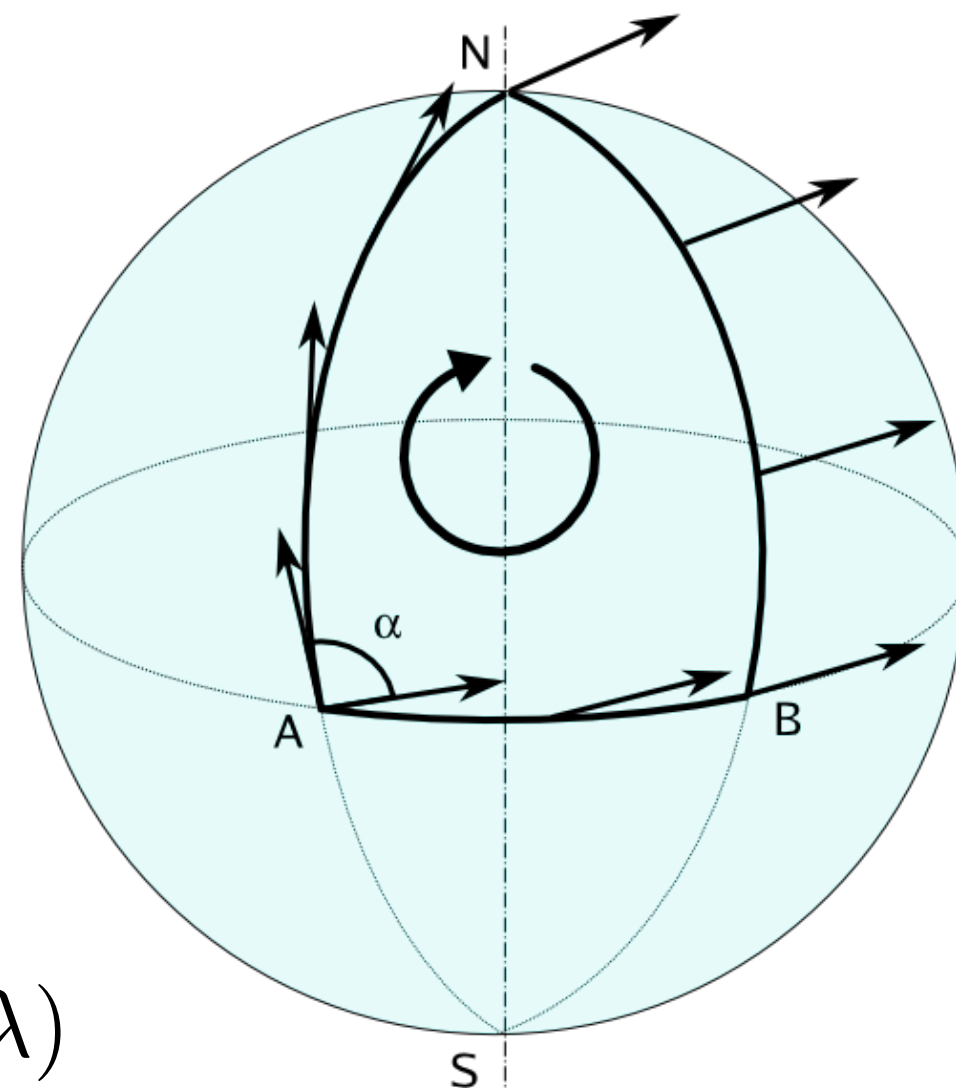
That is, the directional covariant derivative is equal to zero along the curve x^μ parameterized by λ . For a vector this can be simply written as:

$$\nabla_\mu V^\nu = \partial_\mu V^\nu + \Gamma_{\mu\lambda}^\nu V^\lambda = 0$$

