

```
function [r,v] = oe2states(h,e,RA,incl,w,TA,mu)
%OE2STATES Converts orbital elements to state vector
% Inputs are:
% h      :a scalar specific angular momentum in m^2/s
% e      :a scalar orbital eccentricity
% RA     :a scalar right ascension of the ascending node in rad
% incl   :a scalar orbital inclination
% W      :a scalar argument of perigee in rad
% TA     :a scalar true anomaly in rad
% mu     :an optional scalar gravitational parameter in m^3/s^2 (default
%         earth)
%
% Outputs are:
% r      :a numeric array of 3x1 current position vector in m
% v      :a numeric array of 3x1 current velocity vector in m/s

arguments
    h {mustBeScalarOrEmpty, mustBeNumeric, mustBeReal}
    e {mustBeScalarOrEmpty, mustBeNumeric, mustBeReal}
    RA {mustBeScalarOrEmpty, mustBeNumeric, mustBeReal}
    incl {mustBeScalarOrEmpty, mustBeNumeric, mustBeReal}
    w {mustBeScalarOrEmpty, mustBeNumeric, mustBeReal}
    TA {mustBeScalarOrEmpty, mustBeNumeric, mustBeReal}
    mu {mustBeScalarOrEmpty, mustBeNumeric, mustBeReal} = 3.98600442e14
end

% r and v in perifocal frame
rp = (h^2/mu)*(1/(1 + e*cos(TA)))*[cos(TA);sin(TA);0];
vp = (mu/h)*[-sin(TA);e*cos(TA);0];

% Rotate perifocal coordinates to geocentric equatorial frame
Q = euler_rot(RA,incl,w,'313')';
r = Q*rp;
v = Q*vp;
end
```