

INTRODUCTION TO QFT

Peter Tinyakov - N2.112 - PHYS-F410

→ Remark on units:

We're going to set $\hbar = c = 1$ (natural units)

ex: $[L] = [T] = eV^{-1}$ $[LT^{-1}] = eV^0$

$[M] = [LT^{-2}] = eV$

1 CLASSICAL FIELDS

→ example: electromagnetic fields $\vec{E}(\vec{x}, t)$, $\vec{B}(\vec{x}, t)$. They're given by the potentials $A^\mu(\vec{x}, t) \equiv (\varphi, A^i)$. It's then a vector field.

→ Constructing Lagrangians:

→ Hermitian $A^\dagger = A \Leftrightarrow \langle \vec{u}, A \vec{v} \rangle = \langle A \vec{u}, \vec{v} \rangle$

A square matrix is Hermitian \Leftrightarrow it is unitarily diagonalizable with real eigenvalues $A = U D U^\dagger$ with $U U^\dagger = \mathbb{1}$

→ Lorentz invariance $A^{\mu\nu} A_{\mu\nu}$ is a Lorentz scalar for ex.

→ eqs not higher the 2nd order in time

→ energy bounded from below

1.1 Consider a scalar field

→ In general, the Lagrangian density is the following:

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - V(\phi)$$

where $V(\phi) = \sum_i \kappa_i \phi^i(x)$ (Lorentz invariant)

⊙ Equations of motion:

→ We have to solve $\delta S = 0$ with $S \equiv \int dt L = \int d^4x \mathcal{L}$

$$\rightarrow \delta S = S[\phi + \delta\phi] - S[\phi]$$

$$= \int d^4x \left\{ \frac{1}{2} \partial_\mu (\phi + \delta\phi) \partial^\mu (\phi + \delta\phi) - V(\phi + \delta\phi) - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi + V(\phi) \right\}$$

$$= \int d^4x \left\{ \frac{1}{2} \partial_\mu \phi \partial^\mu \phi + \partial_\mu \phi \partial^\mu \delta\phi - V(\phi) - \delta\phi \frac{\partial V}{\partial \phi} - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi + V(\phi) \right\} + \mathcal{O}(\delta\phi^2)$$

$$= \int d^4x \left\{ \partial_\mu \phi \partial^\mu \delta\phi - V'(\phi) \delta\phi \right\}$$

$$= \int d^4x \left\{ -\partial_\mu \partial^\mu \phi \cdot \delta\phi - V'(\phi) \delta\phi \right\} + \int d^4x \left\{ \partial_\mu [\partial^\mu \phi \delta\phi] \right\}$$

$$\Leftrightarrow \partial_\mu \partial^\mu \phi + V'(\phi) = 0$$

↳ By choosing $V = \frac{1}{2} m^2 \phi^2$, we get the Klein-Gordon eq.:

$$\partial_\mu \partial^\mu \phi + m^2 \phi = 0$$

(Similar to eq. for harmonic oscillator $\partial_x^2 x + \omega^2 x = 0$)

↳ Free field: which satisfy linear equations (no interaction)

1.2 Symmetries and conservation laws: Noether thm

→ Conserved qties play an important role. There is a deep relation between those and symmetries of the theory.

③ Example: complex scalar field:

→ Let the action be: $S = \int d^4x \{ \partial_\mu \phi^* \partial^\mu \phi - V(\phi^* \phi) \}$

with $\phi \triangleq \phi_1 + i\phi_2$

↳ The action is invariant under $\phi \mapsto \phi e^{i\alpha}$. This is a global transformation because α doesn't depend on x^μ .

↳ Proof of the invariance:

$$\phi^* \mapsto \phi^* e^{-i\alpha}; \quad \phi^* \phi \mapsto \phi^* e^{-i\alpha} \phi e^{i\alpha} = \phi^* \phi$$

$$\partial_\mu \phi^* \partial^\mu \phi \mapsto \partial_\mu \phi^* e^{-i\alpha} \partial^\mu \phi e^{i\alpha} = \partial_\mu \phi^* \partial^\mu \phi \quad \square$$

DEF This is the $U(1)$ symmetry.

