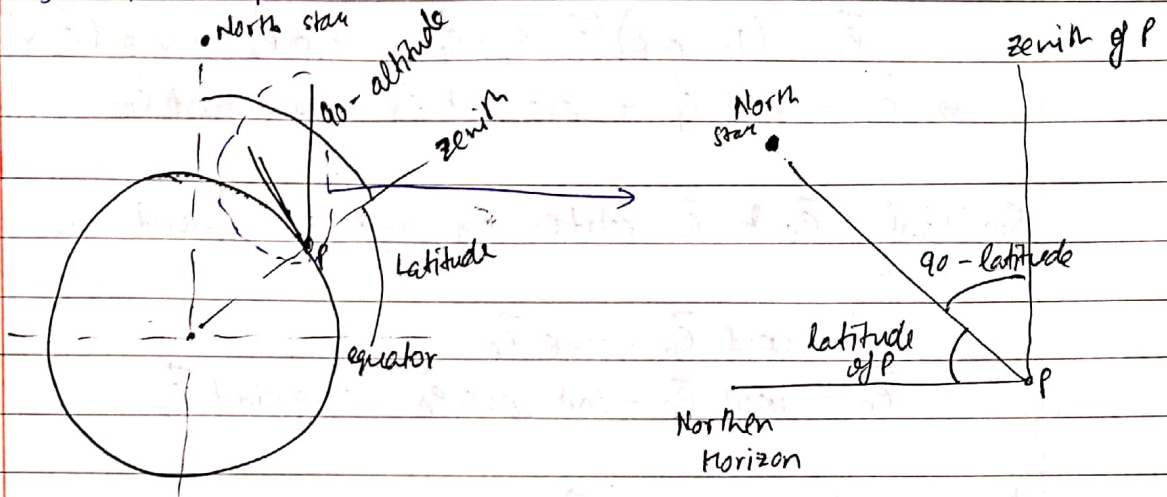


Q.1) $\vec{e}_p \equiv$ pole star
 $\vec{e}_A =$ first point of Aries
 $\vec{e}_B \equiv$ ~~to~~ \perp to e_p & e_A

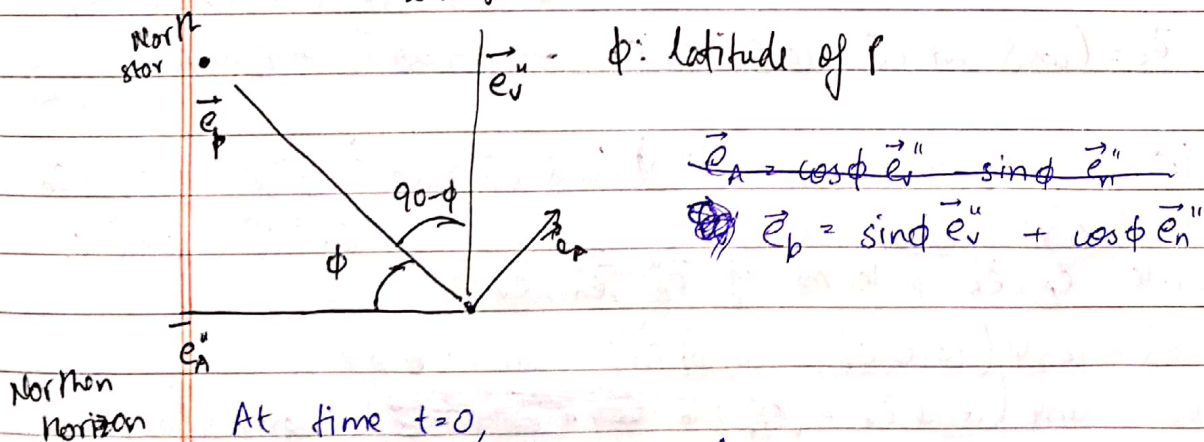
local vectors: e_v , e_n , e_e
 observer north east



Local vector rotates with the earth (e_v, e_e, e_n) and we will apply Rodriguez Rotation Eq. ~~on~~ to find their position.

Rodriguez Rot. eq $\rightarrow \vec{v}_{rot} = (1 - \cos \theta) (\vec{v} \cdot \vec{u}) \vec{u} + \vec{v} \cos \theta + (\vec{u} \times \vec{v}) \sin \theta$
 $[\vec{u}: \text{a vector along the axis of rotation}]$

after rotation by some angle θ ,
 e_v becomes e_v'' ; e_n becomes e_n'' , e_e becomes e_e'' .



$$\vec{e}_A = \cos \phi \vec{e}_v'' - \sin \phi \vec{e}_n''$$

$$\vec{e}_B = \sin \phi \vec{e}_v'' + \cos \phi \vec{e}_n''$$

At time $t=0$,

$$\hat{e}_v = \cos \phi \hat{e}_A + \sin \phi \hat{e}_B$$

$$\hat{e}_e = \hat{e}_B$$

$$\hat{e}_n = \hat{e}_v \times \hat{e}_e = (\cos \phi \hat{e}_A - \sin \phi \hat{e}_B)$$

