Example Sheet 1 Worked Solutions

Question 1

Question. Establish Stirling's Formula: $N \approx \sqrt{2\pi N} N^N e^{-N}$

Proof. We are first given $\int_{0}^{\infty} e^{-x}x^{N}dx = \int_{0}^{\infty} e^{-F(x)}dx$ and since the integrands are equal, we can equate:

$$e^{-x}x^N = e^{-F(x)}$$

$$\Rightarrow F(x) = x - N \ln x.$$
(1)

We then differentiate twice in order to find the approximation $F(x) \approx F(x_0) + F''(x_0)(x - x_0)^2/2$ where x_0 is x value of the minimum of F(x).

$$F'(x) = 1 - \frac{N}{x}$$

$$F''(x) = \frac{N}{x^2}.$$
(2)

The minimum occurs when F'(x) = 0 which implies $x_0 = N$ (we assume this is a minimum as it is the only solution to F'(x) = 0 and the question asks us to find a minimum).

We know have all the information required to determine the approximation and can substitute into the 2nd integral given:

$$F(x) \approx -N + N \ln N - \frac{(x-N)^2}{2N} \tag{3}$$

$$N! = \int_{0}^{\infty} \exp\left(-N + N \ln N - \frac{(x-N)^2}{2N}\right) dx$$
$$= e^{-N} N^N \int_{0}^{\infty} \exp\left(-\frac{(x-N)^2}{2N}\right) dx. \tag{4}$$

The time has come to apply the final approximation alluded to in the question. The term x - N is effectively shifting the function to the right hand side of the y-axis. And since N is very large, the limits of the integral guarantee that we are integrating the entire curve as it tails to 0 at either end. *Note:* The factor of 1/2N does effectively nothing to affect this shifting.

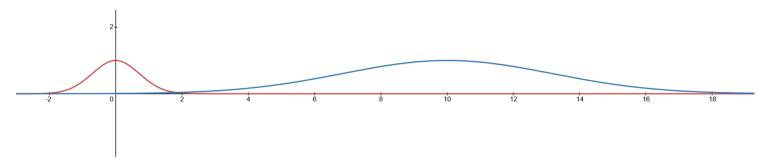


Figure 1: Graph of $\exp -x^2$ in red vs $\exp(-(x-10)^2/20)$ in blue

Thus, we make the approximation:

$$N! \approx e^{-N} N^N \int_{-\infty}^{+\infty} \exp\left(\frac{-x^2}{2N}\right) dx.$$
 (5)

As this is the standard form of the Gaussian Integral:

$$\int_{-\infty}^{+\infty} e^{-ax^2} dx = \sqrt{\frac{\pi}{a}}.$$
 (6)

Thus, with a = 1/2N, we have:

$$N! \approx \sqrt{2\pi N} N^N e^{-N} \tag{7}$$

Question 2i

Question. For a two coupled system in the microcanonical ensemble, show that they maximise their entropy if the heat capacity, C, is positive.

Proof. Considering the system as a whole to have entropy S and equal temperature T, we have the equation:

$$\frac{\partial^2 S}{\partial E^2} = -\frac{1}{T^2 C} \tag{8}$$

If we want the system to have a maximum entropy, the equation on the right hand side must be negative and thus C must be positive.

Question 2ii

Question. Show that the energy fluctuations $\triangle E^2 = \langle E^2 \rangle - \langle E \rangle^2$ are proportional to C_V in the canonical ensemble.

Proof. The fluctuations in the canonical ensemble can be written as:

$$\triangle E^2 = -\frac{\partial \langle E \rangle}{\partial \beta} \tag{9}$$

The definition of the heat capacity C_V in this case is given as:

$$C_{V} = \left(\frac{\partial \langle E \rangle}{\partial T}\right)_{V}$$

$$\Rightarrow \partial \langle E \rangle = C_{V} \cdot \partial T$$

$$\Rightarrow \Delta E^{2} = -C_{V} \cdot \frac{\partial T}{\partial \beta}$$
(10)

where the subscript denotes that the volume is constant.

