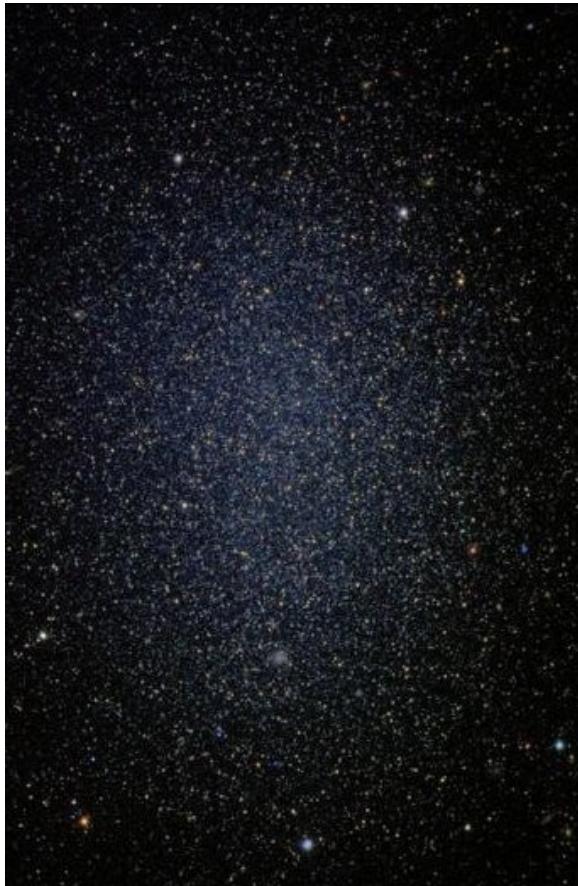


Simultaneous ram pressure and tidal stripping; how dwarf spheroidals lost their gas

Lucio Mayer, Chiara Mastropietro, James Wadsley, Joachim Stadel, Ben Moore

Dwarf Spheroidal



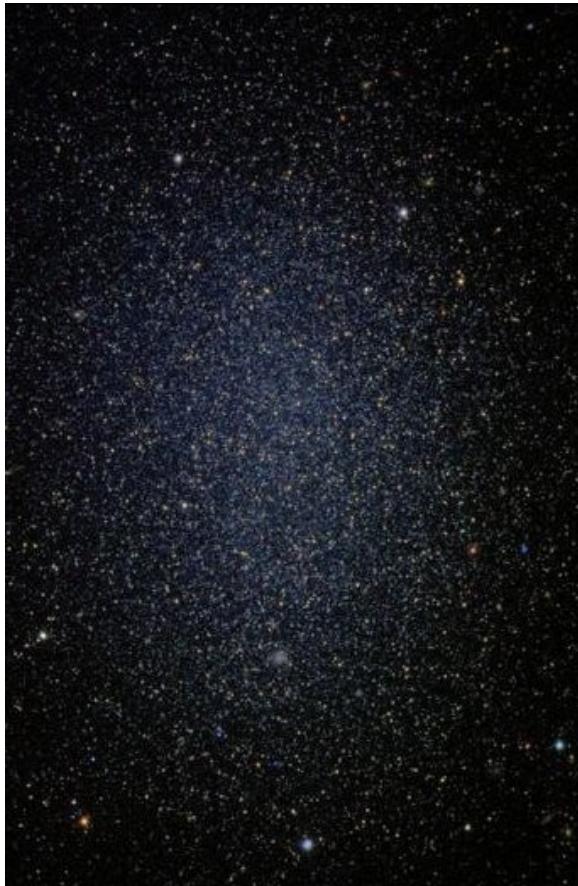
Leo I dSph
250 kpc

Extremely faint galaxies

No significant recent star formation

Gas poor to devoid of gas

Dwarf Spheroidal



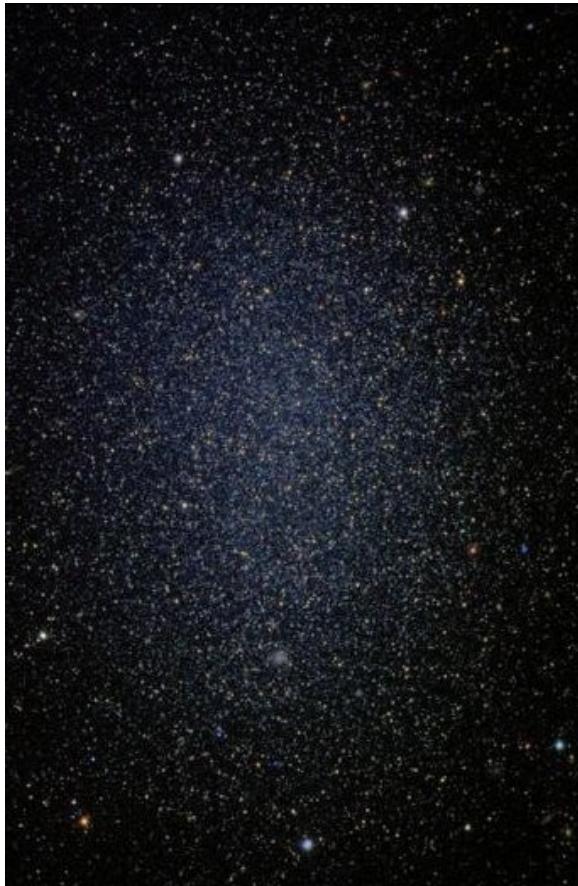
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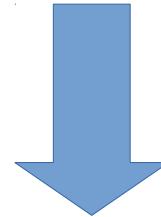


Leo I dSph
250 kpc

Extremely faint galaxies (visual Mag < -14)

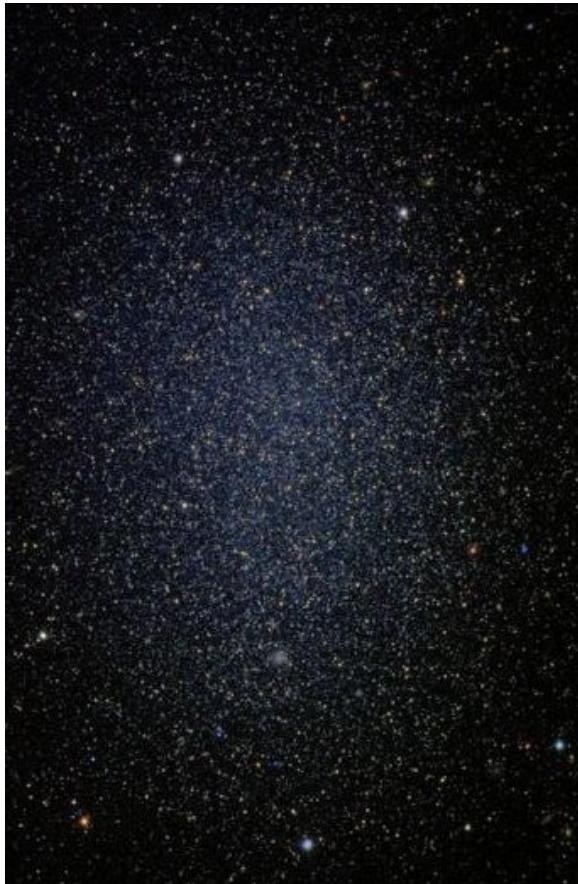
No significant recent star formation

Gas poor to devoid of gas



Possible gas loss during via reionization

Dwarf Spheroidal

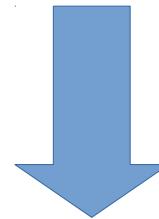


Leo I dSph
250 kpc

Extremely faint galaxies (visual Mag < -14)

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Gas poor to devoid of gas

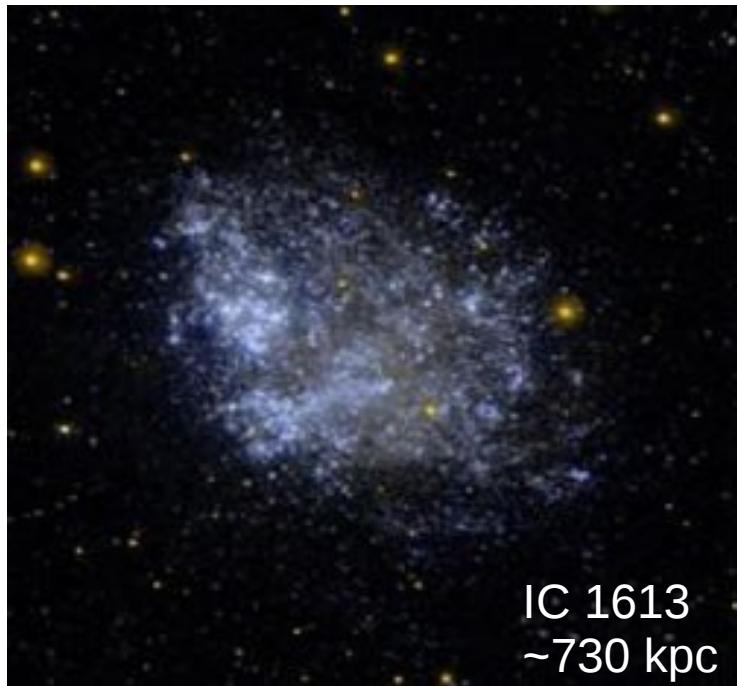


Possible gas loss during via reionization

But what about density morphology relationship?

