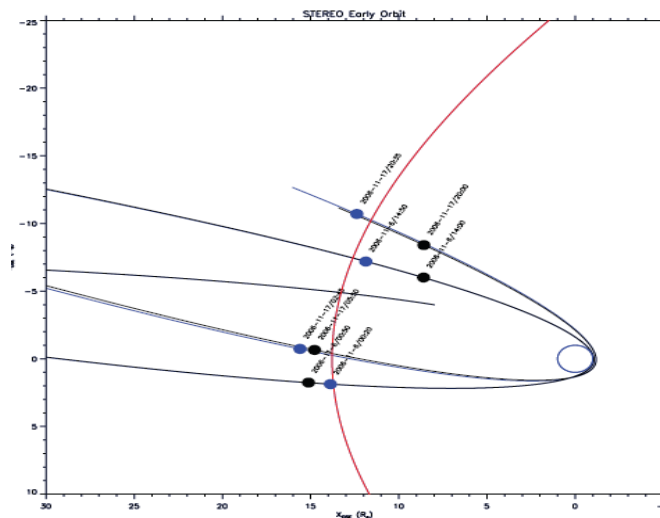
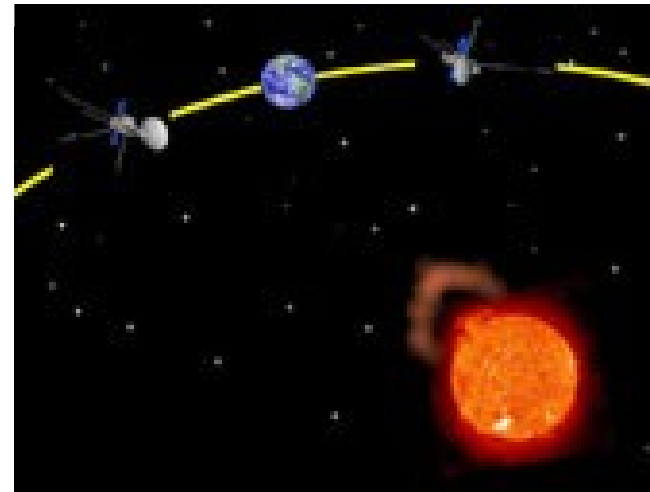
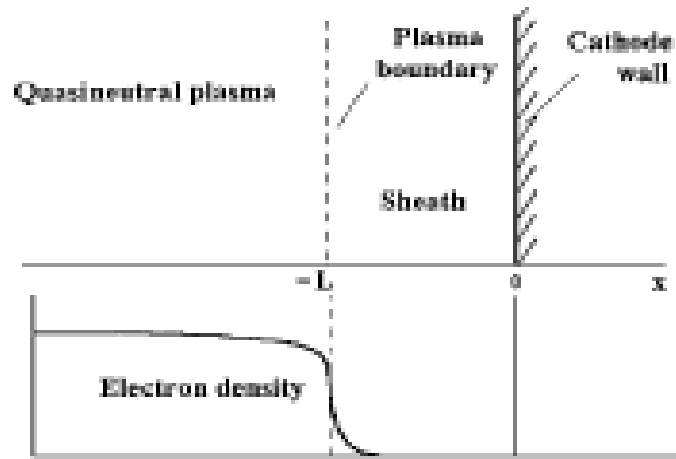


- Direction finding easier when antennas orthogonal
- ...not like on STEREO...
- Direction finding never worked with Cassini, STEREO....
- Cygnus ... for in-flight calibration

The Physics of the plasma sheath



- Antenna behavior is influenced by the surrounding space plasma.
- The influence is highest near the plasma resonance frequencies but also noticeable at the higher parts of the typical radio experiment frequency range.
- Due to the trajectory it could be important for the SOLAR ORBITER RPW antenna calibration.
- A dielectric model of the space plasma can be incorporated in the numerical antenna calibration.

- The effect of plasma is encapsulated in the Green's tensor, which shape depends on the plasma model used.

