

Chapter 15 : Experimental Tests of General Relativity

In contrast to special relativity, there are very few tests of the general theory. *Is this true anymore?*

The main reason is that gravitational fields in our vicinity are weak.

The very first tests were :

- ① Precession of Mercury's perihelion.
- ② Bending of light (gravitational lensing)
- ③ Gravitational Redshift

Other tests :

- ④ The delay of light in a gravitational field
- ⑤ Orbital motion of PSR 1913 + 16 (Hulse - Taylor)
- ⑥ Observation of gravitational waves from Compact Binary Mergers (2015 - present).

Classical Kepler Motion

A mass m moves under the influence of an inverse square law force with centre of attraction at O . The force is:

$$F = -\frac{m\mu}{r^2} \hat{r}$$

$$\Rightarrow$$

$$m\ddot{r} = -\frac{m\mu}{r^2} \hat{r}$$

Angular momentum $\vec{L} = \vec{r} \times (m\dot{\vec{r}})$ and hence,

$$\begin{aligned} \frac{dL}{dt} &= \dot{\vec{r}} \times (m\dot{\vec{r}}) + \vec{r} \times (m\ddot{\vec{r}}) \\ &= 0 \end{aligned}$$

ie. angular momentum is conserved and

$$\vec{L} = m\vec{h} \quad ; \quad \vec{h} \text{ is constant}$$

Classical Kepler Motion

Angular momentum conservation also implies the motion will be restricted to a plane.

To make things convenient, we'll work with polar coordinates (R, ϕ) . With this, the EoM is,

$$(\ddot{R} - R\dot{\phi}^2) \hat{R} + \frac{1}{R} \frac{d}{dt}[R^2\dot{\phi}] \hat{\phi} = -\frac{\mu}{R^2} \hat{R}$$

Dotting with $\hat{\phi}$ immediately shows,

$$R^2\dot{\phi} = h$$

which is indeed conservation of angular momentum.

Classical Kepler Motion

Dotting with \hat{R} gives,

$$\ddot{R} - R\dot{\phi}^2 = -\frac{\mu}{R^2}$$

Substituting $R = \frac{1}{u}$,

$$\frac{d^2 u}{d\phi^2} + u = \frac{\mu}{h^2}$$

← Binet's equation

which has solution,

$$u = \frac{\mu}{h^2} + C \cos(\phi - \phi_0)$$

Classical Kepler Motion

By substituting back $u = 1/R$, and $l = \frac{h^2}{\mu}$, $e = \frac{Ch^2}{\mu}$,

$$\frac{l}{R} = 1 + e \cos(\phi - \phi_0)$$

We immediately notice that this is the polar equation for a conic section!

l : semi-latus rectum

e : eccentricity

ϕ_0 : orientation relative to x -axis

If $e \in (0, 1)$, conic section is an ellipse, and point of closest approach is called perihelion.

Classical Kepler Motion

The motion of a test mass in the field of a massive body is called the one-body problem.

One can map the two-body problem to the one-body problem in the following way,

$$\vec{F}_{12} = m\ddot{\vec{r}} = -\frac{m\mu}{r^2} \hat{r}$$

where,

$$m: \text{reduced mass} = \frac{m_1 m_2}{(m_1 + m_2)}$$

$$\mu = (m_1 + m_2) (G)$$

Advance of Mercury's perihelion

Lets look at the one-body problem, looking at the Schwarzschild solution. (ie. spherically symmetric field)

The "Lagrangian" is,

$$2K = \left(1 - \frac{2m}{r}\right) \dot{t}^2 - \left(1 - \frac{2m}{r}\right)^{-1} \dot{r}^2 - r^2 \dot{\theta}^2 - r^2 \sin^2 \theta \dot{\phi}^2 = 1$$