Dwarf Spheroidal

Dwarf Irregular

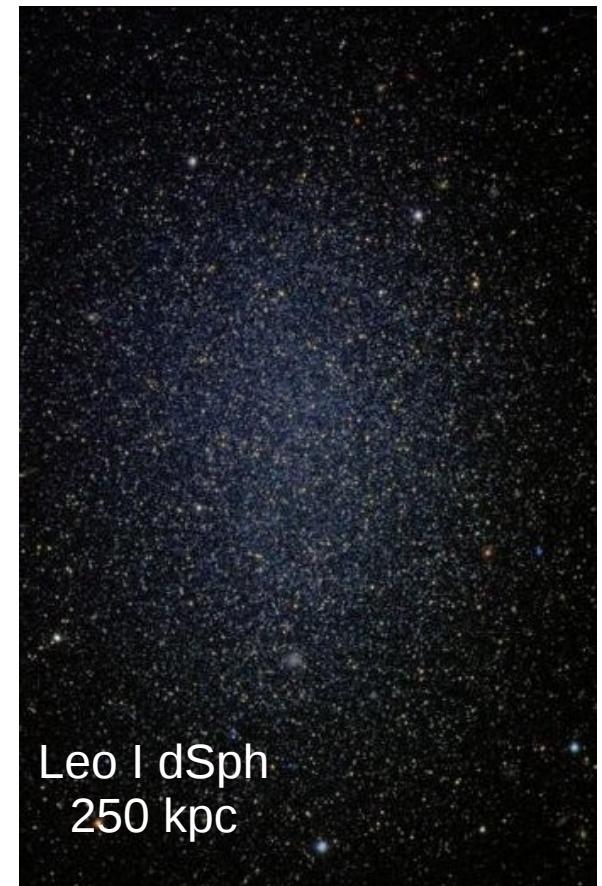


Faint, like dSph... but:

Gas-rich with recent SF

Disc-like

Exponential surface brightness



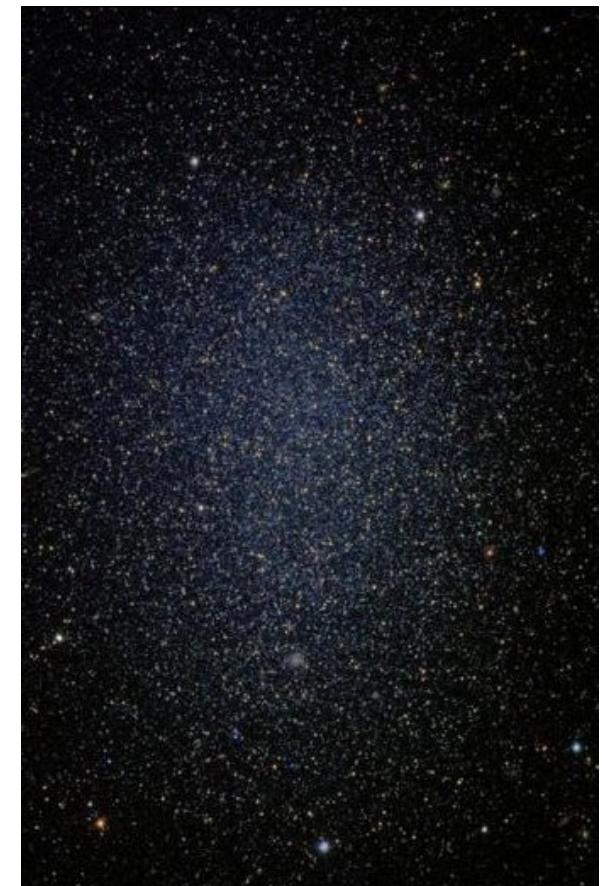
Dwarf Irregular



Preferentially found in
low density regions

Significant stellar rotation

Dwarf Spheroidal



Found as satellites close to a host

Pressure supported

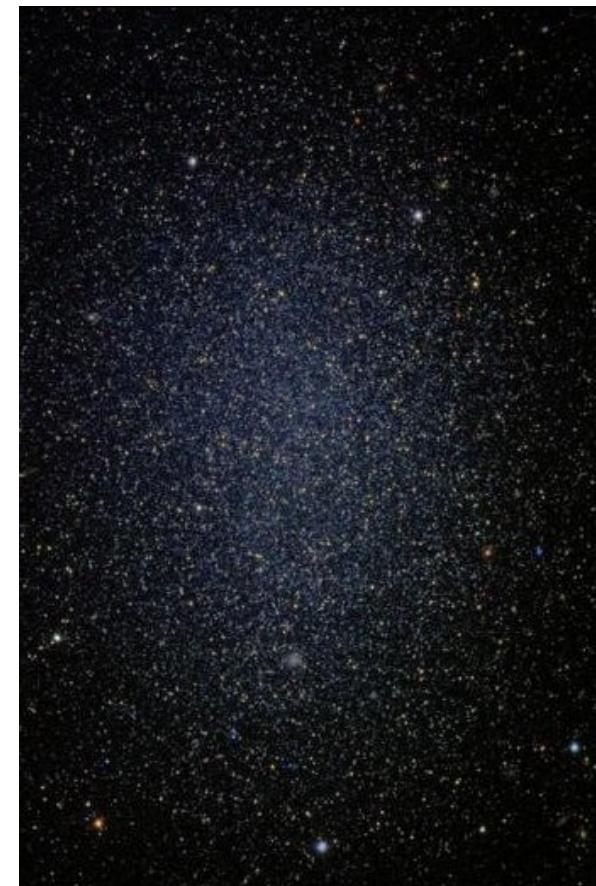
Dwarf Irregular



Is there an
evolutionary path



Dwarf Spheroidal

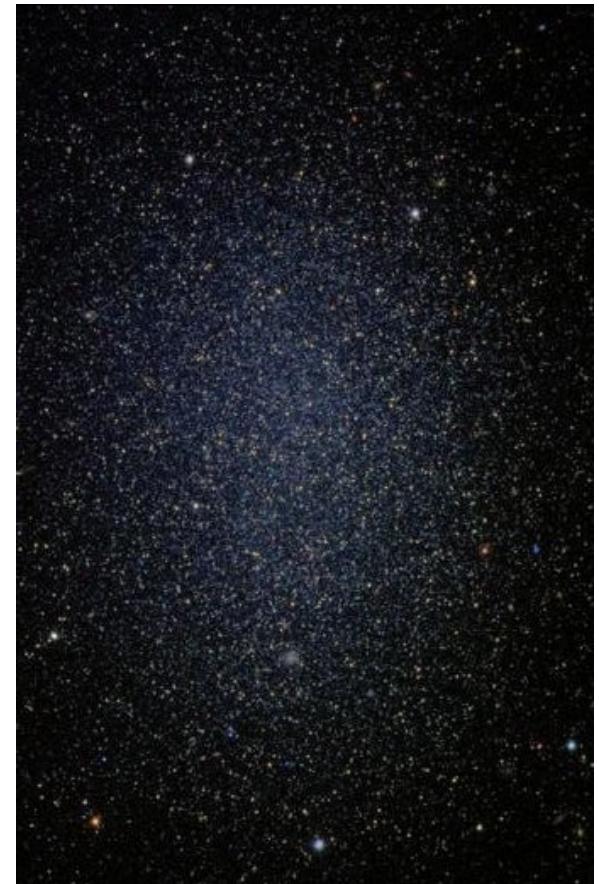
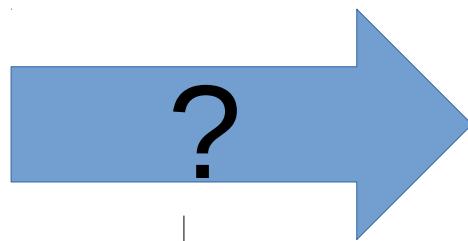