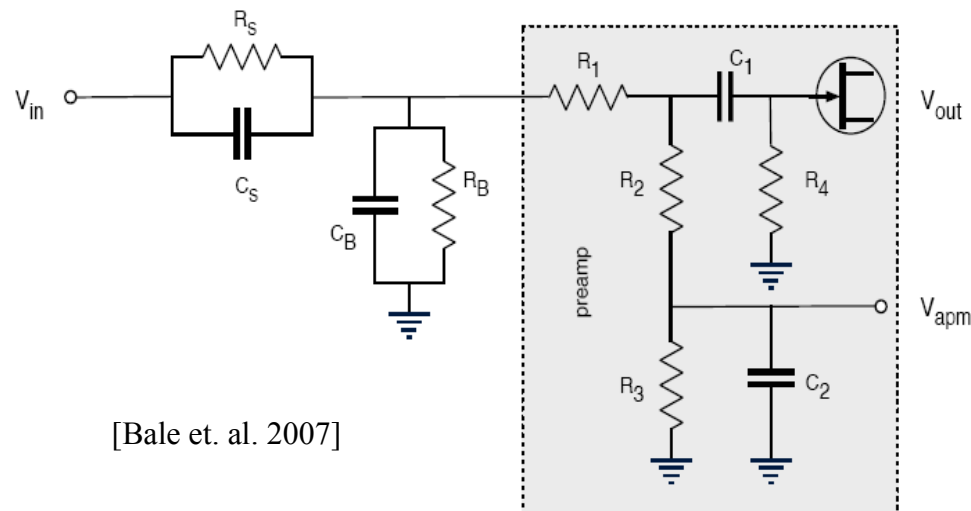
$$\mathbf{E}(\mathbf{r}, \omega) = \int_{V'} \mathbf{G}(\mathbf{r}, \mathbf{r}') \mathbf{j}_{ant}(\mathbf{r}', \omega) dV'$$

- The general form of the response tensor is

$$\mathbf{G}(\mathbf{r}, \mathbf{r}') = \frac{\mu_0 \omega}{(2\pi)^3} \int_{-\infty}^{\infty} \frac{adj(\lambda)}{\det(\lambda)} e^{i(\mathbf{k} \cdot (\mathbf{r} - \mathbf{r}'))} d\mathbf{k}$$

- The resulting tensor can be incorporated into the electric field integral equation.

- Spacecraft antennas are usually coupled to the surrounding space plasma.
- The electromagnetic coupling can be modeled by a system of a resistance and a capacitance.
- In rarefied plasma the coupling can not take place without photoelectrons.
- The sheath thickness must not be larger than the antennas.



[Bale et. al. 2007]

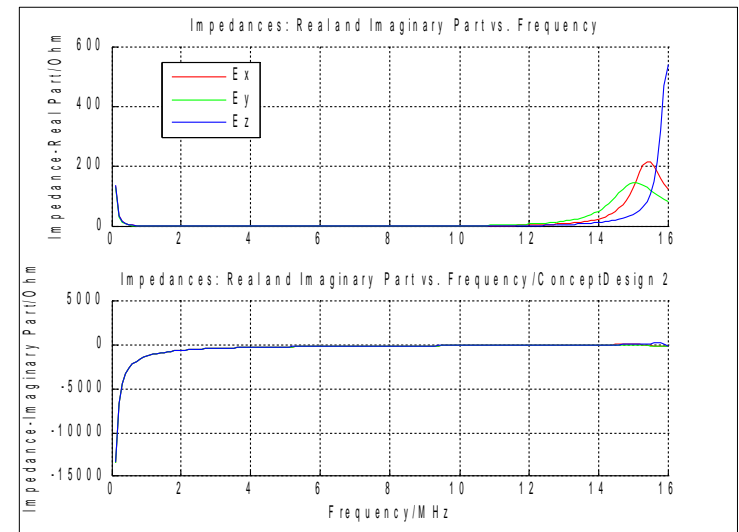
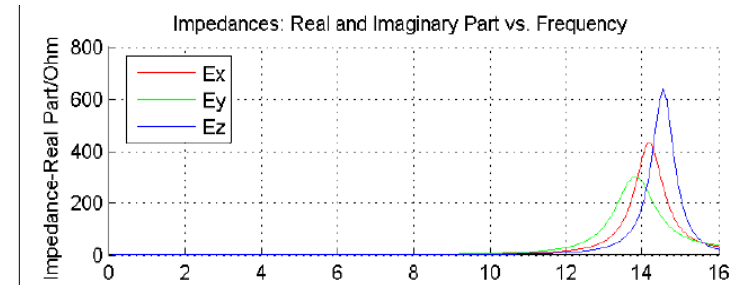
- STEREO operates in solar wind conditions at 1AU.
- The photo-electron production rate is higher than the thermal electron impact rate. --> positive charge.
- $i_{ph} \dots 10^{-4} \text{Am}^{-2}$
- $A_{rel} \dots 0.5$
- $l = 6 \text{m}$
- $d = 1 \text{in}$ (0.0254m) on average
- Mean energy of photo-electrons = 1.5eV
- Mean energy of thermal electrons = 10eV

