Explainable Artificial Intelligence in Deep Learning-Based Solar Storm Predictions

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Abstract

A deep learning model is often considered a black-box model, as its internal workings tend to be opaque to the user. Because of the lack of transparency, it is challenging to understand the reasoning behind the model's predictions. Here, we present an approach to making a deep learning-based solar storm prediction model interpretable, where solar storms include solar flares and coronal mass ejections (CMEs). This deep learning model, built based on a long short-term memory (LSTM) network with an attention mechanism, aims to predict whether an active region (AR) on the Sun's surface that produces a flare within 24 hours will also produce a CME associated with the flare. The crux of our approach is to model data samples in an AR as time series and use the LSTM network to capture the temporal dynamics of the data samples. To make the model's predictions accountable and reliable, we leverage post hoc model-agnostic techniques, which help elucidate the factors contributing to the predicted output for an input sequence and provide insights into the model's behavior across multiple sequences within an AR. To our knowledge, this is the first time that interpretability has been added to an LSTM-based solar storm prediction model.

Introduction

Solar storms include solar flares and coronal mass ejections (CMEs). A solar flare is a large explosion in the Sun's atmosphere caused by tangling, crossing, or reorganizing of magnetic field lines. CMEs are intense bursts of magnetic flux and plasma that are ejected from the Sun into interplanetary space (Liu et al., 2020). CMEs are often associated with solar flares and originate from active regions (ARs) in the Sun's photosphere, where magnetic fields are strong and evolve rapidly. Major CMEs and their associated flares can have severe influences on the near-Earth environment, resulting in potentially life-threatening consequences (Baker et al., 2004). Therefore, substantial efforts are being invested in the development of new technologies for the early detection and prediction of flares and CMEs (Bobra and Ilonidis, 2016; Inceoglu et al., 2018).

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Previously, Liu et al. (2020) designed a long short-term memory (LSTM) network and applied the LSTM network to physical features, or predictive parameters, derived from the photospheric vector magnetic field data provided by the Helioseismic and Magnetic Imager (HMI) on board the Solar Dynamics Observatory (SDO) to predict whether an AR that produces a large flare within 24 hours will also produce a CME associated with the flare. In this paper, we extend the work of Liu et al. (2020) by adding interpretability to the LSTM network using two tools, namely SHapley Additive exPlanations (SHAP) and Local Interpretable Modelagnostic Explanations (LIME) (Molnar, 2020). In what follows, we first survey related work. Then we present the results of SHAP, followed by the results of LIME. Finally, we conclude the paper and point out some directions for future work.

Related Work

Interpretable machine learning, or explainable artificial intelligence (XAI), has recently drawn much attention in the AI field as researchers and practitioners seek to provide more transparency to their models. Miller (2019) defined the interpretability of a model as the degree to which an observer can understand the cause of a decision. The terms "interpretable" and "explainable" are often used interchangeably, and the basic idea behind them is the same. One may wonder, when a machine learning model gives a prediction with high accuracy, why should we bother about the reason behind this prediction? The answer to that question is that it depends on what we are trying to predict. If the prediction task is simple and the only thing we are concerned about is the performance, then explanations may not be needed. On the other hand, for certain complex prediction tasks, it is very important and valuable to understand the reason behind a prediction and to be able to explain the prediction (Molnar, 2020).

Recently, XAI has been incorporated into solar flare predictions. For example, Feldhaus and Carande (2021) implemented LIME in a solar flare prediction model built using SVMs. Li et al. (2022) developed a model named FAST-CF, which provides intuitive post-hoc counterfactual explanations for solar flare predictions. Pandey et al. (2023) performed a post hoc analysis on a deep learning-based full-disk solar flare prediction model using three methods, in-

cluding class activation mapping, deep SHAP, and integrated gradients. In contrast to the above studies, our work focuses on incorporating SHAP and LIME into an LSTM network for solar storm predictions using photospheric magnetic field parameters.

Preliminaries

XAI Background

XAI consists of a set of tools and libraries designed to explain predictions for machine learning algorithms. It is commonly used when dealing with black-box models such as LSTM. Its importance lies in the opaque nature of the black-box models, where only the inputs and outputs are known. However, their inner workings remain enigmatic. Incorporating XAI is a key factor in the deployment of responsible AI. XAI helps to ensure that the results of a model can be explained and understood. Adding interpretability to a black-box model will result in beneficial outcomes such as mitigating risk, reducing bias, and increasing trust.