The equations are:

$$t \text{ eqn} : \frac{d}{dz} \left[\left(1 - \frac{2m}{r}\right) \dot{t} \right] = 0$$

$$\theta \text{ eqn} : \frac{d}{dz} [r^2 \dot{\theta}] - r^2 \sin \theta \cos \theta \dot{\phi}^2 = 0$$

$$\phi \text{ eqn} : \frac{d}{dz} (r^2 \sin \theta \dot{\phi}) = 0$$

} + 1 eqn of
the "Lagrangian"

Advance of Mercury's perihelion

Lets look if planar orbits are possible. Let's fix $\theta = \pi/2$, and $\dot{\theta} = 0$. Putting this in the θ eqn, we see that the higher derivatives are zero.

Using this in the ϕ eqn, we get,

$$\frac{d}{dc} (r^2 \dot{\phi}) = 0 \Rightarrow r^2 \dot{\phi} = h$$

Similarly, r eqn gives,

$$\left(1 - \frac{2m}{r}\right) \ddot{r} = k$$

Advance of Mercury's perihelion

Now, putting this all in the "Lagrangian",

$$k^2 \left(1 - \frac{2m}{r}\right)^{-1} + \dot{r}^2 \left(1 - \frac{2m}{r}\right)^{-1} - r^2 \dot{\phi}^2 = 1$$

Substituting $r = 1/u$,

$$\frac{d^2 u}{d\phi^2} + u = \frac{m}{h^2} + 3mu^2$$

corresponds to
Binet's equation
modulo $3mu^2$.
This term is pretty
small.

We'll solve this using perturbation analysis using,

$$\epsilon = \frac{3m^2}{h^2}$$

Advance of Mercury's perihelion

Then,

$$u'' + u = \frac{m}{h^2} + \epsilon \frac{h^2 u^2}{m}$$

Assume $u = u_0 + \epsilon u_1 + O(\epsilon^2)$, we get,

$$u_0 = \frac{m}{h^2} (1 + e \cos \phi)$$

$$u_1 = A + B\phi \sin \phi + C \cos 2\phi$$

where

$$A = \frac{m}{h^2} \left(1 + \frac{1}{2} e^2\right) ; B = \frac{me}{h^2} ; C = -\frac{me^2}{6h^2}$$

Advance of Mercury's perihelion

u_0 is just the Newtonian result. Let's look at correction u_1 .

$$u_1 = A + B\phi \sin\phi + C\cos 2\phi$$

In this term, $\phi \sin\phi$ is a continuously growing term, and hence are the most relevant.

$$\therefore u = \frac{m}{h^2} (1 + e\cos\phi + \epsilon e\phi \sin\phi)$$

$$\therefore u \approx \frac{m}{h^2} [1 + e\cos\{\phi(1-\epsilon)\}]$$

The new period is,

$$\frac{2\pi}{1-\epsilon} \approx 2\pi(1+\epsilon)$$

Advance of Mercury's perihelion

This means that the planet will travel in an ellipse, but the axis of the ellipse will rotate by an amount $2\pi\epsilon$ between two successive perihelia. In non-relativistic units,

$$2\pi\epsilon \approx \frac{24\pi^3 a^2}{c^2 T^2 (1-e^2)}$$

There is precession even in the Newtonian n -body problem, but there still was a 43 arcseconds per century discrepancy!

General Relativity explained this discrepancy.

Bending of Light

For null geodesics, the relativistic version of Binet's equation is,

$$\frac{d^2 u}{d\phi^2} + u = 3mu^2 \rightarrow \textcircled{1}$$

In the limit of special relativity,

$$\frac{d^2 u}{d\phi^2} + u = 0 \rightarrow \textcircled{2}$$

$$\Rightarrow u = \frac{1}{D} \sin(\phi - \phi_0)$$

This is actually just the equation for a straight line as from ϕ to ϕ_0 .

① can be thought of as a perturbation around

②.