Dwarf Spheroidal

Dwarf Irregular



Is there an
evolutionary path

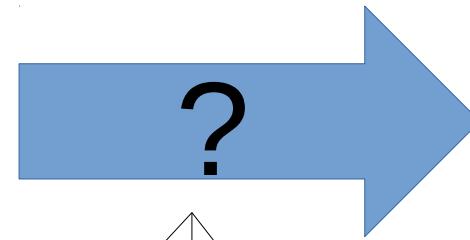


↓
Supernova / Stellar Feedback

Dwarf Irregular

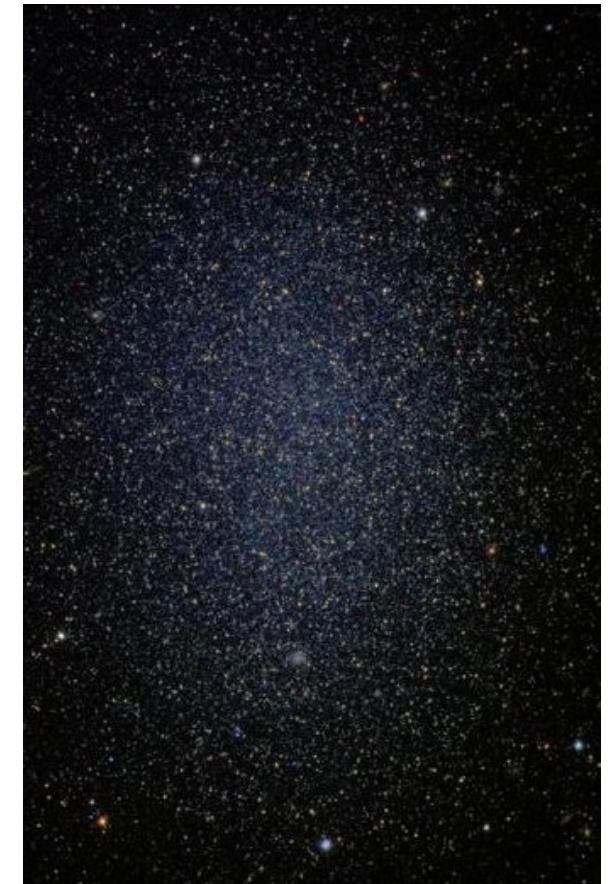


Is there an
evolutionary path



Ram Pressure Stripping

Supernova / Stellar Feedback



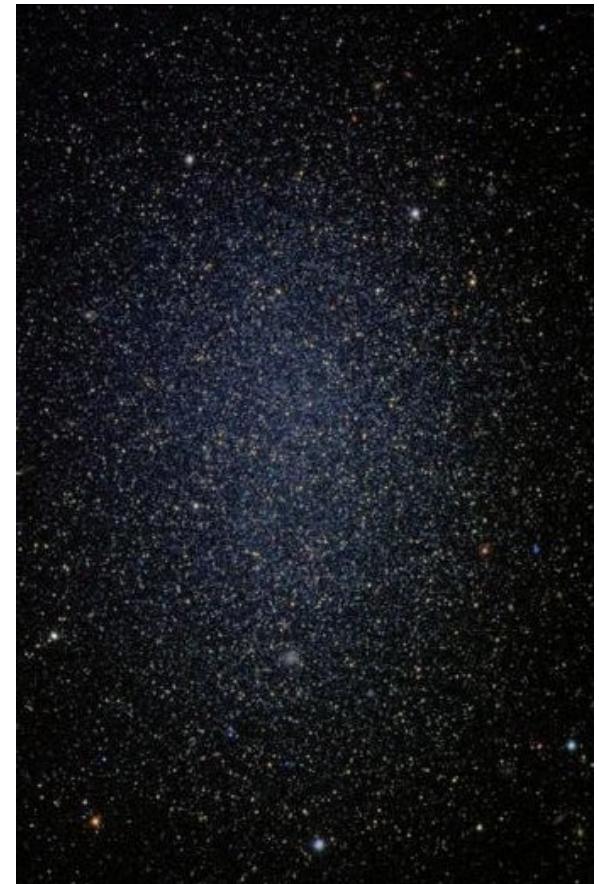
Dwarf Spheroidal

Dwarf Spheroidal

Dwarf Irregular



Is there an
evolutionary path



Ram Pressure Stripping

Tidal Stripping and Interactions

Supernova / Stellar Feedback

Possible Associated SF

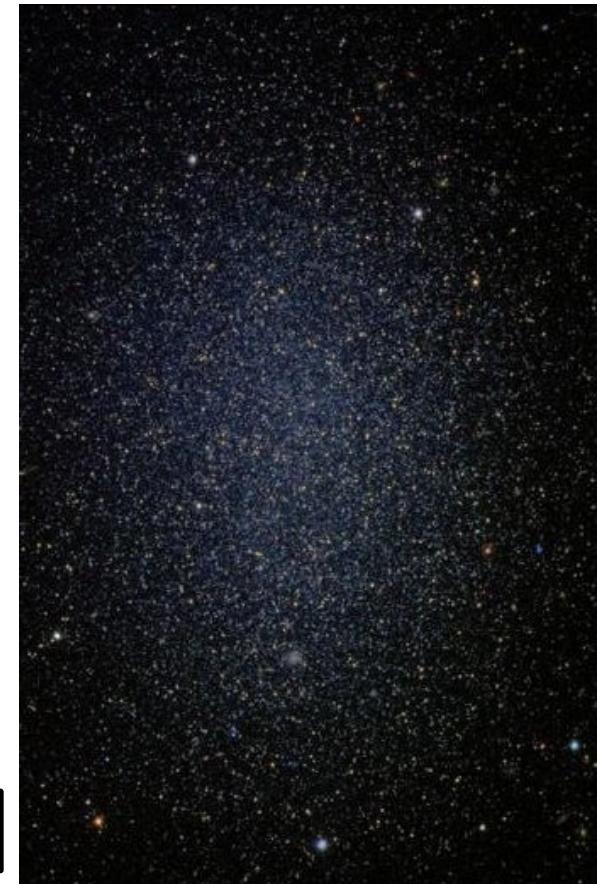


Dwarf Irregular



Is there an
evolutionary path

?



Focus of Mayer et. al. 2006

Ram Pressure Stripping

Tidal Stripping and Interactions

Supernova / Stellar Feedback

► Possible Associated SF

Simulation Design: Galaxy Model

Dwarf: Gas + stars in disc + live NFW halo

Host: Live NFW halo (no disk)

Hot Corona - 10^6 K and 8×10^{-5} cm⁻³ at 50 kpc

3 Dwarfs: Peak rotation velocity = 28, 42, 62 km/s

Concentration = 4, 20, 4

Disc mass ~ 4% of total mass

Simulation Design: Galaxy Model

Dwarf: Gas + stars in disc + live NFW halo

Host: Live NFW halo (no disk)

Hot Corona - 10^6 K and 8×10^{-5} cm⁻³ at 50 kpc

3 Dwarfs: Peak rotation velocity = 28, 42, 62 km/s
Concentration = 4, 20, 4

Disc mass ~ 4% of total mass

Resolution:

Dwarf: 3×10^5 (DM) + 3×10^4 (stars + gas)

Host : 5×10^5 (DM) + 5×10^5 (hot gas)

Dwarf Galaxy Orbits

Bound, eccentric orbits with $r_{\text{apo}}/r_{\text{peri}} = 5$

250 kpc / 50 kpc OR 150 kpc / 30 kpc

Orbital times ~ 3.2 OR 1.2 Gyr

Simulate 3 – 5 orbits

Disc begin at 60 degree incline to orbit

Simulation Code

GASOLINE – Tree+SPH code

Physics Varied in Simulations:

Radiative Cooling: On or Off (adiabatic cooling)