By definition:

$$\beta = \frac{1}{k_B T}, T \neq 0$$

$$\Rightarrow \frac{\partial \beta}{\partial T} = -\frac{1}{k_B T^2}$$

$$\Rightarrow \frac{\partial T}{\partial \beta} = -k_B T^2, T \neq 0$$

$$\Rightarrow \Delta E^2 = C_V k_B T^2, T \neq 0$$
(11)

Question 2iii

Question. Show that the Gibbs entropy from the canonical ensemble can be written as

$$S = k_B \frac{\partial}{\partial T} (T \ln Z) \tag{12}$$

Proof. The entropy in the canonical system, derived using the standard definition of S and Stirlings formula is given as:

$$S = -k_B \sum_{n} p(n) \ln p(n). \tag{13}$$

The probability of the system being in a state $|n\rangle$ is given by the Boltzmann distribution:

$$p(n) = \frac{e^{-\beta E_n}}{Z} \tag{14}$$

Substituting into our form of S:

$$S = -k_B \sum_{n} \frac{e^{-\beta E_n}}{Z} \ln \left(\frac{e^{-\beta E_n}}{Z} \right)$$

$$= -k_B \sum_{n} \frac{e^{-\beta E_n}}{Z} \left[-\beta E_n - \ln Z \right]$$

$$= k_B \sum_{n} \frac{\beta E_n e^{-\beta E_n}}{Z} + \frac{e^{-\beta E_n}}{Z} \ln Z$$

$$= k_B \left[\sum_{n} \frac{\beta E_n}{Z} e^{-\beta E_n} + \ln Z \right],$$
(15)

Where the last step is made as $\sum_n e^{-\beta E_n}/Z = \sum_n p(n) = 1$ as something must happen.

Now, we work backwards from the Gibbs entropy:

$$S = k_b \frac{\partial}{\partial T} (T \ln Z)$$

$$= k_B \left[T \cdot \frac{1}{Z} \cdot \frac{\partial Z}{\partial T} + \ln Z \right]$$
(16)

Z is given as:

$$Z = \sum_{n} e^{-\beta E_n}$$

$$= \sum_{n} e^{-E_n/k_B T}$$
(17)

Differentiating with respect to T:

$$\frac{\partial Z}{\partial T} = \sum \frac{\partial}{\partial T} (e^{-E_n/k_B T})$$

$$= \sum_{n} \frac{\partial}{\partial T} \left(-\frac{E_n}{k_B T} \right) e^{-E_n/k_B T}$$

$$= \sum_{n} \frac{E_n}{k_B T^2} e^{-E_n/k_B T}$$

$$= \sum_{n} \frac{E_n}{k_B T^2} e^{-\beta E_n}.$$
(18)

We can now substitute this into the Gibbs entropy equation:

$$S = k_B \left[\frac{T}{Z} \sum_n \frac{E_n}{k_B T^2} e^{-\beta E_n} + \ln Z \right]$$

$$= k_B \left[\sum_n \frac{\beta E_n}{Z} e^{-\beta E_n} + \ln Z \right]$$

$$= k_B \frac{\partial}{\partial T} (T \ln Z)$$
(19)

Question 3

Question. In a system of N spin 1/2 particles, each with energy of $\pm \mu B/2$ where μ is the magnetic moment, show that the partition function is

$$Z = 2^N \cosh^N \left(\frac{\beta \mu B}{2}\right),\tag{20}$$

Proof. A single particle in the system has two states $E_1 = -\mu B/2$ or $E_2 = +\mu B/2$, thus:

$$Z = \sum_{n} e^{-\beta E_n}$$

$$= e^{\beta \mu B/2} + e^{-\beta \mu B/2}$$

$$= 2 \cosh\left(\frac{\beta \mu B}{2}\right).$$
(21)

Since Z is multiplicative, i.e. $Z = Z_1 Z_2$ upon combining systems, we have:

$$Z = 2^N \cosh^N \left(\frac{\beta \mu B}{2}\right) \tag{22}$$

Question. Find the average energy $\langle E \rangle$ and entropy S. Show that the result makes sense when T = 0 and $T \to \infty$. Solution. For the average energy:

$$\langle E \rangle = -\frac{\partial}{\partial \beta} \ln Z$$

$$= -\frac{\partial}{\partial \beta} \ln \left(2^N \cosh^N \left(\frac{\mu B}{2} \beta \right) \right)$$

$$= -\frac{\partial}{\partial \beta} \left(\ln 2^N + N \ln \cosh \left(\frac{\mu B}{2} \beta \right) \right)$$
(23)

Since 2^N is independent of β , its derivative evaluates to 0 and using the fact that:

$$\frac{d}{dx}\ln\cosh(ax) = a\tanh(ax),\tag{24}$$

we have:

$$\langle E \rangle = -\frac{N\mu B}{2} \tanh\left(\frac{\mu B}{2}\beta\right) \tag{25}$$

This is physically viable (thankfully) since when $T \to 0$, we have $\langle E \rangle \to -N\mu B/2$ which is the lowest possible energy of the system. And, when $T \to \infty$ we have $\langle E \rangle \to 0$ which is expected as when the system has infinite heat, none of the energy in the system remains accessible.

