Physics 2211 - Lab 3 Black Hole

Advaith Menon¹

¹Computer Engineering Georgia Institute of Technology

February 29th, 2024





Outline

Introduction

Aim of the experiment

Prerequisites

Newton's Second Law Newton's Universal law of Gravitation Initial Conditions

Experiment

Formulae

Code

Results

Conclusion

What does it mean? What if...





Outline

Introduction

Aim of the experiment

Prerequisites

Newton's Second Law Newton's Universal law of Gravitation Initial Conditions

Experiment

Formulae

Code

Results

Conclusion

What does it mean? What if...





Aim

Purpose of this lab assignment

- Given the mass and trajectory of a star that orbits around a black hole, find the mass of the back hole.
- Find the velocity of a spaceship of a given mass, which is on the planet, such that:
 - The spaceship just passes the black hole
 - The spaceship orbits around the black hole





Outline

Introduction

Aim of the experiment

Prerequisites

Newton's Second Law Newton's Universal law of Gravitation Initial Conditions

Experiment

Formulae

Code

Results

Conclusion

What does it mean? What if...





Newton's Second Law

Quantitative analysis of motion

"The net force acting on a body is defined as the change in its momentum per unit time."

$$\vec{F}_{net} = \frac{\mathrm{d}\vec{p}}{\mathrm{d}t} \tag{1}$$

In most daily life scenarios, mass doesn't change with respect to time, hence this equation is better known as:

$$\vec{F}_{net} = m \cdot \vec{a}$$

where,

- \vec{F}_{net} = Net force acting on a body
- *m* = Mass of the body
- $\vec{a} = \frac{d\vec{v}}{dt} = \frac{d^2\vec{r}}{dt^2} =$ Net acceleration on body
- $\Delta \vec{p} = m \times \Delta \vec{v} =$ Change in momentum





Newton's Universal law of Gravitation

"The gravitational force acting of an object is directly proportional to the mass of the two objects, and inversely proportional to the square of the distance between the planets."

$$\vec{F}_{g,21} = \frac{Gm_1m_2}{|\vec{r}_{21}|^2} \cdot \hat{r}_{12} \tag{2}$$

where,

- $\vec{F}_{g,21}$ = The gravitational force on object 1 due to object 2.
- *m*₁ = Mass of the first body
- m₂ = Mass of the second body
- $\vec{r}_{21} = \vec{r}_2 \vec{r}_1 =$ The vector "distance" between the planets Change in momentum





Initial Conditions - Part 1

- Position of the black hole, $\vec{r}_{BH} = \langle 0, 0, 0 \rangle$ m
- Position of the star, \vec{r}_{star} given to us in a Python List at given time intervals.
- Mass of the star, $m_{star} = 2 \times 10^{30} \text{ kg}$
- $\Delta t = 86400 \text{ s}$



Initial Conditions - Part 2

- Position of the black hole, $\vec{r}_{BH} = \langle 0, 0, 0 \rangle$ m
- Mass of the blackhole, $m_{BH}=1.35641\times 10^{34}~{\rm kg}$ (from part 1)
- Initial position of the planet, $\vec{r}_{planet} = \langle 1 \times 10^{10}, 0, 0 \rangle$ m
- Radius of the planet, $a = 1 \times 10^8 \text{ m}$
- Initial position of the ship, $\vec{r}_{ship} = \langle 1 \times 10^{10}, 1 \times 10^8, 0 \rangle$ m





System and Surroundings - Part 1

- **System:** The star (hereafter referred to 'object')
- Surroundings: Everything else (black hole etc.)





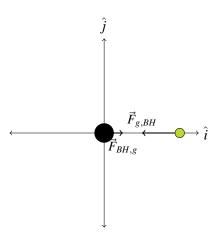
System and Surroundings - Part 2

- **System:** The ship (hereafter referred to 'object')
- Surroundings: Everything else (black hole, planet etc.)





Diagram



Outline

Introduction

Aim of the experiment

Prerequisites

Newton's Second Law Newton's Universal law of Gravitation Initial Conditions

Experiment

Formulae Code

Results

Conclusion

What does it mean? What if...