Bending of Light

So then, we want to seek an approximate solution

$$u = u_0 + 3m u_1$$

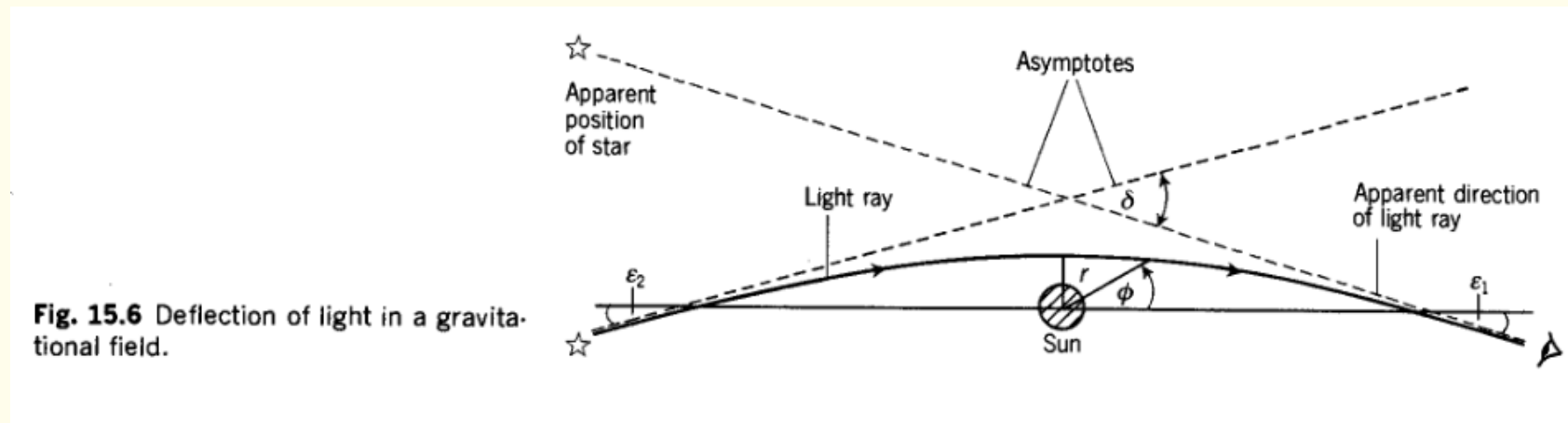
Substituting this approximation, we get,

$$u_1'' + u_1 = u_0^2 = \frac{\sin^2 \phi}{D^2} \quad \leftarrow \phi_0 = 0 \text{ Assumption}$$

And the general solution is,

$$u \approx \frac{\sin \phi}{D} + \frac{m}{D^2} [1 + C \cos \phi + \cos^2 \phi]$$

Bending of Light



We note that as $r \rightarrow \infty$, $u \rightarrow 0$, For $u \rightarrow 0$,
let's take angles $-\epsilon_1$, $\pi + \epsilon_2$, then,

