Quantum field theory I

ULB MA | 2023–2024 | Prof. Petr TINIAKOV

Chapter 1: Classical Field

Handwritten notes (scanned)

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Heads up: only Chapter 1 here. This DocHub upload contains only the first chapter. The full set of chapters, personal notes, exercise corrections, and a reference-book list are on my website.

- All chapters: see the course page
- Exercise corrections & personal work: see the main page.
- Reference books: see the book section.

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INTRODUCTION TO QFT Peter Tinyakov-N7.112-PHYS-F410

- Remark on units: We're going to set h = c = 1 (natural units) α: [L]=[T]=eV-1 [LT-1]=eV° [M] = [LT-2] = eV

M CLASSICAL FIELDS

-> example: electromagnetic fields E(x,t), B(x,t). They're given by the potentials Am (x,t) = (q, Ai). It's then a vector field.

- Constructing Lagrangians: → Hermitean A+= A (vi, A v) = (A vi, v) A square natrix is Hermitian () it is unitarity diagonalisable

with real eigenvalues A= UDU+ with UU+= 11 -> Lorentz invariance AAV April is a Lovetz scalar for ex.

- egs not higher the 2nd order in time - P energy bounded from below

Consider a scalar field

-> In general, the Lagrangian density is the following:

R = 1 2 p & 2 m & - V(\$) where V(\$) = E x o (x) (Lorentz invariant)

· Equations of motion:

→ We have to solve SS=0 with S= Sdt L=Sd*x R

-> SS= S[+ Sq]-S[d]

= 194x { \frac{5}{7} \(\phi \) \(\ph

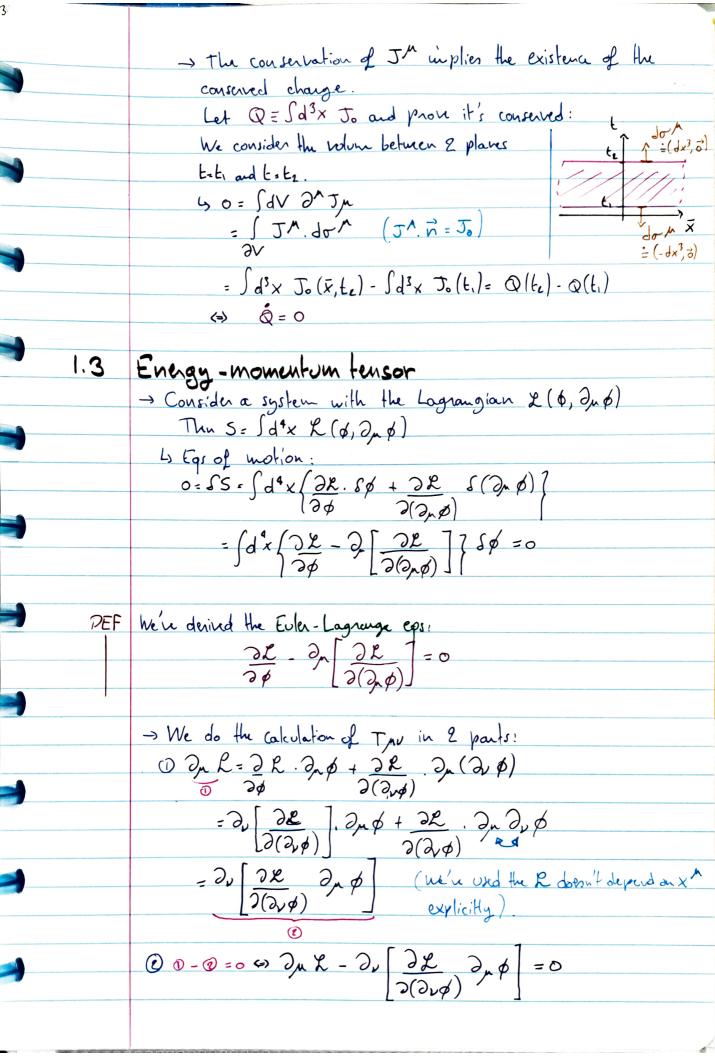
= \d'x \fi \frac{1}{2}, \phi \gamma^{\phi} \phi \quad \quad

= J11x { 2, domsø - V'(ø) sø}

= [dox f-2m omp. Sp - V'(p) Sp + Sp + Sp + Sm [2m 6 m 6] }

20 20 0 + V'(\$) = 6

Lo By choosing V= 2 m2 62, we get the Klein-gerden eq.: 2 2 pm + m2 p = 0 (Similar to eq. for harmonic oscillator 32x+w2x=0) Li Free field: wich satisfy linear equations (no interaction) 1.2 Symmetries and conservation laws: Noether thm - Conserved ofter play an important role. There is a deep relation between those and symmetries of the theory. Example: complex scalar field: → Let the action be: S= Sd*x { 2mp*2mp-V(p* \$)} with \$ = \$ tipe Lo The action is irraniant under \$ De ix. This is a global transformation because & doesn't depend on x. F Li Proof of the invariance: φ* +> φ*e-iα, φ*φ -> φ*e-iα φ eia = φ*φ 2, o * 2 m o H o > 2, o * e - ix 2 m o e ix = 2, o * 2 m o B DEF This is the U(1) symmetry. Lets consider a local transformation $\phi \mapsto \phi e^{i\kappa(x)} = \phi(1+i\kappa(x))^{+o(x^2)}$ > SS= Sd*x!(),(\$e-ix) 2^(6eix)-10,000 0 x of? = (d4x 13" (Hix) \$ 0 (Hix) \$ - = 5 2 \$ 2 4 } = 1 Sd*x f () p + -ia m p + -i ma. px) B Sd4x { in 2, 6 + 0 n 0 + i 2, 0 + 2 mx. \$ -ia 2 p & m & - i 2 mx . 6 x 2 mp } SS = \don'x \G; \don'x (\man \par \par * \par - \par * \don' \par \par) \f We then set Jn = -i (2n pt. p - pt 2m p) and we get SS = Jd x { 2m x , Jn } In gunal, Ju = Jr (6, 20, 20, -) -> SS has to vanish if fields satisfy eas of motion: 85 = - Sdax 600 x. Ju? = + Jdtx x(x). Dh Jh = 0 (=) 2/1/n =0