Formulae - Experiment 1

Resolving the net forces on each object

- $\vec{p}_i = m_{star} \cdot \vec{v}_{init}, \, \vec{p}_i = m_{star} \cdot \vec{v}_{final}$
- $\Delta \vec{v} = \vec{v}_{final} \vec{v}_{init}$, $\Delta \vec{p} = \vec{p}_{final} \vec{p}_{init}$
- $\vec{F}_{net} = \frac{\Delta \vec{p}}{\Delta t}$
- $\left(\frac{\mathrm{d}\vec{p}}{\mathrm{d}t}\right)_{\parallel} = |\vec{F}_{net}||\vec{p}|\cos(\theta) = \vec{F}_{net} \cdot \hat{p} \times \hat{p}$

•
$$\left(\frac{\mathrm{d}\vec{p}}{\mathrm{d}t}\right)_{\perp} = \vec{F}_{net} - \left(\frac{\mathrm{d}\vec{p}}{\mathrm{d}t}\right)_{\parallel}$$





Formulae - Experiment 2

Resolving the net forces on each object

$$\vec{F}_{g,21} = \frac{Gm_1m_2}{|\vec{r}_{21}|^2} \cdot \hat{r}_{12} \tag{3}$$

Finding mass of black hole

```
seed="amenon301@gatech.edu"
                                              t = t + deltat
# ... kepler's formula
                                              idx=1
scene.background = color.white
                                             mvballs = 0
star = sphere(color=color.yellow,
                                             while idx < (len(X)-1):

→ radius=3e11)

                                                  rate (50)
trail = curve(color=color.green)
origin = sphere(pos=vector(0,0,0),
                                                 v init =

→ color=color.black, radius=3e11)

    vector(Xvel[idx-1], Yvel[idx-1], 0)

                                                 v final =
arrowFnet = arrow(pos=star.pos,
                                                  → vector(Xvel[idx], Yvel[idx], 0)
\hookrightarrow axis=vector(0,0,0),

→ color=color.orange)

                                                  # EDIT THESE TWO LINES
arrowFnet_par = arrow(pos=star.pos,
                                                  p init = star.m * v init
\hookrightarrow axis=vector(0,0,0),
                                                  p final = star.m * v final # $\vec{p}

→ color=color.magenta)

                                                  \hookrightarrow = m \setminus vec\{v\}$
arrowFnet_perp = arrow(pos=star.pos,

→ axis=vector(0,0,0), color=color.cyan)
                                                  # EDIT THESE TWO LINES
                                                  deltav = v final - v init
G = 6.7e - 11
                                                  deltap = p_final - p_init
# Mass of the star -- EDIT THIS LINE
                                                  # EDIT THIS LINE
star.m = 2e30
                                                  dpdt = deltap/deltat
# Time interval between imported data
                                                  # EDIT THIS LINE

→ points -- EDIT THIS LINE
```

EDIT THIS LINE

Finding mass of black hole

```
Fnet perp = dpdt perp # yes.
# Calculate the mass of the black
→ hole
# EDIT AND ADD LINES AS NEEDED
# or, F = (G*star.m*mBH)/(r^2)
\# mBH = (F * r^2)/(G * star.m)
mBH = (mag(Fnet) * (X[idx]**2 +

→ Y[idx]**2))/(G * star.m)
myballs += mBH
print ("Mass of Black Hole", mBH)
# print("Velocity of ball", v_init)
# Update current position and

→ velocity and show the object's

→ current track

star.pos = vector(X[idx],Y[idx],0)
trail.append(pos=star.pos)
# EDIT THE NEXT SEVEN LINES
```

```
arrowscale = 1E-17 # determines

→ how long to draw the arrows that

→ represent vectors

    arrowFnet.pos = star.pos
    arrowFnet.axis = Fnet*arrowscale
    arrowFnet par.pos = star.pos
    arrowFnet par.axis =

→ Fnet_par*arrowscale

    arrowFnet perp.pos = star.pos
    arrowFnet perp.axis =

→ Fnet_perp*arrowscale

    t = t + deltat
    idx = idx + 1
print ("Average mass of black hole",
\hookrightarrow mvballs/(idx-1))
print("|Fnet|=", mag(Fnet))
print("|Fnet par|=", mag(Fnet par))
print("|Fnet perp|=",mag(Fnet perp))
print("All done!")
```