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After time t ,
local vector \vec{e}_v changes to \vec{e}_v'' by rotating 'o' angle about \vec{e}_p .
Using Rodrigues' rotation,

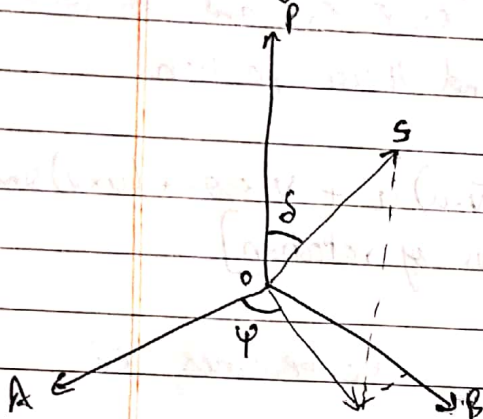
$$\begin{aligned}\vec{e}_v'' &= (1 - \cos\phi)(\vec{e}_v, \vec{e}_p)\vec{e}_p + \cos\phi\vec{e}_v + \sin\phi(\vec{e}_p \times \vec{e}_v) \\ \Rightarrow \vec{e}_v'' &= \sin\phi\vec{e}_p + \cos\phi\cos\phi\vec{e}_A + \sin\phi\cos\phi\vec{e}_B\end{aligned}$$

Similarly, \vec{e}_e & \vec{e}_n rotates by angle 'o' about \vec{e}_p .

$$\vec{e}_e'' = \cos\phi\vec{e}_B - \sin\phi\vec{e}_A$$

$$\vec{e}_n'' = \cos\phi\vec{e}_p - \sin\phi\cos\phi\vec{e}_A - \sin\phi\sin\phi\vec{e}_B$$

for any fixed vector \vec{e}_s in space,



$$\vec{e}_s = \vec{e}_p \cos\delta + \sin\delta \cos\psi \vec{e}_A + \sin\delta \sin\psi \vec{e}_B$$

Substituting value of \vec{e}_p ,

$$\vec{e}_s = (\cos\delta \sin\phi \vec{e}_v'' + \cos\delta \cos\phi \vec{e}_n'') + \sin\delta \cos\psi \vec{e}_A + \sin\delta \sin\psi \vec{e}_B \quad (*)$$

where $r_1 \vec{v}_1 = (\sin\delta \cos\psi \vec{e}_v'')$
 $r_2 \vec{v}_2 = (\sin\delta \sin\psi \cos\phi \vec{e}_v'' - \sin\delta \cos\psi \vec{e}_n'')$
write \vec{e}_A, \vec{e}_B in terms of $\vec{e}_e'', \vec{e}_n'', \vec{e}_v''$

$$\vec{e}_A = \cos\psi (\cos\phi \vec{e}_v'' - \sin\phi \vec{e}_n'') + \sin\psi (\vec{e}_e'')$$

$$\vec{e}_B = \sin\psi (\cos\phi \vec{e}_v'' - \sin\phi \vec{e}_n'') - \cos\psi (\vec{e}_e'')$$

put values of \vec{e}_A & \vec{e}_B in (*) to get

$$\vec{e}_s = (\cos\delta \sin\phi \vec{e}_v'' + \cos\delta \cos\phi \vec{e}_n'') + r_1 \cos\psi \vec{v}_1 + r_2 \sin\psi \vec{v}_2$$

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where $\theta = \psi$

$$r_1 \vec{v}_1 = \sin \delta \left[\cos \psi (\cos \phi \vec{e}_v'' - \sin \phi \vec{e}_n'') + \sin \psi (\vec{e}_e'') \right]$$

and

$$r_2 \vec{v}_2 = \sin \delta \left[\sin \psi (\cos \phi \vec{e}_v'' - \sin \phi \vec{e}_n'') - \cos \psi (\vec{e}_e'') \right]$$

Q.2) Planet: Mars

- i) Altitude: $+54^\circ 12' 51.9'' = 54.2^\circ$
Azimuth: $+245^\circ 55' 17.4'' = 245.92^\circ$

Star: Rigel Kentarus

- ii) Altitude: $-54^\circ 42' 17.4'' = -54.7^\circ$
Azimuth: $+170^\circ 26' 20.8'' = 170.43^\circ$

Mars dist = 0.748 AU

Star dist: 4.3 ly

$$\text{North comp.} = (\text{distance}) \cdot \cos(\text{altitude}) \cos(\text{azimuth})$$

$$\text{East comp} = (\text{distance}) \cdot \cos(\text{altitude}) \sin(\text{azimuth})$$

$$\text{Vertical comp} = (\text{distance}) \cdot \sin(\text{elevation})$$