Now, for entropy:

$$S = k_{B} \frac{\partial}{\partial T} (T \ln Z)$$

$$= k_{b} \left[\frac{T}{Z} \frac{\partial Z}{\partial T} + \ln Z \right]$$

$$= k_{b} \left[\frac{T}{Z} \frac{\partial}{\partial T} \left(2^{N} \cosh^{N} \left(\frac{\mu B}{2k_{B}T} \right) \right) + \ln 2^{N} \cosh^{N} \left(\frac{\mu B}{2k_{B}T} \right) \right]$$

$$= k_{b} \left[\frac{T}{Z} \cdot 2^{N} \cdot N \cosh^{N-1} \left(\frac{\mu B}{2k_{B}T} \right) \cdot \left(-\frac{\mu B}{2k_{B}T^{2}} \right) \sinh \left(\frac{\mu B}{2k_{B}T} \right) + N \ln 2 \cosh \left(\frac{\mu B}{2k_{B}T} \right) \right]$$

$$(26)$$

Luckily, the cosh and sinh terms cancel with the factor of 1/Z and turn to a single tanh term:

$$S = k_B \left[-\frac{N\mu B}{2k_B T} \tanh\left(\frac{\mu B}{2k_B T}\right) + \ln 2 \cosh\left(\frac{\mu B}{2k_B T}\right) \right]. \tag{27}$$

When $T \to 0$, $S \to 0$ which makes complete sense physically as when there is no heat, there is perfect order. When $T \to \infty$, $S \to N \ln 2$ which is also physically accurate as it is the maximum value of S (fact checked by true Desmos patriots).

Question. Compute the magnetisation $M=N_{\uparrow}-N_{\downarrow}$ and show that $\chi=\partial M/\partial B$ satisfies $\chi\sim 1/T$ at high temperatures (Curie's Law)

Proof. The energy of the system in total is given by:

$$E = \frac{\mu B}{2} N_{\uparrow} - \frac{\mu B}{2} N_{\downarrow} \tag{28}$$

And this total energy will be the average energy of the system as the fluctuations of energy tend to 0 as $N \to \infty$ which is the typical case for a classic system such as the Oxygen that Curie experimented with. Thus,

$$\frac{\mu B}{2} N_{\uparrow} - \frac{\mu B}{2} N_{\downarrow} = -\frac{N\mu B}{2} \tanh\left(\frac{\mu B}{2}\beta\right)$$

$$\Rightarrow N_{\uparrow} - N_{\downarrow} = -N \tanh\left(\frac{\mu B}{2}\beta\right)$$

$$= M.$$
(29)

Thus we now have a differentiable function of M:

$$M = -N \tanh\left(\frac{\mu\beta}{2}B\right)$$

$$\Rightarrow \frac{\partial M}{\partial B} = -\frac{N}{2} \cdot \frac{\mu\beta}{2} \operatorname{sech}^{2}\left(\frac{\mu\beta}{2}B\right)$$

$$= -\frac{N\beta\mu}{4} \operatorname{sech}^{2}\left(\frac{\mu\beta}{2}B\right) \cdot \frac{1}{k_{B}T}$$

$$\Rightarrow \chi = -\frac{N\mu}{4k_{B}} \operatorname{sech}^{2}\left(\frac{\mu B}{2}\beta\right) \cdot \frac{1}{T}.$$

$$\Rightarrow \chi(T \to \infty) = -\frac{N\mu}{4k_{B}} \cdot \frac{1}{T}$$
(30)

At high temperatures, ${\rm sech}^2(\mu B\beta/2) \to 1$ and thus, $\chi \sim 1/T$ as required. The negative demonstrates an antialignment with the magnetic field and the magnitudes of N and k_B are the same ($\approx 10^{23}$) and thus cancel.

${\it Question}~5$

Question. Compute Z for a one dimensional quantum harmonic oscillator with frequency ω and:

$$E_n = \hbar\omega \left(n + \frac{1}{2} \right) \tag{31}$$

Find the average energy and entropy as a function of T. Creating a system of N particles for a solid, show that at high temperatures

Solution.