Finding trajectory of the ship

```
Glowscript 3.2 VPython
                                             ##
## PHYS 2211
## Lab 3: Black Hole - Part 2
                                             ## SYSTEM PROPERTIES & INITIAL CONDITIONS
## 2211-lab3workingPART2.pv
## Last updated: 2022-05-14 EAM
                                             # Mass of the black hole
                                             # EDIT THIS LINE with your result from
                                             → Part 1 of the lab
                                             bh.m = 1.35641e+34
   VISUALIZATION & GRAPH INITIALIZATION
                                             # Other constants - DO NOT CHANGE THESE
# White background so the black hole is

→ FIVE LINES

→ visible

                                             G = 6.7e-11
                                                                      # gravitational
scene.background = color.white

→ constant

                                             planet.m = 9e28
                                                                     # mass of planet
# Visualization (object, trail, origin)
                                             ship.m = 7e10
                                                                      # mass of
bh = sphere(pos=vec(0,0,0))

→ spaceship

→ color=color.black, radius=2e8)
                                             deltat = 1
                                                                      # deltat (in
planet = sphere(pos=vec(1e10,0,0),

→ seconds)

→ color=color.cyan, radius=1e8)

                                                                      # initial time
trailplanet = curve(color=color.cvan)
                                             # DO NOT CHANGE THE ABOVE FIVE LINES
ship =
```

rmag = r

if isinstance(r, vec):

⇒ sphere (pos=vec (planet.pos.x, planet.radiudef) f_q (m_1, m_2, r, debuq=False):

→ color=color.magenta, radius=3e7)

trailship = curve(color=color.magenta)

Finding trajectory of the ship

```
# Maximum time to run the simulation
                                           def resetConds():
# Default is ONE HOUR, but you can change
                                              bh.pos = vec(0,0,0)

→ it to a longer

                                              planet.pos = vec(1e10,0,0)
# amount (e.g., 2hrs, 5hrs, etc) if you
                                              trailplanet.clear()

→ need more time to

                                              ship.pos =
# visualize a full orbit

→ vec(planet.pos.x,planet.radius,0)
tmax = 60 * 60 * 3
                                              trailship.clear()
                                              planet_speed = sqrt(G * bh.m /
                                              → mag(planet.pos) )
# Initial conditions for the planet
                                              planet.vel = vec(0, planet speed, 0)
#planet speed = sqrt(G * bh.m /
                                              print ("Planet's orbital speed:",
→ mag(planet.pos) )
                                              → planet_speed, "m/s")
#planet.vel = vec(0,planet_speed,0)
                                              print ("Planet's orbital period:",
                                              #print("Planet's orbital speed:",
→ planet_speed, "m/s")
                                              \hookrightarrow "hrs")
#print("Planet's orbital period:",
                                              print("---")

→ "hrs")
#print ("---")
                                               # Initial conditions for the

→ spaceship

                                               # EDIT THESE LINES as needed to
## Initial conditions for the spaceship
                                              ## EDIT THESE LINES as needed to change

→ the ship

→ the initial velocity of the ship

                                               # Remember that your spaceship cannot
## Remember that your spaceship cannot
                                              → move faster than the speed of
                                                                                    Georgia
Tech

→ move faster than the speed of light

                                              → light
#ship speed x = 0
                                               ship speed x = 0
                                               ship\_speed\_V = 0 \circlearrowleft V \triangleleft E \lor \triangleleft E \lor
```

#ship speed v = 0

Finding trajectory of the ship

```
# Maximum time to run the simulation
                                           def resetConds():
# Default is ONE HOUR, but you can change
                                              bh.pos = vec(0,0,0)

→ it to a longer

                                              planet.pos = vec(1e10,0,0)
# amount (e.g., 2hrs, 5hrs, etc) if you
                                              trailplanet.clear()

→ need more time to

                                              ship.pos =
# visualize a full orbit

→ vec(planet.pos.x,planet.radius,0)
tmax = 60 * 60 * 3
                                              trailship.clear()
                                              planet_speed = sqrt(G * bh.m /
                                              → mag(planet.pos) )
# Initial conditions for the planet
                                              planet.vel = vec(0, planet speed, 0)
#planet speed = sqrt(G * bh.m /
                                              print ("Planet's orbital speed:",
→ mag(planet.pos) )
                                              → planet_speed, "m/s")
#planet.vel = vec(0,planet_speed,0)
                                              print ("Planet's orbital period:",
                                              #print("Planet's orbital speed:",
→ planet_speed, "m/s")
                                              \hookrightarrow "hrs")
#print("Planet's orbital period:",
                                              print("---")

→ "hrs")
#print ("---")
                                               # Initial conditions for the

→ spaceship

                                               # EDIT THESE LINES as needed to
## Initial conditions for the spaceship
                                              ## EDIT THESE LINES as needed to change

→ the ship

→ the initial velocity of the ship

                                               # Remember that your spaceship cannot
## Remember that your spaceship cannot
                                              → move faster than the speed of
                                                                                    Georgia
Tech

→ move faster than the speed of light

                                              → light
#ship speed x = 0
                                               ship speed x = 0
                                               ship\_speed\_V = 0 \circlearrowleft V \triangleleft E \lor \triangleleft E \lor
```