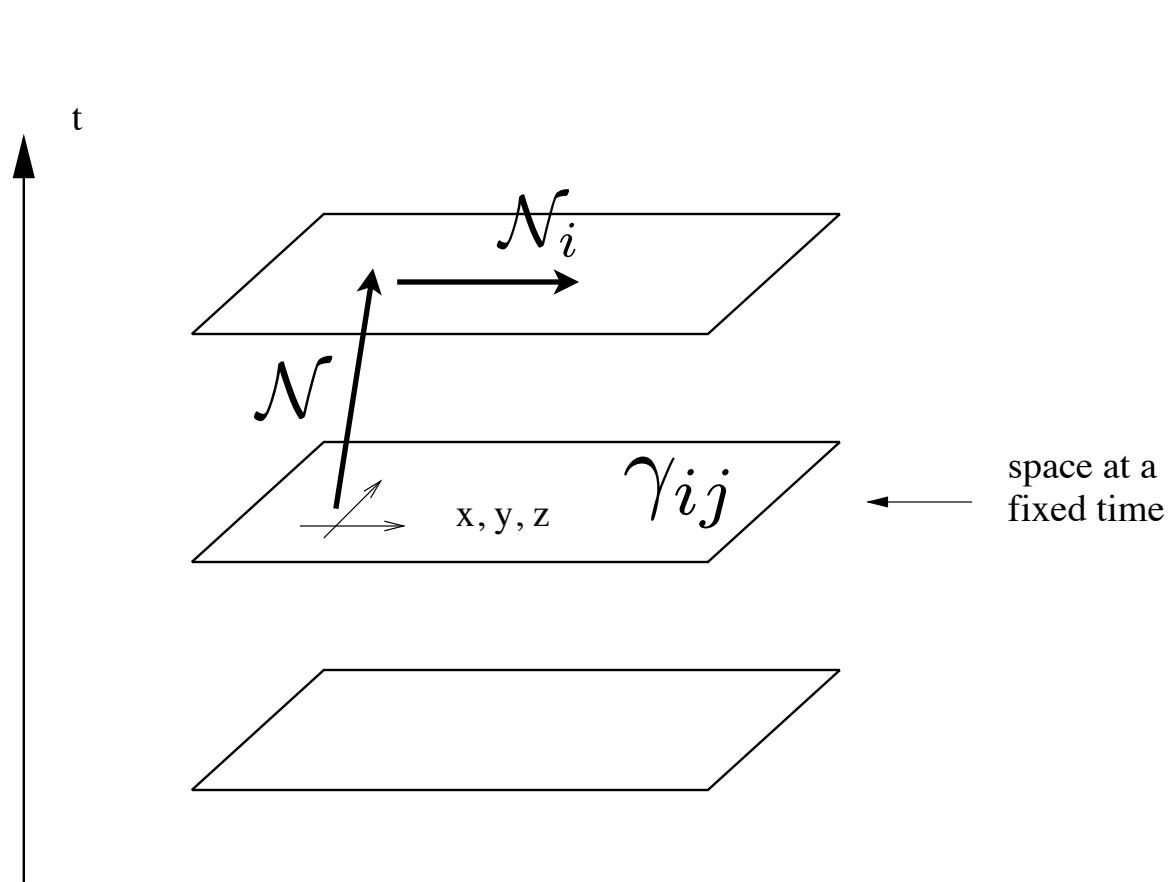
Parallel Transport of the tangent vector

$$\frac{D}{d\lambda} \frac{dx^\mu}{d\lambda} = \frac{d^2 x^\mu}{d\lambda^2} + \Gamma_{\rho\sigma}^\mu \frac{dx^\rho}{d\lambda} \frac{dx^\sigma}{d\lambda} = 0$$

Geodesic equation



ADM 3+1 formalism



$$g_{\mu\nu} = \begin{pmatrix} \mathcal{N} & \mathcal{N}_1 & \mathcal{N}_2 & \mathcal{N}_3 \\ 0 & \gamma_{11} & \gamma_{12} & \gamma_{13} \\ 0 & 0 & \gamma_{22} & \gamma_{23} \\ 0 & 0 & 0 & \gamma_{33} \end{pmatrix}$$

$$\mathcal{N}_i = {}^{(4)}g_{0i}$$

$$\mathcal{N} = \left(-{}^{(4)}g_{00} \right)^{-\frac{1}{2}}$$

Quantum gravity Hamilton-Jacobi equations

$$H + \frac{\partial S}{\partial t} = 0$$

$$S_H = \frac{1}{2\kappa} \int \sqrt{-g} R d^4x$$

$$\hat{g}_{ij}(t, x^k) \rightarrow g_{ij}(t, x^k)$$

$$\hat{\pi}_{ij}(t, x^k) \rightarrow -i \frac{\delta}{\delta g_{ij}(t, x^k)}$$

Wheeler-DeWitt equation

$$\left[-G_{ijkl} \frac{\delta^2}{\delta \gamma_{ij} \delta \gamma_{kl}} - {}^3 R(\gamma) \gamma^{1/2} + 2\Lambda \gamma^{1/2} \right] \Psi[\gamma_{ij}] = 0$$

$$G_{ijkl} = \frac{1}{2} \gamma^{-1/2} (\gamma_{ik} \gamma_{jl} + \gamma_{il} \gamma_{jk} - \gamma_{ij} \gamma_{kl})$$