We propose to incorporate SHAP and LIME into the LSTM model that originates from Liu et al. (2020) for solar storm predictions. SHAP handles global explanations, while LIME is dedicated to local explanations. Global explanations are used to understand the model as a whole on the basis of the entire test set. It is useful to see which feature plays the biggest role in the model's output. Local explanations, on the other hand, take a look at specific instances in the test set. Through local explanations, we can better understand each feature and why certain behaviors lead to specific predictions. For both SHAP and LIME, we adopt a random seed function. Specifically, we use np.random.seed(42), where the seed number, 42, is commonly used in the machine learning community.

Data and Feature Space

This study included 33,604 data samples in the training set and 540 data samples in the test set. The dataset size was predetermined by a forecast horizon (24 hrs) to predict whether a solar active region (AR) that produces a flare within 24 hrs will also produce a CME associated with the flare.

We used 12 magnetic field parameters or features, listed in Table 1, from the Space-weather HMI Active Region Patches (SHARPs; Bobra et al., 2014) to make the predictions. Each data sample contains the 12 features, along with a label of "P" indicating that the data sample is in the positive class (i.e., the corresponding AR that produces a flare within 24 hours will produce a CME associated with the flare), or a label of "N" indicating that the data sample is in the negative class (i.e., the corresponding AR that produces a flare within 24 hours will not produce a CME associated with the flare). The 12 features in Table 1 produced the highest mean True Skill Statistics (TSS), which was 0.562, for the 24-hour forecast horizon (Liu et al., 2020).

Global Explanations with SHAP

SHAP incorporates game theory to give each feature a SHAP value. Positive SHAP values have a positive effect,

Table 1: Features Used for Solar Storm Predictions

Feature	Description
TOTUSJZ	Total unsigned vertical current
USFLUX	Total unsigned flux
TOTPOT	Total magnetic free energy density
SAVNCPP	Sum of the net current per polarity
ABSNJZH	Absolute value of net current helicity
MEANPOT	Mean magnetic free energy
MEANSHR	Mean shear angle
SHRGT45	Area fraction with shear $\geq 45^{\circ}$
MEANJZH	Mean current helicity
MEANGAM	Mean angle of field from radial
MEANALP	Mean characteristic twist parameter
MEANGBZ	Mean gradient of vertical field

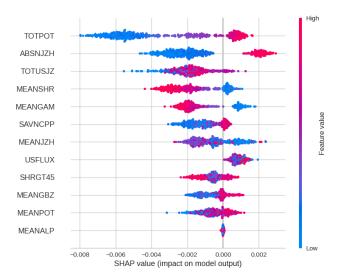


Figure 1: Beeswarm plot to visualize the positive or negative effect of a feature for each test sample, represented by a color dot, on the model's predictions.

while negative SHAP values have a negative effect. A positive effect increases the probability of predicting that a test sample is in the positive class, whereas a negative effect increases the probability of predicting that a test sample is in the negative class. We adopted shap.GradientExplainer in this study.

Figure 1 presents the beeswarm plot for the LSTM model. In the beeswarm plot, we can observe, for each feature, the distribution of SHAP values for all test samples. Each test sample corresponds to a dot in each feature row. The placement of each dot on the X-axis is determined by the SHAP value that the corresponding test sample receives. When dots cluster, it shows common test samples based on their SHAP values. The color of a dot depends on the value of a feature in the corresponding test sample. Red indicates a high feature value, blue indicates a low feature value, and purple indicates an average or moderate feature value.

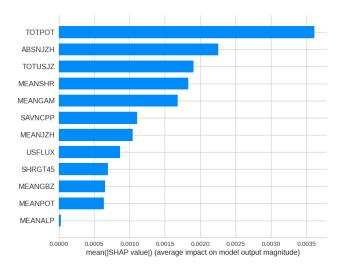


Figure 2: Bar plot to display the global importance of each feature on the model's predictions.

Now, focus on the TOTPOT feature in Figure 1. There are more test samples with negative SHAP values than with positive SHAP values. This means that the TOTPOT feature is more likely to push the model's predictions towards a negative class. Furthermore, low feature values (blue dots) tend to lead to negative predictions, as these low feature values have negative SHAP values. High feature values (red dots) tend to lead to positive predictions, as these high feature values have positive SHAP values.

Next, focus on the MEANALP feature in Figure 1. If we compare the MEANALP feature with the TOTPOT feature, MEANALP may appear to have fewer dots. The reason for this appearance is that most of the SHAP values for MEANALP are zero or near zero. This causes the dots to cluster up and makes MEANALP seem to have fewer dots. Furthermore, since most of the SHAP values for MEANALP are zero or near zero, the MEANALP feature neither pushes the model's predictions towards a negative class nor pushes the model's predictions towards a positive class, implying that the MEANALP feature plays an unimportant role in the model's predictions.