$$0 = -\frac{\epsilon_1}{D} + \frac{m}{D^2} [2 + C]$$

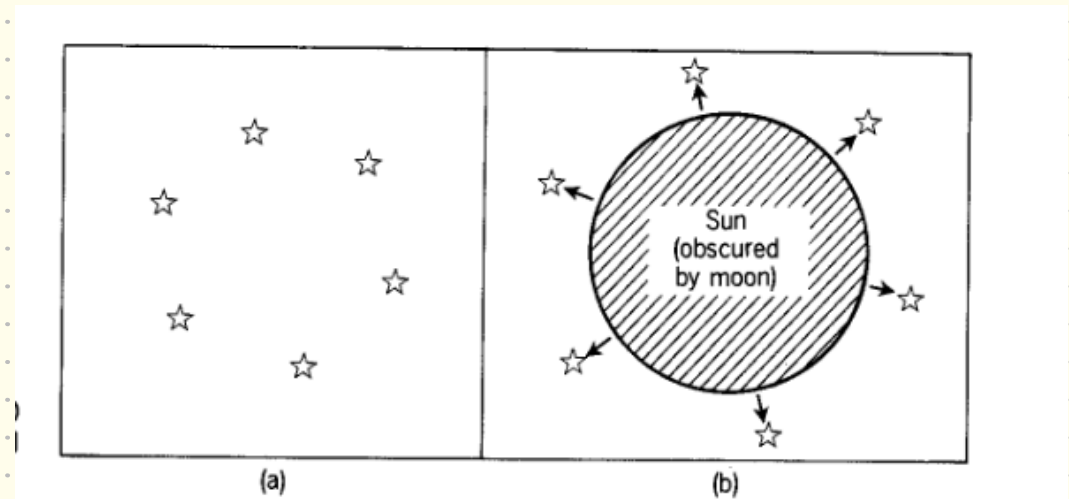
$$0 = -\frac{\epsilon_2}{D} + \frac{m}{D^2} [2 - C]$$

$$\delta = \epsilon_1 + \epsilon_2 = \frac{4m}{D}$$

$$\Rightarrow \delta = \frac{4GM}{c^2 D}$$

Bending of Light

Deflection produced by a light ray that just grazes the sun is 1.75 arcsec .



If one compares positions of stars when they are behind the sun and when they are not, one can measure this deflection. Eddington and team did, in 1919.

Lensing is now a very important probe of physics,

Gravitational Redshift

Consider a thought experiment of an endless chain running between earth and sun,

This experiment results in the construction of a perpetual motion machine, hence contradictory!

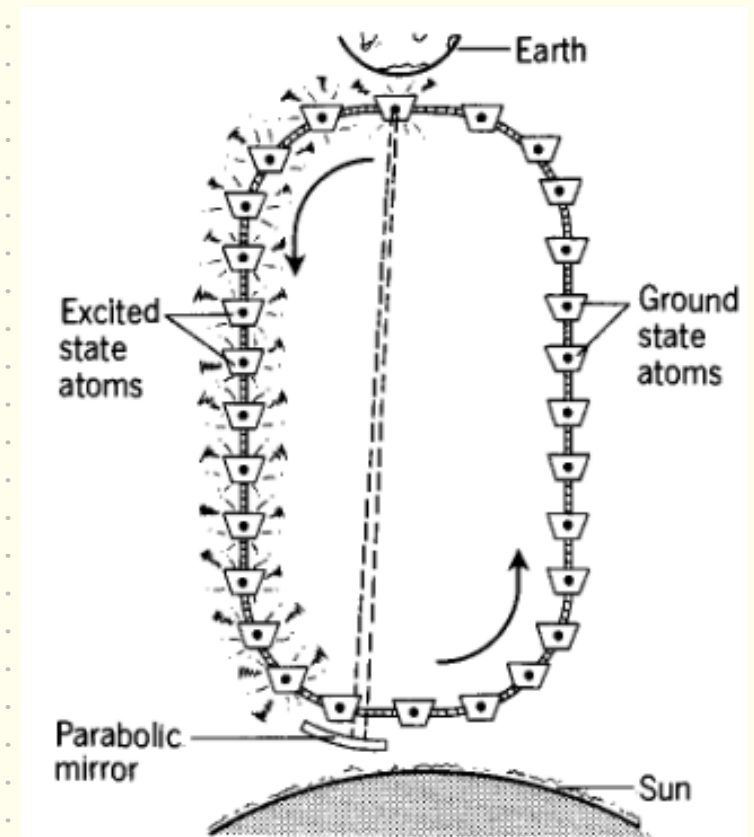
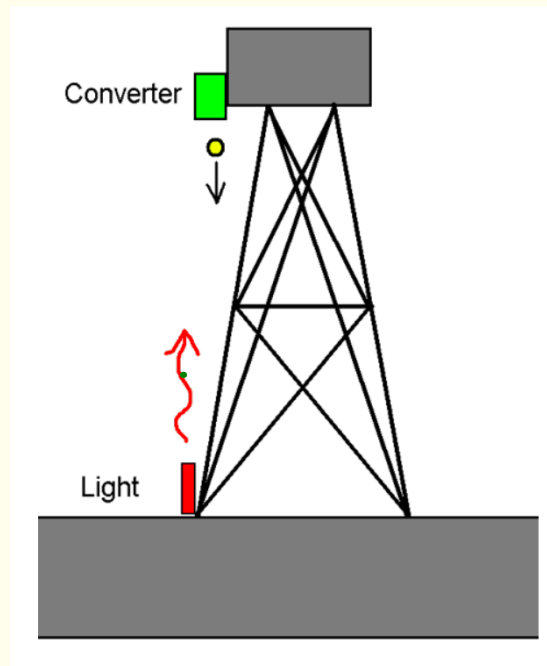
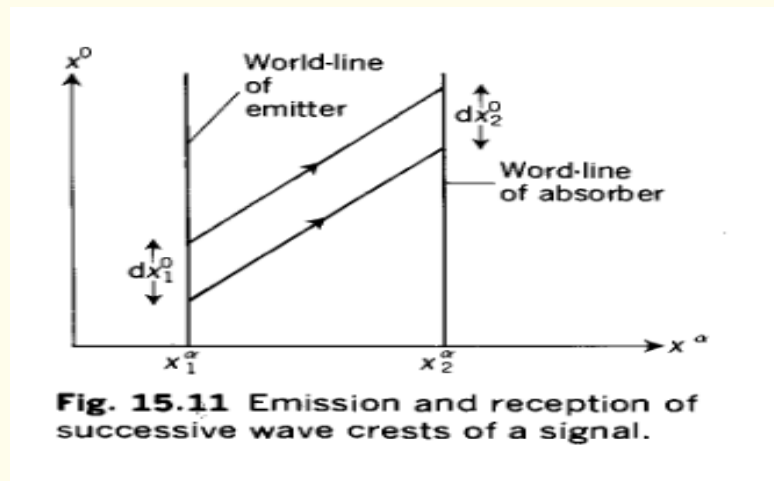