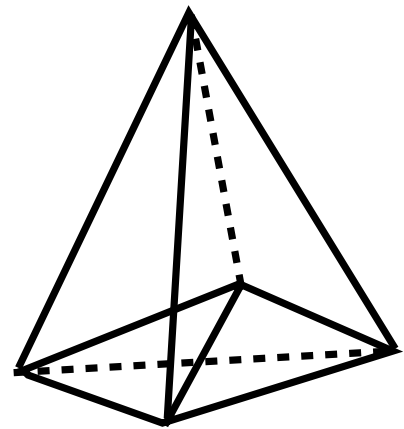
Semiclassical Universe from First Principles

- Ambjorn, Jurkiewicz, and Loll get quantum cosmology without the Wheeler-DeWitt equation using the sum-over-histories approach of Causal Dynamical Triangulations
- They “observe” a bounce, or state of vanishing spatial extension which tunnels to a universe of finite linear extension
- What is Causal Dynamical Triangulations?

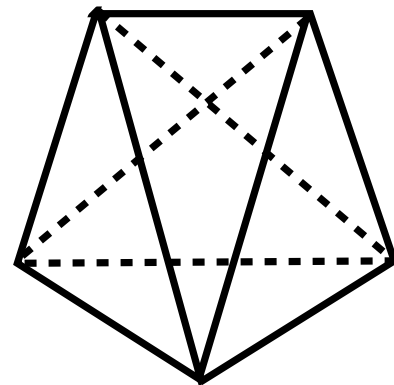
Causal Dynamical Triangulations

- Causal, or Lorentzian, signature for the metric (which avoids problems of degenerate quantum geometries which occur with Euclidean metrics)
- Non-perturbative path integral without a fixed background. Spacetime is determined dynamically as part of the theory
- Spacetime is approximated by tiling flat simplicial building blocks (Regge calculus) without coordinates (simplicial manifold)
 1. Any face of a simplex in T is a simplex in T
 2. Two simplices in T are either disjoint or share a common face
- Well-defined Wick rotation (analytic continuation)

