

# A (partially implemented) hydrodynamics code

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Physics 6810

April 2023

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## 1 Overview

Here, I describe some of the structure of the code here, and since the program is heavy on the equations, I also note the critical equations. See the README.md file for a description of how to execute the program.

A note about Julia (since I am not sure how much you have used/seen the language). Julia uses unicode (as a language feature even), so the source code does use unicode symbols for variables (like  $\rho$ ,  $\phi$ , etc.). Also, Julia is compiled at runtime so there is no need to worry about makefiles. Julia also has some very abbreviated syntax for arithmetic (like  $2x^2$ ) and vectorization (just add a `.`).

## 2 Structure

The main body of the code is in the `src/` directory. This directory includes the files

- `GalaxySim.jl`. This just imports and exports other pieces of the project.
- `evolve.jl` contains the main loop of the simulation, including the leapfrog integration scheme and time-step criteria
- `gal_files.jl` writes the simulation outputs to files. (Unfortunately, other io to files for testing are scattered through the project)
- `density.jl` contains routines for density estimation.
- `gravity.jl` calculates the gravity
- `physics.jl` all the rest of the physics (hydrodynamics, viscosity, etc.)
- `particles.jl` definition of the `Particle` struct

- `params.jl` struct to read in Params (stored in `init/` directory)
- `constants.jl` Physical constants in cgs (which the code uses internally)

### 3 Physics

The hydrodynamic equations using lagrangian derivatives are

$$\nabla \rho = ? \quad (1)$$

Smoothed Particle Hydrodynamics (SPH) is a variant of the Lagrangian method for solving hydrodynamics equations. Instead of dividing space up into a grid, each particle is followed and the physical properties are calculated in the frame of each particle.

First, we need a way to estimate the density at any given point from the distribution of nearby points. We do this using a weighted sum over the nearby points of a particle. The density of particle  $j$  is

$$\rho_j = \sum_i m_i W(r_{i,j}, h_j), \quad (2)$$

where the sum index  $i$  is over all points and  $W(r, h)$  is the kernel weight at a distance  $r$  with a smoothing length  $h$ .

### 4 Implementation

I follow a variety of sources to use standard smoothed particle hydrodynamics (SPH) to implement the physics (Monaghan, 1992, 2005; Price & Monaghan, 2007; Price et al., 2018; Pasetto et al., 2010; Price, 2012; Springel et al., 2001). The idea (as you probably know) is to estimate the density with a kernel. The kernel also has a smoothing length  $h$ , which should represent the mass inside the smoothing sphere, i.e.

$$h = \eta \left( \frac{m}{\rho} \right)^{1/3} \quad (3)$$

where  $\eta$  is density parameter. This system can be solved using Newton-Raphsons method. I follow Monaghan (2005) and use the function

$$f(h) = \rho - \rho_{\text{new}} \quad (4)$$

where  $\rho$  is calculated from the current value of  $h$  in Eq. 3, and  $\rho_{\text{new}}$  is calculated from the summation above.

So each new  $h$  is found with

$$h_{\text{new}} = h - \frac{f(h)}{f'(h)} \quad (5)$$

Other physics (like the change in density, position, etc.) are the standard SPH equations (Monaghan, 2005, 1992), except gravity is done following SPH using a smoothed kernel as described in Price & Monaghan (2007).

Integration is leapfrog, so

- $x \rightarrow x + v \, dt/2$
- $v \rightarrow v + a \, dt/2$
- calculate  $a$  from current half-step  $v$  and  $x$
- $v \rightarrow v + a \, dt/2$
- $x \rightarrow x + v \, dt/2$

## 5 Bibliography

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