$$A_{rel} i_{ph} l d \sim 7.6 \cdot 10^{-6} \text{A}$$

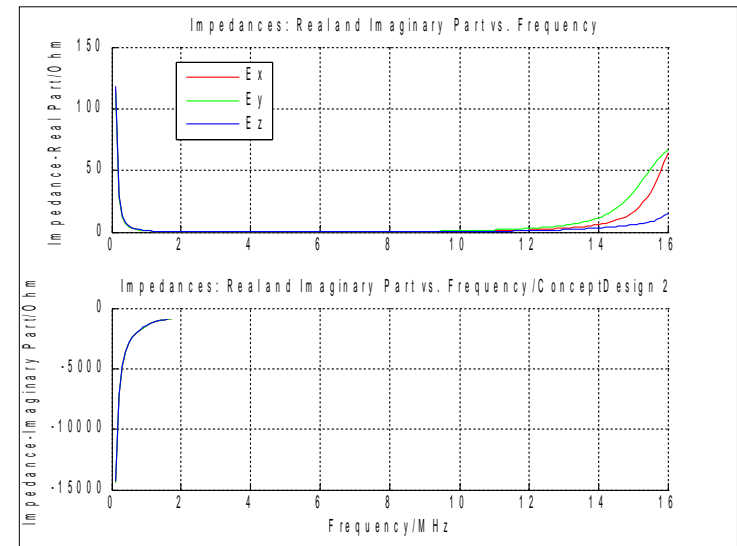
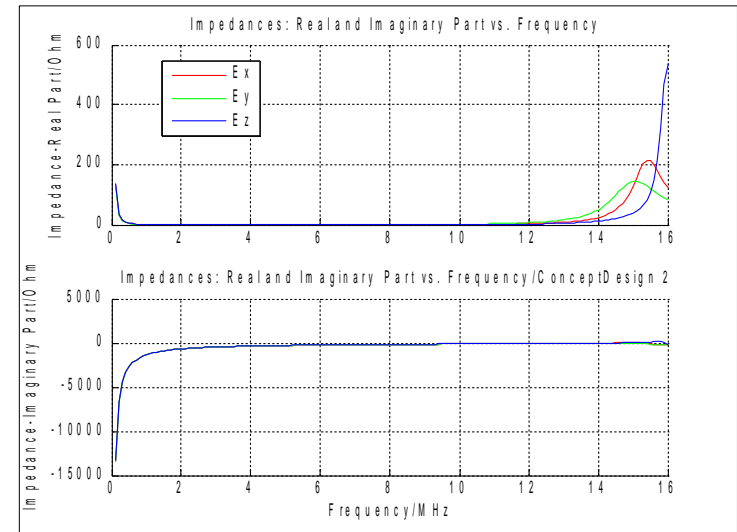
$$I_e = -e n_e d \pi \sqrt{\frac{\kappa T_e}{2\pi m_e}} \sim -2 \cdot 10^{-7} \text{A}$$

- --> $V = 5.5 \text{V}$
- $\bar{n}_e = 10^6 \text{m}^{-3}$
- --> $n_{ph}(0) = 2 \times 10^8 \text{m}^{-3}$
- --> 2.5% of the photo-electrons reach the plasma.
- --> $\lambda_{sh} = 85 \text{cm}$, using the photo-electron Debye length at the surface.

- Using the appropriate equations, one finds:
 - $R_s = 0.2 \text{ M}\Omega$
 - $C_s = 87 \text{ pF}$
- Via these parameters the sheath can be included into the numerical antenna calibration (wire-grid).
- No calculations for the effective length vectors were done so far.
- Computation of the impedances show that the inclusion of the sheath has an effect.



- Stuart Bale estimates
 - $R_s = 0.75 \text{ M}\Omega$
 - $C_s = 40 \text{ pF}$
- The figures show the impedances using Bale's values in relation to the results using our theoretical model.



- Comparing the location of the second resonance of antenna E_z with measured data shows that the inclusion of the sheath capacitance has a corrective effect.
- Further and more accurate data is needed to verify the model.