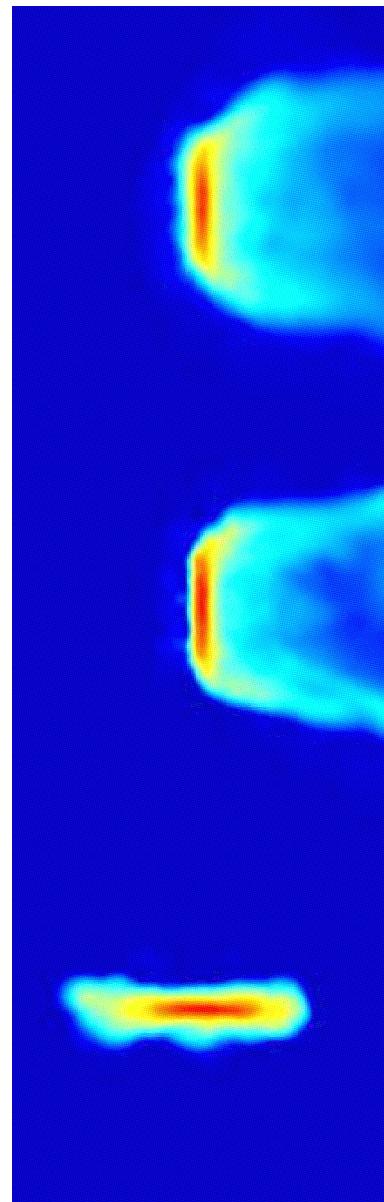
UV Background : On in some RC runs

Wind Tunnel : Some RPS only runs

Wind Tunnel Tests

0.05 Gyr

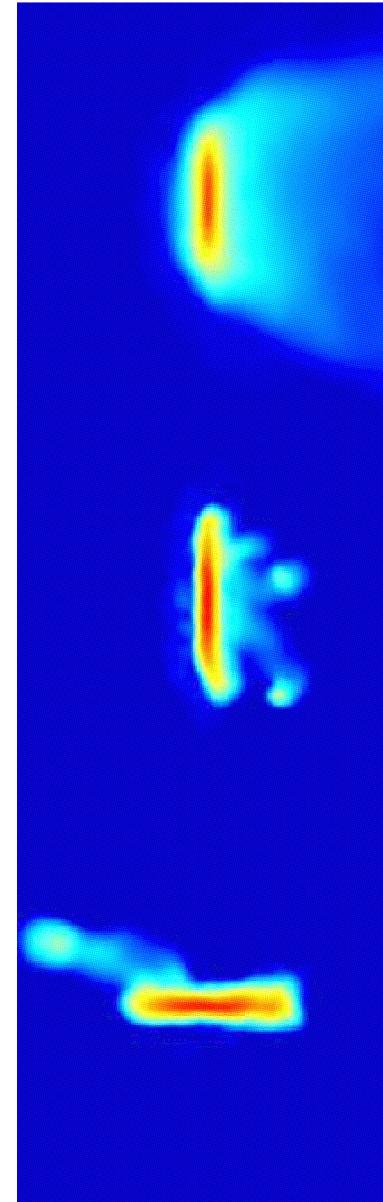
0.3 Gyr



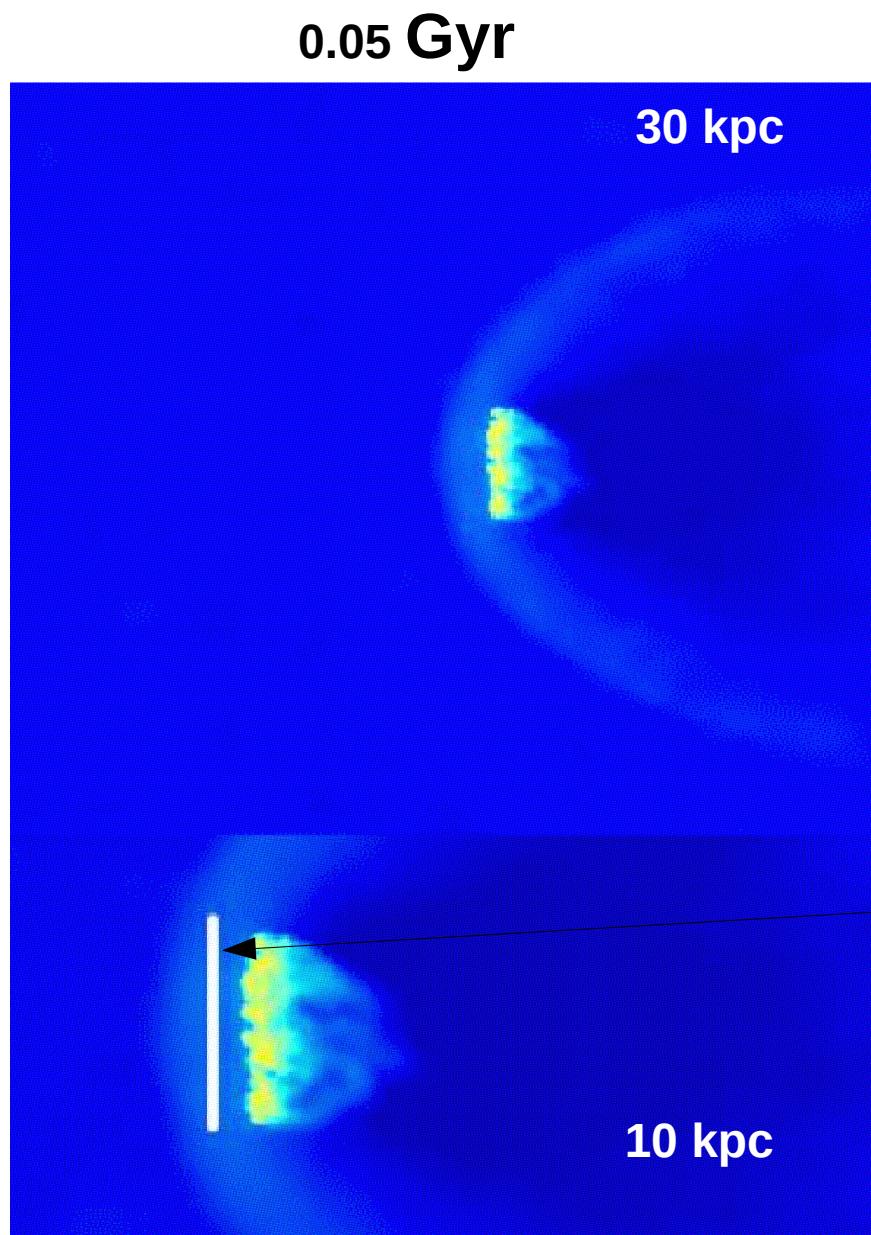
NO RC

RC

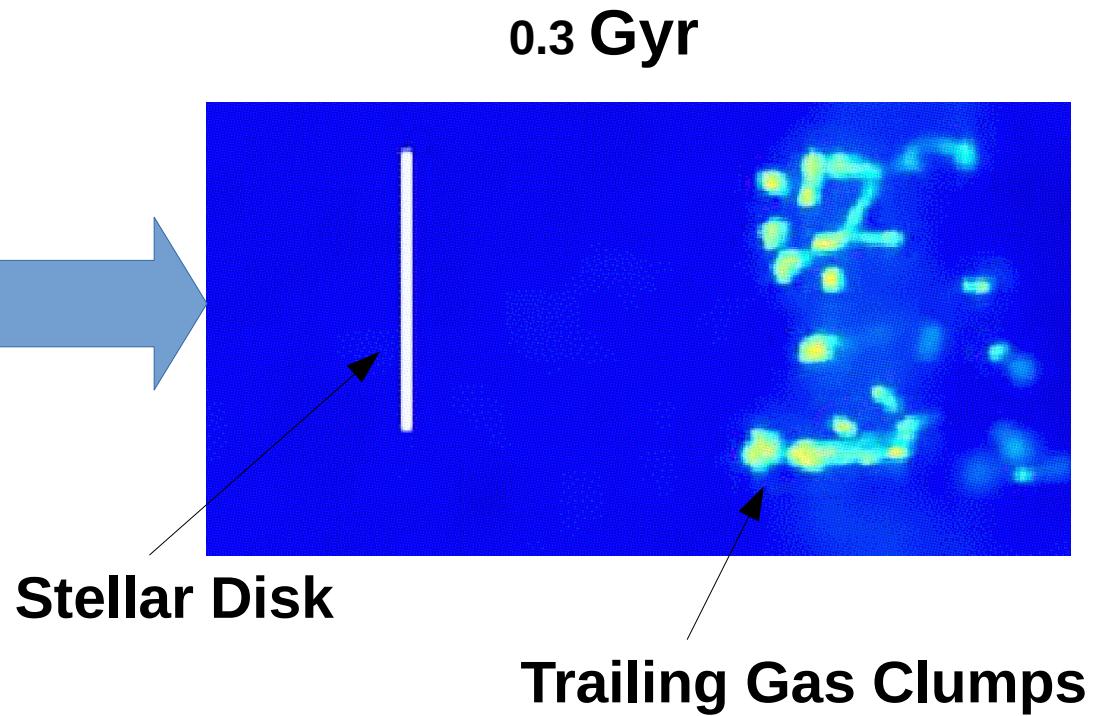
RC, $i = 90$



