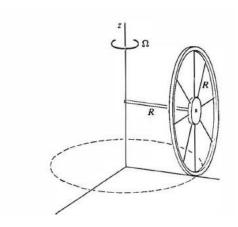
Tutorial 9 Solutions

PHY 101

Q1. A thin hoop of mass M and radius R rolls without slipping about the z axis. It is supported by an axle of length R through it's centre, as shown. The hoop circles around the z axis with angular velocity Ω .

- (a) What is the instantaneous angular velocity ω of the hoop?
- (b) What is the angular momentum L of the hoop? Is L parallel to ω ?

(Note: the moment of inertia of the hoop for an axis along its diameter is $\frac{1}{2}MR^2$.)



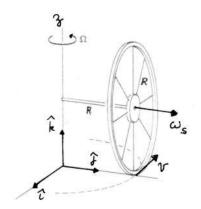
Soln:

8.1 Rolling hoop

(a)
$$\omega_s = \frac{v}{R} = \frac{\Omega R}{R} = \Omega$$

$$\omega = \omega_s + \Omega = \Omega(\hat{\mathbf{j}} + \hat{\mathbf{k}})$$

(b)
$$\begin{split} \mathbf{L} &= \mathbf{L}_s + \mathbf{L}_\omega = I_s \omega_s + I_z \mathbf{\Omega} \\ I_s &= MR^2 \qquad I_z = I_0 + MR^2 = \frac{3}{2} MR^2 \\ \mathbf{L} &= MR^2 \left(\omega_s + \frac{3}{2} \mathbf{\Omega} \right) = MR^2 \, \Omega(\hat{\mathbf{j}} + \frac{3}{2} \hat{\mathbf{k}}) \end{split}$$

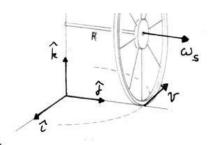


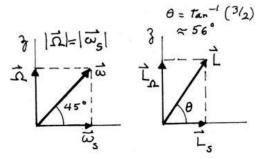
(b)
$$\mathbf{L} = \mathbf{L}_{s} + \mathbf{L}_{\omega} = I_{s}\omega_{s} + I_{z}\mathbf{\Omega}$$

$$I_{s} = MR^{2} \qquad I_{z} = I_{0} + MR^{2} = \frac{3}{2}MR^{2}$$

$$\mathbf{L} = MR^{2}\left(\omega_{s} + \frac{3}{2}\mathbf{\Omega}\right) = MR^{2}\Omega(\hat{\mathbf{j}} + \frac{3}{2}\hat{\mathbf{k}})$$

The lower sketches show that ω and \mathbf{L} are not parallel. We treated \mathbf{L} as the angular momentum of a body with moment of inertia from the parallel axis theorem. \mathbf{L} can also be viewed as the sum of orbital angular momentum $MR^2\Omega$ plus spin angular momentum $(1/2)MR^2\Omega$.





Q2. A bowling ball of mass 4.0kg, a moment of inertia of 1.6×10^{-2} kgm² and a radius of 0.10m. If it rolls down the lane without slipping at a linear speed of 4msec⁻¹, what is its total energy?

Solution

The total (kinetic) energy of an object which rolls without slipping is given by

$$\omega = \frac{v_{\rm CM}}{R} = \frac{(4.0 \, \frac{\rm m}{\rm s})}{(0.10 \, \rm m)} = 40.0 \, \frac{\rm rad}{\rm s}$$

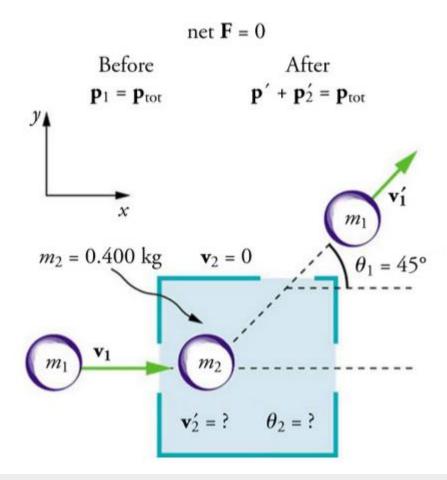
$$K_{\text{roll}} = \frac{1}{2} I_{\text{CM}} \omega^2 + \frac{1}{2} M v_{\text{CM}}^2$$

= $\frac{1}{2} (1.6 \times 10^{-2} \,\text{kg} \cdot \text{m}^2) (40.0 \,\frac{\text{rad}}{\text{s}})^2 + \frac{1}{2} (4.0 \,\text{kg}) (4.0 \,\frac{\text{m}}{\text{s}})^2$
= 44.8 J

Q3. A 0.250-kg object (m_1) is slid on a frictionless surface into a dark room, where it strikes an initially stationary object with mass of 0.400 kg (m_2). The 0.250-kg object emerges from the room at an angle of 45.0° with its incoming direction. The speed of the 0.250-kg object is originally 2.00 m/s and is 1.50 m/s after the collision. Calculate the magnitude and direction of the velocity (v_2 and v_2) of the 0.400-kg object after the collision.