↳ Lets consider a local transformation $\phi \mapsto \phi e^{i\alpha(x)} = \phi(1 + i\alpha(x)) + \mathcal{O}(\alpha^2)$

$$\rightarrow \delta S = \int d^4x \left\{ \frac{1}{2} (\partial_\mu (\phi^* e^{-i\alpha}) \partial^\mu (\phi e^{i\alpha}) - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi \right\}$$

$$= \int d^4x \left\{ \frac{1}{2} \partial_\mu (i\alpha) \phi^* \partial^\mu (1 + i\alpha) \phi - \frac{1}{2} \partial_\mu \phi^* \partial^\mu \phi \right\}$$

$$= \frac{1}{2} \int d^4x \left\{ (\partial_\mu \phi^* - i\alpha \partial_\mu \phi^* - i\partial_\mu \alpha \cdot \phi^*) \right.$$

$$\left. (\partial_\mu \phi + i\alpha \partial_\mu \phi + i\partial_\mu \alpha \cdot \phi) \right\} - \frac{1}{2} \int d^4x \partial_\mu \phi^* \partial^\mu \phi \}$$

$$\Rightarrow \int d^4x \left\{ i\alpha \partial_\mu \phi^* \partial^\mu \phi + i\partial_\mu \phi^* \partial^\mu \alpha \cdot \phi \right. \\ \left. - i\alpha \partial_\mu \phi^* \partial^\mu \phi - i\partial_\mu \alpha \cdot \phi^* \partial^\mu \phi \right\}$$

$$\delta S = \int d^4x \left\{ i\partial^\mu \alpha (\partial_\mu \phi^* \cdot \phi - \phi^* \partial_\mu \phi) \right\}$$

We then set $T_\mu \equiv -i (\partial_\mu \phi^* \cdot \phi - \phi^* \partial_\mu \phi)$ and we get

$$\delta S = \int d^4x \left\{ \partial^\mu \alpha \cdot T_\mu \right\}$$

In general, $T_\mu = T_\mu(\phi, \partial\phi, \partial^2\phi, \dots)$

→ δS has to vanish if fields satisfy eqs of motion:

$$\delta S = - \int d^4x \left\{ \partial^\mu \alpha \cdot T_\mu \right\}$$

$$= + \int d^4x \alpha(x) \cdot \partial^\mu T_\mu = 0$$

$$\Leftrightarrow \partial^\mu T_\mu = 0$$

→ The conservation of J^μ implies the existence of the conserved charge.

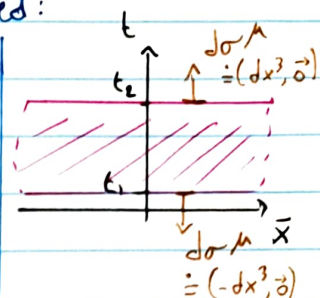
Let $Q \equiv \int d^3x J_0$ and prove it's conserved:

We consider the volume between 2 planes

$t=t_1$ and $t=t_2$.

$$\hookrightarrow 0 = \int dV \partial^\mu J_\mu$$

$$= \int \frac{dV}{\partial V} J^\mu \cdot d\sigma_\mu \quad (J^\mu \cdot \vec{n} = J_0)$$



$$= \int d^3x J_0(\vec{x}, t_2) - \int d^3x J_0(\vec{x}, t_1) = Q(t_2) - Q(t_1)$$

$$\Rightarrow \dot{Q} = 0$$

1.3 Energy-momentum tensor

→ Consider a system with the Lagrangian $\mathcal{L}(\phi, \partial_\mu \phi)$

$$\text{Then } S = \int d^4x \mathcal{L}(\phi, \partial_\mu \phi)$$

↳ Eqs of motion:

$$0 = \delta S = \int d^4x \left\{ \frac{\partial \mathcal{L}}{\partial \phi} \delta \phi + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \delta (\partial_\mu \phi) \right\}$$

$$= \int d^4x \left\{ \frac{\partial \mathcal{L}}{\partial \phi} - \partial_\mu \left[\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \right] \right\} \delta \phi = 0$$

DEF We've derived the Euler-Lagrange eqs:

$$\frac{\partial \mathcal{L}}{\partial \phi} - \partial_\mu \left[\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \right] = 0$$

→ We do the calculation of $T_{\mu\nu}$ in 2 parts:

$$\textcircled{1} \quad \partial_\mu \mathcal{L} = \frac{\partial \mathcal{L}}{\partial \phi} \cdot \partial_\mu \phi + \frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \cdot \partial_\mu (\partial_\nu \phi)$$

$$= \partial_\nu \left[\frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \right] \cdot \partial_\mu \phi + \frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \cdot \partial_\mu \partial_\nu \phi$$

$$= \underbrace{\partial_\nu \left[\frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \right] \partial_\mu \phi}_{\textcircled{2}} \quad (\text{we've used the } \mathcal{L} \text{ doesn't depend on } x^\mu \text{ explicitly}).$$

$$\textcircled{1} - \textcircled{2} = 0 \Leftrightarrow \partial_\mu \mathcal{L} - \partial_\nu \left[\frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \right] \partial_\mu \phi = 0$$