Figure 2 presents the bar plot for the LSTM model. In the bar plot, each feature is given the mean of the absolute SHAP values across all test samples. This mean of the absolute SHAP values is how the importance of each feature is measured in the bar plot. The longer the bar in the bar plot, the more important the corresponding feature is to the model's predictions. We can see in Figure 2 that the TOT-POT feature is of the highest importance to the model's predictions, while the MEANALP feature is of the lowest importance to the model's predictions.

In contrast to the bar plot in Figure 2, the beeswarm plot in Figure 1 displays a separate SHAP value for each test sample, and shows the variability of the test samples, represented by color dots, for each feature. The fewer clusters and the more spread out from the zero in the beeswarm plot indicate higher SHAP values (positive or negative), causing

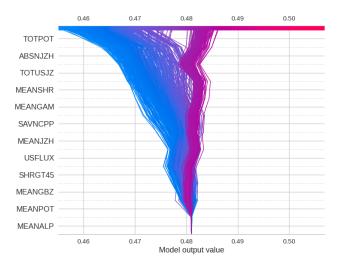


Figure 3: Decision plot to understand how the model produces its predictions.

a higher mean of absolute SHAP values and ultimately a higher importance. Now, consider again the appearance of MEANALP in the beeswarm plot in Figure 1. The result of having most of its SHAP values equal or near zero reflects how short its bar is in the bar plot in Figure 2. On the other hand, the nature of high variability of the TOTPOT feature in the beeswarm plot in Figure 1 causes its bar to be significantly longer than those of the other features in the bar plot in Figure 2.

Figure 3 presents the decision plot for the LSTM model. On the Y-axis of the decision plot, the features are displayed, from top to bottom, based on their importance, with the most important feature displayed at the top and the least important feature displayed at the bottom. Each prediction/test sample is represented by a line in the decision plot. The predictions for the test samples begin at the bottom with the same base value, which is approximately 0.48. This base value is the mean of the model's prediction values across all data samples in the training set. It is used as a starting point before considering any feature contribution. As a prediction/line moves from bottom to top, the SHAP value for each feature is added to the base value. This can help to understand the contribution of each feature in the prediction. The features can have positive or negative contributions, pushing the corresponding line to the right or left, respectively.

The prediction value (i.e., the model output value) of a test sample is the probability that the test sample belongs to the positive class. For example, a prediction value of 0.4 indicates that there is a 40% chance that the test sample belongs to the positive class. At the top, each prediction/line reaches its prediction value. The prediction value determines the corresponding line color of the prediction. A blue line indicates a low prediction value, while a red line indicates a high prediction value with respect to the base value. Specifically, the features in blue reduce the model output value compared to the base value. The deeper the blue, the stronger the negative contribution of that feature. The features in red increase

the model's output value compared to the base value. The deeper the red, the stronger the positive contribution of that feature.

The results shown in the beeswarm, bar, and decision plots are consistent. Based on these results, we conclude that the TOTPOT feature is the most important feature in the model's predictions, while the MEANALP feature has the least importance in the model's predictions. Other features such as MEANPOT and MEANGBZ are indicated as features of low importance, while features such as ABSNJZH and TOTUSJZ are of high importance.

Local Explanations with LIME

Local explanations look at individual predictions made by the LSTM model. With LIME, we can better understand each feature and why certain behaviors lead to a specific prediction. Another characteristic of LIME is that it is model-agnostic. This means that it can be applied to any machine learning model dealing with any input data such as text, image, or, in the case of our study, multivariate time series. The flexibility of the LIME tool allows it to be implemented no matter the complexity of the inner workings of a deep learning model. We adopted lime.lime_tabular.LimeTabularExplainer in this study.

Figure 4 presents the LIME plot for a test sample predicted to be in the positive class. Figure 5 presents the LIME plot for a test sample predicted to be in the negative class. The X-axis shows LIME values. The features on the Y-axis are ranked from highest to lowest importance. We need to distinguish the negative and positive bars with respect to the LIME values on the X-axis. When a feature favors one side over the other, it means that the model is identifying the class to which it believes the test sample should belong. Negative LIME values fall into the negative class, while positive LIME values fall into the positive class.

Consider, for example, Figure 4. Most of the features in the figure have positive bars and the positive magnitude outweighs the negative magnitude. Thus, the model predicts that the test sample in Figure 4 is in the positive class. On the other hand, in Figure 5, most of the features have negative bars and the negative magnitude outweighs the positive magnitude, which strongly contributes to the negative class. Consequently, the model predicts that the test sample in Figure 5 is in the negative class.

In both Figures 4 and 5, each feature on the Y-axis is associated with a range. For example, in Figure 4, ABSNJZH with a positive bar has a range greater than 1.34; this means that ABSNJZH values greater than 1.34 increase the chances that a test sample belongs to the positive class. On the other hand, in Figure 5, TOTUSJZ with a negative bar has a range less than or equal to -0.81; this means that TOTUSJZ values less than or equal to -0.81 decrease the chances that a test sample belongs to the positive class (or increase the chances that the test sample belongs to the negative class).