Fig. 15.10 A gravitational *perpetuum mobile*?

Gravitational Redshift



Redshift in static spacetime; let $x^a = (x^0, x^a)$

Consider two observers with atomic clocks with worldlines x_1^a, x_2^a respectively. The metric cannot depend on x^0 's.

Assume that the first observer is sending radiation to the second. The first observer measures dz i.e. separation between successive wave crests. Second observer measures αdz .

Gravitational Redshift

$$dz^2 = g_{00}(x_1^\alpha) (dx_1^0)^2$$

$$\alpha^2 dz^2 = g_{00}(x_2^\alpha) (dx_2^0)^2$$

Since spacetime is static, $dx_2^0 = dx_1^0$. Hence

$$\alpha = \left(\frac{g_{00}(x_2^\alpha)}{g_{00}(x_1^\alpha)} \right)^{1/2}$$

means that frequency
will be redshifted

The second observer will hence measure frequency,

$$\bar{\nu}_0 = \nu_0 \left[\frac{g_{00}(x_1^\alpha)}{g_{00}(x_2^\alpha)} \right]^{1/2} \Rightarrow \bar{\nu}_0 < \nu_0 \quad \text{if} \quad g_{00}(x_1^\alpha) < g_{00}(x_2^\alpha)$$

Gravitational Redshift

$$\frac{\Delta \nu}{\nu} = \frac{\bar{\nu}_0 - \nu_0}{\nu_0} = \frac{\phi_1 - \phi_2}{c^2}$$

in weak field

This has been confirmed to very good precision by terrestrial expts, such as the ones by Pound and Rebka, and numerous other follow-up's.