Recall: $\partial_\nu \mathcal{L} - \partial_\nu \left[\frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \cdot \partial_\mu \phi \right] = 0$

$$\Leftrightarrow \partial_\nu (\delta^\nu_\mu \mathcal{L}) - \partial_\nu \left[\frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \cdot \partial_\mu \phi \right] = 0$$

$$\Leftrightarrow \partial_\nu \left[\delta^\nu_\mu \mathcal{L} - \frac{\partial \mathcal{L}}{\partial (\partial_\nu \phi)} \cdot \partial_\mu \phi \right] = 0$$

$$\Leftrightarrow \partial^\nu \left[\frac{\partial \mathcal{L}}{\partial (\partial^\nu \phi)} \partial_\mu \phi - \eta_{\mu\nu} \mathcal{L} \right] = 0$$

DEF

We define the energy-momentum tensor $T_{\mu\nu}$ as:

$$T_{\mu\nu} = \frac{\partial \mathcal{L}}{\partial (\partial^\nu \phi)} \partial_\mu \phi - \eta_{\mu\nu} \mathcal{L}$$

By construction, $\partial^\nu T_{\mu\nu} = 0$. It's the conservation of the energy-impulsion.

→ By the divergence theorem, we've 4 conserved quantities:

DEF

The 4-vector energy-momentum P_μ is defined by

$$P_\mu \equiv \int d^3x T_{\mu 0}$$

They are integrals of motion. We have:

$E = \int d^3x T_{00}$ (then T_{00} is the energy density) and

$P_i = \int d^3x T_{i0}$ is the momentum.

① Noether theorem.

THM

Let the action of a system be invariant under a continuous set of transformations with N parameters w_a such that for small w_a :

$$x^\mu \rightarrow x'^\mu = x^\mu + \chi^\mu_a w_a + \mathcal{O}(w_a^2)$$

$$\phi_i(x) \rightarrow \phi'_i(x') = \phi_i(x) + \psi_{ia} w_a + \dots$$

Then the following N currents are conserved:

$$J^\mu_a = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_i)} (\partial_\nu \phi_i \cdot \chi^\nu_a - \psi_{ia}) - \chi^\mu_a \cdot \mathcal{L}$$

→ Example:

① Space-time translations:

$$x'^\mu = x^\mu + \omega^\mu ; \phi'_i(x') = \phi_i(x) \text{ (scalar field)}$$

$$\text{We've: } \chi^\mu_\nu = \delta^\mu_\nu \text{ and } \psi_{ia} = 0$$

The conserved current is:

$$J^{\mu\nu} = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_i)} \partial_\nu \phi_i - \delta^\mu_\nu \mathcal{L} \equiv T^{\mu\nu}$$

② Angular momentum and spin:

→ Consider Lorentz transformations: $x'^{\mu} = x^{\mu} + \omega^{\mu\nu} x_{\nu}$

↳ Recall: Lie algebra of $SO(1,3)$, developed around the identity element of the Lorentz group:

$$\Lambda^{\mu}{}_{\nu} = \delta^{\mu}{}_{\nu} + \omega^{\mu}{}_{\nu} + \mathcal{O}(\omega^2) \quad \text{with}$$

$$\eta_{\mu\nu} = \eta_{\alpha\beta} \Lambda^{\alpha}{}_{\mu} \Lambda^{\beta}{}_{\nu} \rightarrow \eta_{\mu\nu} = \eta_{\alpha\beta} (\delta^{\alpha}{}_{\mu} + \omega^{\alpha}{}_{\mu}) (\delta^{\beta}{}_{\nu} + \omega^{\beta}{}_{\nu}) + \mathcal{O}(\omega^2)$$

$$\Rightarrow \eta_{\mu\nu} = \eta_{\mu\nu} + \eta_{\mu\alpha} \omega^{\alpha}{}_{\nu} + \eta_{\alpha\nu} \omega^{\alpha}{}_{\mu} = \eta_{\mu\nu} + \omega_{\mu\nu} + \omega_{\nu\mu}$$

$\Rightarrow \omega_{\mu\nu} = -\omega_{\nu\mu}$. Since $\omega_{\mu\nu}$ is antisymmetric 4×4 matrix
 $\frac{n(n-1)}{2} \Big|_{n=4} = 6$ independent parameters of transformation.