Recall: 22 - 2, [20,0). 2,0] =0 (3) 2, (SmR) - Du [DR . Dmp] =0 (=) 2 [S / 2 - 2e . 2 p f] = 0 (=) 2 [DL 2 p f - 1 p L] = 0 We define the energy-momentum tensor Tox as: Tru = 22 2/2 / 2/2 / - May & By construction, D' Tru = 0. It's the conservation of the energy - impulsion. → By the divergence theorem, we've 4 conserved quantities: The 4-vector energy-mountum Pu is defined by DEF Pm = Jd3x Tno They are integrals of motion. We have: E = Jd3x Too (then Too is the energy density) and Pi = Sd3x Tio is the momentum. 1 Noether theorem. THM Let the action of a system be invariant under a continuous set of transformations with N parameter wa such that for small wa: $x^{M} \rightarrow x^{2}M = x^{M} + x^{M} = x^{M} + x^{$ $\phi_i(x) \rightarrow \phi_i'(x') = \phi_i(x) + \mathcal{V}_{ia} \omega_a + \dots$ Then the following Nouverts are conserved: Ja= 22 (2) p. x - +:) - x . 2 -> Example: 1) Space-time translations: x'm=xm+wm; p.(x)=p.(x) (scolar field) We've: X "v = SMV and His =0 The conserved cornect is: TMU = DR ON PE - SMUR = TAN >(Dupi)

@ Auguler mouentour and spin: -> Consider Lorentz transformations: x'M = x M + w MU XV Ly Recall: Lie algebra of SO(1,3), developped oround the identity elevent of the Lorentz group: Mr= Smr + w Mr + & (we) with Pro = Trp 1 x 1 p v -> Tru = Pro (Sx + w x) (Sx + w v) + o(w) => 1 = 1 mv + 1 m w b + 1 m m = 1 mv + wmv + wm m(n-1)/2 | = 6 independent parameters of transformation. Lo Our parameter are $\omega^{\mu\nu}$ s.t. $\mu>\nu$ $S \times^{\mu} = \omega^{\mu\nu} \times_{\nu} = \chi^{\mu} \times_{\nu} \omega^{\alpha}$ $= \sum_{\lambda < \rho} \chi^{\mu} \times_{\nu} \omega^{\lambda\rho} \quad \text{with } \lambda < \rho \quad \alpha = (\mu\nu), \mu < \nu$ b= = wh (sh xp-sh xx) Jaclov 12 9 We find: Xxp = Sixp - Sipxx Ly Because the field is realar! Sφ:=0 → V:(Ap) = 0 Ly The conserved corrent is then: Mary = DR Dr di (Sxxp-Spxx)-(shxp-spxx)& = 22 2, 0; xp - 5 12 xp & -(2 P Dp x 2 - 5 m x2 P) = The xp-The xx -> Mip are 6 consured corrects, and therefore me'u o changes: Sd3x Moxp. For lip = i,; the conserved pty is the orbital momentum: (d3x {Pi x; -P; xi}=IL= feish Mij dr3 DEF

DEF The spin is
$$S_{Ap}^{(0)} = \int_{A}^{3} \left(\frac{\partial \mathcal{L}}{\partial A} \right) \left(\frac{\partial \mathcal{L}}{\partial A} \right)$$

Let's consider how a vector field A^{M}

I have spin and $A_{A}^{0} = \int_{A}^{3} \left(\frac{\partial \mathcal{L}}{\partial A} \right) \left(\frac{\partial \mathcal{L}}{\partial A} \right) \left(\frac{\partial \mathcal{L}}{\partial A} \right)$

Thus,

$$A^{M} = \int_{A}^{(0)} \left(\frac{\partial \mathcal{L}}{\partial A} \right) \left(\frac{\partial \mathcal{L}}{\partial A} \right)$$

The spin is $S_{Ap}^{M} = \int_{A}^{3} \left(\frac{\partial \mathcal{L}}{\partial A} \right) \left(\frac{\partial \mathcal{L}}{\partial A} \right)$

Or $S_{Ap} = \int_{A}^{3} \left(\frac{\partial \mathcal{L}}{\partial A} \right) \left(\frac{\partial \mathcal{L}}{\partial A} \right)$

And the vector of spin is $A_{A} = \frac{\partial \mathcal{L}}{\partial A} \cdot \frac{\partial \mathcal{L}}{\partial A} = \frac{\partial \mathcal{L}}{\partial A} = \frac{\partial \mathcal{L}$

2 QUANTIZATION OF A FREE SCALAR FIELD

2.1 Reminder: quantization in QM