#ship speed v = 0

Finding trajectory of the ship

```
def calcLoop():
    t = 0
    while t < tmax:
        rate (1000)
        # Net force on the planet
        # EDIT THESE LINES (and/or add
       → more if needed) to find:
        # 1) Fgrav on the planet by the
       → black hole
        # 2) Fgrav on the planet by the

→ spaceship

        # 3) the net force on the planet
        r bh to planet = planet.pos
        F bh ON planet = f g(planet.m,
       ⇔ bh.m, -r_bh_to_planet)
        r ship to planet =

→ planet.pos-ship.pos

        F_ship_ON_planet = f_g(ship.m,
       → planet.m, -r ship to planet)
        Fnet_ON_planet = F_bh_ON_planet +

→ F_ship_ON_planet
```

CALCULATION LOOP

```
# 2) Fgrav on the spaceship by

→ the planet

# 3) the net force on the

→ spaceship

r_bh_to_ship = ship.pos
F_bh_ON_ship = f_q(ship.m, bh.m,
\hookrightarrow -r bh to ship)
r_planet_to_ship =

→ ship.pos-planet.pos

F planet ON ship = f q(ship.m,
→ planet.m, -r_planet_to_ship)
Fnet_ON_ship = F_bh_ON_ship +

→ F_planet_ON_ship

# print (Fnet_ON_ship)
# MAKING THINGS MOVE
# Apply Newton's 2nd law and the
→ position update formula
# to make the planet and

→ spaceship move

planet.vel = planet.vel +

→ (Fnet_ON_planet/ planet.m)

→ deltat

planet.pos = planet.pos +
→ planet.vel * deltat
```

Finding trajectory of the ship

```
# Adding break conditions, in
                                         # for i in range(-int(3e8), int(3e8),
case of a crash
                                         → int(1e5)):
# Do not edit these lines
                                              # for j in range(-int(3e7),
                                             \hookrightarrow int(3e7), int(1e6)):
if mag(r bh to ship) < bh.radius:
    # print("Oh no, vou got
                                                 i = -1e8 \# -1e1 \# -1e8

→ sucked into the black

                                                  i = 1e7 # 1e7 # 1e7
    → hole!")
                                                  # reset conds
    # break
                                                 resetConds()
    return False
                                                 vel = vec(i, i, 0)
                                                  ship.vel = vec(i, j, 0)
if mag(r ship to planet) <
                                                 print("Try! v =", vel)
→ planet.radius:
    # print("Oh no, you crashed
                                                 if calcLoop():
                                                      print("This works! v =",

    into the planet!")

    # break

→ vel)

    return False
```

bruteForceIsHowNasaLaunchesSatellites()





Results

- Mass of the black hole, $m_{BH} = 1.35641 \cdot 10^{34} \text{ kg}$
- Velocity to visit the black hole, $\vec{v}_{ship} = \langle -1 \times 10^8, 1 \times 10^7, 0 \rangle \text{ ms}^{-1}$
- Velocity to orbit the black hole, $\vec{v}_{ship} = \langle -1 \times 10^1, 1 \times 10^7, 0 \rangle \text{ ms}^{-1}$

Outline

Introduction

Aim of the experiment

Prerequisites

Newton's Second Law Newton's Universal law of Gravitation Initial Conditions

Experiment

Formulae Code

Results

Conclusion

What does it mean? What if...





What does it mean?

Parallel and perpendicular components

- From Newton's Second Law, we now that the net force is the change in net momentum per unit time.
- However, $\frac{\Delta \vec{p}}{\Delta t} = \frac{\Delta m \times \Delta \vec{v}}{\Delta t}$, which means that we need mass, velocity and time taken to calculate rate of change of momentum.
- Force is an abstraction that can be used in other contexts, for example, the Work-Energy Theorem:

$$\Delta K = W(= \vec{F} \cdot \Delta s)$$





What if...

...we wanted to orbit the black hole?

Sure thing! By using logical brute-forcing, I got this answer:

$$\vec{v}_{ship} = \langle -1 \times 10^1, 1 \times 10^7, 0 \rangle \text{ ms}^{-1}$$