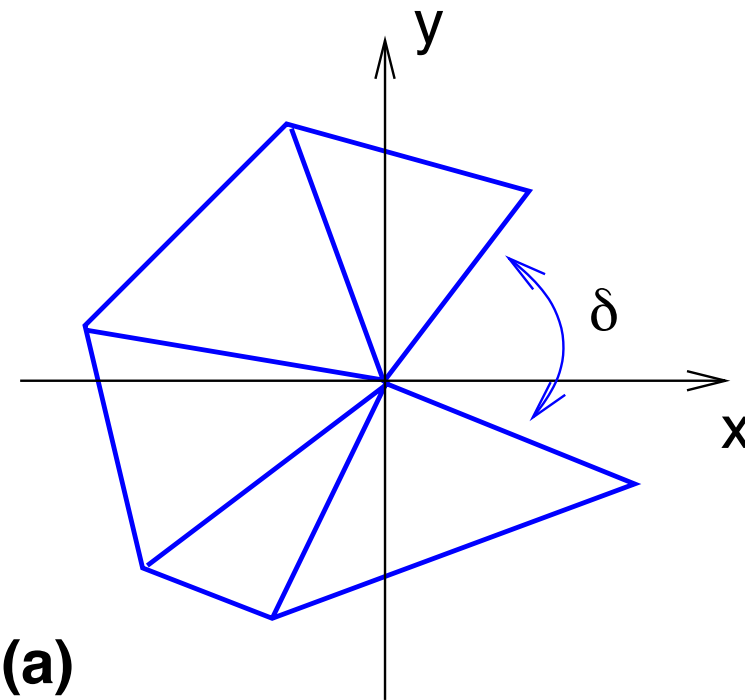
Causal Dynamical Triangulations



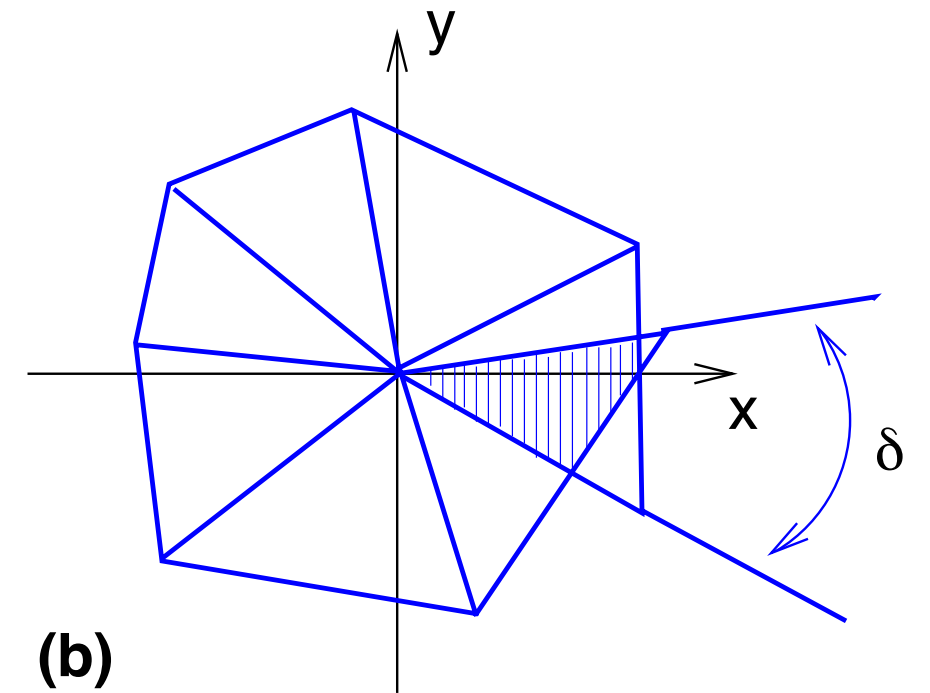
(a)



(b)



(a)



(b)

Name	Dim	0-faces	1-faces	2-faces	3-faces	4-faces	Causal Structure
Vertex	0	1					
Edge	1	2	1				$\{1,1\}$
Triangle	2	3	3	1			$\{2,1\} \{1,2\}$
Tetrahedron	3	4	6	4	1		$\{3,1\} \{2,2\} \{1,3\}$
Pentatope	4	5	10	10	5	1	$\{4,1\} \{3,2\} \{2,3\} \{1,4\}$

Table 1: Types and causal structures of simplices

Causal Dynamical Triangulations

$$G(g_i, g_f; t_i, t_f) = \int \mathcal{D}[g] e^{iS_{EH}}$$

$$S_H = \frac{1}{2\kappa} \int \sqrt{-g} (R - 2\Lambda) d^4x$$

$$S_R = \frac{1}{8\pi G} \sum_h V_h \delta_h - \frac{\Lambda}{8\pi G} \sum_d V_d$$

$$\mathbf{CDT}$$

$$S=\frac{1}{8\pi G}\sum_{space-like\;h}Vol(h)\frac{1}{i}\left(2\pi-\sum_{d-simplices}\Theta\right)-\frac{1}{8\pi G}\sum_{time-like\;h}Vol(h)\left(2\pi-\sum_{d-simplices}\Theta\right)-\frac{\Lambda}{8\pi G}\sum_{time-like}$$

$$\alpha \rightarrow -\alpha, \alpha > \frac{7}{12}$$

$$\int \mathcal{D}[g] e^{iS} \rightarrow \sum_T \frac{1}{C(T)} e^{iS(T)}$$

$$Z=\sum_Texp\left((k_0+6\Delta)N_0-k_4N_4-\Delta(2N_4^{(4,1)}+N^{(3,2)})\right)$$

$$k_0=\frac{\sqrt{3}}{8G}, k_4=\frac{\Lambda\sqrt{5}}{768\pi G}+\frac{\sqrt{3}}{8G}\left(\frac{5}{2\pi}cos^{-1}(1/4)-1\right)$$

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