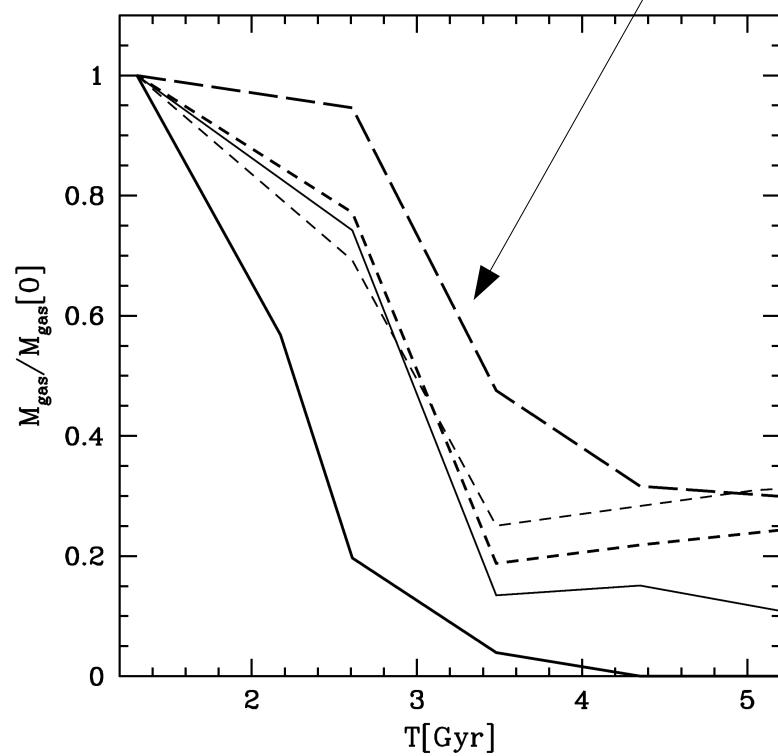
Example of Complete RPS



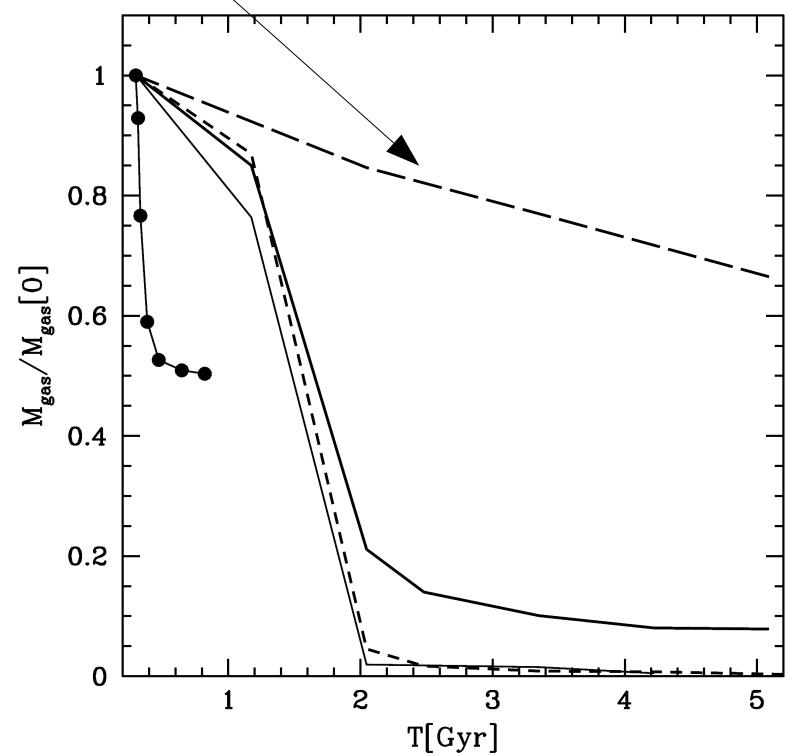
Peak Velocity = 28 km/s
 $c = 4$
Radiative Cooling ON



Tides Alone



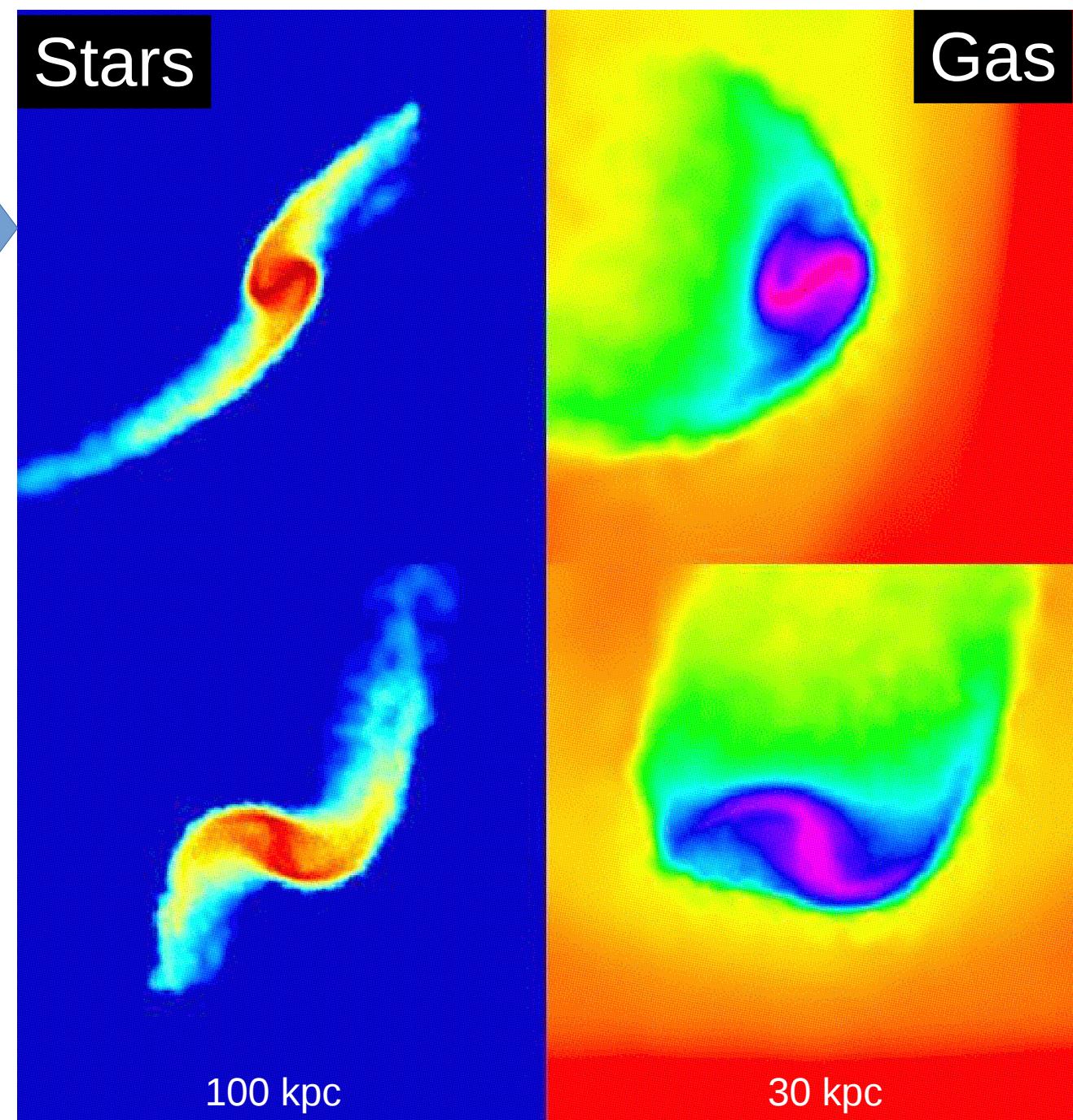
$v_{\text{peak}} = 40 \text{ km/s}$



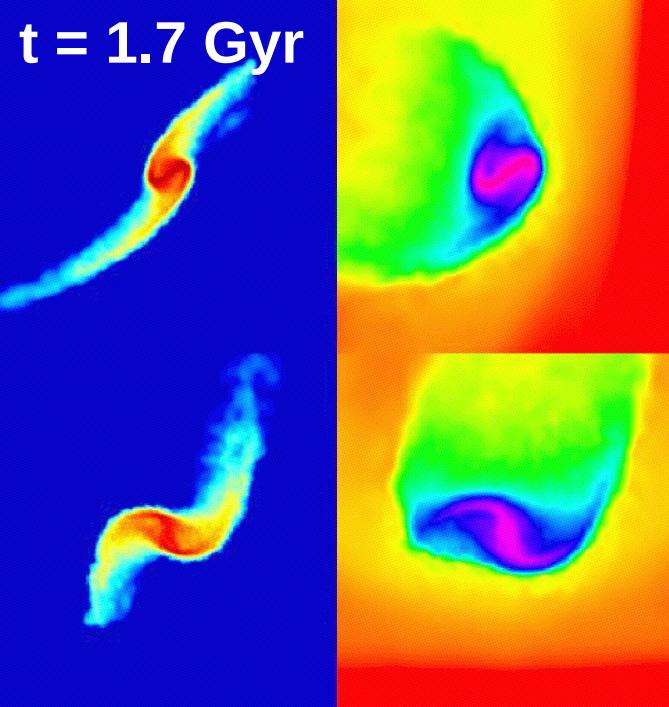
t = 0 Gyr:

- Adiabatic
- Peak Vel ~ 62 km/s
- Concen. $c = 4$
- Gas Mass = 450 Msun

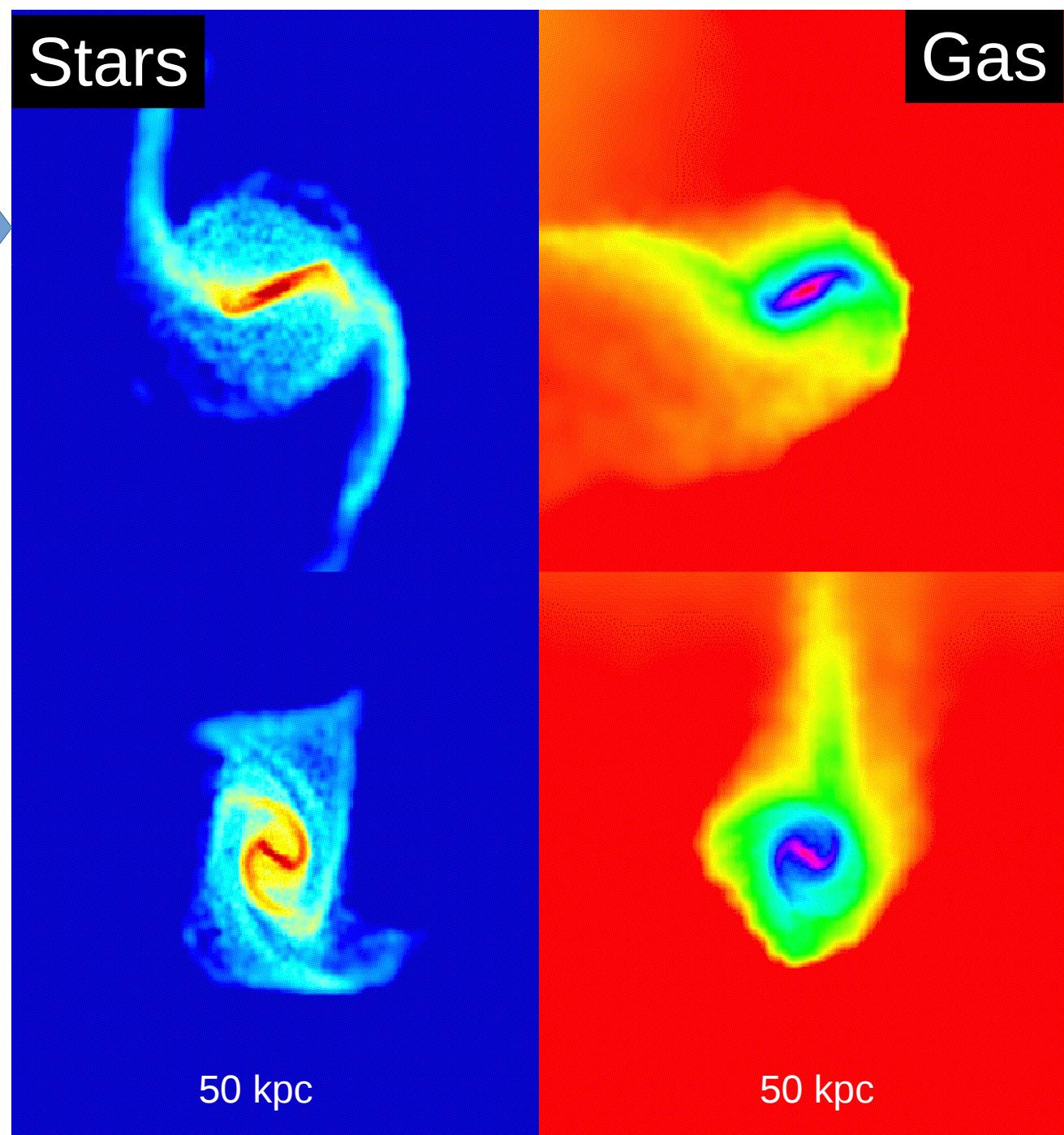
1.7 Gyr after First Pericenter Passage



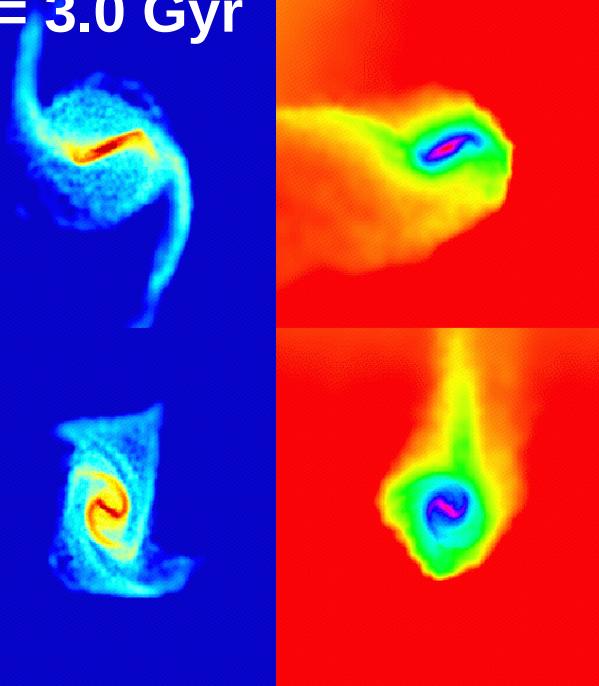
$t = 1.7 \text{ Gyr}$



3.0 Gyr after First Pericenter Passage

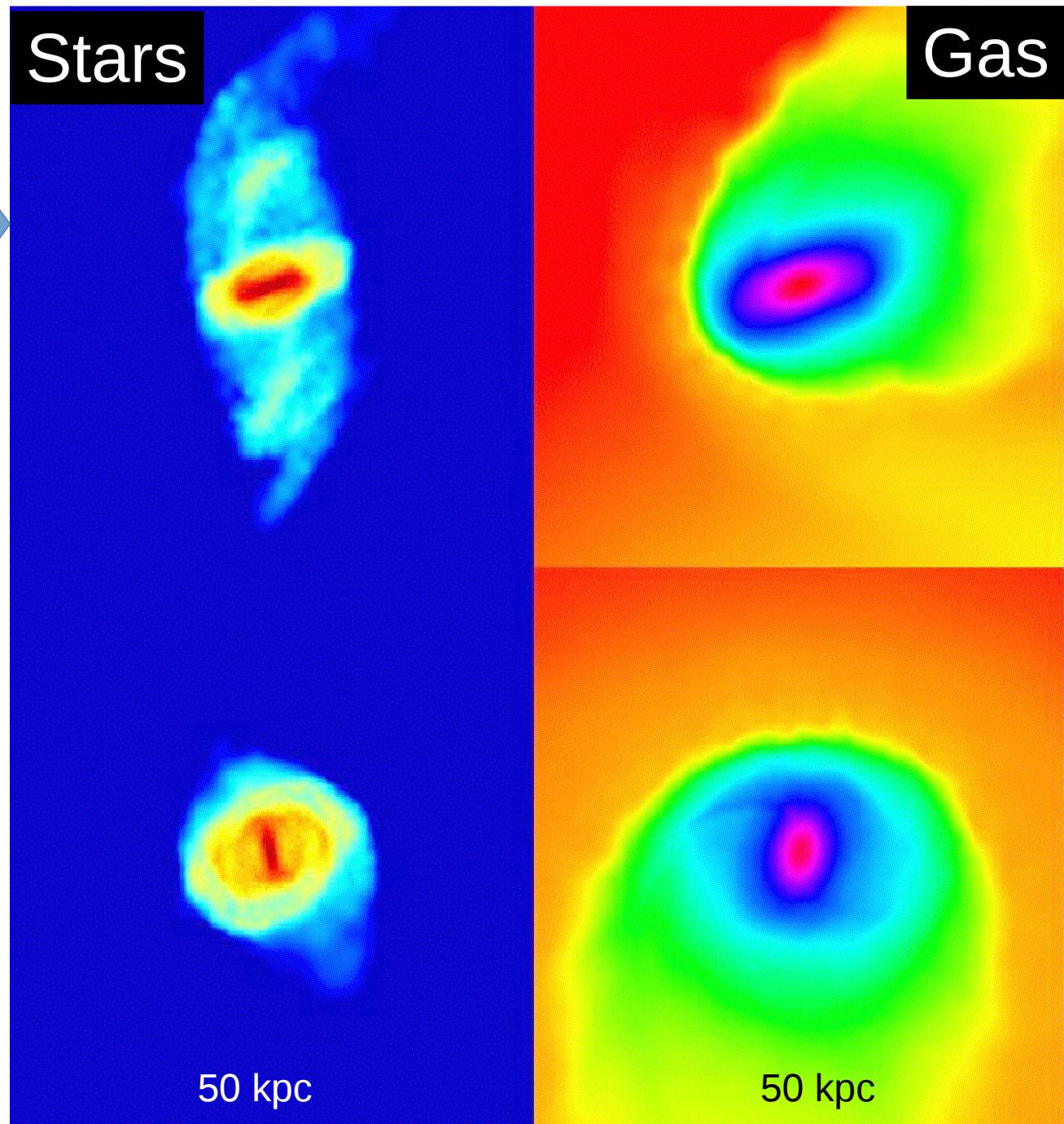


$t = 3.0$ Gyr



6.0 Gyr after Second Pericenter Passage

Stars



Final Gas Mass:
60 Msun

(450 Msun Initial)