↳ Our parameters are $\omega^{\mu\nu}$ s.t. $\mu > \nu$

$$\delta x^{\mu} = \omega^{\mu\nu} x_{\nu} = \chi^{\mu}{}_{\alpha} \omega^{\alpha}$$

$$= \sum_{\lambda < \rho} \chi^{\mu}{}_{\lambda\rho} \omega^{\lambda\rho} \quad \text{with } \lambda < \rho \quad \alpha = (\mu\nu), \mu < \nu$$

$$\Big| \Rightarrow \sum_{\lambda < \rho} \omega^{\lambda\rho} (\delta^{\mu}{}_{\lambda} x_{\rho} - \delta^{\mu}{}_{\rho} x_{\lambda})$$

? factor 1/2?

$$\text{We find: } \chi^{\mu}{}_{\lambda\rho} = \delta^{\mu}{}_{\lambda} x_{\rho} - \delta^{\mu}{}_{\rho} x_{\lambda}$$

↳ Because the field is scalar:

$$\delta \phi_i = 0 \Rightarrow \chi_i(\lambda\rho) = 0$$

↳ The conserved current is then:

$$M^{\mu}{}_{\lambda\rho} = \frac{\partial \mathcal{L}}{\partial(\partial_{\mu} \phi_i)} \partial_{\nu} \phi_i (\delta^{\nu}{}_{\lambda} x_{\rho} - \delta^{\nu}{}_{\rho} x_{\lambda}) - (\delta^{\mu}{}_{\lambda} x_{\rho} - \delta^{\mu}{}_{\rho} x_{\lambda}) \mathcal{L}$$

$$= \frac{\partial \mathcal{L}}{\partial(\partial_{\mu} \phi_i)} \partial_{\lambda} \phi_i x_{\rho} - \delta^{\mu}{}_{\lambda} x_{\rho} \mathcal{L}$$

$$- \left(\frac{\partial \mathcal{L}}{\partial(\partial_{\mu} \phi_i)} \partial_{\rho} \phi_i x_{\lambda} - \delta^{\mu}{}_{\rho} x_{\lambda} \mathcal{L} \right)$$

$$= T^{\mu}{}_{\lambda} x_{\rho} - T^{\mu}{}_{\rho} x_{\lambda}$$

→ $M^{\mu}{}_{\lambda\rho}$ are 6 conserved currents, and therefore we've 6 charges:
 $\int d^3x M^0{}_{\lambda\rho}$. For $\lambda, \rho = i, j$, the conserved qty is the

orbital momentum: $\int d^3x \{ p_i x_j - p_j x_i \} = I_k = \int \epsilon_{ijk} M^0{}_{ij} d^3x$

DEF

→ Let's consider now a vector field A^μ

↳ We still have $M_{\lambda\rho}^\mu = \delta_\lambda^\mu x_\rho - \delta_\rho^\mu x_\lambda$

↳ $\delta A_\mu = \omega_{\mu\nu} A^\nu$

$$= \sum_{\sigma\rho} \Psi_{\mu\sigma\rho} \omega^{\sigma\rho} \quad ("_{\mu\sigma\rho} \leftrightarrow "_{\sigma\rho\mu})$$

?

One reads: $\Psi_{\mu\sigma\rho} = \eta_{\mu\sigma} A_\rho - \eta_{\mu\rho} A_\sigma$

Thus,

$$\begin{aligned} M_{\lambda\rho}^\mu &= M_{\lambda\rho}^{(0)\mu} - \frac{\partial \mathcal{L}}{\partial (\partial_\mu A_\nu)} \Psi_{\sigma\lambda\rho} \\ &= \underbrace{M_{\lambda\rho}^{(0)\mu}}_{\text{orbital momentum}} - \frac{\partial \mathcal{L}}{\partial (\partial_\mu A_\nu)} (\eta_{\mu\lambda} A_\rho - \eta_{\mu\rho} A_\lambda) \end{aligned}$$

DEF

The spin is $S_{\lambda\rho}^\mu = \int d^3x \left\{ \frac{\partial \mathcal{L}}{\partial (\partial_\mu A_\nu)} (\eta_{\nu\lambda} A_\rho - \eta_{\nu\rho} A_\lambda) \right\}$

or $S_{\lambda\rho} = \int d^3x \left\{ \frac{\partial \mathcal{L}}{\partial \dot{A}^\lambda} A_\rho - \frac{\partial \mathcal{L}}{\partial \dot{A}^\rho} A_\lambda \right\}$

And the vector of spin is $\mathcal{S}_i \equiv \epsilon_{ijk} S_{jk}$

2 QUANTIZATION OF A FREE SCALAR FIELD

2.1 Reminder: quantization in QM

→ We use the Lagrangian of an harmonic oscillator:

$$L = \frac{1}{2} m \dot{\tilde{q}}^2 - \frac{1}{2} k \tilde{q}^2$$

\tilde{q} - position
 $\dot{\tilde{q}}$ - velocity

↳ $S = \int L dt = \int dt \left\{ \frac{1}{2} m \dot{\tilde{q}}^2 - \frac{1}{2} k \tilde{q}^2 \right\}$

↳ We rescale our variables: $q \equiv \sqrt{m} \tilde{q}$ and write $\omega^2 = k/m$

Then:

$$S = \int dt \left\{ \frac{1}{2} \dot{q}^2 - \frac{1}{2} \omega^2 q^2 \right\}$$

→ Equations of motion (classical):

$$\delta S = 0$$