

Graph Convolutional Neural Networks for Position Reconstruction in the XENON1T Experiment

Alejandro Oranday

Advisor: Christopher Tunnell

aeo3@rice.edu

March 17, 2021

1	Contents	
2	Introduction	2
3	Background	2
4	Dark Matter	2
5	XENON1T Detector	2
6	Machine Learning	2
7	References	3

Introduction

Dark matter constitutes 85% of the matter in our Universe and was made evident by observations on galaxy formation, gravitational lensing, and the cosmic microwave background [3]. However, detecting dark matter experimentally is exceedingly difficult with particle physics detectors. These particles are suspected to be weakly interacting [2] such that any detection attempt would need to be sensitive to recoils at keV levels of energy.

The XENON1T detector operated as the most sensitive dark matter detector [1], and the soon-to-be-active XENONnT detector plans to overtake that title [4]. In these detectors, there are two important elements that need to be reconstructed: energy and position. By reconstructing these key elements, we are able to accept or reject numerous observations if the reconstructed position is within the detector’s fiducial volume and if the reconstructed energy is within a rejection threshold [2].

Previous machine learning implementations for position reconstruction perform well enough [5], but still run into issues with reconstruction outside the bounds of the detector as well as having an inward reconstruction bias. Finding the type of machine learning algorithm that is the most appropriate for these problems is essential for the best use of the detector’s fiducial volume. This project produced the first application of graph convolutional neural networks (GCNNs) for position reconstruction in the dark-matter field, and one of the first applications of a GCNN for use in regression.

Background

Dark Matter

XENON1T Detector

The XENON1T detector is a dual phase xenon time projection chamber (TPC) located in the *Laboratori Nazionali del Gran Sasso* (LNGS) in central Italy. The detector aimed to observe weakly interacting massive particles (WIMPs) as the primary candidate for dark matter particles. It was a requirement for the detector to be sensitive to keV energy levels in order to observe these particles and was made possible through a combination of the stable xenon 136 isotope, water shielding, and depth within the Gran Sasso massif.

Machine Learning

References

- [1] E. Aprile, J. Aalbers, F. Agostini, et al. First dark matter search results from the xenon1t experiment. *Physical Review Letters*, 119(18), Oct 2017. ISSN 1079-7114. doi: 10.1103/physrevlett.119.181301. URL <http://dx.doi.org/10.1103/PhysRevLett.119.181301>.
- [2] E. Aprile, J. Aalbers, F. Agostini, et al. Xenon1t dark matter data analysis: Signal reconstruction, calibration, and event selection. *Physical Review D*, 100(5), Sep 2019. ISSN 2470-0029. doi: 10.1103/physrevd.100.052014. URL <http://dx.doi.org/10.1103/PhysRevD.100.052014>.
- [3] Gianfranco Bertone and Dan Hooper. History of dark matter. *Reviews of Modern Physics*, 90(4), Oct 2018. ISSN 1539-0756. doi: 10.1103/revmodphys.90.045002. URL <http://dx.doi.org/10.1103/RevModPhys.90.045002>.
- [4] The XENON collaboration, E. Aprile, J. Aalbers, et al. Projected wimp sensitivity of the xenonnt dark matter experiment, 2020.
- [5] B. E. J. Pelssers. Position reconstruction and data quality in xenon. Master's thesis, University of Utrecht, July 2015. URL <https://dspace.library.uu.nl/handle/1874/322783>.