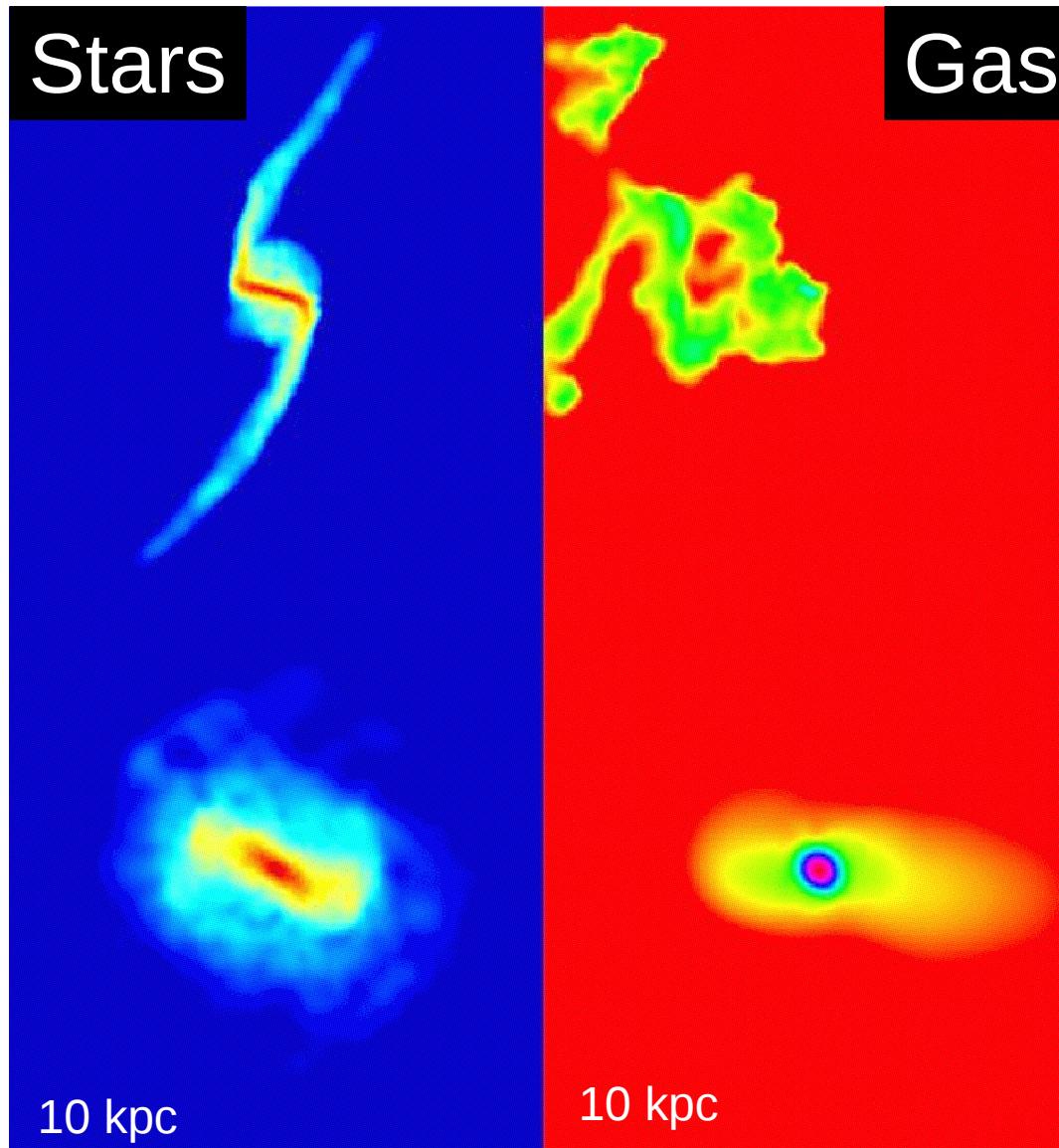
Complete RPS Dwarf + Tidal Effects

After 2nd
Pericenter

t = 1.7 Gyr

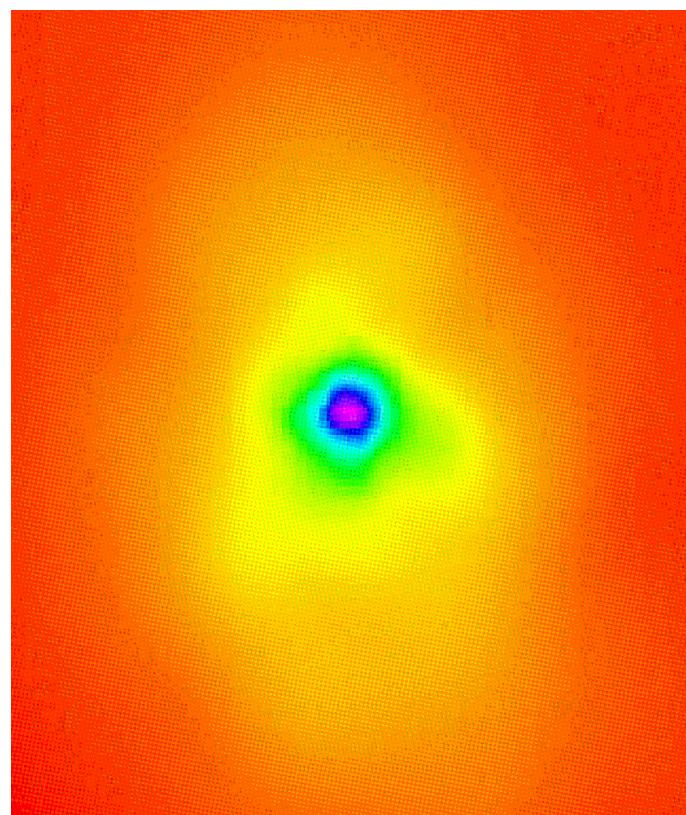
After 2nd
Pericenter

t = 6.0 Gyr

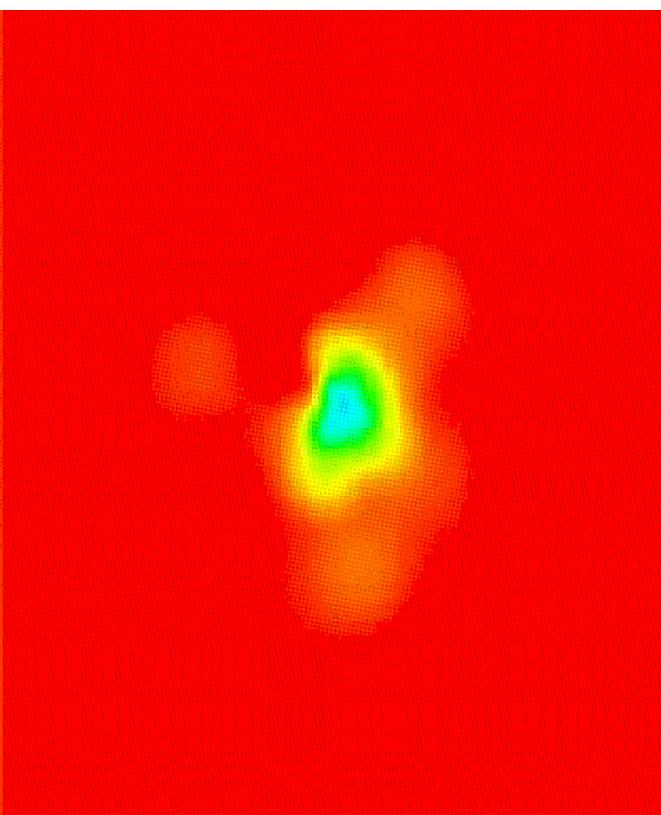


RC + Photoionization

RC – No UV



RC – UV Background



5 Gyr after 3rd pericenter passage

Results Summary

RPS + Tids \rightarrow 50 – 100% gas mass loss

Greater loss with closer pericenters

UV ionizes gas ($M_{\text{HI}} < 0.01 M_{\text{star}}$)

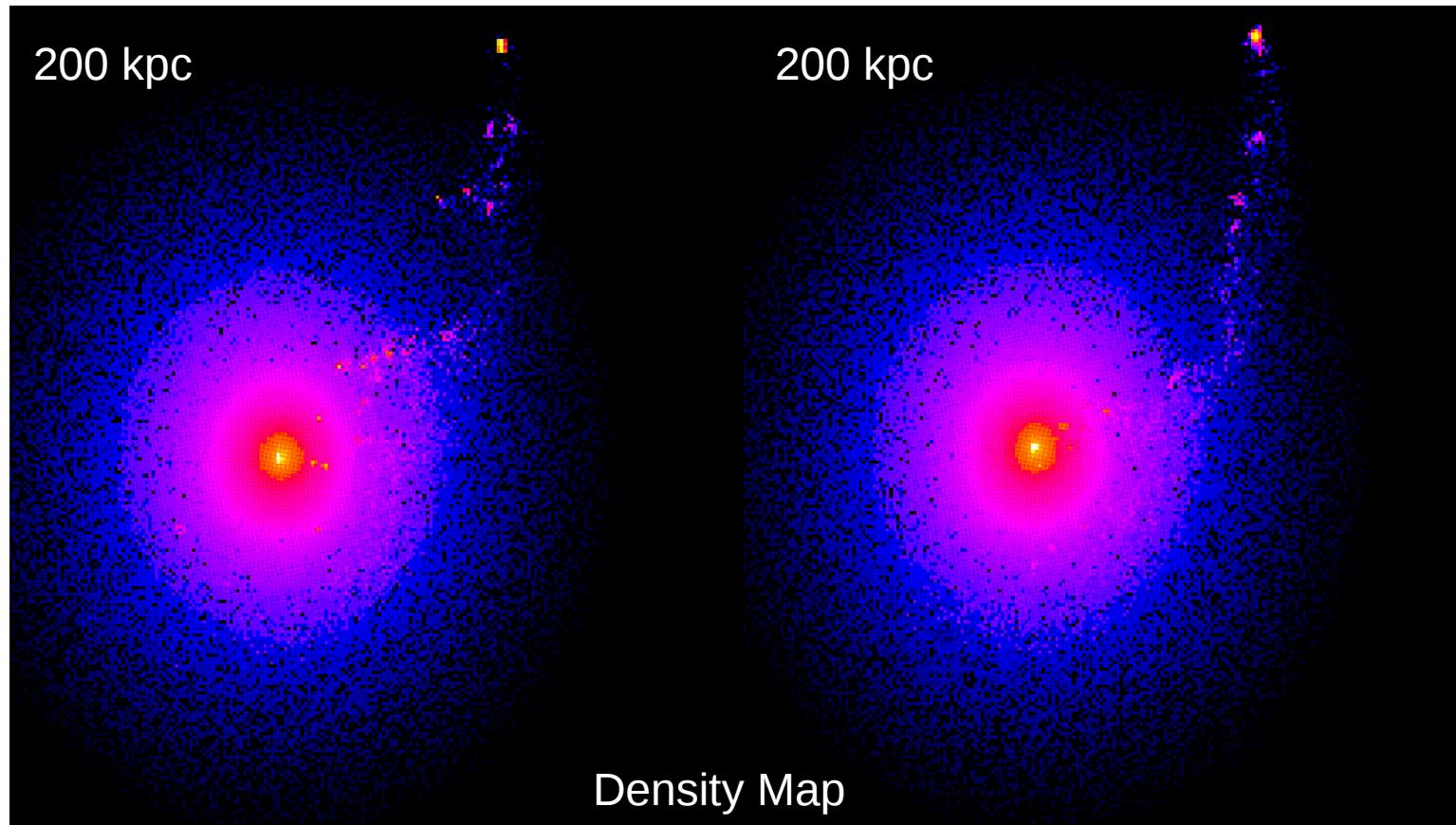
$V_{\text{peak}} < 30 \text{ km/s}$ always stripped