When comparing the LIME plots to the SHAP plots, it is important to recall that SHAP is designed for global explanations considering all data samples in the test set, while LIME is designed for local explanations inspecting specific test samples. However, comparing the two plots can provide

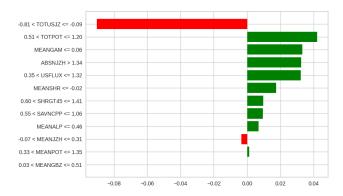


Figure 4: LIME plot for a test sample predicted to be in the positive class.

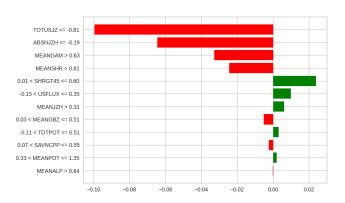


Figure 5: LIME plot for a test sample predicted to be in the negative class.

valuable insight to identify possible overlaps in importance of the features. TOTUSJZ, for example, is of high importance in both LIME plots and SHAP plots.

Conclusion and Future Work

In this paper, we present an interpretable LSTM model for solar storm predictions by incorporating two XAI tools (SHAP and LIME) into the model. The SHAP tool helps us understand the features that have the largest impact on the overall predictions of the model. The LIME tool helps us explore and explain the individual predictions made by the model.

In the future we plan to extend our work to build interpretable deep learning models for other space weather events such as solar energetic particles, interplanetary shocks and geomagnetic storms. Another direction of future work includes the quantitative evaluation of the XAI tools. Quantitative metrics help to move away from subjective evidence in assessing interpretability (Nauta et al., 2023). Many metrics have been suggested in the literature (Pawlicki et al., 2024), including AUC, cover, disagreement, infidelity, and monotonicity metric, to name a few. We plan to explore these metrics in the context of space weather forecasting.

References

- Baker, D. N.; Daly, E.; Daglis, I.; Kappenman, J. G.; and Panasyuk, M. 2004. Effects of space weather on technology infrastructure. *Space Weather* 2:S02004.
- Bobra, M. G., and Ilonidis, S. 2016. Predicting coronal mass ejections using machine learning methods. *The Astrophysical Journal* 821:127.
- Bobra, M. G.; Sun, X.; Hoeksema, J. T.; and et al. 2014. The Helioseismic and Magnetic Imager (HMI) vector magnetic field pipeline: SHARPs space-weather HMI active region patches. *Solar Physics* 289(11):3549–3578.
- Feldhaus, C., and Carande, W. 2021. Explainable artificial intelligence for solar flare prediction. In *AGU Fall Meeting Abstracts*, volume 2021, NG45B–0576.
- Inceoglu, F.; Jeppesen, J. H.; Kongstad, P.; Hernández Marcano, N. J.; Jacobsen, R. H.; and Karoff, C. 2018. Using machine learning methods to forecast if solar flares will be associated with CMEs and SEPs. *The Astrophysical Journal* 861:128.
- Li, P.; Bahri, O.; Boubrahimi, S. F.; and Hamdi, S. M. 2022. Fast counterfactual explanation for solar flare prediction. In *Proceedings of the 21st IEEE International Conference on Machine Learning and Applications*, 1238–1243.
- Liu, H.; Liu, C.; Wang, J. T. L.; and Wang, H. 2020. Predicting coronal mass ejections using SDO/HMI vector magnetic data products and recurrent neural networks. *The Astrophysical Journal* 890(1):12.
- Miller, T. 2019. Explanation in artificial intelligence: Insights from the social sciences. *Artificial intelligence* 267:1–38.
- Molnar, C. 2020. Interpretable Machine Learning. lulu.com.
- Nauta, M.; Trienes, J.; Pathak, S.; Nguyen, E.; Peters, M.; Schmitt, Y.; Schlötterer, J.; van Keulen, M.; and Seifert, C. 2023. From anecdotal evidence to quantitative evaluation methods: A systematic review on evaluating explainable AI. ACM Comput. Surv. 55(13s):295:1–295:42.
- Pandey, C.; Angryk, R. A.; Georgoulis, M. K.; and Aydin, B. 2023. Explainable deep learning-based solar flare prediction with post hoc attention for operational forecasting. In *Proceedings of the 2023 International Conference on Discovery Science*, 567–581.
- Pawlicki, M.; Pawlicka, A.; Uccello, F.; Szelest, S.; D'Antonio, S.; Kozik, R.; and Choras, M. 2024. Evaluating the necessity of the multiple metrics for assessing explainable AI: A critical examination. *Neurocomputing* 602:128282.