SN 2020jgb

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ABSTRACT

Keywords: keywords

- 1. INTRODUCTION
- 2. OBSERVATIONS

2.1. Detection and Classification

SN 2020jgb was first discovered by the Zwicky Transient Facility (ZTF; Bellm et al. 2019; Graham et al. 2019) on 2020 May 03.463 UT (MJD 58972.463) with the 48-inch Samuel Oschin Telescope (P48) at Palomar Observatory. The internal designation is ZTF20aayhacx. It was detected at a magnitude of 19.86 in ZTF gband, and J2000 coordinates $\alpha=17^{\rm h}53^{\rm m}12^{\rm s}.651,~\delta=-00^{\circ}51'21''.81$. The last non-detection was on 2020 April 27.477 (MJD 58966.477; 5.99 days before the first detection) up to a limiting magnitude of 20.7 in ZTF r-band.

Classification, ...

2.2. Optical Photometry

We obtained gr-band photometry of SN 2020jgb with the ZTF camera. A Galactic extinction of E(B-V)=0.404 is reported by the maps of (Schlafly & Finkbeiner 2011), for which we correct all our photometry. We do not account for any additional host extinction due to the lack of any Na I D absorption in our spectra (Is it in the outskirt?).

2.3. Optical Spectroscopy

2.4. Near-infrared (NIR) Spectroscopy

We obtained one NIR spectrum of the transient using the Gemini near-infrared spectrometer (GNIRS; Elias et al. 1998) on the Gemini North telescope on 2020 June 9 (\approx 22 days after r-band peak), for an integration time of 2400 s. The spectra were reduced with the PypeIt Python package (Prochaska et al. 2020; Prochaska et al. 2020).

3. ANALYSIS

- 3.1. Photometric Properties
- \bullet sub-luminous
- first light time, peak time
- color evolution
 - 3.2. Spectroscopic Properties
- infrared Ca II triplet (Ca II IRT)
- tentative He I absorption at $\approx 9900 \text{ Å}$
 - 3.3. Optical Spectroscopy
 - 4. HOST GALAXY
 - 5. MODEL COMPARISONS
 - 6. DISCUSSION AND CONCLUSION

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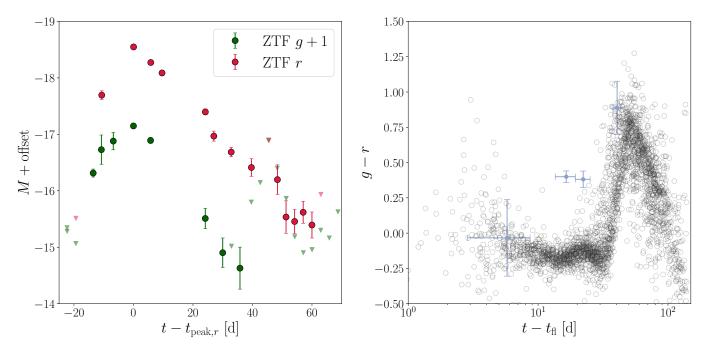


Figure 1. Left: multi-color (ZTF g and r bands) light curves of SN 2020jgb after extinction correction. The inverted triangles are 5- σ upper limits. Right: g-r color evolution of SN 2020jgb (blue dots), accompanied by 62 normal SNe Ia (open circles) with prompt observations within 5 days of first light by ZTF (Bulla et al. 2020). The shaded region denotes the 1- σ credible interval of the color of SN 2020jgb, estimated using Gaussian process.

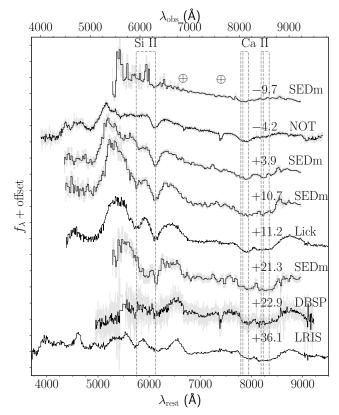


Figure 2. Optical spectroscopic sequence of SN 2020jgb.

Prochaska, J. X., Hennawi, J. F., Westfall, K. B., et al. 2020, Journal of Open Source Software, 5, 2308, doi: 10.21105/joss.02308

Prochaska, J. X., Hennawi, J., Cooke, R., et al. 2020, pypeit/PypeIt: Release 1.0.0, v1.0.0, Zenodo, doi: 10.5281/zenodo.3743493 SN 2020JGB 3

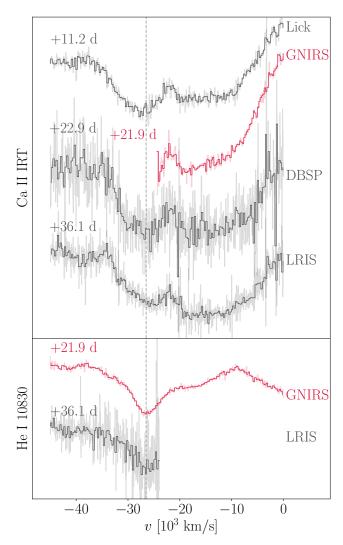


Figure 3. Line velocity comparison between the high-velocity component of Ca II IRT and the infrared absorption feature at ≈ 9900 Å assuming it is associated with He I at 10830 Å.