$V_{\text{peak}} > 60 \text{ km/s}$ always retains gas

Potential dwarf elliptical progenitors

Fate of Stripped Gas

$t = 1.5 \text{ Gyr}$

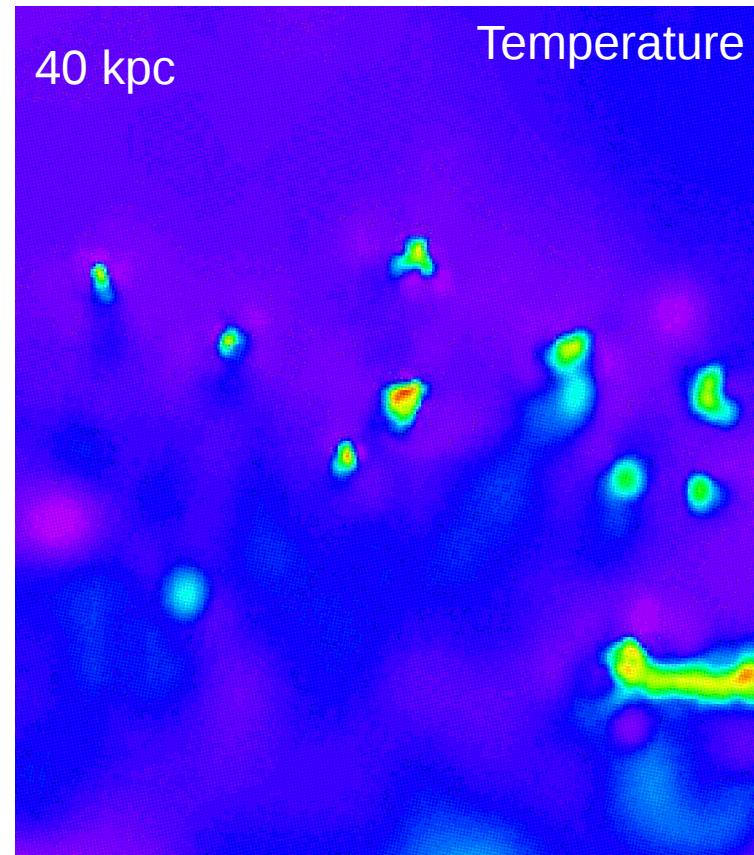


$$V_{\text{peak}} = 40 - RC$$

$$V_{\text{peak}} = 40 - RC + UV$$

Fate of Stripped Gas: Clumps

$t = 1.5 \text{ Gyr}$



Zooming in on
the gas clumps

$$V_{\text{peak}} = 40 - RC$$

Conclusions

RPS alone in some small dwarfs

Compressional heating enhances RPS w/o RC

Tides + RPS enhance total stripping

Tidal bar formation limits RPS

Realistic baryon content and M/L

Cloud fragments with RC realistic