Momentum is conserved because the surface is frictionless. The coordinate system shown in the figure is one in which m_2 is originally at rest and the initial velocity is parallel to the x-axis, so that conservation of momentum along the x- and y-axes is applicable.

Everything is known in these equations except v'_2 and ϑ_2 , which are precisely the quantities we wish to find. We can find two unknowns because we have two independent equations: the equations describing the conservation of momentum in the x- and y-directions.



Solving $m_1v_1=m_1v_1'\cos\theta_1+m_2v_2'\cos\theta_2$ for $v_2'\cos\theta_2$ and $0=m_1v_1'\sin\theta_1+m_2v_2'\sin\theta_2$ for $v_2'\sin\theta_2$ and taking the ratio yields an equation (in which θ_2 is the only unknown quantity. Applying the identity $\left(\tan\theta=\frac{\sin\theta}{\cos\theta}\right)$, we obtain:

$$an heta_2 = rac{v_1' \sin heta_1}{v_1' \cos heta_1 - v_1}.$$

Entering known values into the previous equation gives

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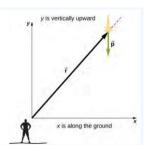
$$\tan \theta_2 = \frac{(1.50 \text{ m/s})(0.7071)}{(1.50 \text{ m/s})(0.7071) - 2.00 \text{ m/s}} = -1.129.$$
 8.69

Thus,

$$heta_2 = an^{-1}(-1.129) = 311.5^{\circ} pprox 312^{\circ}.$$

Angles are defined as positive in the counter clockwise direction, so this angle indicates that m_2 is scattered to the right in Figure 8.11, as expected (this angle is in the fourth quadrant). Either equation for the x- or y-axis can now be used to solve for v_2 , but the latter equation is easiest because it has fewer terms.

Q4. Q4. A meteor enters Earth's atmosphere as shown in fig. and is observed by someone on the ground before it burns up in the atmosphere. The vector $\vec{r} = 25 \text{ km } \hat{\imath} + 25 \text{ km } \hat{\jmath}$ gives the position of the meteor with respect to the observer. At the instant the observer sees the meteor, it has linear momentum $\vec{p} = (15.0 \text{ kg})(-2.0 \text{ km/s } \hat{\jmath})$, and it is accelerating at a constant 2.0 m/s² (- $\hat{\jmath}$) along its path, which for our purposes can be taken as a straight line. What is the angular momentum of the meteor about the origin, which is at the location of the observer?



We resolve the acceleration into x- and y-components and use the kinematic equations to express the velocity as a function of acceleration and time. We insert these expressions into the linear momentum and then calculate the angular momentum using the cross-product. Since the position and momentum vectors are in the xy-plane, we expect the angular momentum vector to be along the z-axis. To find the torque, we take the time derivative of the angular momentumThe meteor is entering Earth's atmosphere at an angle of 90.0° below the horizontal, so the components of the acceleration in the x- and y-directions are

$$a_x = 0$$
, $a_y = -2.0 \ m/s^2$.

We write the velocities using the kinematic equations.

$$v_x = 0, \ v_y = (-2.0 \times 10^3 \ m/s) - (2.0 \ m/s^2)t.$$

a. The angular momentum is

$$egin{aligned} ec{l} &= ec{r} imes ec{p} = (25.0 \; km \; \hat{i} + 25.0 \; km \; \hat{j}) imes (15.0 \; kg) (0 \, \hat{i} + v_y \, \hat{j}) \ &= 15.0 \; kg [25.0 \; km (v_y) \hat{k}] \ &= 15.0 \; kg \{ (2.50 imes 10^4 \; m) [(-2.0 imes 10^3 \; m/s) - (2.0 \; m/s^2) t] \hat{k} \}. \end{aligned}$$

At t = 0, the angular momentum of the meteor about the origin is

$$\vec{l}_0 = 15.0 \; kg[(2.50 \times 10^4 \; m)(-2.0 \times 10^3 \; m/s)\hat{k}] = 7.50 \times 10^8 \; kg \cdot m^2/s(-\hat{k}).$$

This is the instant that the observer sees the meteor.